

# A Higgs but no sparticles yet: what it means for the (p)MSSM

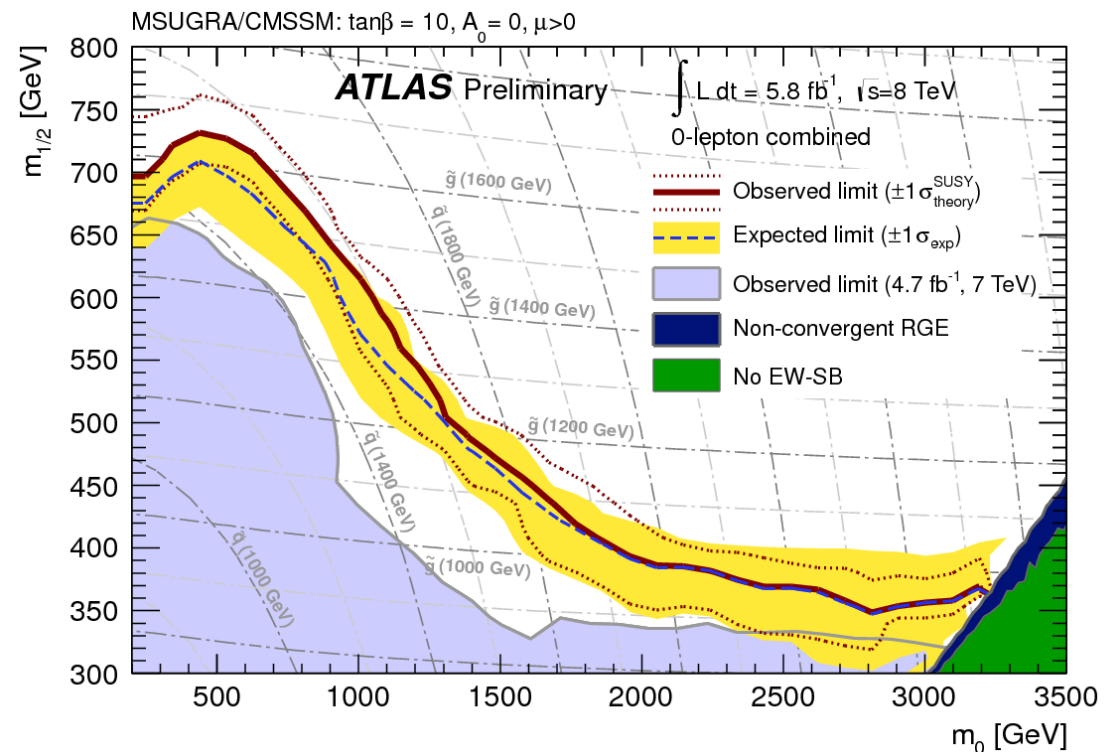
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1206.4321, 1206.5800, 1211.XXXX

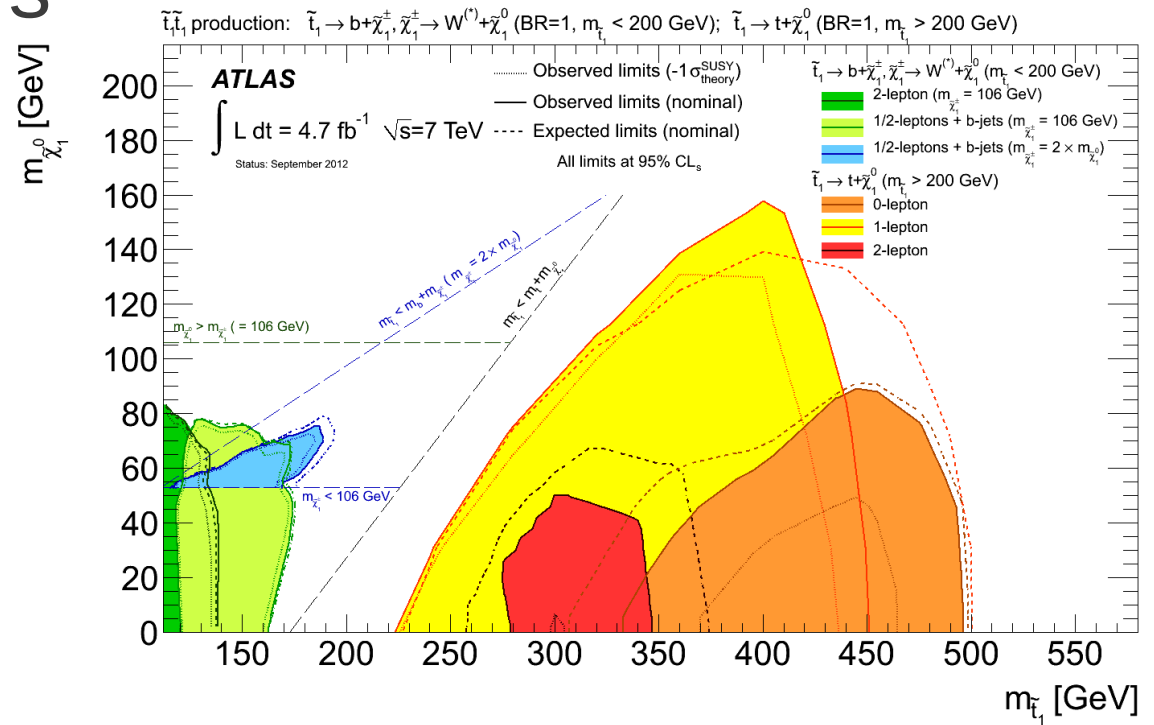
# How well is SUSY being limited?

- LHC data continues to constrain new physics, including SUSY
- MSSM has many free parameters, so search limits are often presented in less general frameworks, e.g. mSUGRA



# How well is SUSY being limited?

- Simplified models use more search-relevant parameters like new particle masses
- Assume rest of spectrum is decoupled



# Another approach

- Instead: can scan over MSSM parameter space, searching for spectra that are consistent with existing experimental bounds (Berger et al., 0812.0980)
- Large number of parameters; sacrifice full coverage for more generality
- Results are not to be interpreted as hard limits on parameters, but examples of wide array of available MSSM phenomenology

# Outline

- The phenomenological MSSM
- LHC search results
- Higgs discovery implications
- Fine-tuning
- Outlook

# The phenomenological MSSM

- The full MSSM has 105 new free parameters, many of which are very strongly constrained by flavor data
- Minimal flavor violation decreases scan dimensionality without losing much generality
- Take sparticle mass matrices to be flavor diagonal, with first two generations degenerate
- No new sources of CP violation

# The phenomenological MSSM

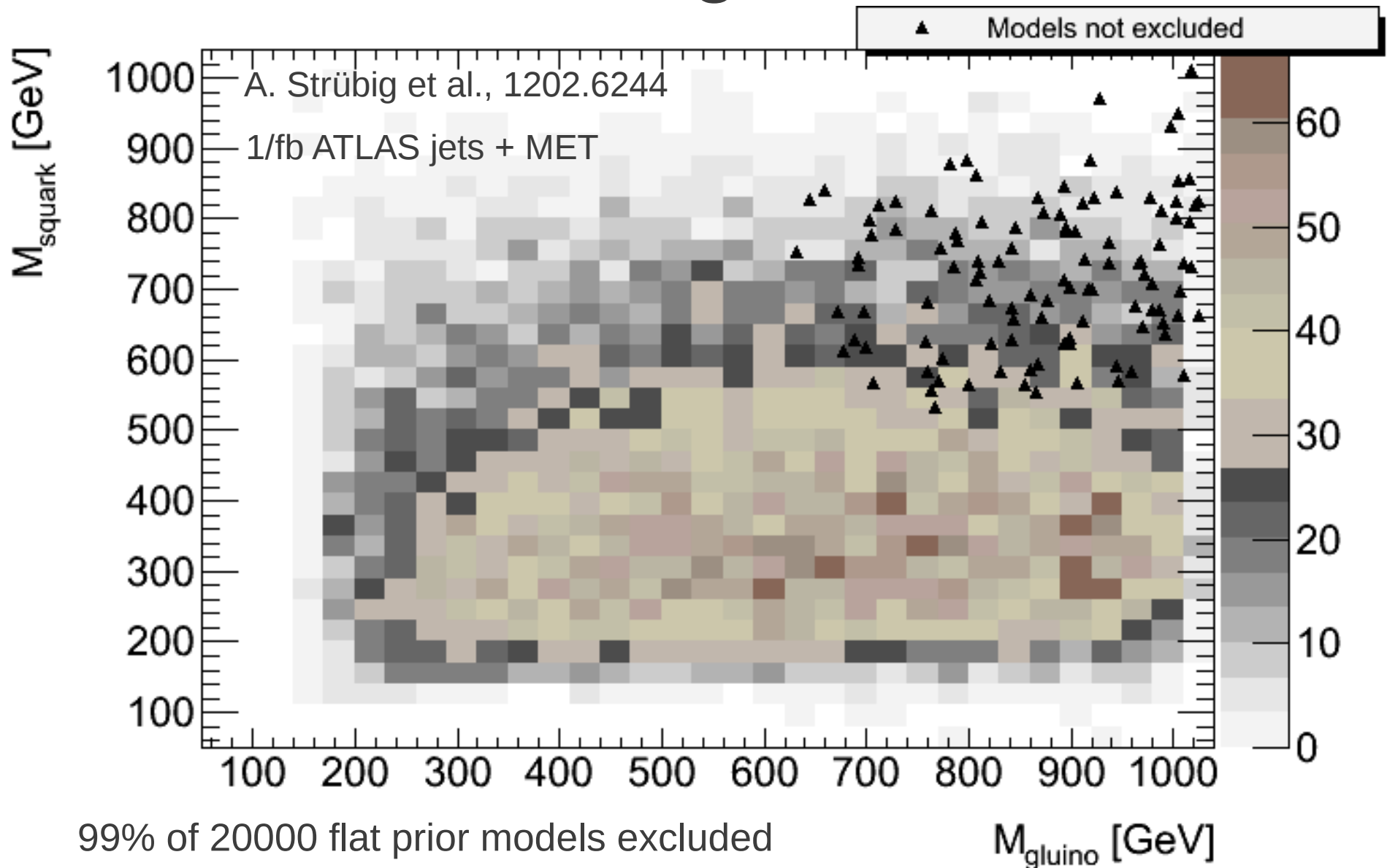
- Together, these assumptions leave us with the 19 free parameters of the *phenomenological MSSM*
- $M_1, M_2, M_3, \mu, \tan \beta, M_A, q_{1,3}, u_{1,3}, d_{1,3}, l_{1,3}, e_{1,3}, A_{t,b,\tau}$
- Can also add gravitino, with mass  $m_{3/2}$
- Generate random points in this parameter space, and test vs. experimental constraints
- Investigate properties of resulting models

# Parameter scan ranges

- $50 \text{ GeV} \leq |M_1| \leq 4 \text{ TeV}$
- $100 \text{ GeV} \leq |M_2, \mu| \leq 4 \text{ TeV}$  Compare with Berger et al.
- $400 \text{ GeV} \leq M_3 \leq 4 \text{ TeV}$
- $50 \text{ GeV} \leq |M_{1,2}, \mu| \leq 1 \text{ TeV}$
- $1 \leq \tan \beta \leq 60$
- $100 \text{ GeV} \leq M_3 \leq 1 \text{ TeV}$
- $100 \text{ GeV} \leq M_A, l, e \leq 4 \text{ TeV}$
- $1 \leq \tan \beta \leq 50$
- $400 \text{ GeV} \leq q_1, u_1, d_1 \leq 4 \text{ TeV}$
- $43.5 \text{ GeV} \leq M_A \leq 1 \text{ TeV}$
- $200 \text{ GeV} \leq q_3, u_3, d_3 \leq 4 \text{ TeV}$
- $100 \text{ GeV} \leq q, u, d, l, e \leq 1 \text{ TeV}$
- $|A_{t,b,t}| \leq 4 \text{ TeV}$
- $|A_{t,b,\tau}| \leq 1 \text{ TeV}$
- $1 \text{ eV} \leq m_{3/2} \leq 1 \text{ TeV}$  (log prior)



# Model set generation



Most models from old scan now ruled out!

