# An Assessment of Radiative Forcing from CMM5 Models

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#### **Collaborators:**



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y Dan Schwarzkopf L) (NOAA/GFDL) Gabe Vecchi (NOAA/GFDL)

### **Take Home Points**

**Dominant sources of uncertainty in CMIP3 climate projections :** 

(1) <u>Direct</u> radiative forcing from aerosols

(2) Low (stratocumulus/cumulus) cloud feedback

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#### (1) <u>Direct</u> radiative forcing from aerosols

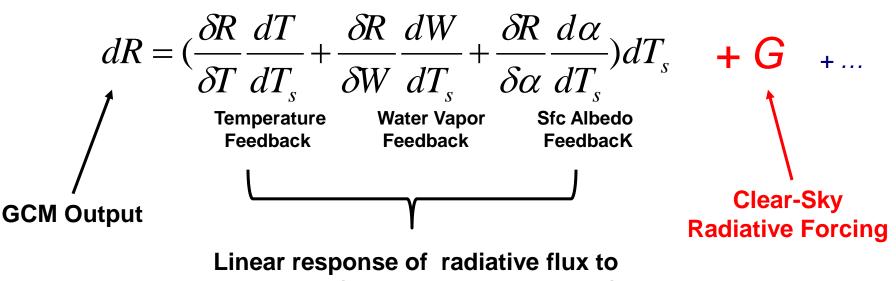
(2) Low (stratocumulus/cumulus) cloud feedback

### Motivation

- Most modeling centers do not provide (calculate?) the radiative forcing for different emission scenarios.
- Those that do calculate the radiative forcing usually do so differently from one group to the next.
- This leads to attempts to estimate the radiative forcing from available output (e.g., "Gregory Method")
- We use "radiative kernels" (Soden et al., 2008) to estimate clear-sky radiative forcing.

### **Estimating Radiative Forcing using "Kernels"**

Consider the Change in Net Clear-sky Flux at TOA: dR

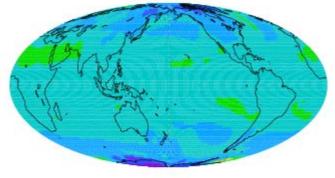


feedbacks (computed from kernels)

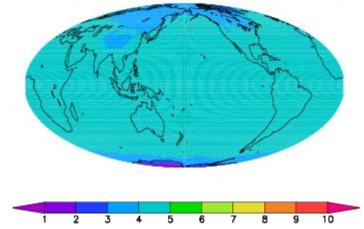
### **Radiative Forcing: Kernel vs. Direct Calculation**

2x CO2

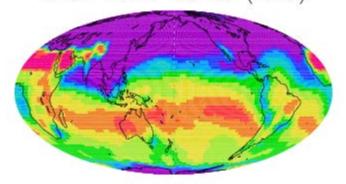
GFDL CM2.0 Kernel (4.20)



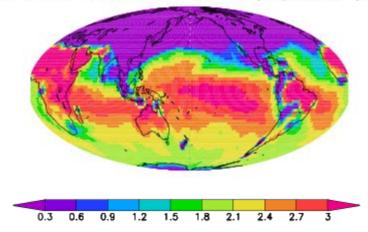
GFDL AM2p12b Instant Tropopause (4.27)



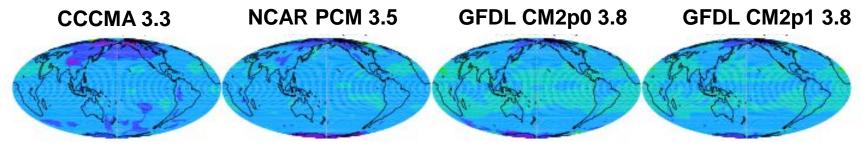
20C3M GFDL CM2.0 Kernel (0.76)

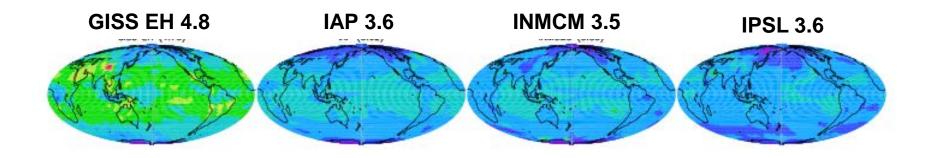


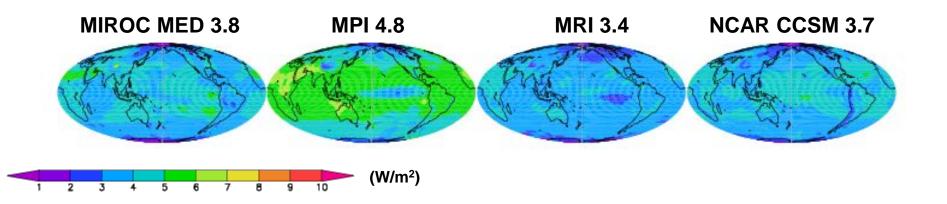
GFDL AM2 Instantaneous Tropopause (0.85)



