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NICS

Star Formation with Turbulence and Magnetic Fields

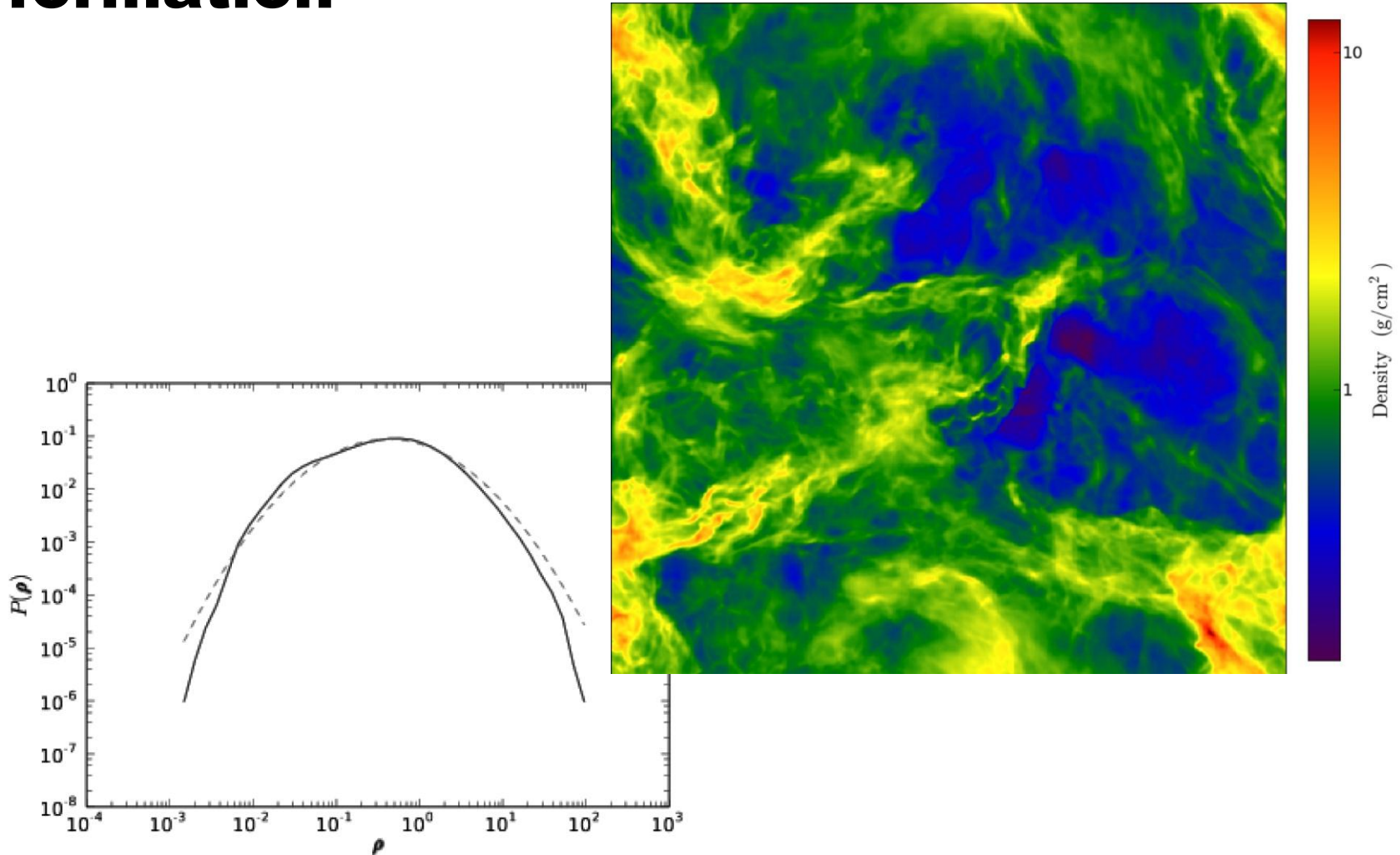
Numerical models with
EnzoMHD on Kraken