# Model set generation

- Two separate scans
- Neutralino LSP: generate spectra for  $3 \cdot 10^6$  points in 19 dimensional parameter space, requiring lightest neutralino to be LSP
- Gravitino LSP: add gravitino mass and scan over 20 dimensional space using  $7 \cdot 10^5$  points, with gravitino as LSP
- Spectra are generated with SOFTSUSY and SuSpect, and tossed if there are problems (tachyons, color/charge breaking minima, unbounded scalar potentials) or the generators disagree significantly
- Decay tables are calculated with modified versions of SDECAY, HDECAY, MadGraph, and CalcHEP

# Model set generation

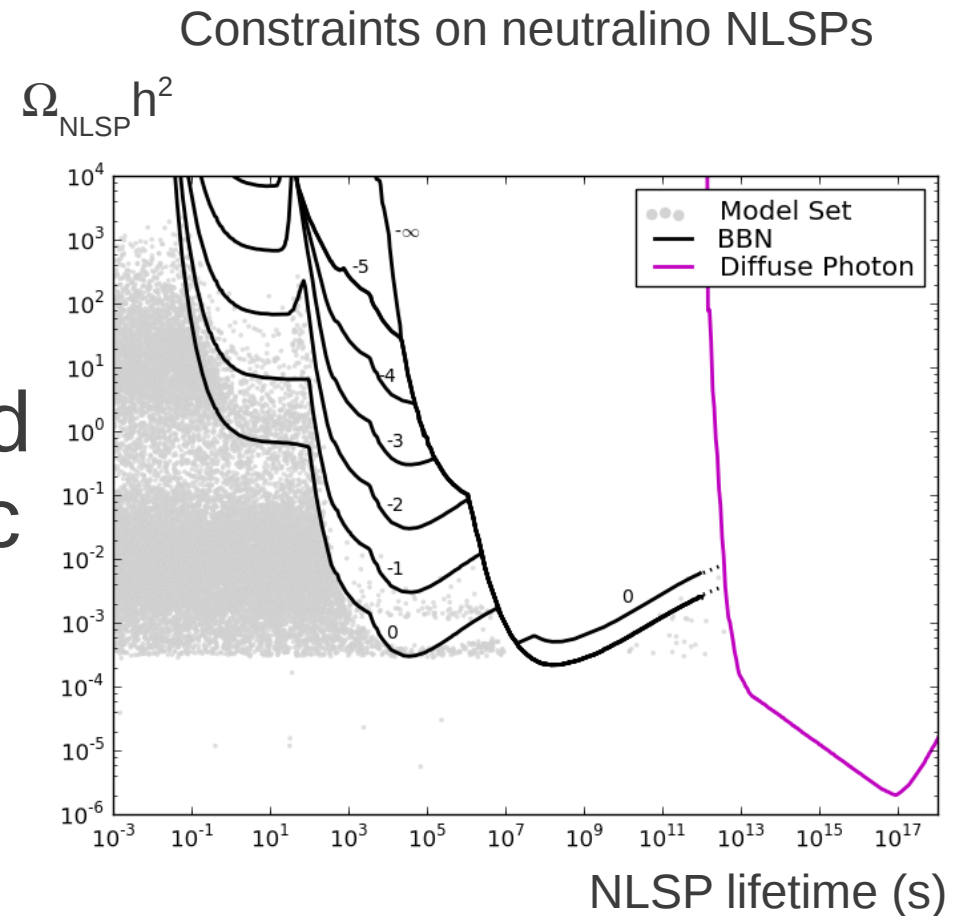
- Neutralino LSP set: impose WMAP as upper bound on thermal relic density of lightest neutralino, and check against DM direct detection constraints
- Gravitino LSP set: assume the NLSP is quasi-stable and reaches its relic density, decaying to the gravitino after freezeout; impose WMAP, cosmological constraints
- Precision EW constraints:  $g - 2$ , invisible width of Z,  $\Delta\rho$
- Flavor constraints:  $b \rightarrow s\gamma$ ,  $B_s \rightarrow \mu\mu$ ,  $B \rightarrow \tau\nu$
- Require all charged sparticles  $> 100$  GeV
- Impose LHC stable particle,  $\phi \rightarrow \tau\tau$  constraints as of 12/2011
- $2 \cdot 10^5$  models left in each set; computationally demanding!

# Gravitino LSP cosmology

- No assumptions about early universe gravitino cosmology, e.g. reheating temperature or entropy production
- NLSP freezes out later
- Assume NLSP reaches its thermal relic density, and consider out-of-equilibrium decays to gravitino
- Gravitino LSP has very weak couplings, so no dark matter detection constraints
- However, for a gravitino LSP, the NLSP can be very long-lived

# Gravitino LSP cosmology

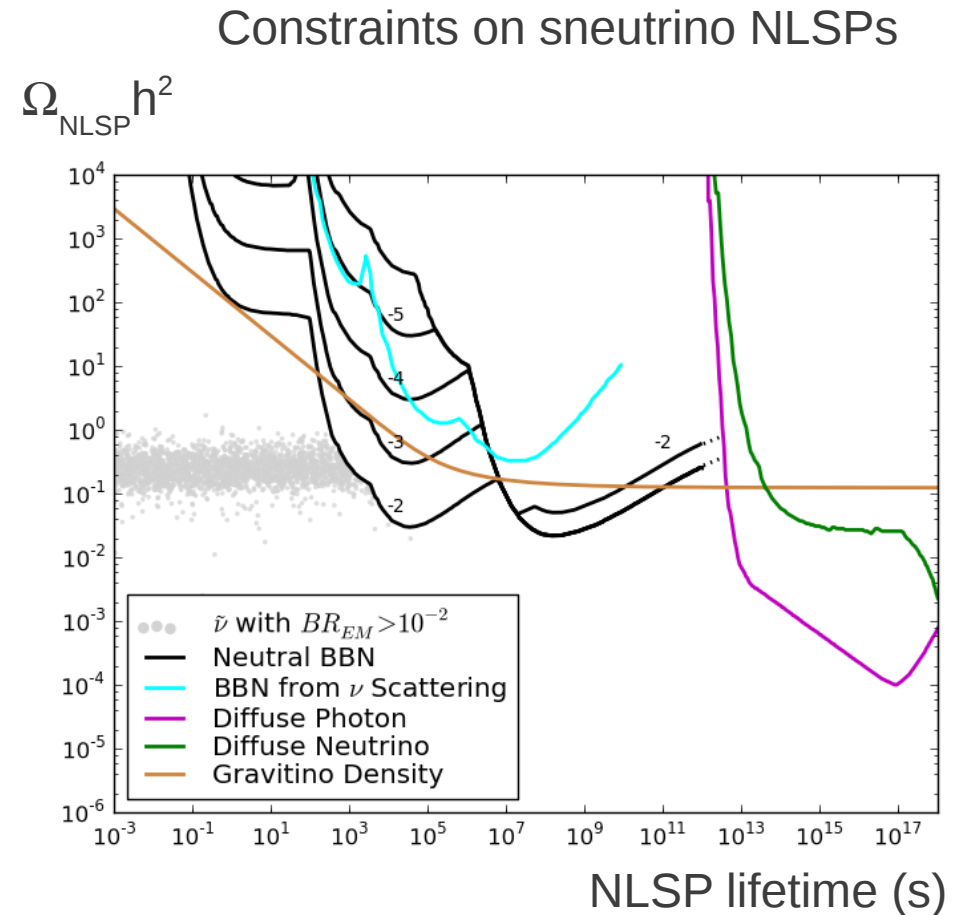
- NLSP lifetimes between  $10^{-2}$  and  $10^5$  s can affect BBN if decay products are hadronic
- For lifetimes from  $10^5$  s to  $10^{12}$  s, BBN is affected even for electromagnetic energy injection
- Diffuse photon constraints become applicable for longer lifetimes



BBN limits from Jedamzik et al., hep-ph/0604251

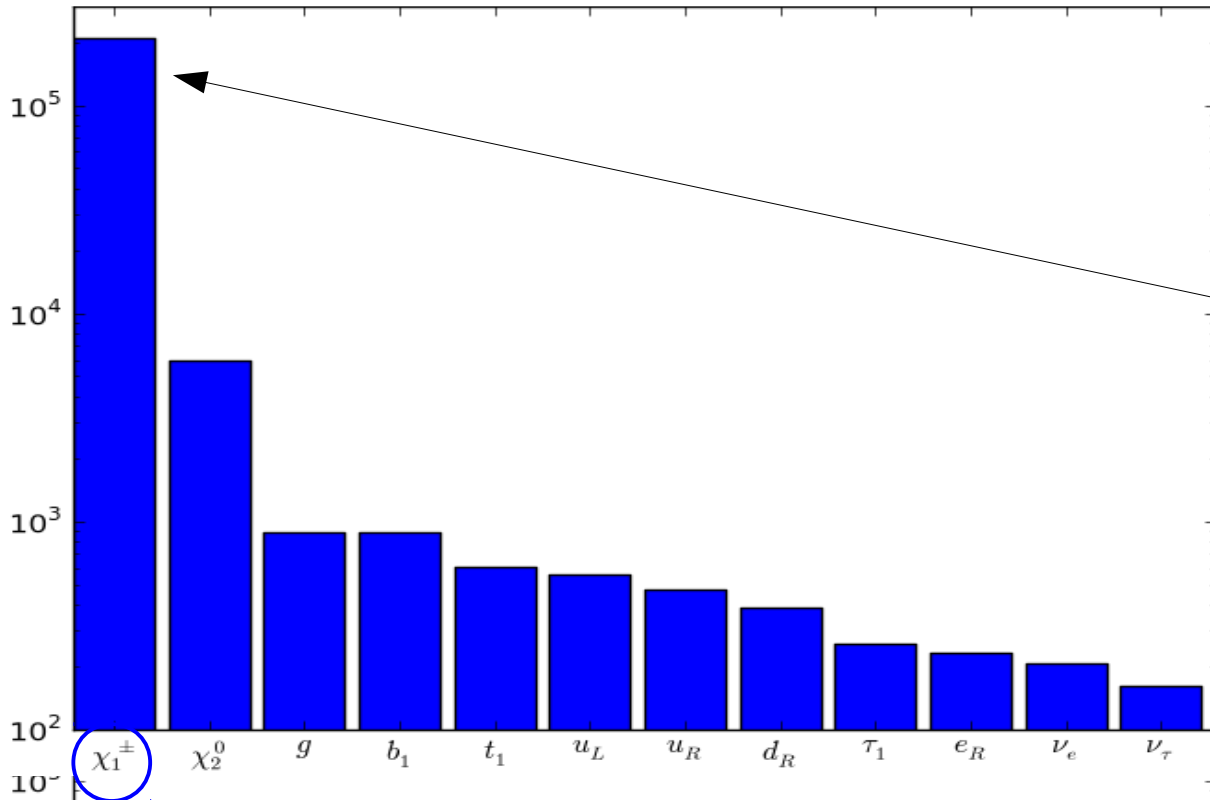
# Gravitino LSP cosmology

- Sneutrino NLSPs have small branching ratios for decays that produce visible SM particles
- Neutrinos resulting from sneutrino NLSP decays can also scatter off leptons, giving leptons/mesons that affect BBN
- Diffuse photon/neutrino flux for longer lifetimes



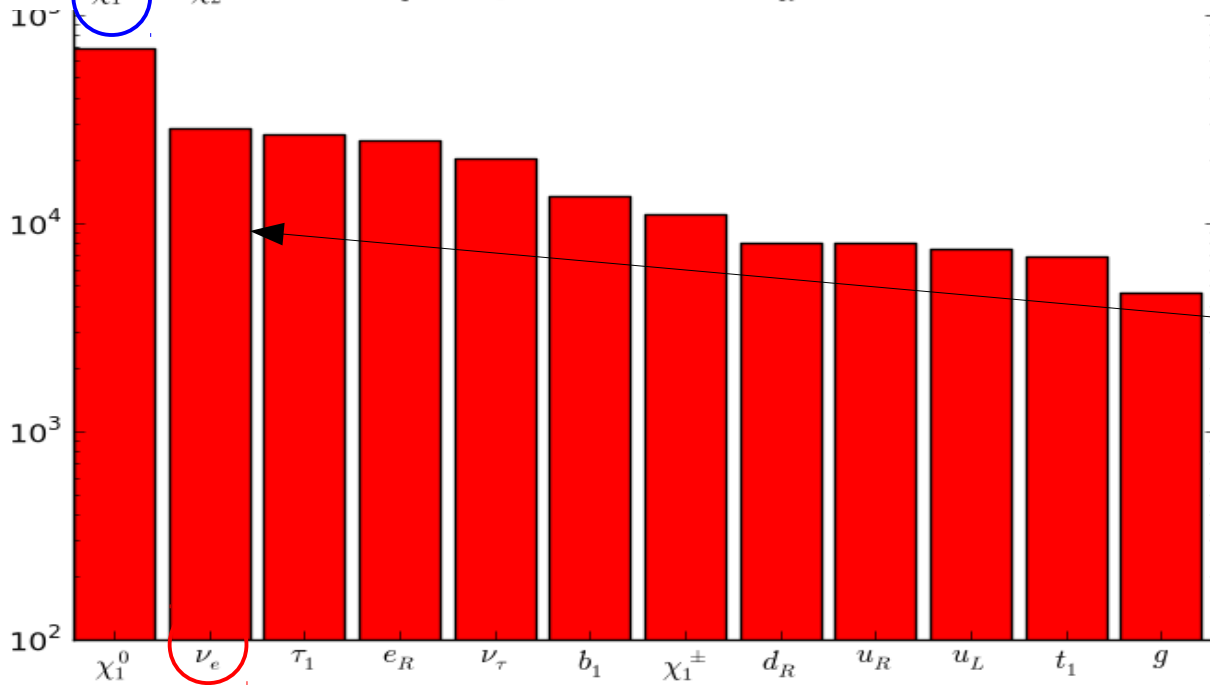
Neutrino scattering limits from  
Kanzaki et al., 0705.1200