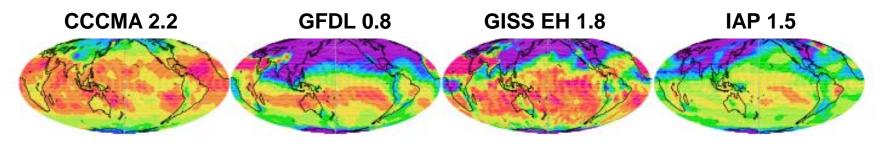
### **Clear-sky Radiative Forcing: IPCC AR4 2xCO2**

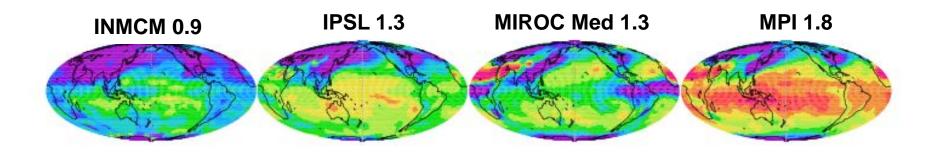


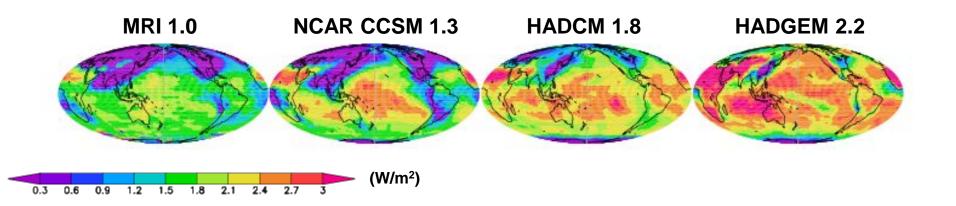




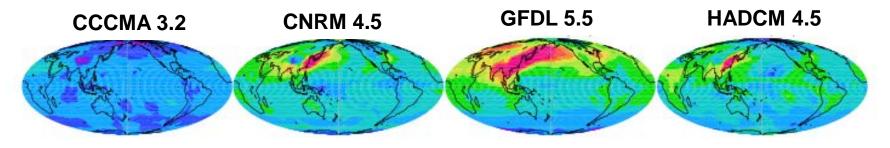
### **Clear-sky Radiative Forcing: IPCC AR4 20C3M**

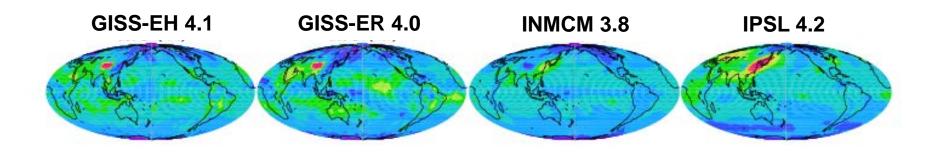


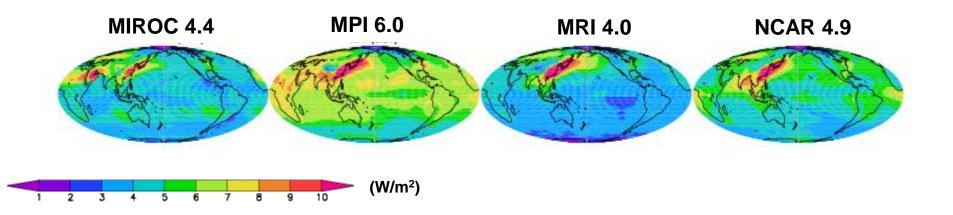




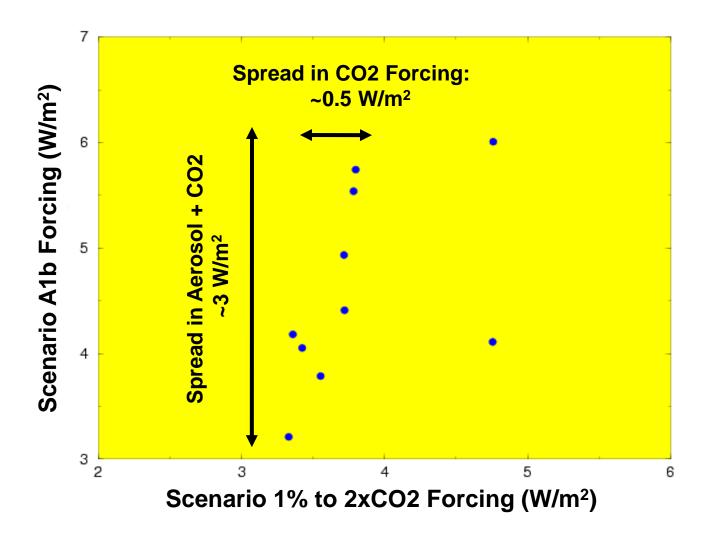
### **Clear-sky Radiative Forcing: IPCC AR4 A1b**







