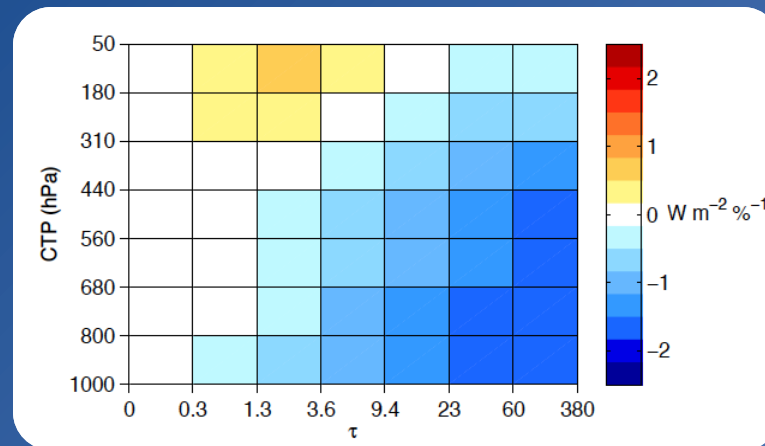


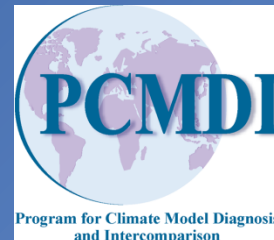
COMPUTING AND PARTITIONING CLOUD FEEDBACKS USING CLOUD PROPERTY HISTOGRAMS



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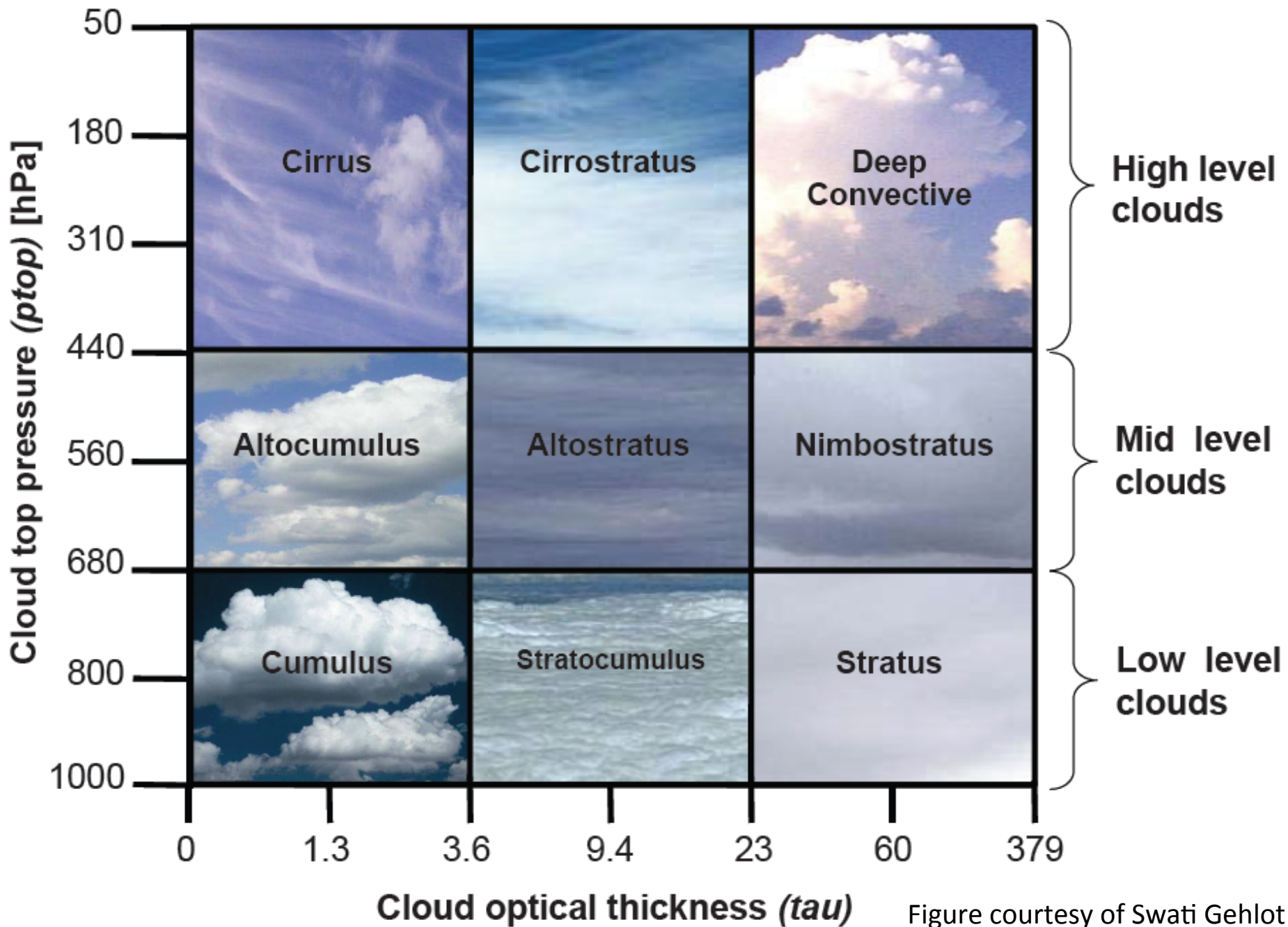


Goals

- To provide a **clean and simple** method of computing cloud feedbacks that is highly **informative**
- **Clean:**
 - compute cloud feedback from ISCCP simulator-interpreted cloud changes directly (not inferred)
 - standard definition of “cloud” and radiation code across models
- **Simple:**
 - no need to correct for non-cloud effects
 - no partial radiative perturbation calculations are needed
 - can use monthly mean model output
- **Informative:**
 - can **quantify** the contribution to cloud feedback from **changing amounts of individual cloud types** (high, middle, low) and from **individual processes** (Δ altitude, Δ optical depth, Δ total amount)

Data & Methodology

- Doubled CO₂ equilibrium slab ocean model simulations from 12 GCMs as part of CFMIP1
- ISCCP simulator (Klein & Jakob 1999) run inline during integration
 - Produce distribution of cloud fraction (as function of CTP and τ) that is consistent with how a satellite-borne passive sensor would “view” the model atmosphere
 - Simulated cloud fractions are defined consistently across models