# NLSP identity



Neutralino LSP model set has many chargino NLSPs

NLSP in neutralino LSP set



NLSP in gravitino LSP set

Sneutrino NLSPs are common in gravitino LSP model set because of lack of stable particle constraints

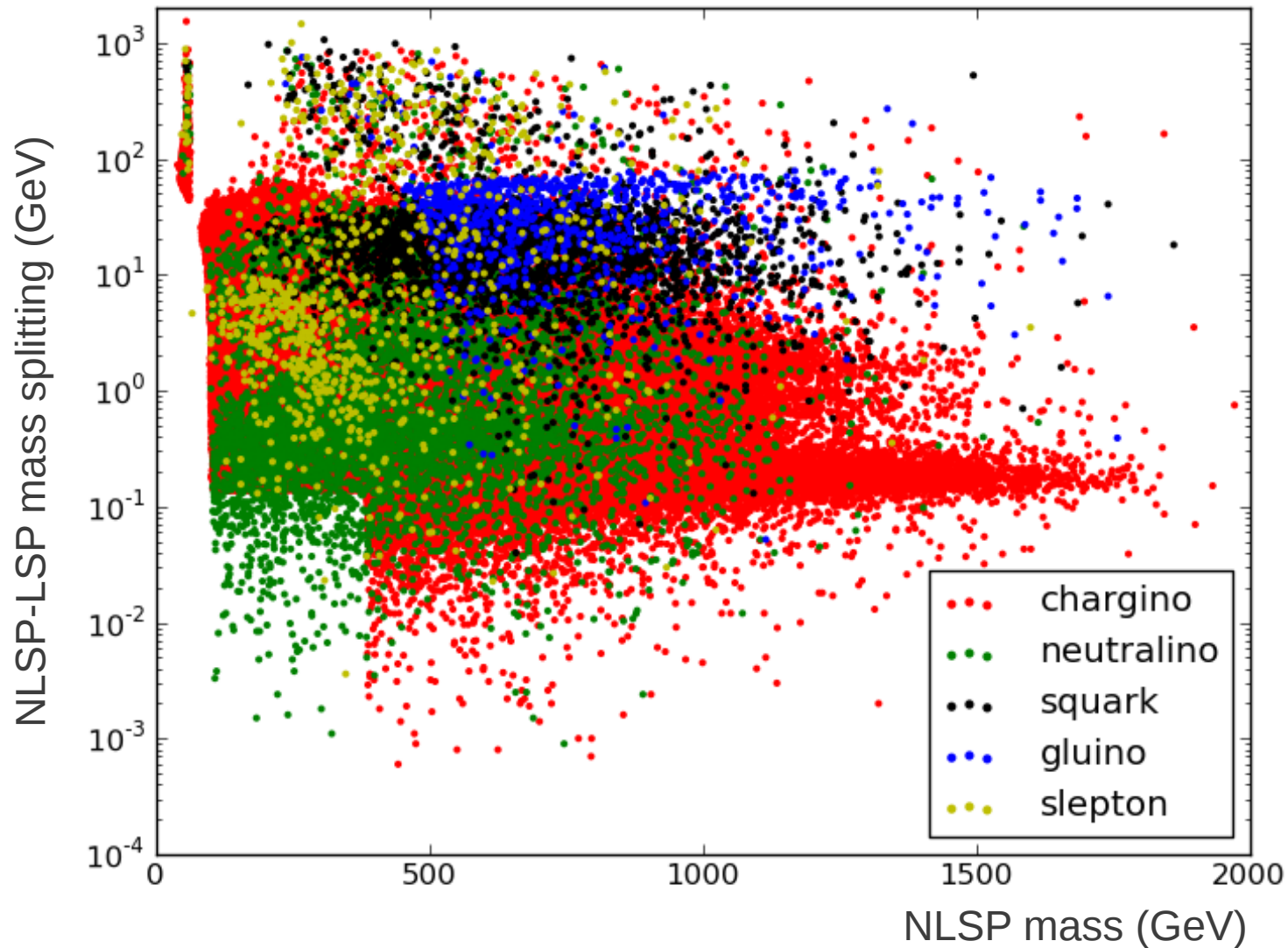
# Lightest neutralino composition

Lightest Neutralino	Definition	$\tilde{\chi}_1^0$ LSP	$\tilde{G}$ LSP
Bino	$ N_{11} ^2 > 0.95$	0.024	0.313
Mostly Bino	$0.80 <  N_{11} ^2 < 0.95$	0.002	0.012
Wino	$ N_{12} ^2 > 0.95$	0.546	0.296
Mostly Wino	$0.80 <  N_{12} ^2 < 0.95$	0.022	0.019
Higgsino	$ N_{13} ^2 +  N_{14} ^2 > 0.95$	0.340	0.296
Mostly Higgsino	$0.80 <  N_{13} ^2 +  N_{14} ^2 < 0.95$	0.029	0.029
All other models	$ N_{11} ^2,  N_{12} ^2,  N_{13} ^2 +  N_{14} ^2 < 0.80$	0.036	0.035

Bino LSPs tend to give high relic densities in neutralino LSP model set  
 In gravitino LSP model set, lightest neutralino does not make up DM

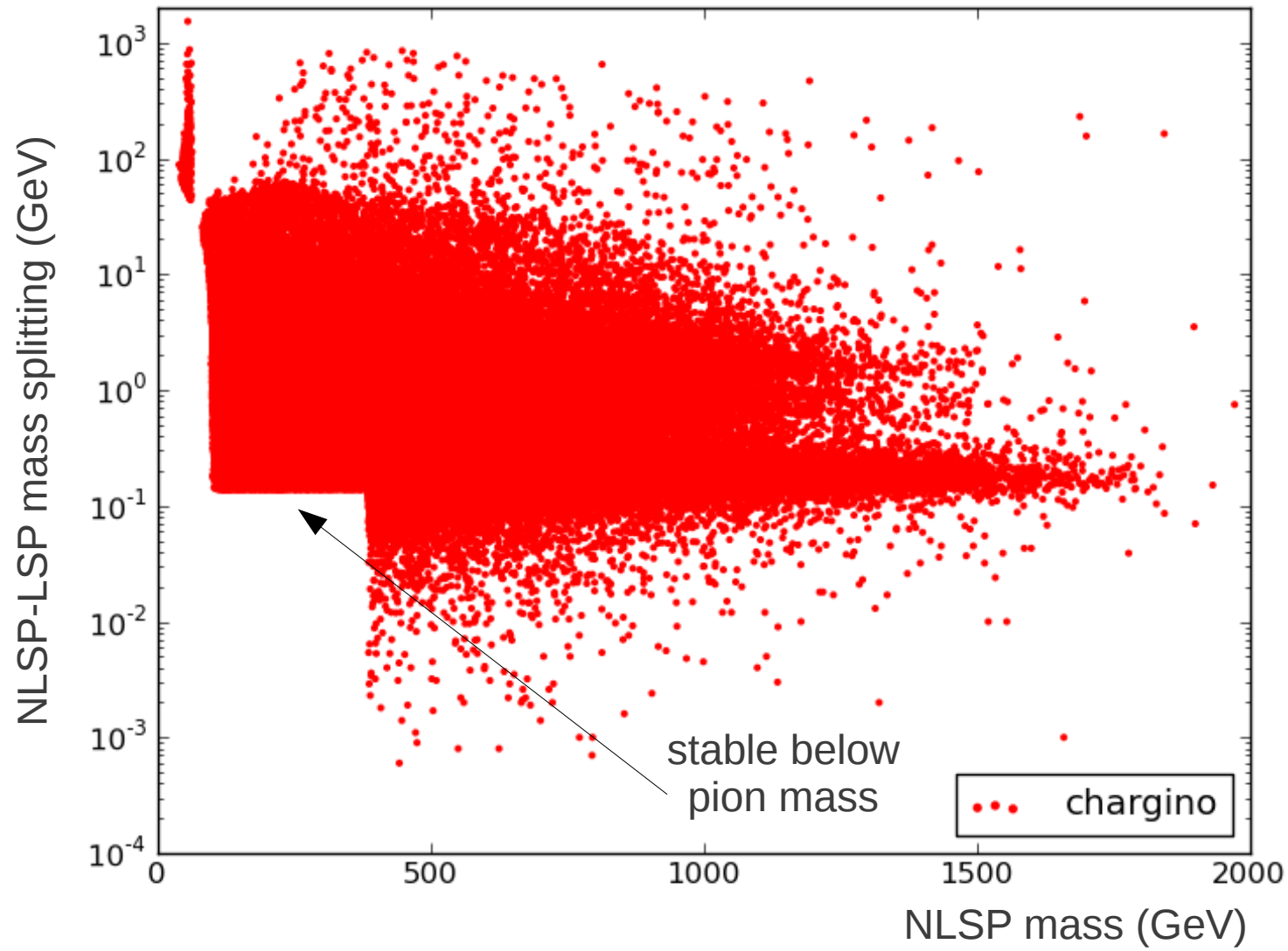


# Stable particles



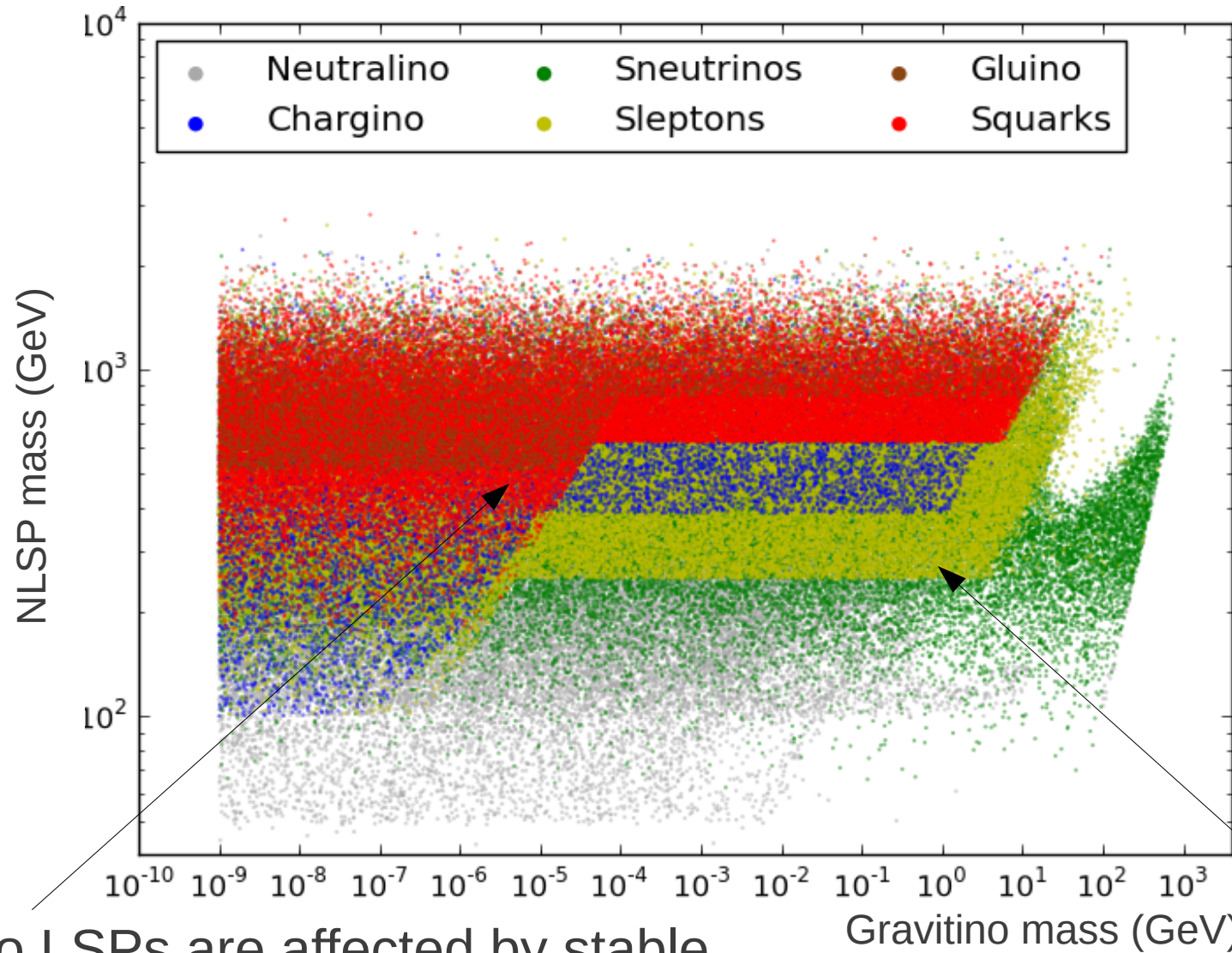
Neutralino LSP models can have stable charginos if LSP is wino