### Clear-sky Radiative Forcing: 2xCO2 vs. A1b



### **Take Home Points**

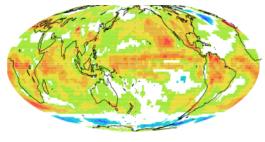
### **Dominant sources of uncertainty in CMIP3 climate projections :**

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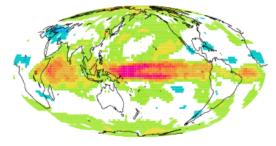
(2) Low (stratocumulus/cumulus) cloud feedback

### **CMIP3 Ensemble Mean Cloud Feedback**

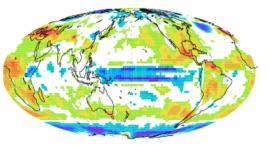
#### **Net Cloud Feedback**



#### LW Cloud Feedback



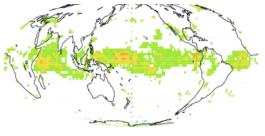
#### **SW Cloud Feedback**



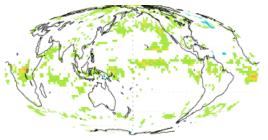
#### 3 2.5 2 1.5 1 0.5 -0.5 -1 -1.5 -2 -2.5

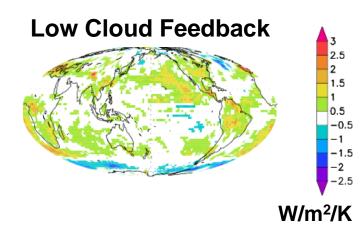
#### W/m²/K

#### **High Cloud Feedback**



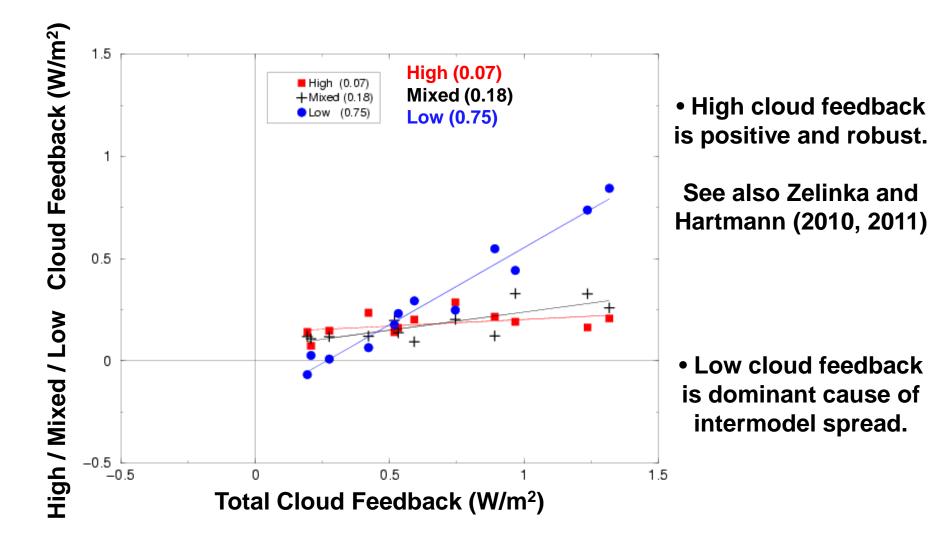
#### **Mixed Cloud Feedback**





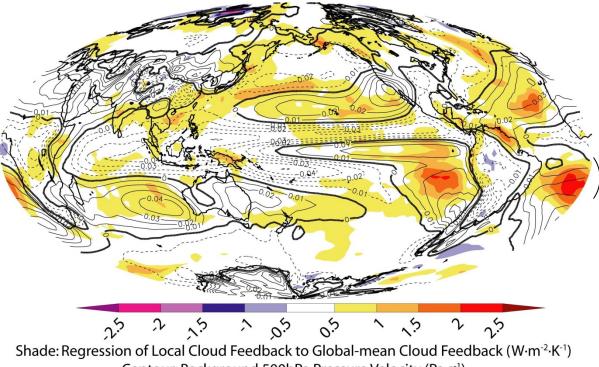
#### Soden and Vecchi (2011)