**David C. Collins, Paolo Padoan,
Michael L. Norman, and Hao Xu**

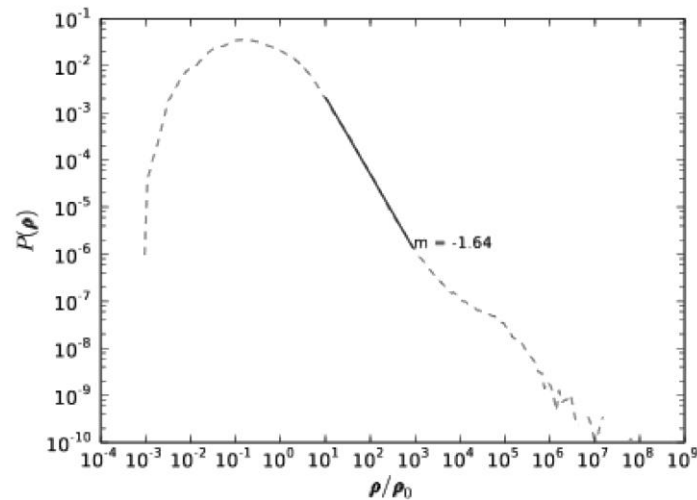
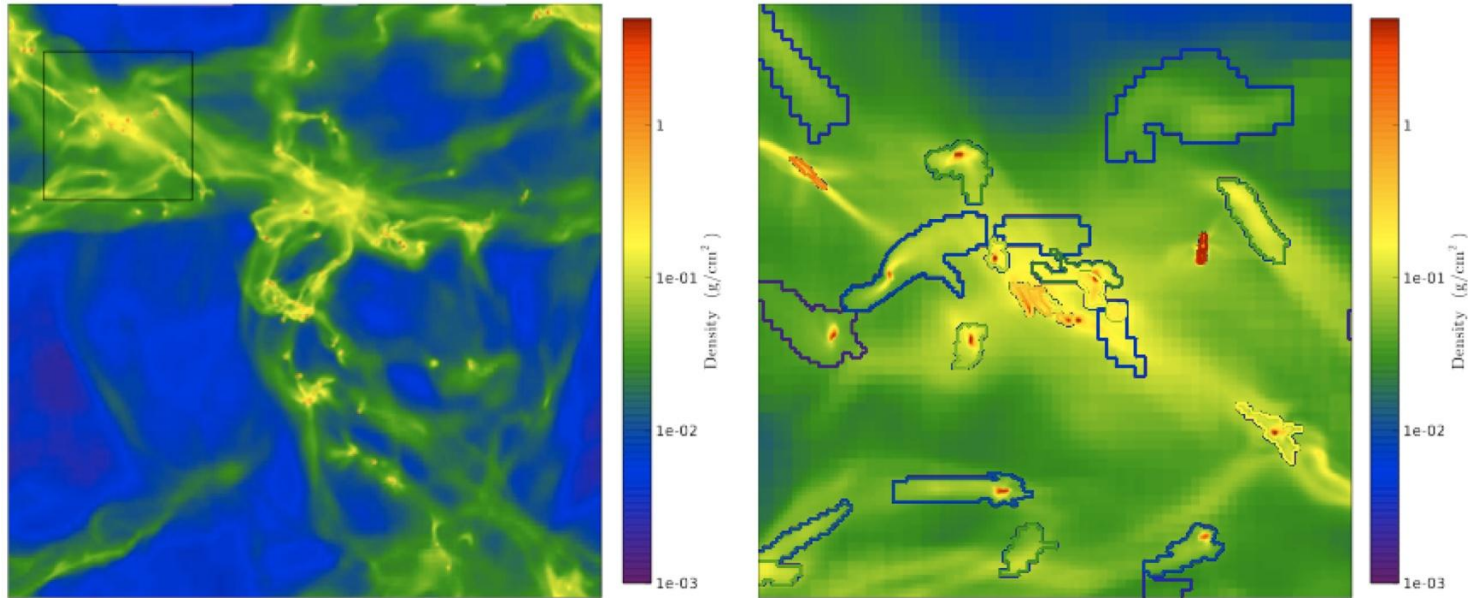
**Center for Astrophysics and Space Science
University of California, San Diego**