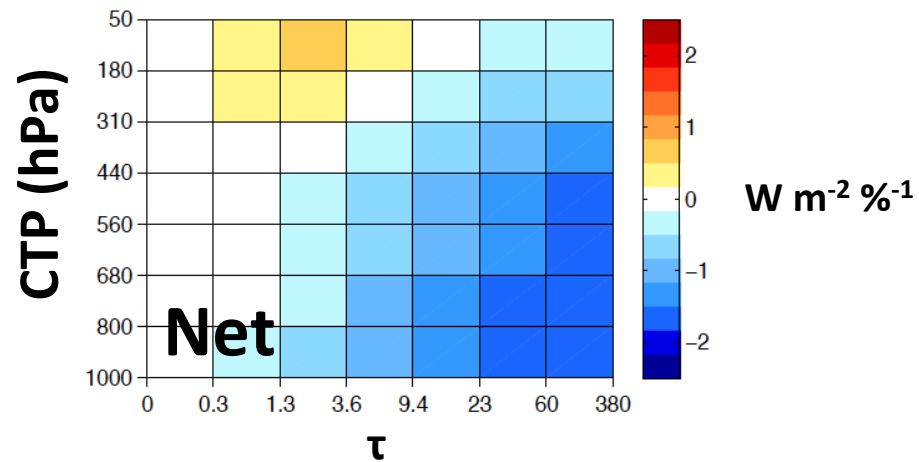
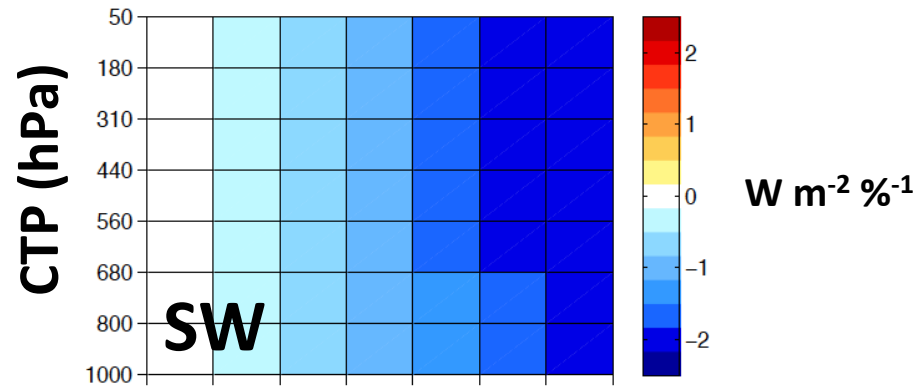
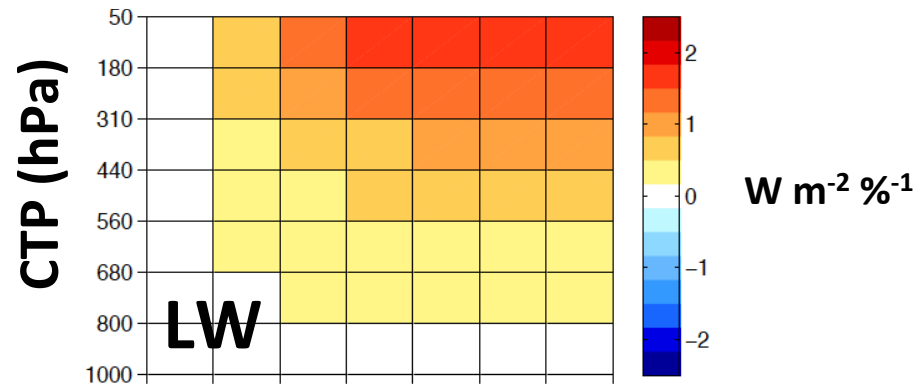
Data & Methodology

- Doubled CO₂ equilibrium slab ocean model simulations from 12 GCMs as part of CFMIP1
- ISCCP simulator run inline during integration
 - Produce distribution of cloud fraction (as function of CTP and τ) that is consistent with how a satellite-borne passive sensor would “view” the model atmosphere
 - Simulated cloud fractions are defined consistently across models
- We compute cloud radiative kernels \rightarrow sensitivity of TOA radiation to cloud fraction changes in each CTP- τ bin
- Cloud feedback = Δ cloud fraction times cloud kernel normalized by ΔT_{sfc}

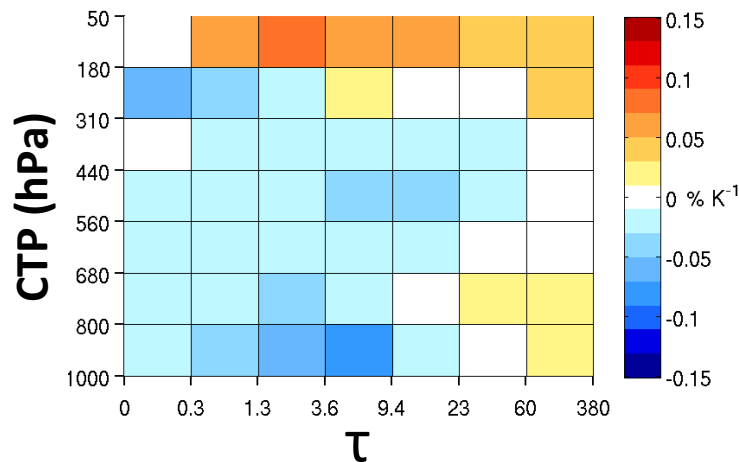
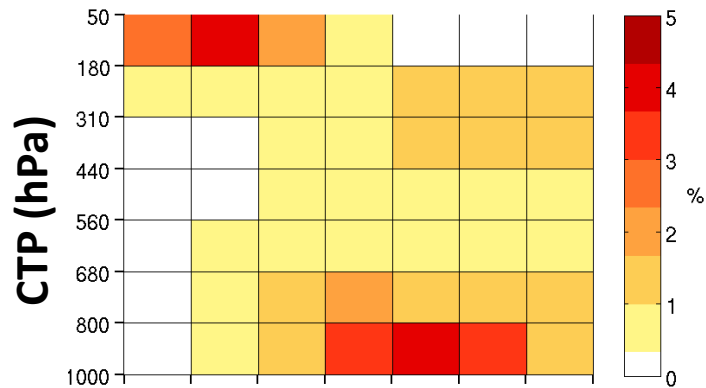
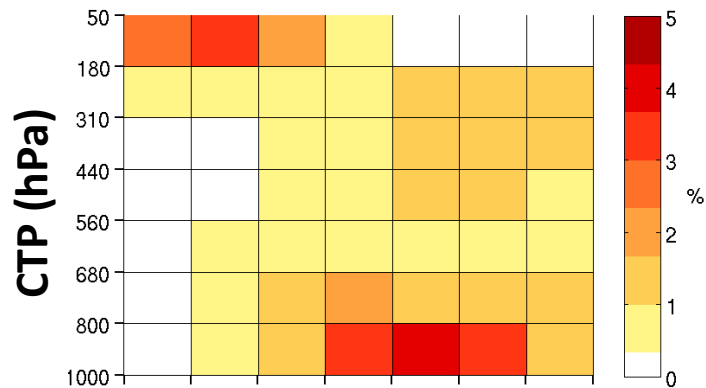
Recipe for Constructing Cloud Radiative Kernels

- ❑ Input model mean zonal mean T and q profiles to Fu-Liou code
- ❑ Compute clear-sky TOA fluxes
- ❑ Compute overcast-sky fluxes for each CTP and τ bin by setting the LWC / IWC profiles to values appropriate for each cloud type
- ❑ Subtract overcast TOA fluxes in each bin from the clear-sky flux to compute a matrix of overcast sky cloud forcing
- ❑ Divide by 100 to get $W m^{-2} \%^{-1}$
- ❑ Repeat every calculation for 24 solar zenith angles, all latitudes, 12 months, and 10 surface albedo bins between 0 and 1