# Stable particles



Neutralino LSP models can have stable charginos if LSP is wino

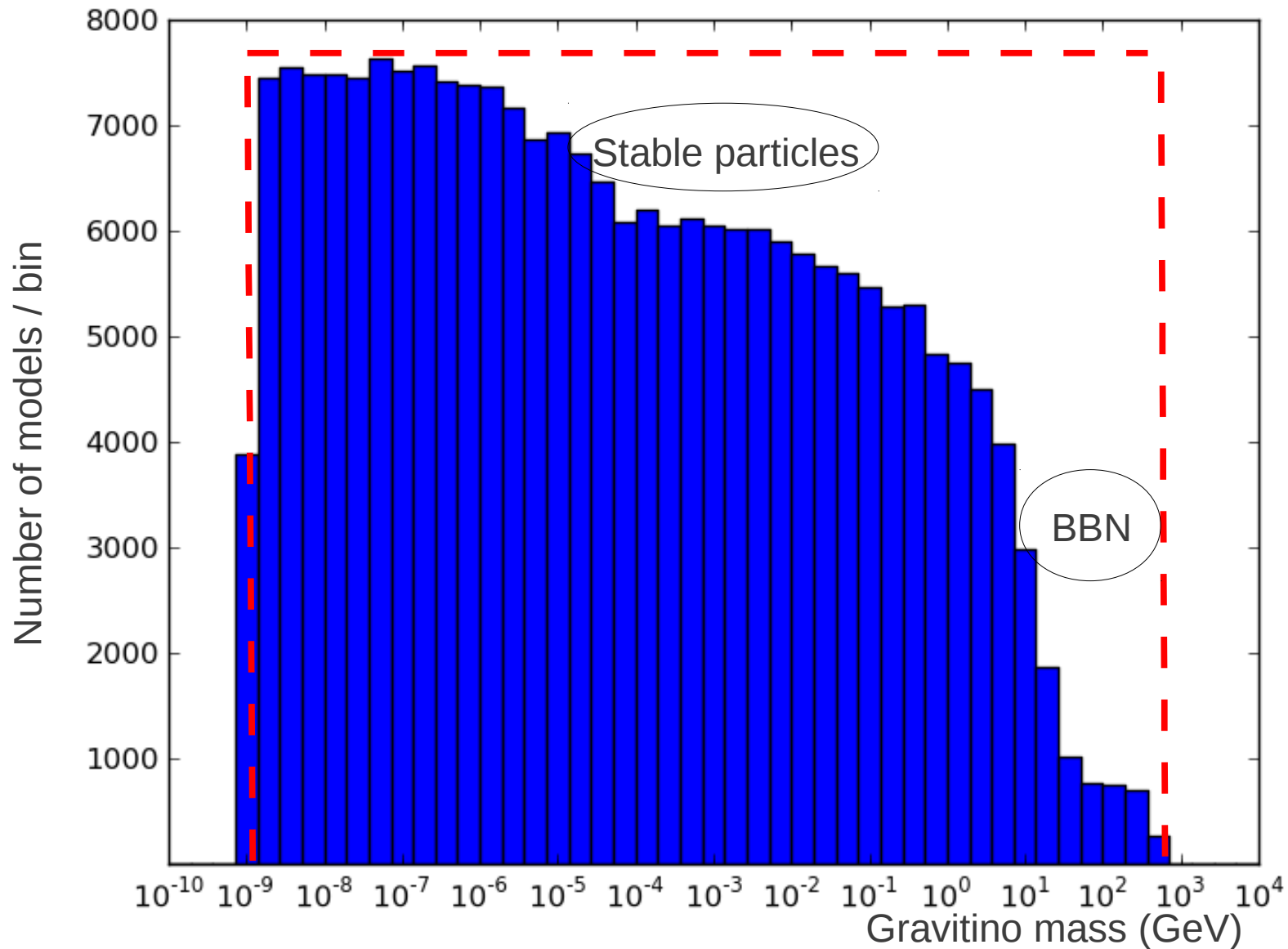
# Stable particles



Gravitino LSPs are affected by stable charged particle searches too

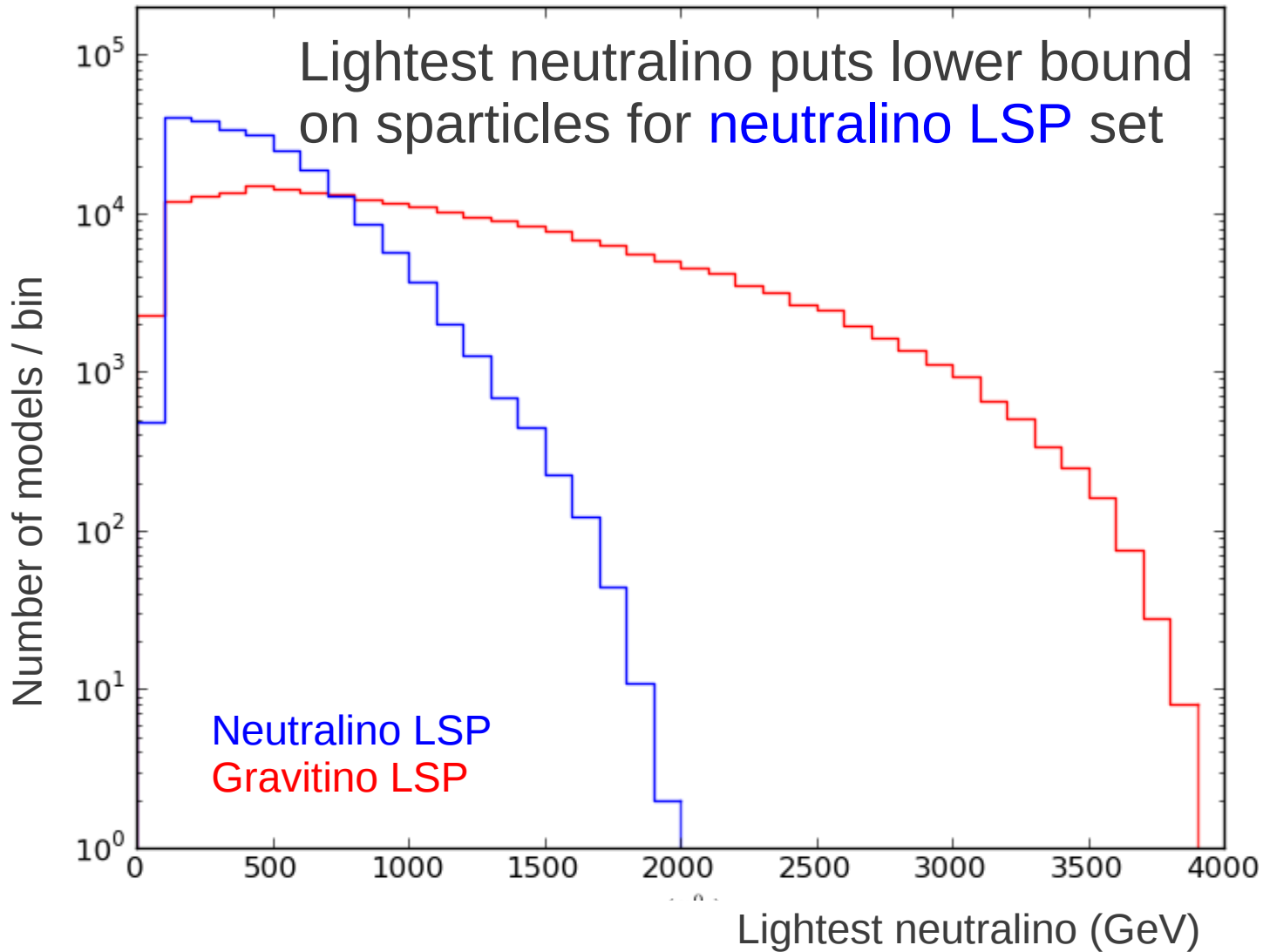
When the NLSP is very long-lived, cosmological constraints come in

# Gravitino LSP mass distribution



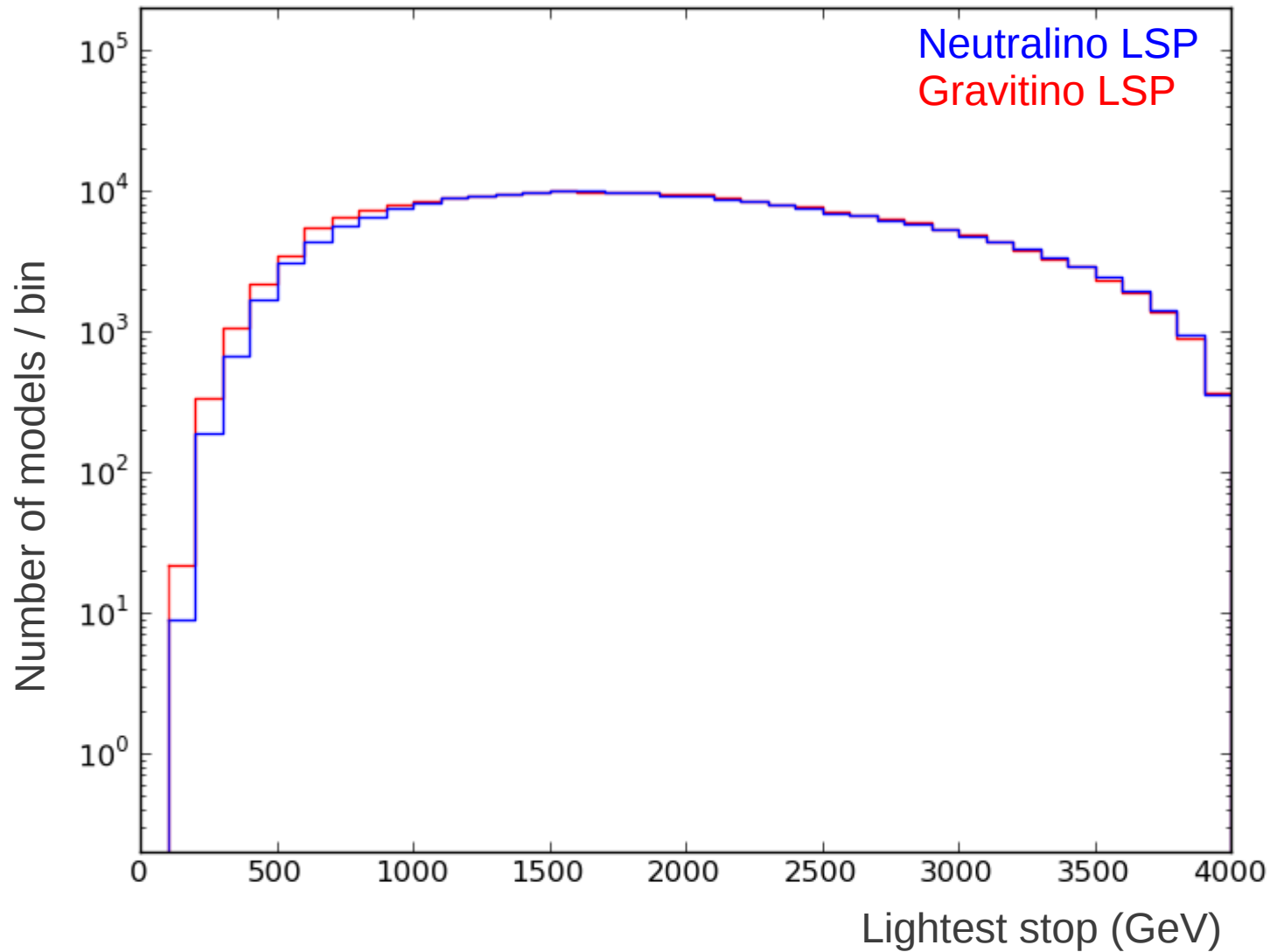
Resulting mass distribution is peaked towards lighter gravitinos