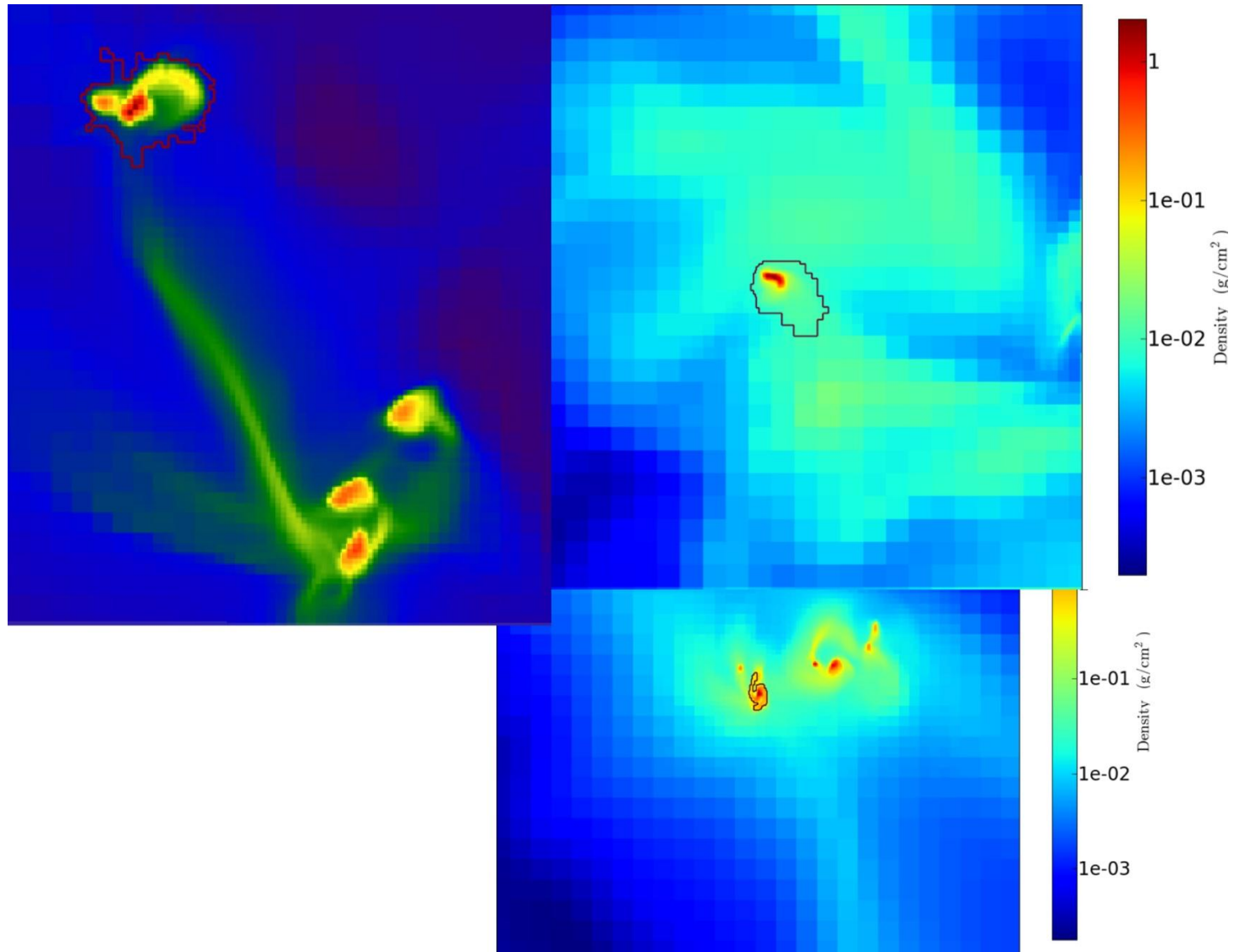
High resolution simulations of MHD turbulence give initial conditions for star formation