Global Annual Mean Cloud Kernels



Cloud Fraction



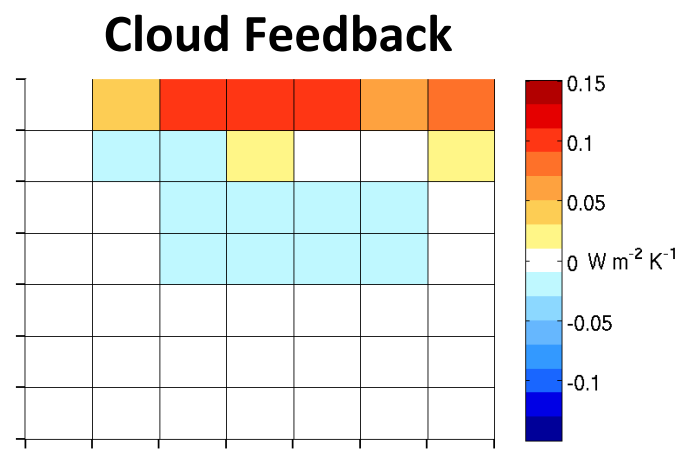
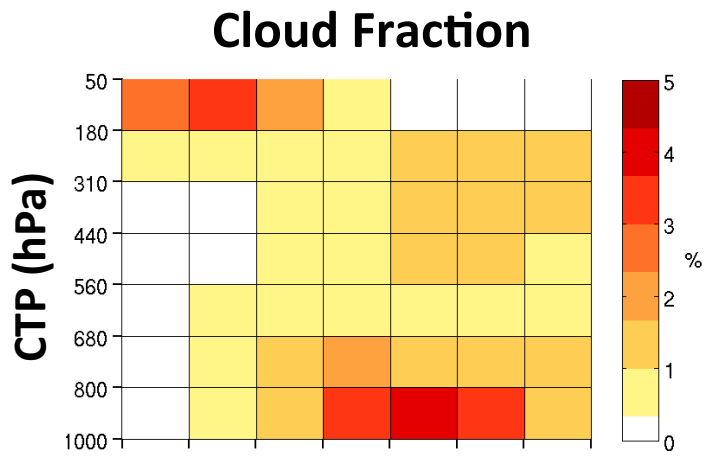
1xCO₂

2xCO₂

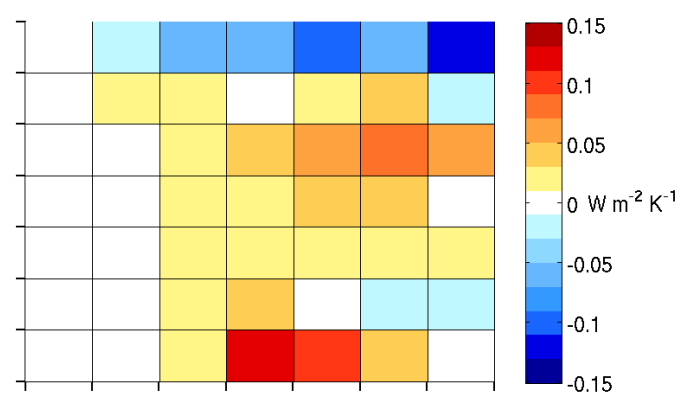
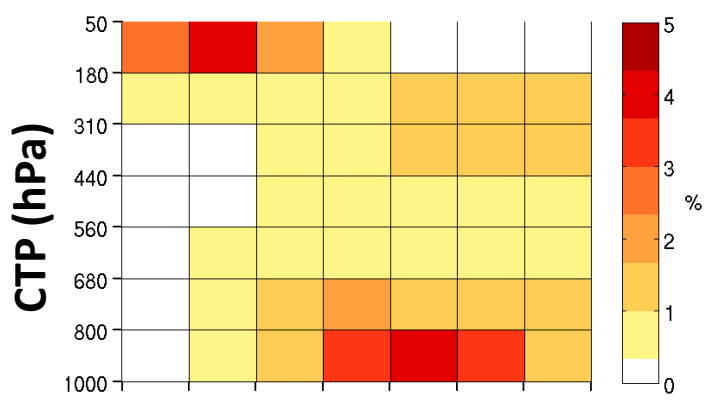
Change
-0.4 % K⁻¹

x Cloud Radiative Kernels
at each location and month,
then averaged annually,
globally, and across models...

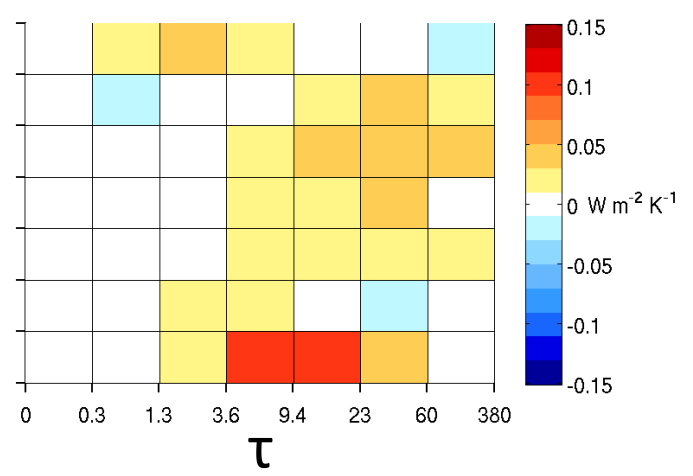
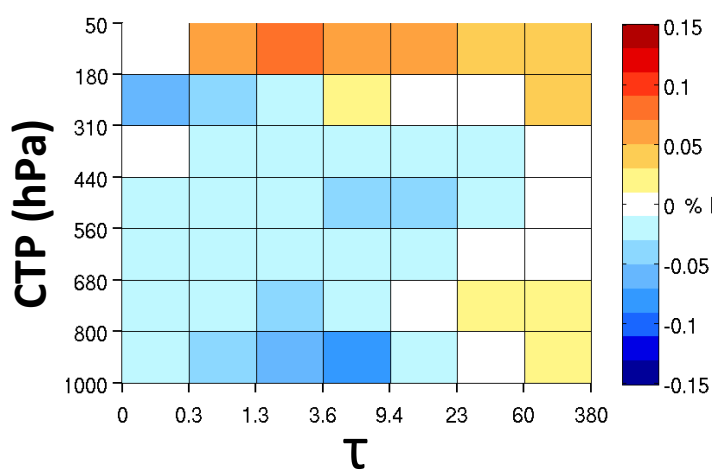
1xCO₂



2xCO₂



Change
-0.4 % K⁻¹



LW
0.27 $\text{W m}^{-2} \text{K}^{-1}$

SW
0.44 $\text{W m}^{-2} \text{K}^{-1}$

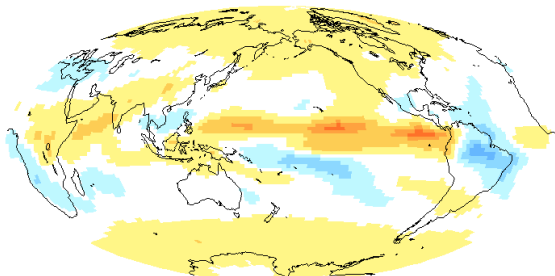
Net
0.71 $\text{W m}^{-2} \text{K}^{-1}$