# Model set comparison



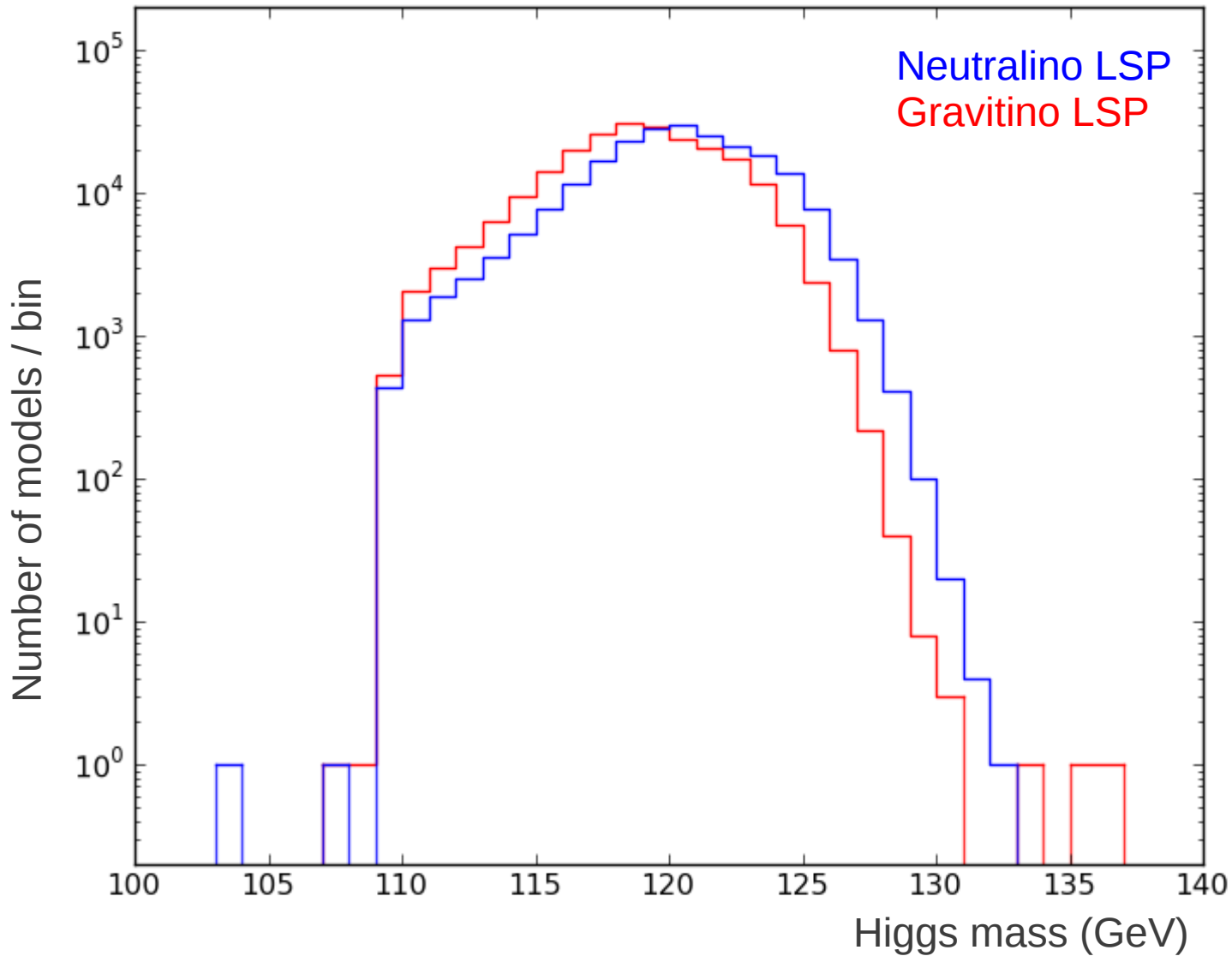
For **gravitino LSP** set, there is no such bound from the lightest neutralino mass, so sparticles end up lighter

# Model set comparison



Stops are pushed up by this lower bound in the **neutralino LSP** set

# Model set comparison



Neutralino LSP set gets heavier Higgses from heavier stops, on average

# Outline

- The phenomenological MSSM
- **LHC search results**
- Higgs discovery implications
- Fine-tuning
- Outlook



# LHC searches

- Generate SUSY events for each of our models with PYTHIA, scale to NLO with Prospino, pass through PGS
  - Analysis suite based on code from previous scans (Conley et al., 1009.2539, 1103.1697)
  - 7 TeV: ATLAS 5/fb (leptons +) jets + MET, stop/sbottom, disappearing tracks; CMS HSCP,  $\phi \rightarrow \tau\tau$ ; LHCb  $B_s \rightarrow \mu\mu$
  - 8 TeV: ATLAS 6/fb (leptons +) jets + MET\*
- \*neutralino LSP set only for now

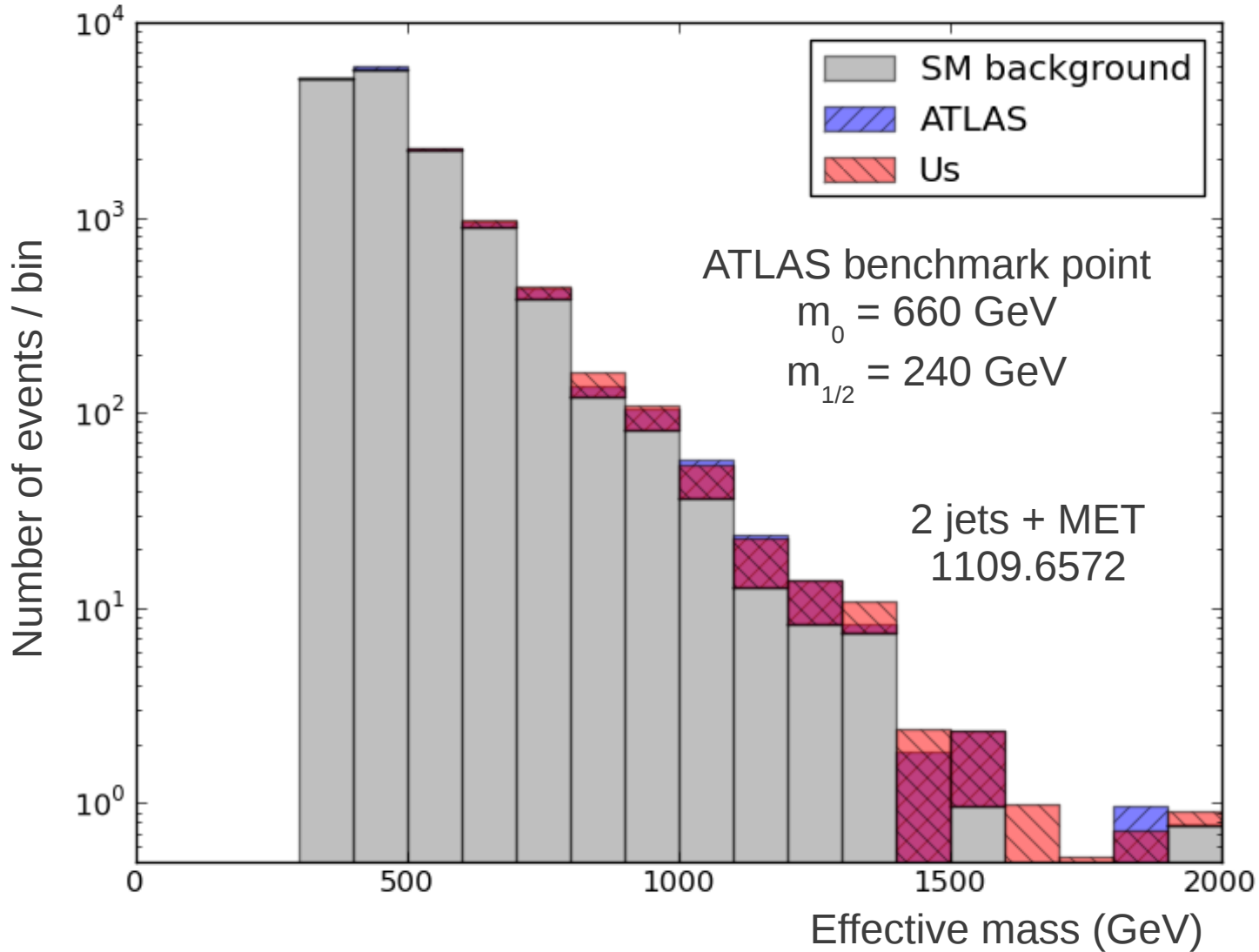
# LHC searches

	Short Title of the CONF note	Date	$\sqrt{s}$ (TeV)	$L$ ( $\text{fb}^{-1}$ )	Document	Plots
✓	0 leptons + $\geq 2$ -6 jets + $E_{\text{miss}}$	08/2012	8	5.8	<a href="#">ATLAS-CONF-2012-109</a>	<a href="#">Link</a>
✓	0 leptons + $\geq 6$ -9 jets + $E_{\text{miss}}$	08/2012	8	5.8	<a href="#">ATLAS-CONF-2012-103</a>	<a href="#">Link</a>
✓	1 lepton + $\geq 4$ jets + $E_{\text{miss}}$	08/2012	8	5.8	<a href="#">ATLAS-CONF-2012-104</a>	<a href="#">Link</a>
✓	2 same-sign leptons + $\geq 4$ jets + $E_{\text{miss}}$	08/2012	8	5.8	<a href="#">ATLAS-CONF-2012-105</a>	<a href="#">Link</a>

## 2011 Data (7 TeV)

	Short Title of the Paper	Date	$\sqrt{s}$ (TeV)	$L$ ( $\text{fb}^{-1}$ )	Document	Plots+Aux. Material	Journal
	Monophoton [ADD, WIMP] <b>NEW</b>	09/2012	7	4.7	<a href="#">1209.4625</a>	<a href="#">Link</a>	Submitted to PRL
✓	2 leptons + jets + $E_{\text{miss}}$ [Medium stop] <b>NEW</b>	09/2012	7	4.7	<a href="#">1209.4186</a>	<a href="#">Link</a>	Submitted to JHEP
	1-2 b-jets + 1-2 leptons + jets + $E_{\text{miss}}$ [Light Stop] <b>NEW</b>	09/2012	7	4.7	<a href="#">1209.2102</a>	<a href="#">Link</a>	Submitted to PLB
	2 photons + $E_{\text{miss}}$ [GGM] <b>NEW</b>	09/2012	7	4.7	<a href="#">1209.0753</a>	<a href="#">Link</a>	Submitted to PLB
✓	1-2 leptons + $\geq 2$ -4 jets + $E_{\text{miss}}$	08/2012	7	4.7	<a href="#">1208.4688</a>	<a href="#">Link</a>	Accepted by PRD
✓	2 leptons + $\geq 1$ jet + $E_{\text{miss}}$ [Very light stop]	08/2012	7	4.7	<a href="#">1208.4305</a>	<a href="#">Link (inc. HEPData)</a>	Submitted to EPJC
✓	3 leptons + $E_{\text{miss}}$ [Direct gauginos]	08/2012	7	4.7	<a href="#">1208.3144</a>	<a href="#">Link (inc. HEPData)</a>	Submitted to PLB
✓	2 leptons + $E_{\text{miss}}$ [Direct gauginos/sleptons]	08/2012	7	4.7	<a href="#">1208.2884</a>	<a href="#">Link</a>	Submitted to PLB
✓	1 lepton + $\geq 4$ jets ( $\geq 1$ b-jet) + $E_{\text{miss}}$ [Heavy stop]	08/2012	7	4.7	<a href="#">1208.2590</a>	<a href="#">Link</a>	Accepted by PRL
✓	0 lepton + 1-2 b-jet + 5-4 jets + $E_{\text{miss}}$ [Heavy stop]	08/2012	7	4.7	<a href="#">1208.1447</a>	<a href="#">Link</a>	Accepted by PRL
✓	0 lepton + $\geq 2$ -6 jets + $E_{\text{miss}}$	08/2012	7	4.7	<a href="#">1208.0949</a>	<a href="#">Link</a>	Submitted to PRD
✓	0 lepton + $\geq 3$ b-jets + $\geq (1-3)$ jets + $E_{\text{miss}}$ [Gluino med. stop/sb.]	07/2012	7	4.7	<a href="#">1207.4686</a>	<a href="#">Link</a>	Accepted by EPJC
✓	0 lepton + $\geq (6-9)$ jets + $E_{\text{miss}}$	06/2012	7	4.7	<a href="#">1206.1760</a>	<a href="#">Link</a>	<a href="#">JHEP 1207 (2012) 167</a>
	Electron-muon continuum [RPV]	05/2012	7	2.05	<a href="#">1205.0725</a>	<a href="#">Link (inc. HEPData)</a>	<a href="#">EPJC 72 (2012) 2040</a>
✓	Z- $\rightarrow$ ll + b-jet + jets + $E_{\text{miss}}$ [Direct stop in natural GMSB]	04/2012	7	2.05	<a href="#">1204.6736</a>	<a href="#">Link (inc. HEPData)</a>	<a href="#">PLB 715 (2012) 44</a>
	$\geq 3$ leptons + $E_{\text{miss}}$ [Direct gauginos]	04/2012	7	2.05	<a href="#">1204.5638</a>	<a href="#">Link (inc. HEPData)</a>	<a href="#">PRL 108 (2012) 261804</a>
	$\geq 1$ tau + jets + $E_{\text{miss}}$ [GMSB]	04/2012	7	2.05	<a href="#">1204.3852</a>	<a href="#">Link (inc. HEPData)</a>	<a href="#">PLB 714 (2012) 197</a>