### **Intermodel Spread in Cloud Feedback**



Soden and Vecchi (2011)

### Local contribution to intermodel spread in cloud feedback



Contour: Background 500hPa Pressure Velocity (Pa·s<sup>-1</sup>)

#### Most of intermodel spread arises from low stratocumulus/cumululs regions

Soden and Vecchi (2011)

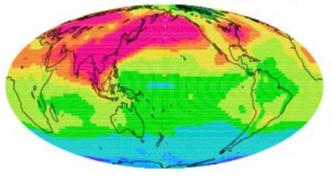
### **Future Work**

- □ Account for cloud masking of radiative forcing
- □ Apply to CMIP5 Models
- **D** Evaluate surface radiative forcings
- □ Isolate "indirect" (fast) forcings

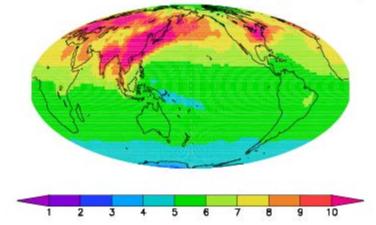
## **Extra Slides**

### **Radiative Forcing: Kernel vs. Direct Calculation**

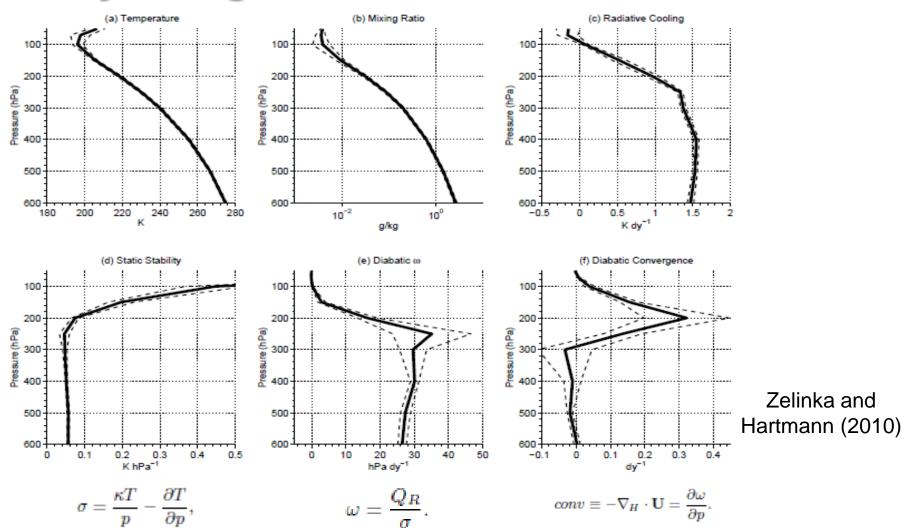
GFDL CM2.0 Kernel (6.91)



GFDL AM2 Instantaneous Tropopause (6.30)



### Why is High Cloud Feedback Positive?



## As climate warms, there is an upward shift in the level of divergence (and $Q_R$ ) due to increased water vapor

### **Observational Evidence for PHAT**

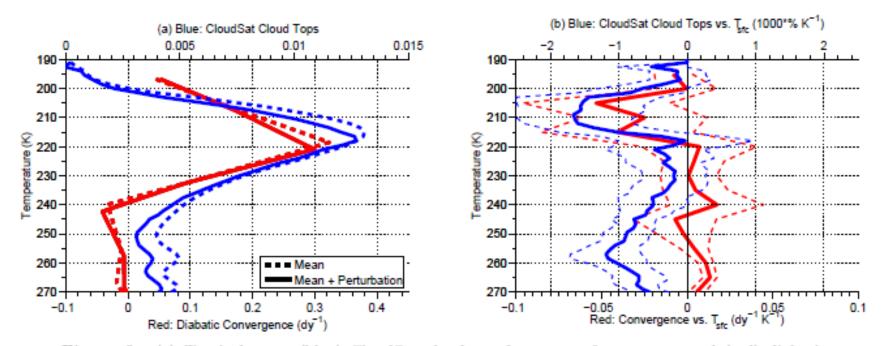


Figure 9. (a) Tropical mean (blue) CloudSat cloud top frequency of occurrence and (red) diabatic convergence. The dashed lines represent the mean profile and the solid lines represent the sum of the mean and perturbation profile shown in panel b. (b) Sensitivity of tropical mean (blue) CloudSat cloud top frequency of occurrence and (red) diabatic convergence to tropical mean surface temperature. The dashed lines represent the  $2\sigma$  range on the regression coefficients computed using a bootstrapping method as described in the text.

#### **Observed interannual changes in tropical high clouds follow FAT/PHAT.**

Zelinka and Hartmann (2011)