Cloud Kernel

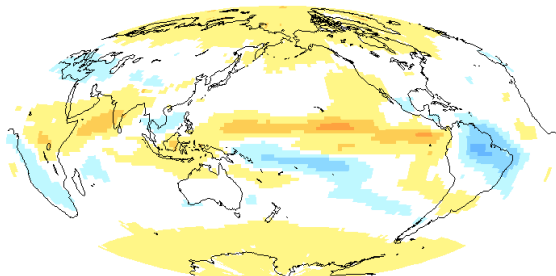
Adjusted Δ CRF

**Kernel minus
Adjusted Δ CRF**

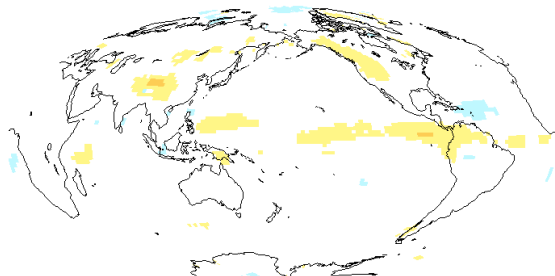
LW



0.27 $\text{W m}^{-2} \text{K}^{-1}$

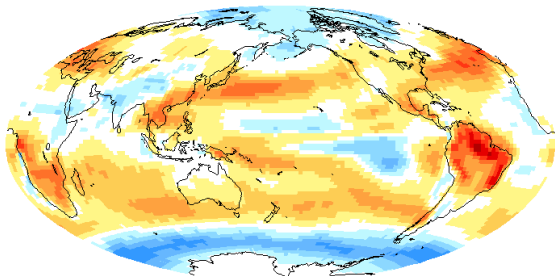


0.21 $\text{W m}^{-2} \text{K}^{-1}$

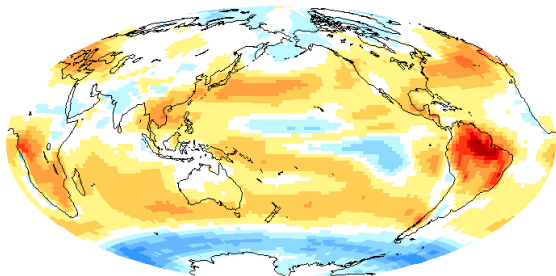


0.06 $\text{W m}^{-2} \text{K}^{-1}$

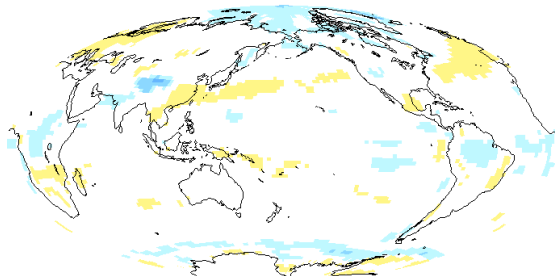
SW



0.44 $\text{W m}^{-2} \text{K}^{-1}$

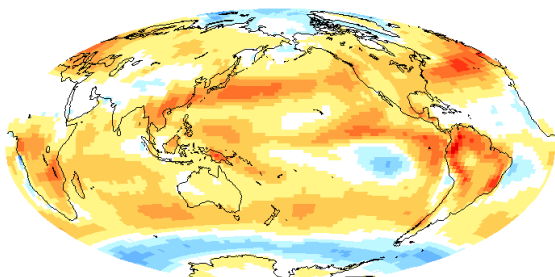


0.44 $\text{W m}^{-2} \text{K}^{-1}$

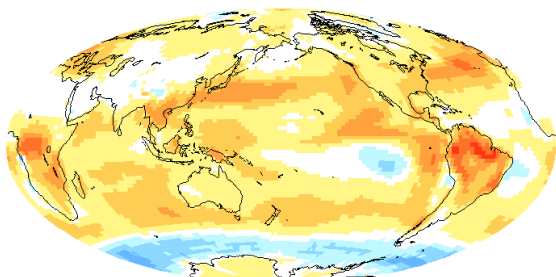


-0.01 $\text{W m}^{-2} \text{K}^{-1}$

Net



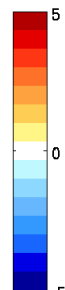
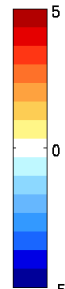
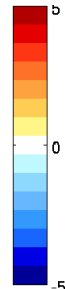
0.70 $\text{W m}^{-2} \text{K}^{-1}$



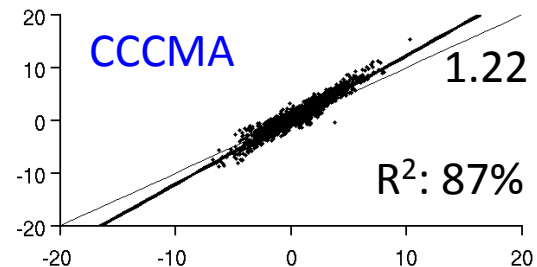
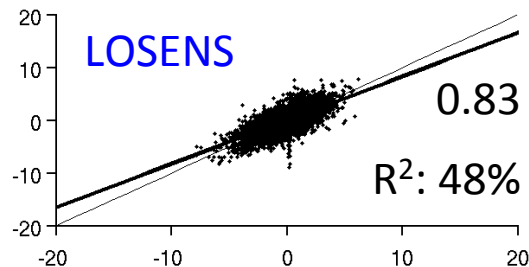
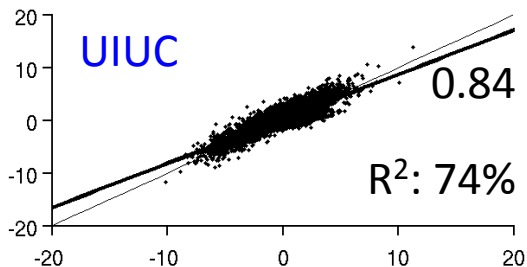
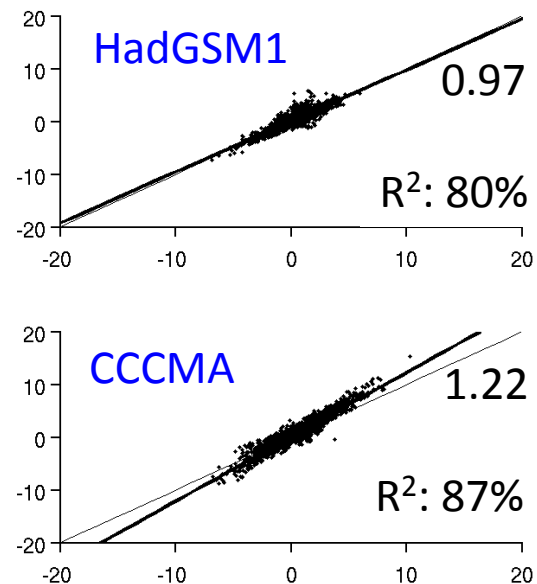
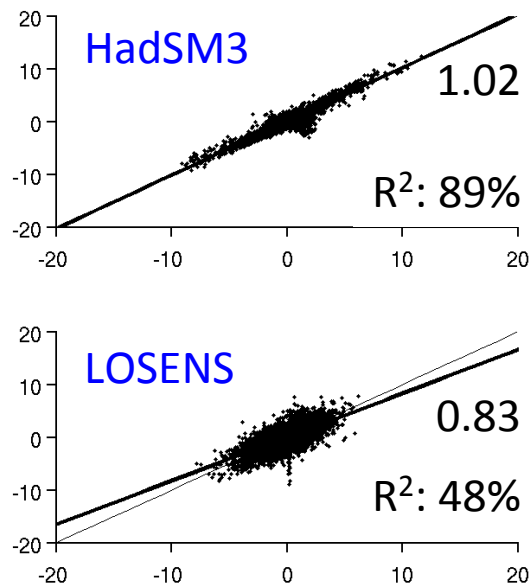
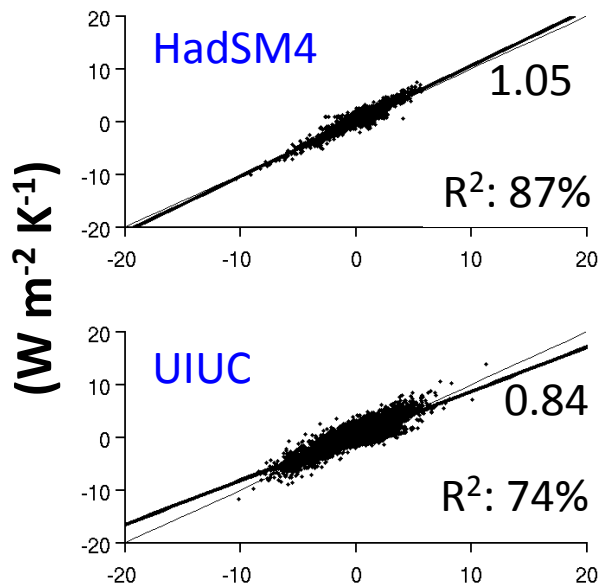
0.65 $\text{W m}^{-2} \text{K}^{-1}$