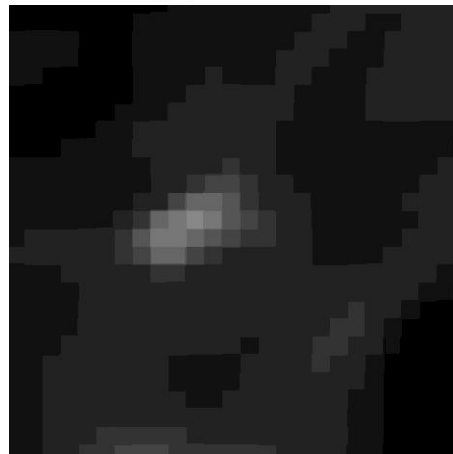
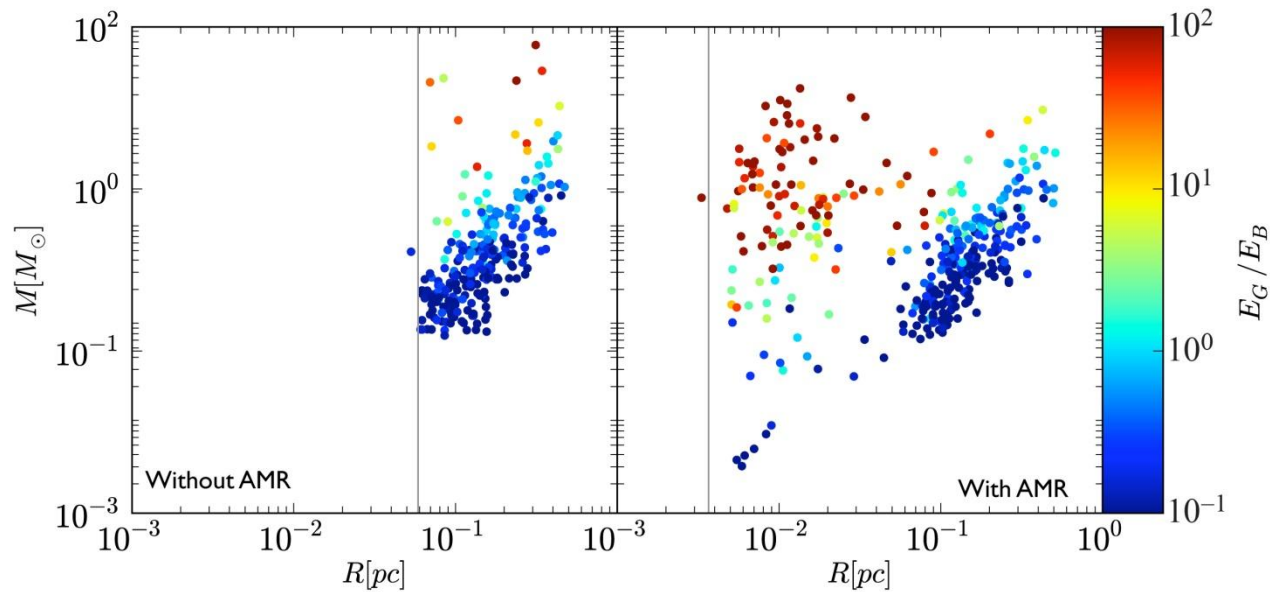
The addition of gravity creates regions of very high density, which creates protostellar cores



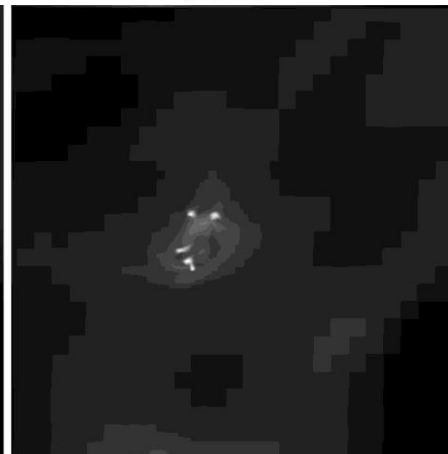
Prestellar cores, stellar birthplace



Adaptive mesh refinement allows us to follow prestellar cores in unprecedented detail