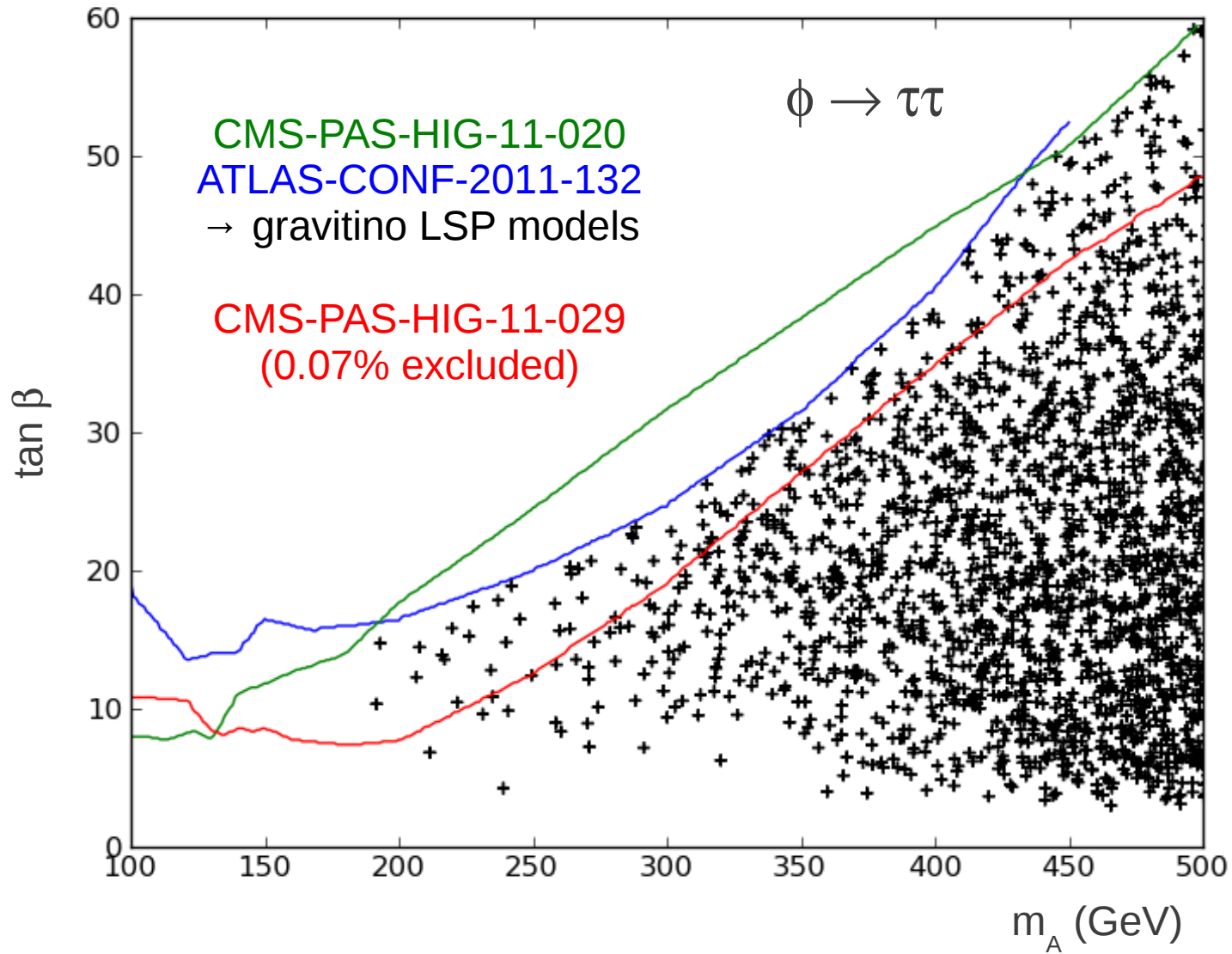
# LHC searches



Shape reproduced well

Normalization good given ~35%  
uncertainty quoted by ATLAS

# LHC searches



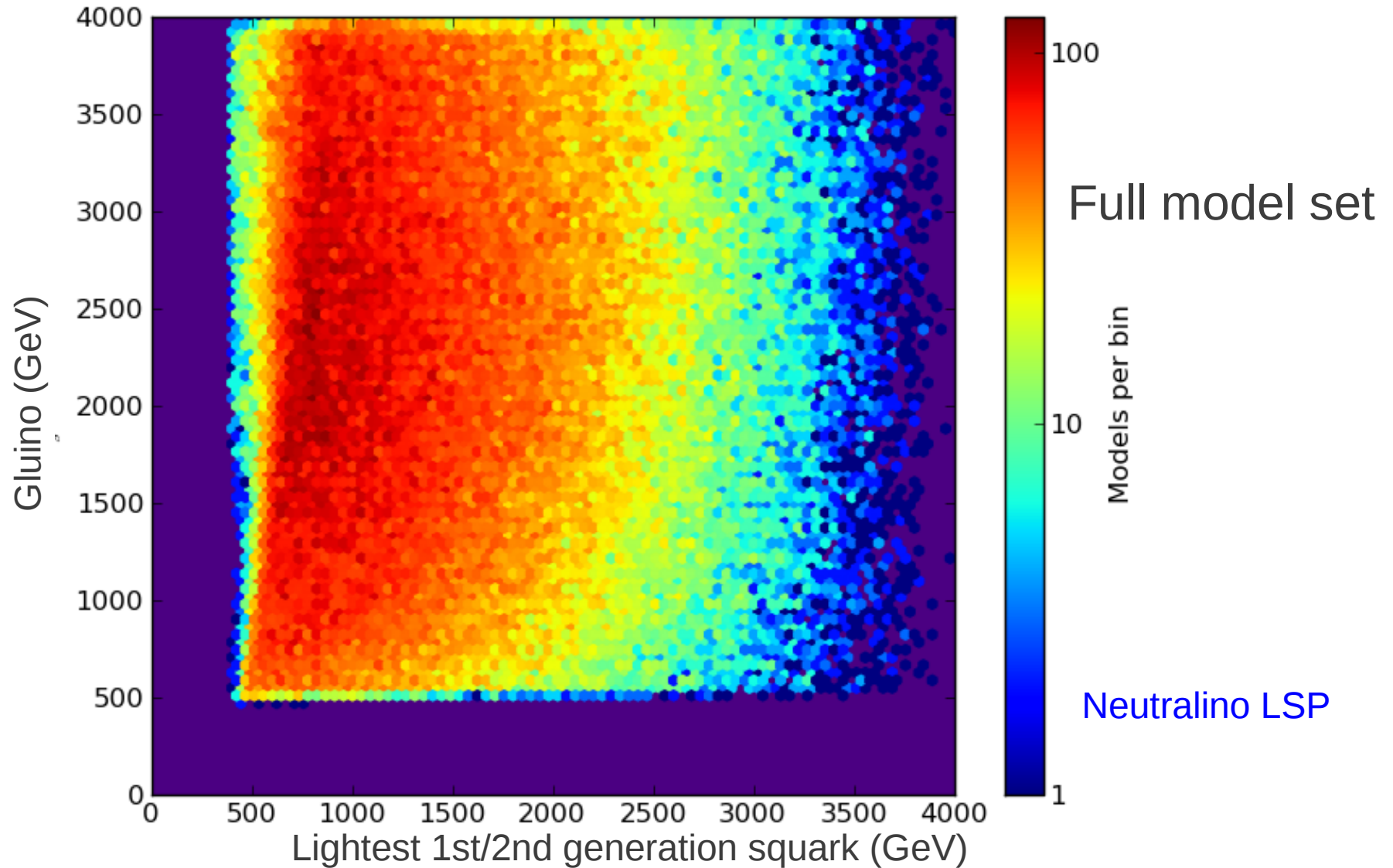
# LHC searches

Neutralino LSP

Analysis	7 TeV 4.7 fb <sup>-1</sup>	8 TeV 5.8 fb <sup>-1</sup>	8 TeV 25 fb <sup>-1</sup>
Jets + MET	21.0%	26.5%	25.3%
Many jets + MET	1.6%	3.3%	3.3%
1 $\ell$ + jets + MET	3.2%	3.3%	3.8%
SSDL	—	4.9%	7.5%
Multi-leptons	4.3%	—	—
Stop/sbottom	7.3%	—	—
HSCP	4.0%	—	—
Disappearing tracks	2.6%	—	—
$B_s \rightarrow \mu\mu, \phi \rightarrow \tau\tau$	4.2%	—	—
Remaining models	66.0%		65.2%

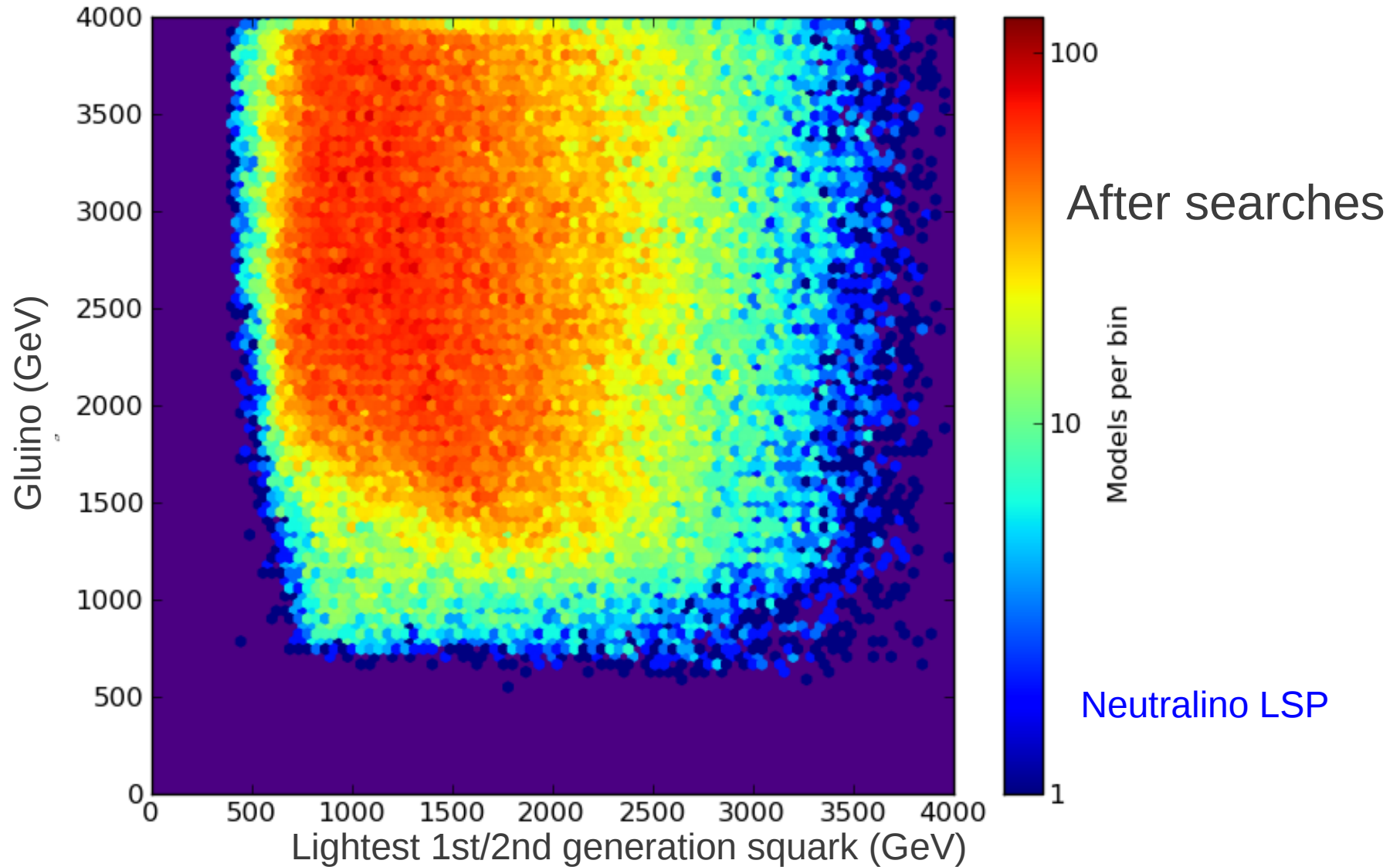
- 0.7% of the model set is excluded by the 7 TeV vanilla SUSY searches but *not* by the corresponding 8 TeV analyses (tighter cuts)
- Going to 25/fb at 8 TeV doesn't gain much!
- Fractions of models killed are ~independent of Higgs mass cut