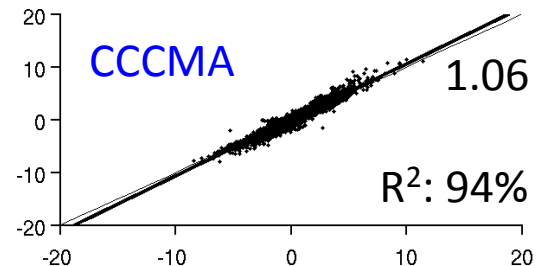
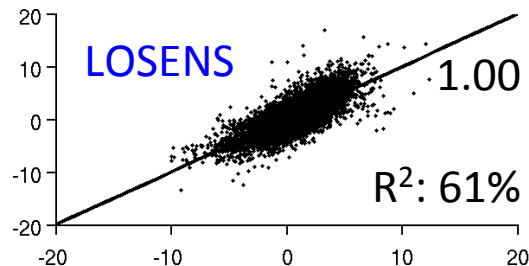
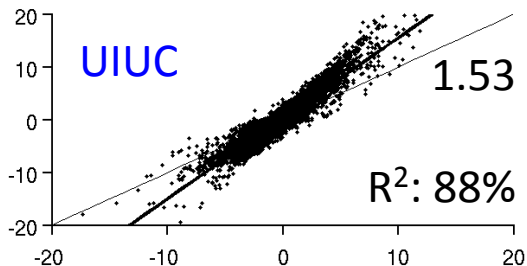
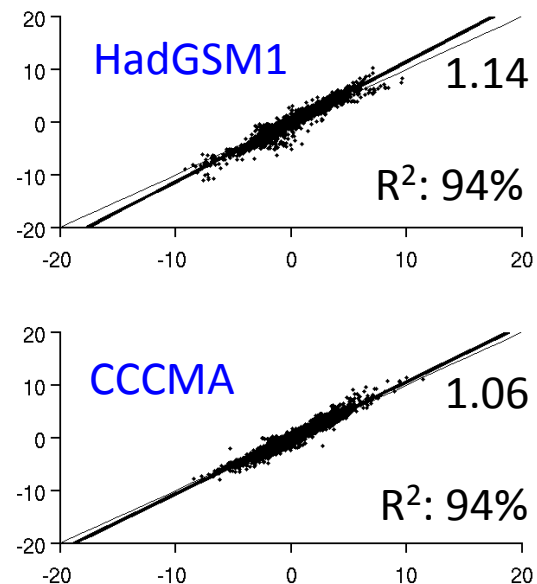
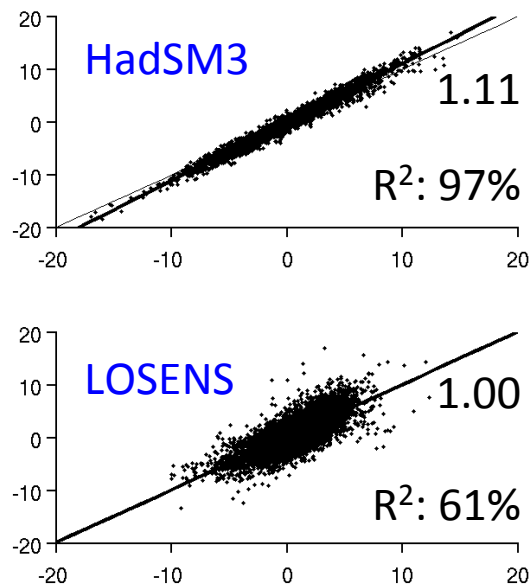
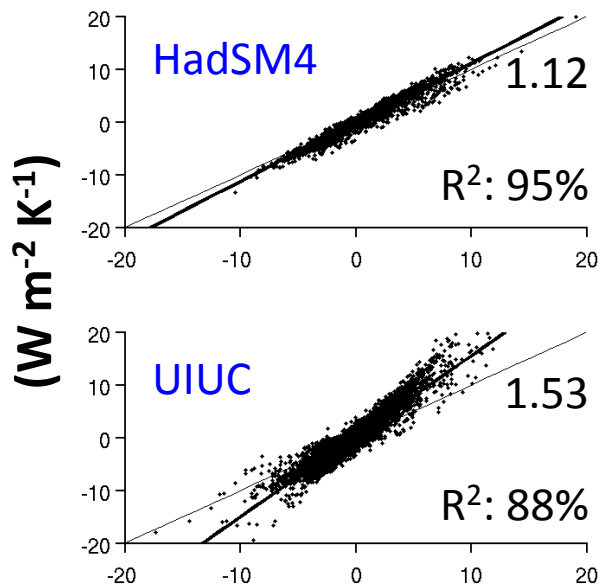
0.04 $\text{W m}^{-2} \text{K}^{-1}$