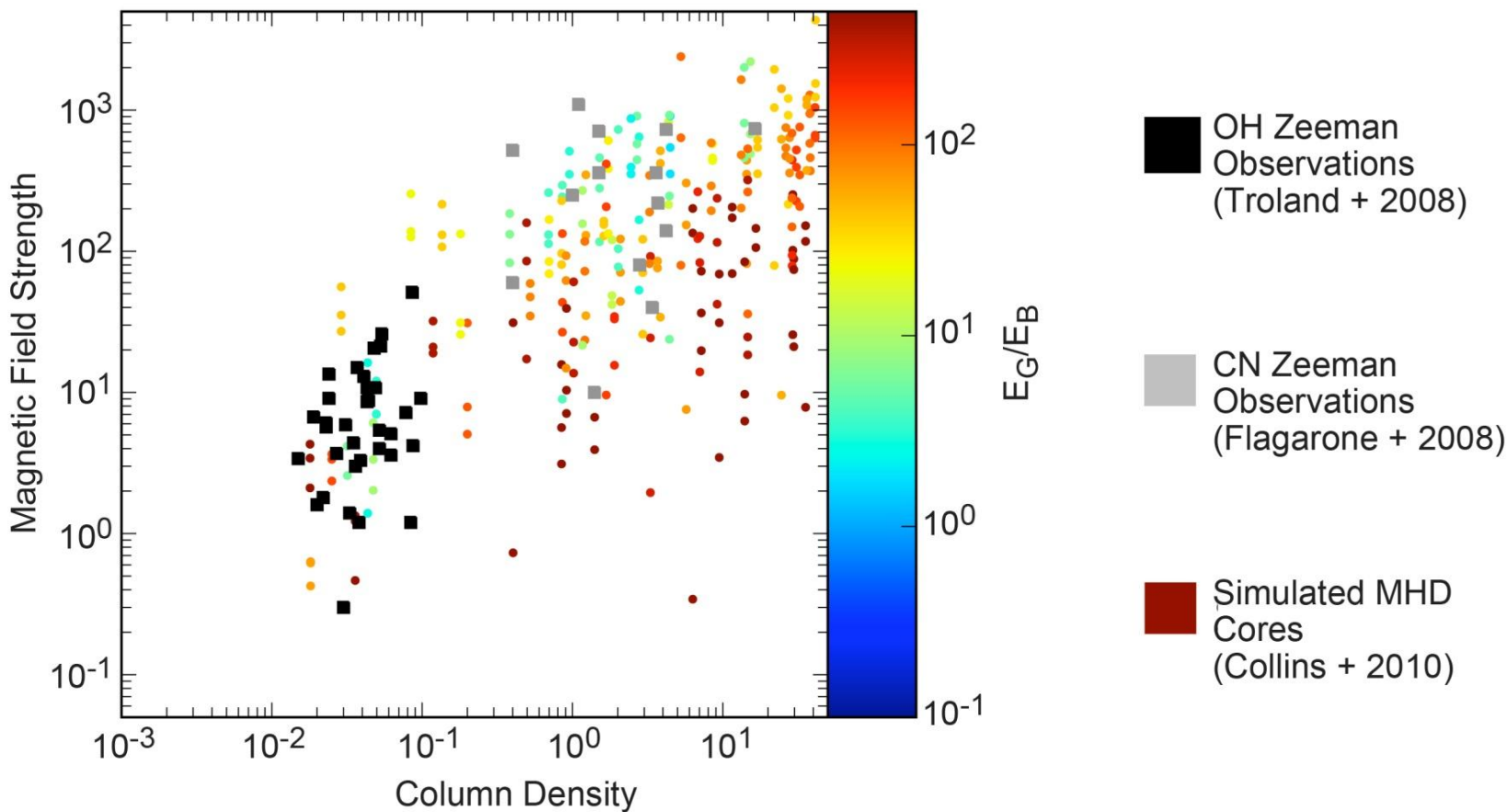


Without AMR



With AMR

Magnetic properties of simulated cores match observations



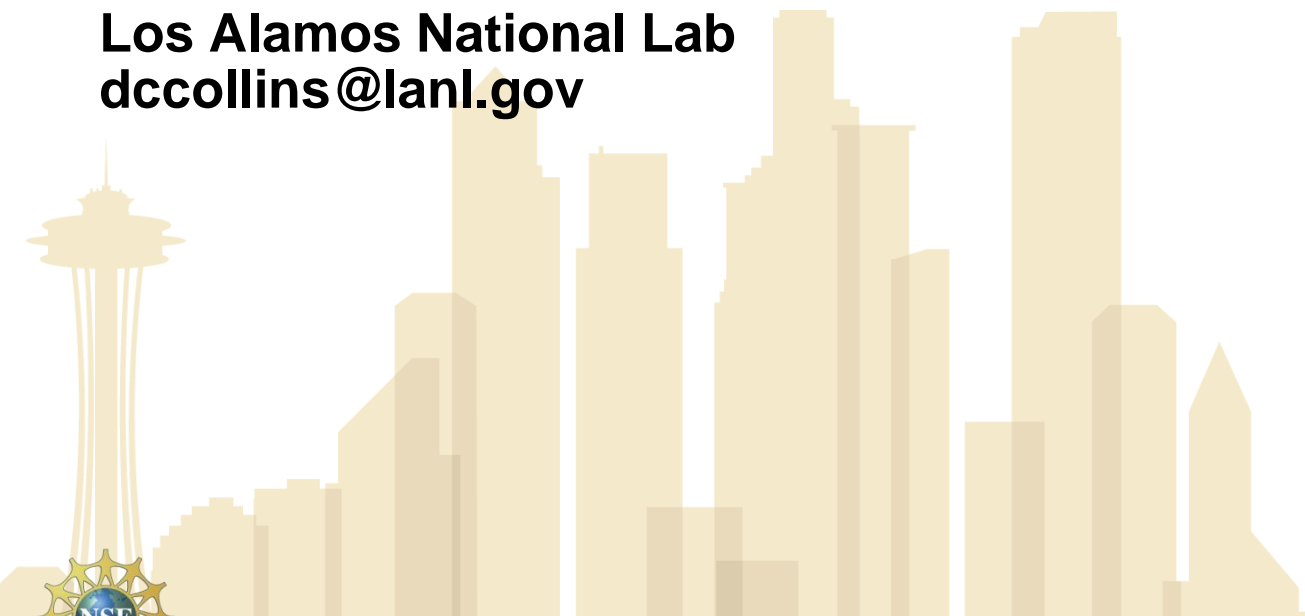


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