# LHC searches



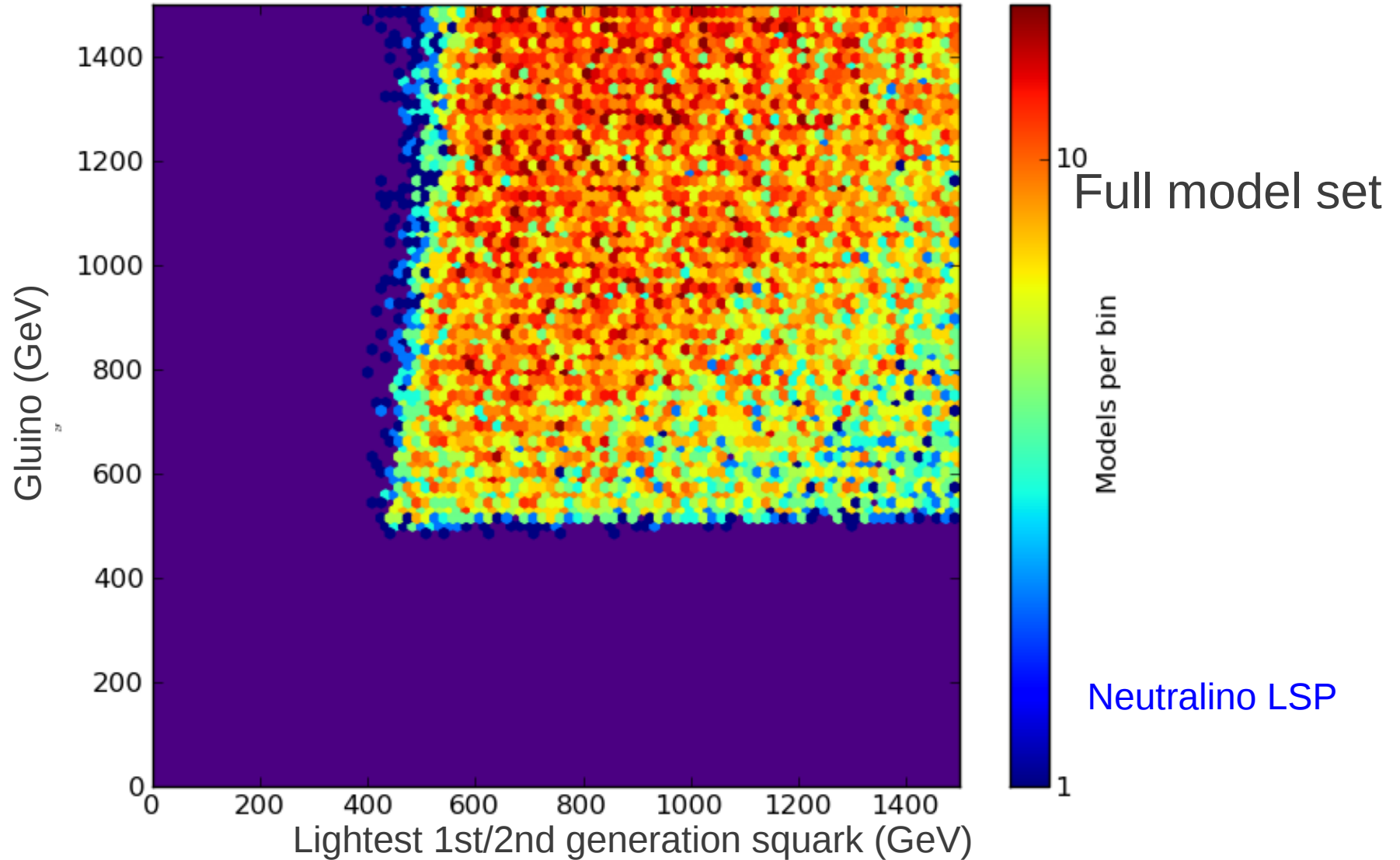
Most SUSY searches designed to exclude models on bottom/left of density plot

# LHC searches



Most SUSY searches designed to exclude models on bottom/left of density plot

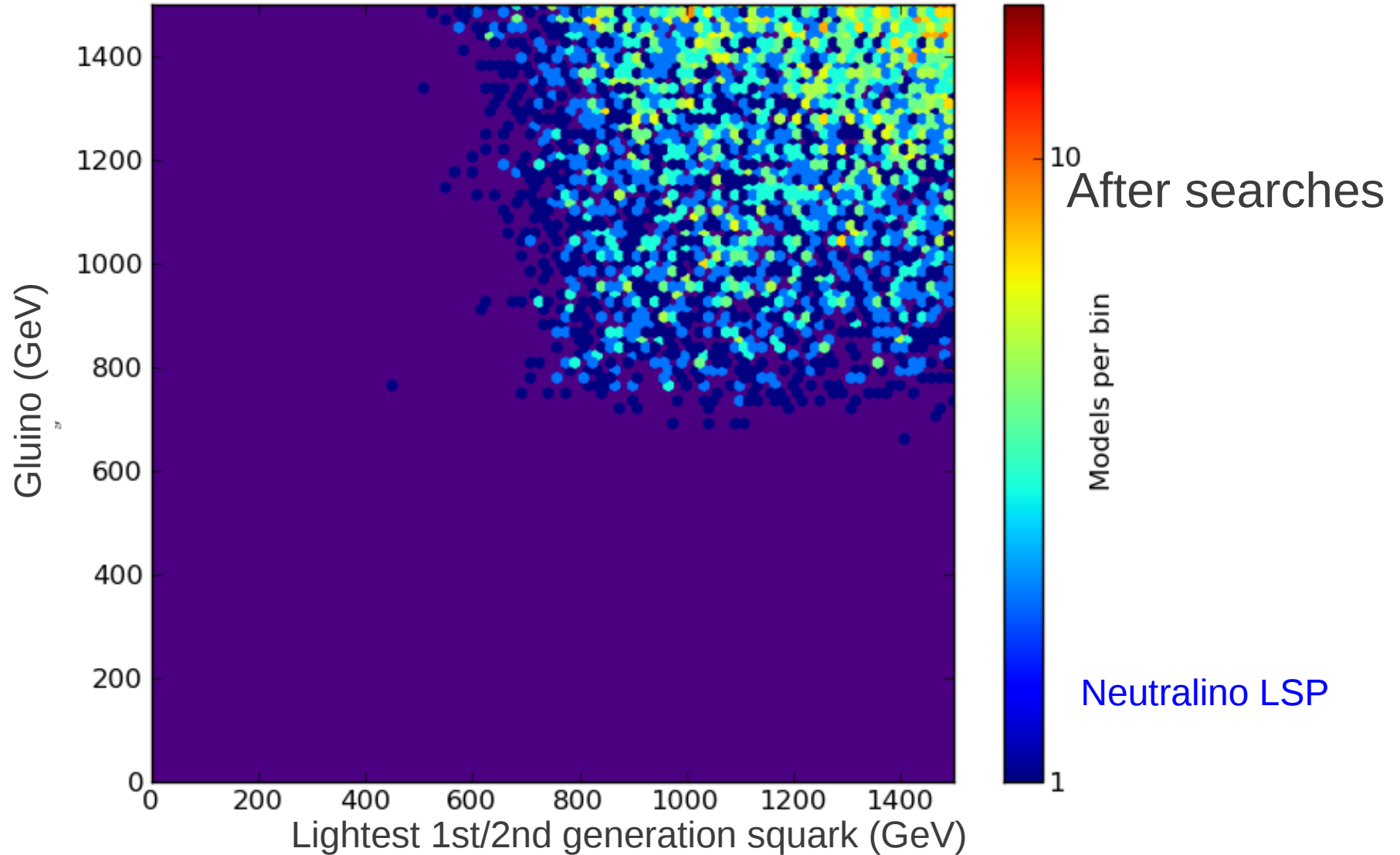
# LHC searches



Models with light squarks and gluinos are constrained,  
but can survive with, e.g., compressed spectra

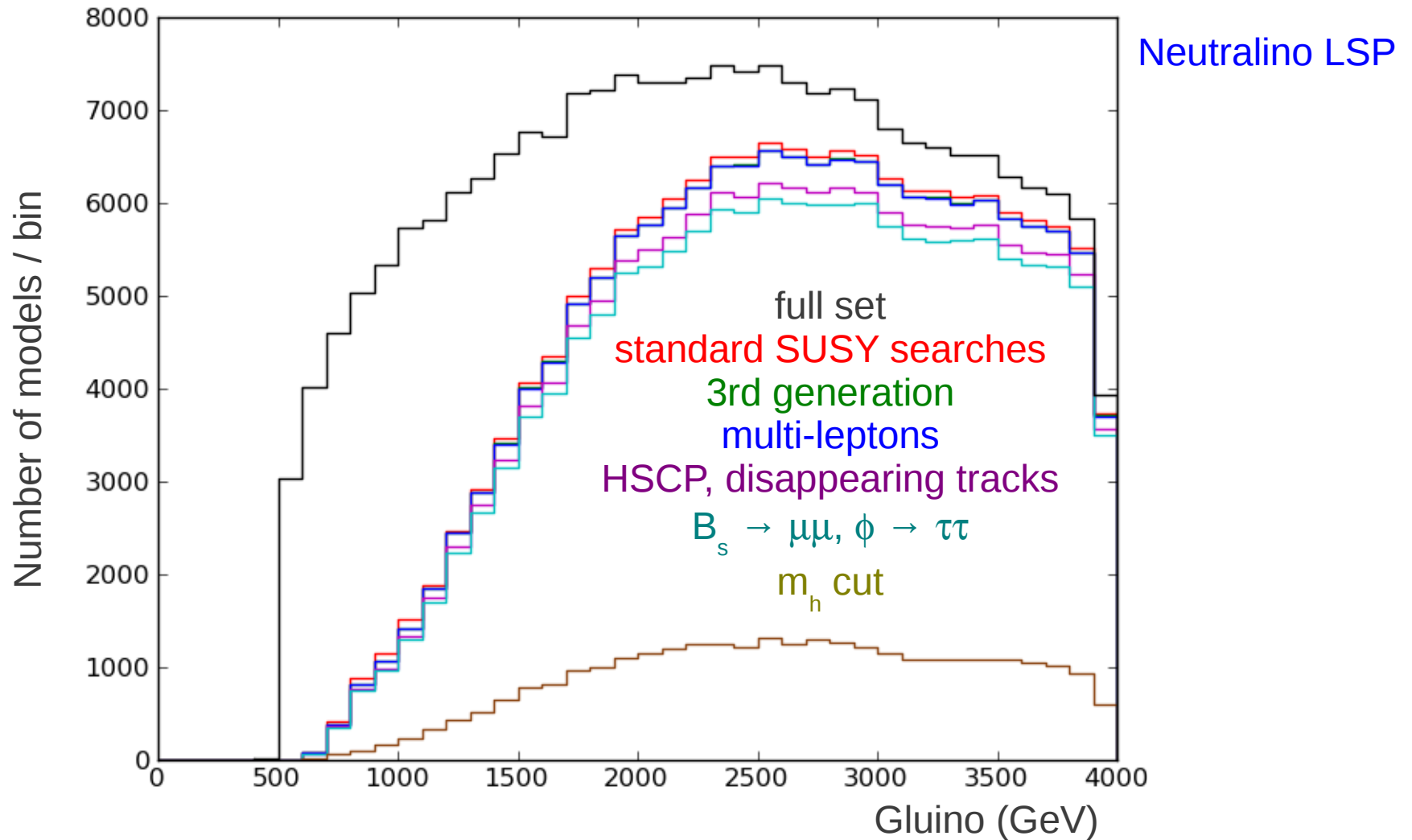


# LHC searches



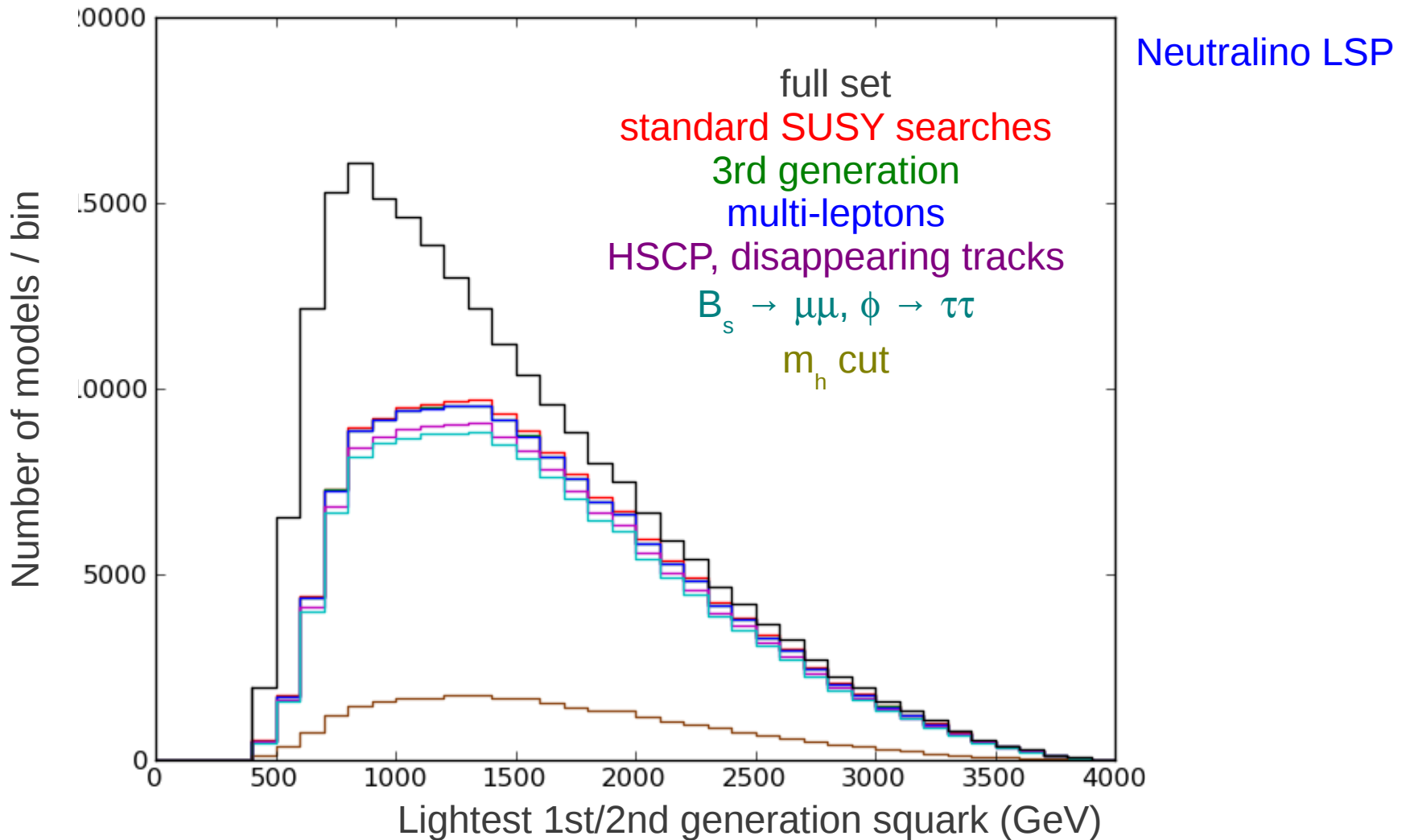
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# LHC searches