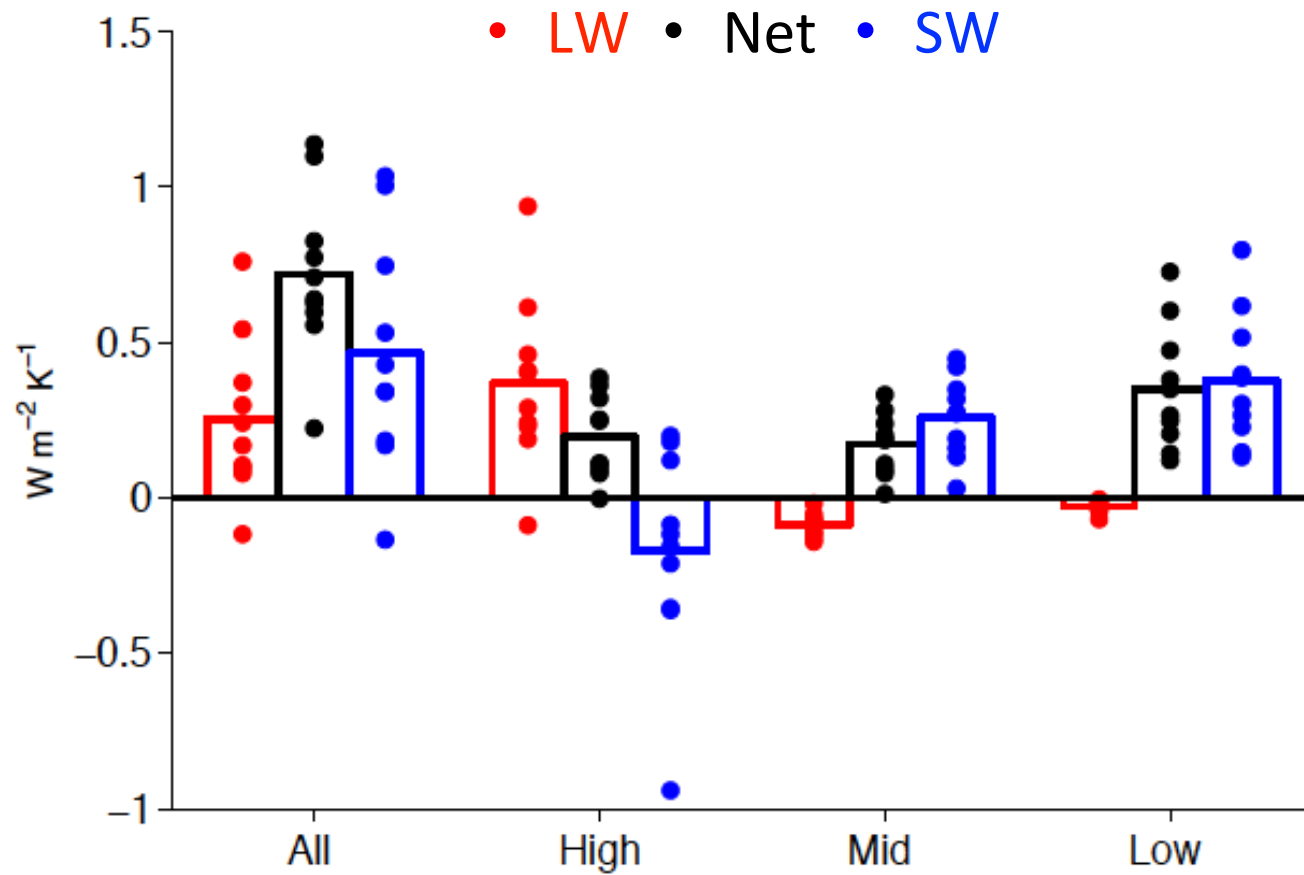
LW Cloud Kernel



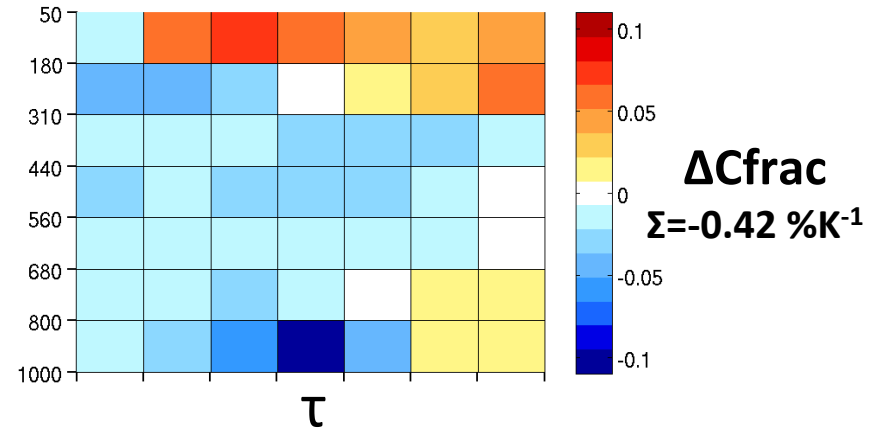
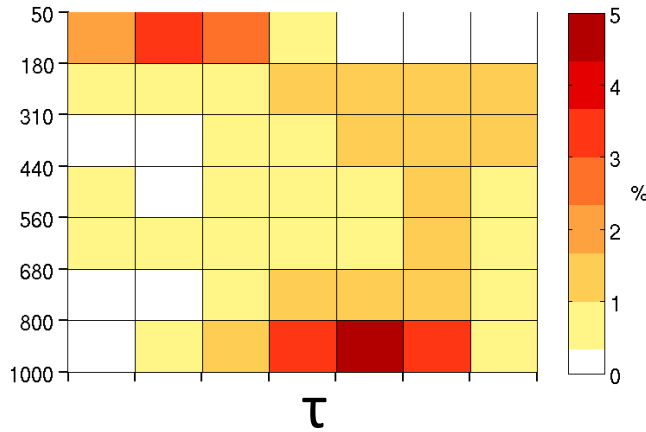
SW Cloud Kernel



Adjusted ΔCRF ($W m^{-2} K^{-1}$)



**Mean
Cfrac
 $\Sigma=52.5\%$**



- Decompose the cloud changes into

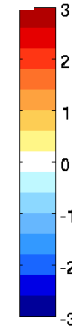
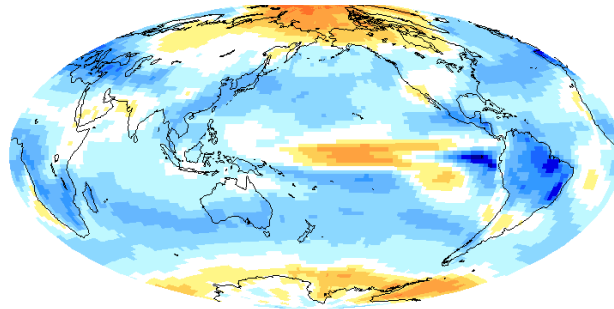
Δ AMOUNT

Δ ALTITUDE

Δ OPTICAL DEPTH

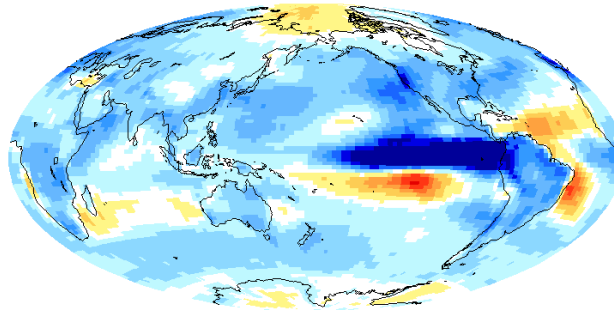
Δ Total Cloud Fraction

$\Sigma = -0.4 \% K^{-1}$



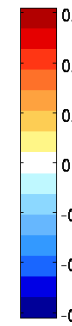
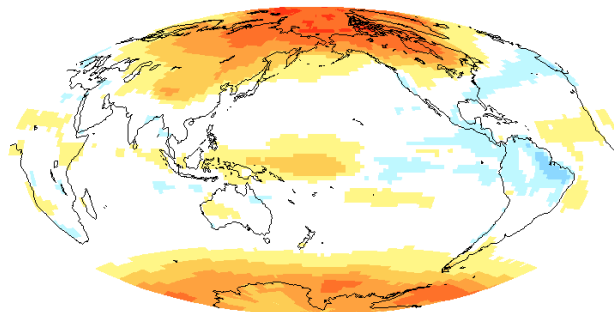
Δ CTP

$\Sigma = -4.05 \text{ hPa } K^{-1}$



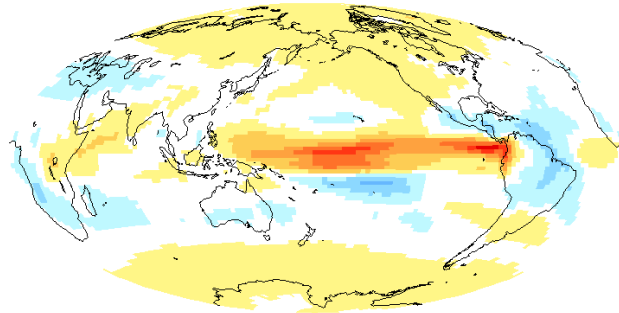
$\Delta \ln(\tau)$

$\Sigma = 3 \% K^{-1}$



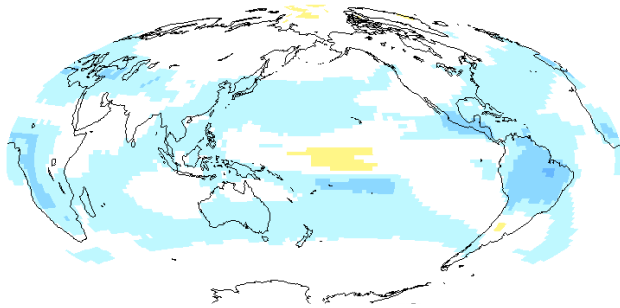
LW Cloud Feedback

$0.26 \text{ W m}^{-2} \text{ K}^{-1}$



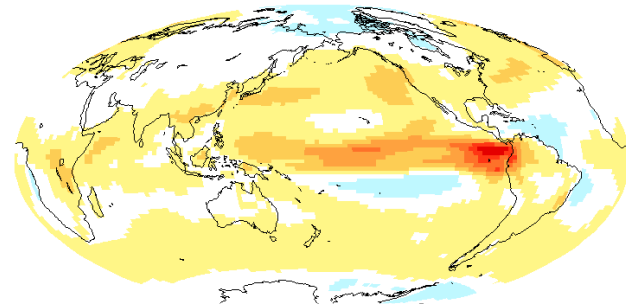
Amount

$-0.30 \text{ W m}^{-2} \text{ K}^{-1}$



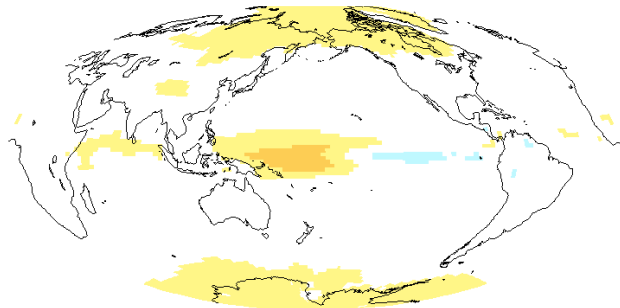
Altitude

$0.44 \text{ W m}^{-2} \text{ K}^{-1}$



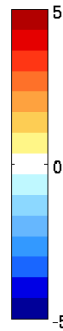
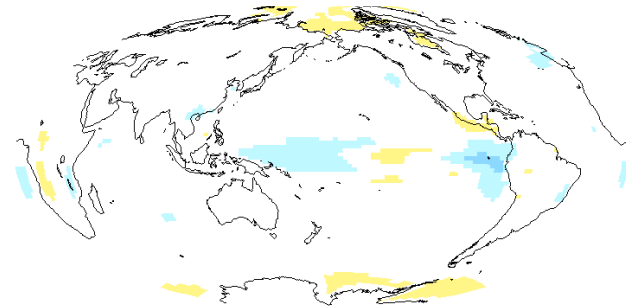
Optical Depth

$0.16 \text{ W m}^{-2} \text{ K}^{-1}$



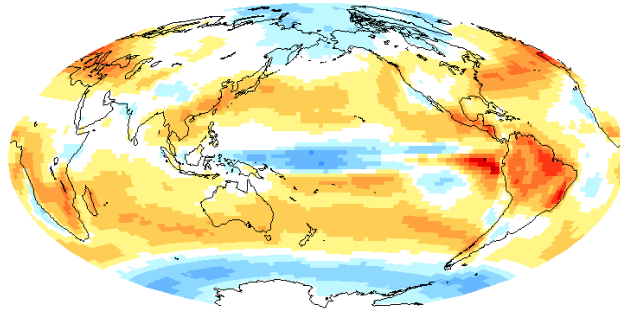
Residual

$-0.04 \text{ W m}^{-2} \text{ K}^{-1}$



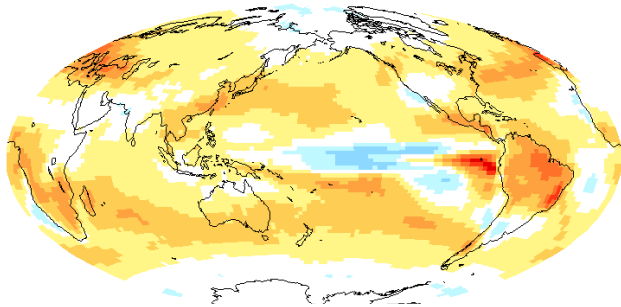
SW Cloud Feedback

$0.46 \text{ W m}^{-2} \text{ K}^{-1}$



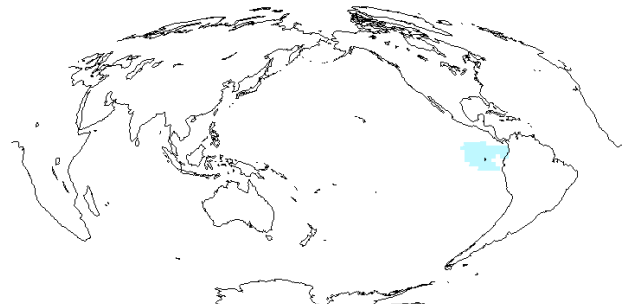
Amount

$0.66 \text{ W m}^{-2} \text{ K}^{-1}$



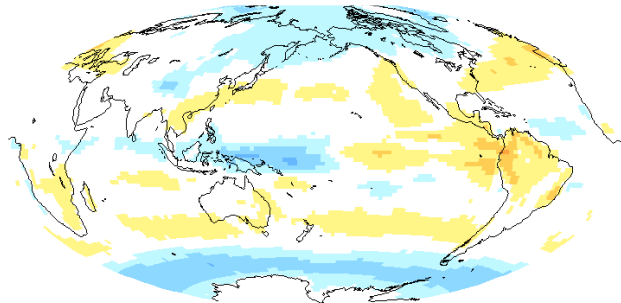
Altitude

$-0.03 \text{ W m}^{-2} \text{ K}^{-1}$



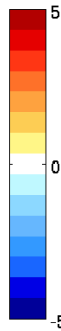
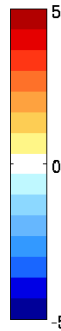
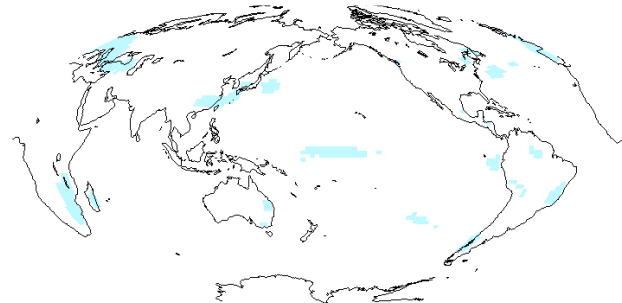
Optical Depth

$-0.05 \text{ W m}^{-2} \text{ K}^{-1}$



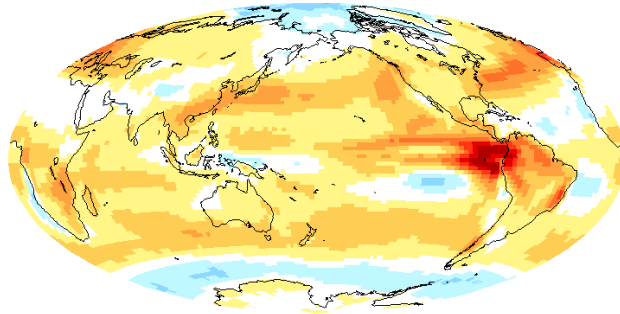
Residual

$-0.11 \text{ W m}^{-2} \text{ K}^{-1}$



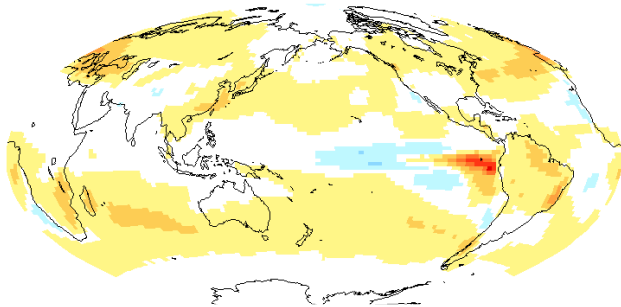
Net Cloud Feedback

$0.71 \text{ W m}^{-2} \text{ K}^{-1}$



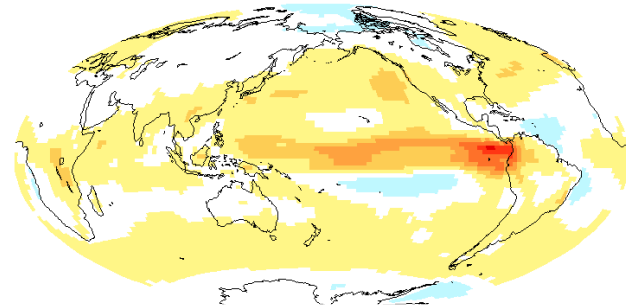
Amount

$0.36 \text{ W m}^{-2} \text{ K}^{-1}$



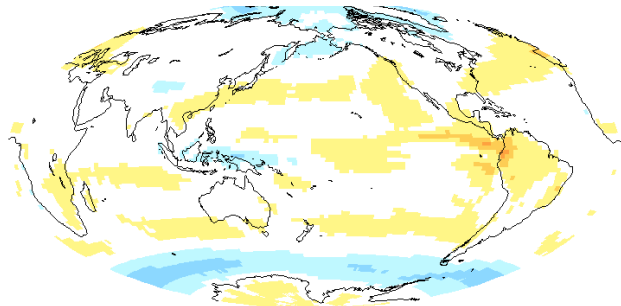
Altitude

$0.41 \text{ W m}^{-2} \text{ K}^{-1}$



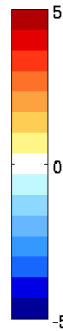
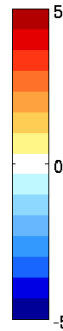
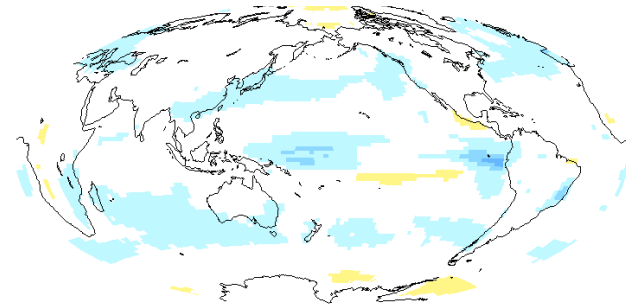
Optical Depth

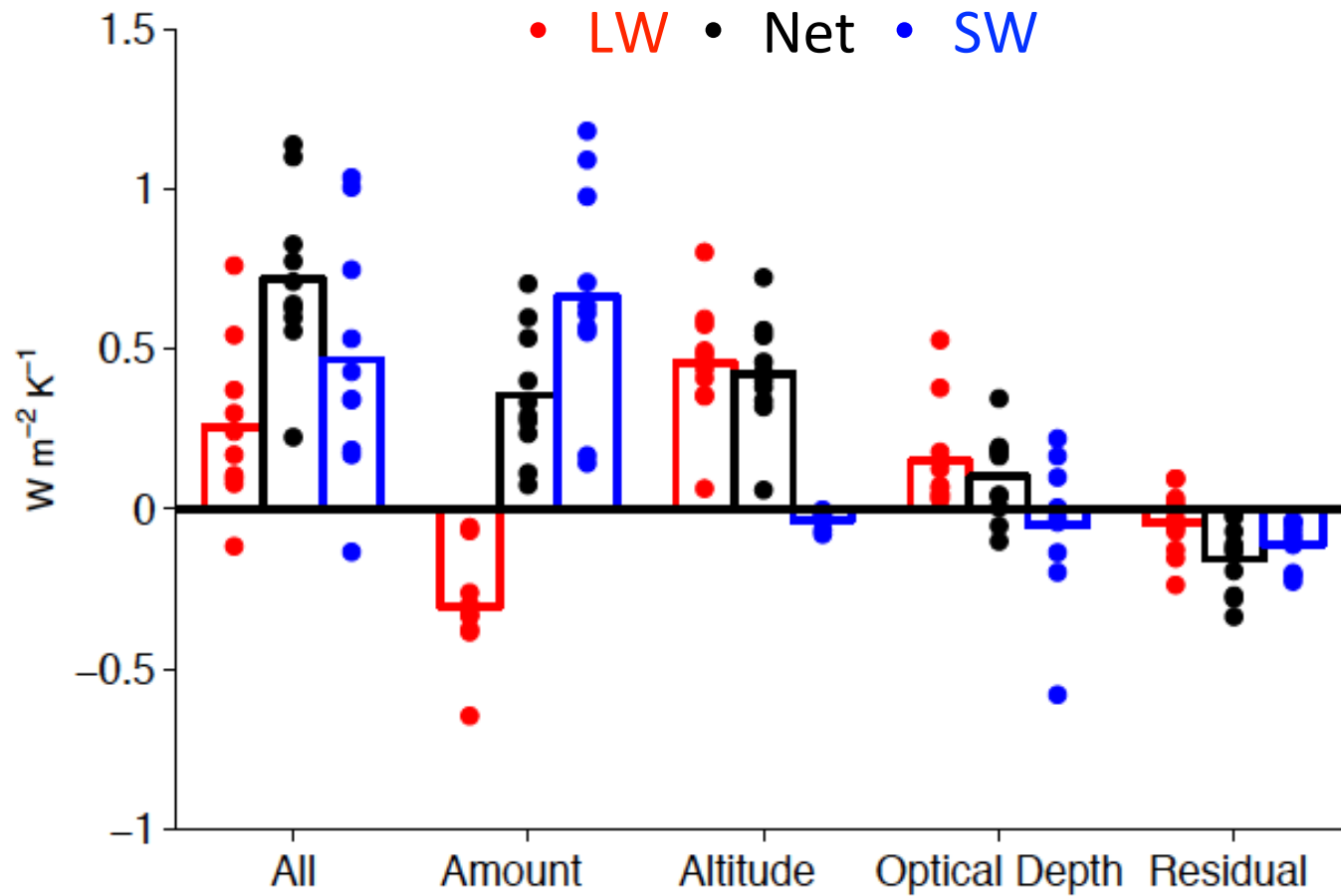
$0.09 \text{ W m}^{-2} \text{ K}^{-1}$



Residual

$-0.15 \text{ W m}^{-2} \text{ K}^{-1}$





Conclusions

- Cloud radiative kernels allow computation of cloud feedback directly from cloud fraction histograms produced by ISCCP simulator
- Feedbacks computed with cloud kernels compare very well with those computed by adjusting the change in cloud forcing
- More than half of the global mean net cloud feedback can be attributed to the combined response of middle- and high-level clouds
- High cloud changes induce wider range of LW and SW cloud feedbacks across models than do low clouds
- Increasing cloud top altitude is dominant contributor to the positive global mean LW and net cloud feedbacks (positive in every model)
- Decreasing total cloud fraction is dominant contributor to global mean positive SW cloud feedback (positive in every model)
- Large negative net cloud feedback at high latitudes is caused by increased optical depth, not increased cloud amount

Thank you!

Drafts of papers: Google “Mark Zelinka”

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