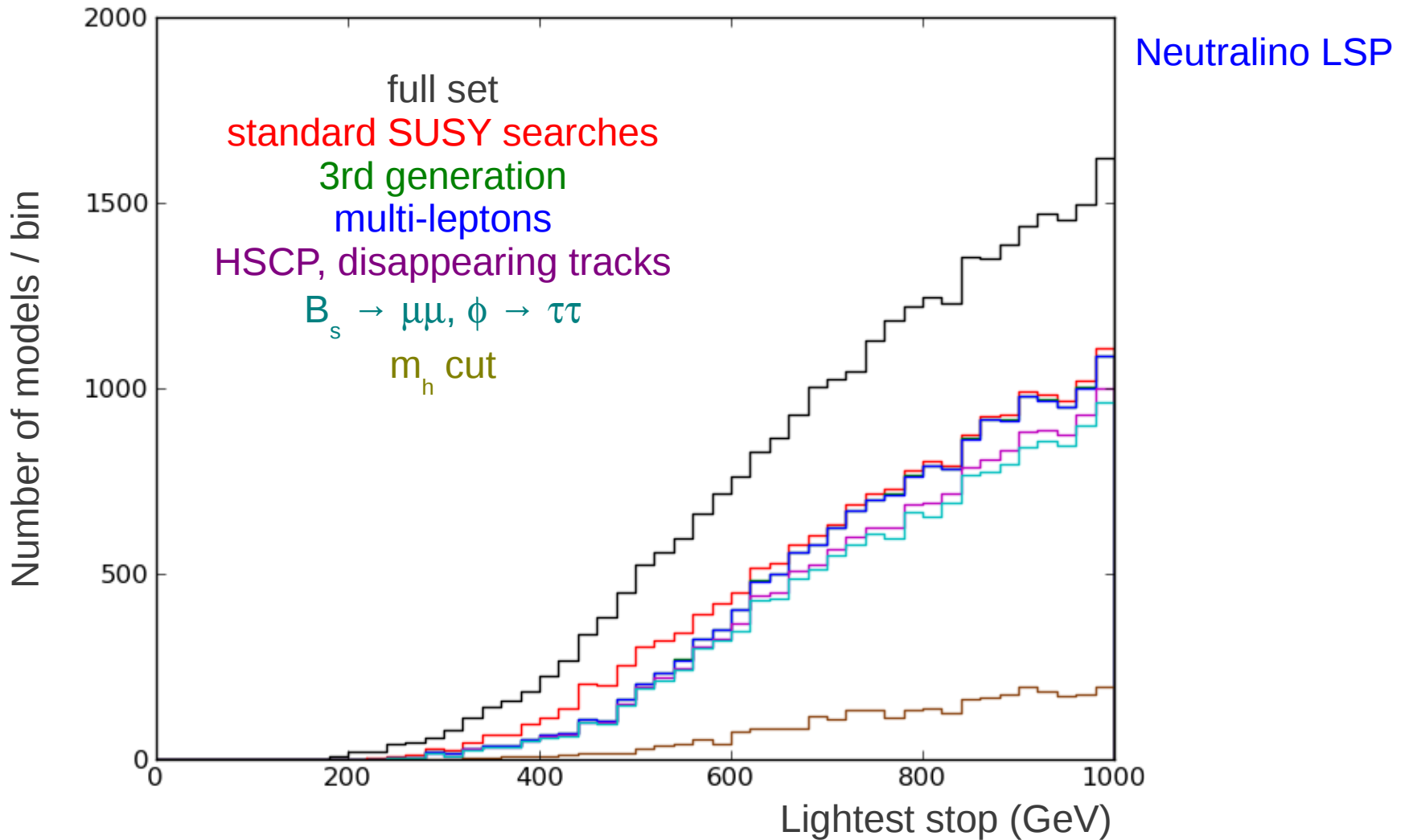
Non-MET searches are orthogonal to MET searches  
Cutting on Higgs mass affects gluino distribution

# LHC searches



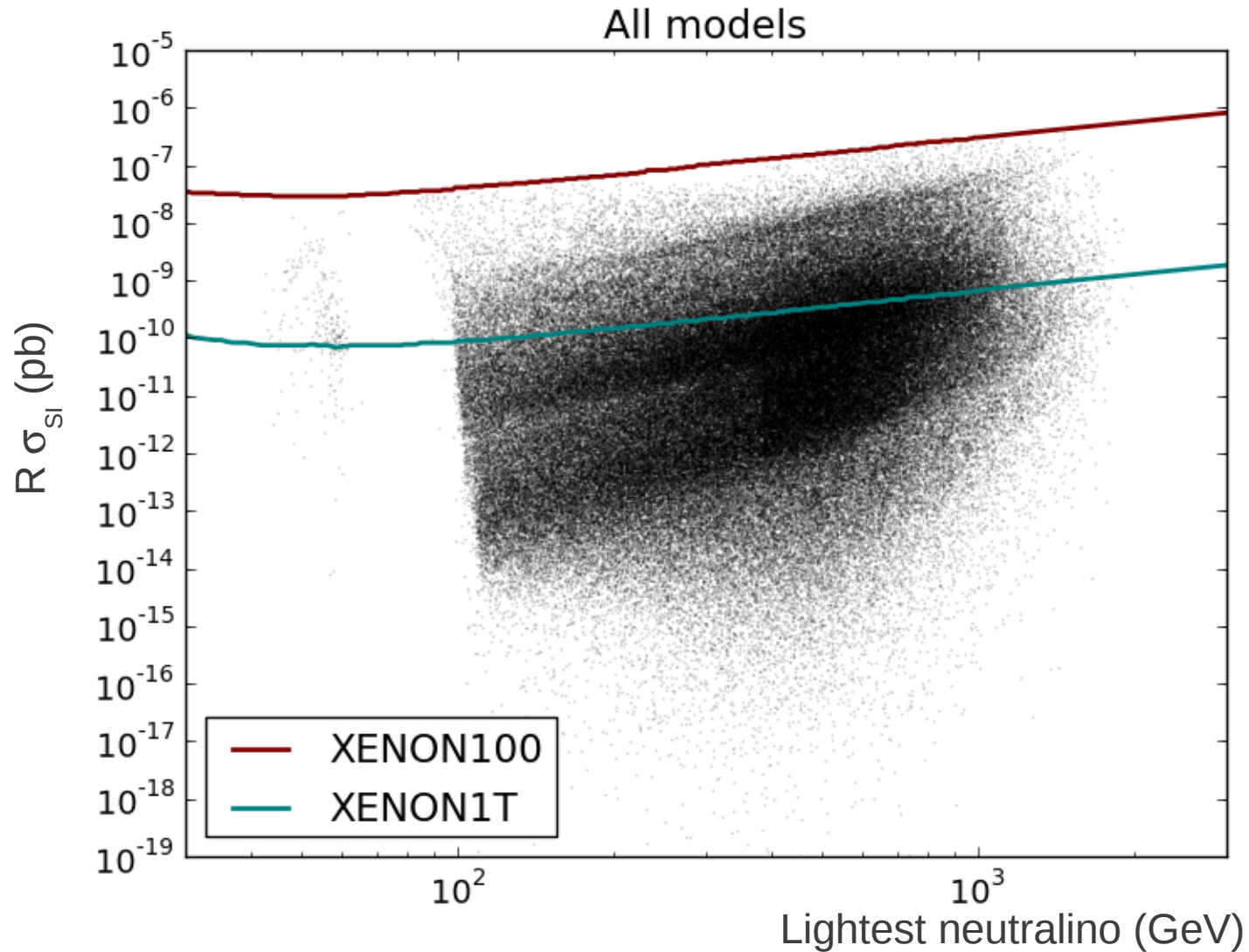
“Vanilla” SUSY searches do well at seeing light squarks, more specific searches are less successful

# LHC searches



Searches for stop/sbottoms work to some extent, but some models have tricky cascade decays

# LHC searches

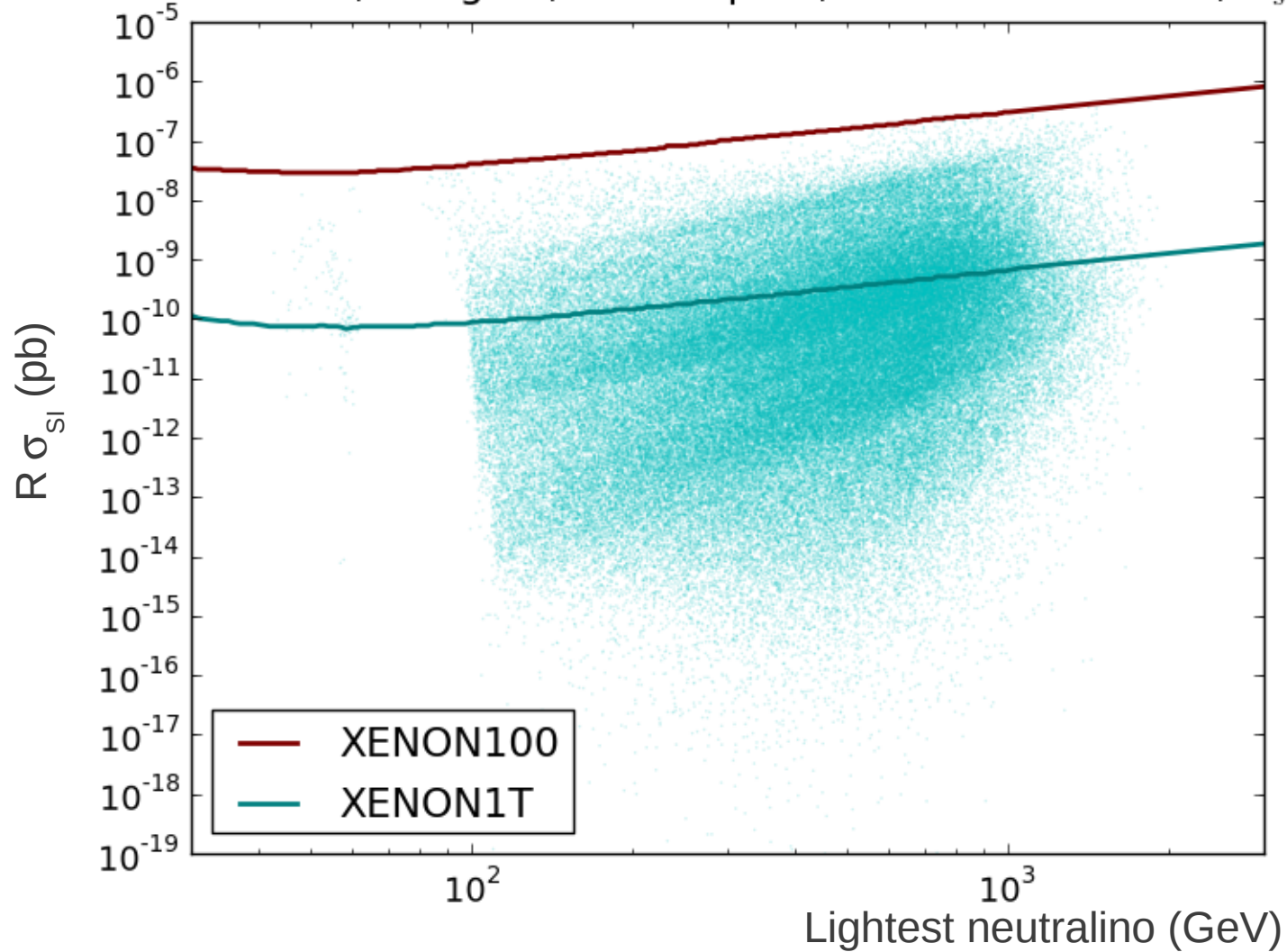


Neutralino LSP

Dark matter and LHC searches for SUSY  
complement each other

# LHC searches

Vanilla SUSY, 3rd gen., multi-lepton, HSCP/DT searches,  $B_s$ ,  $\tau\tau$

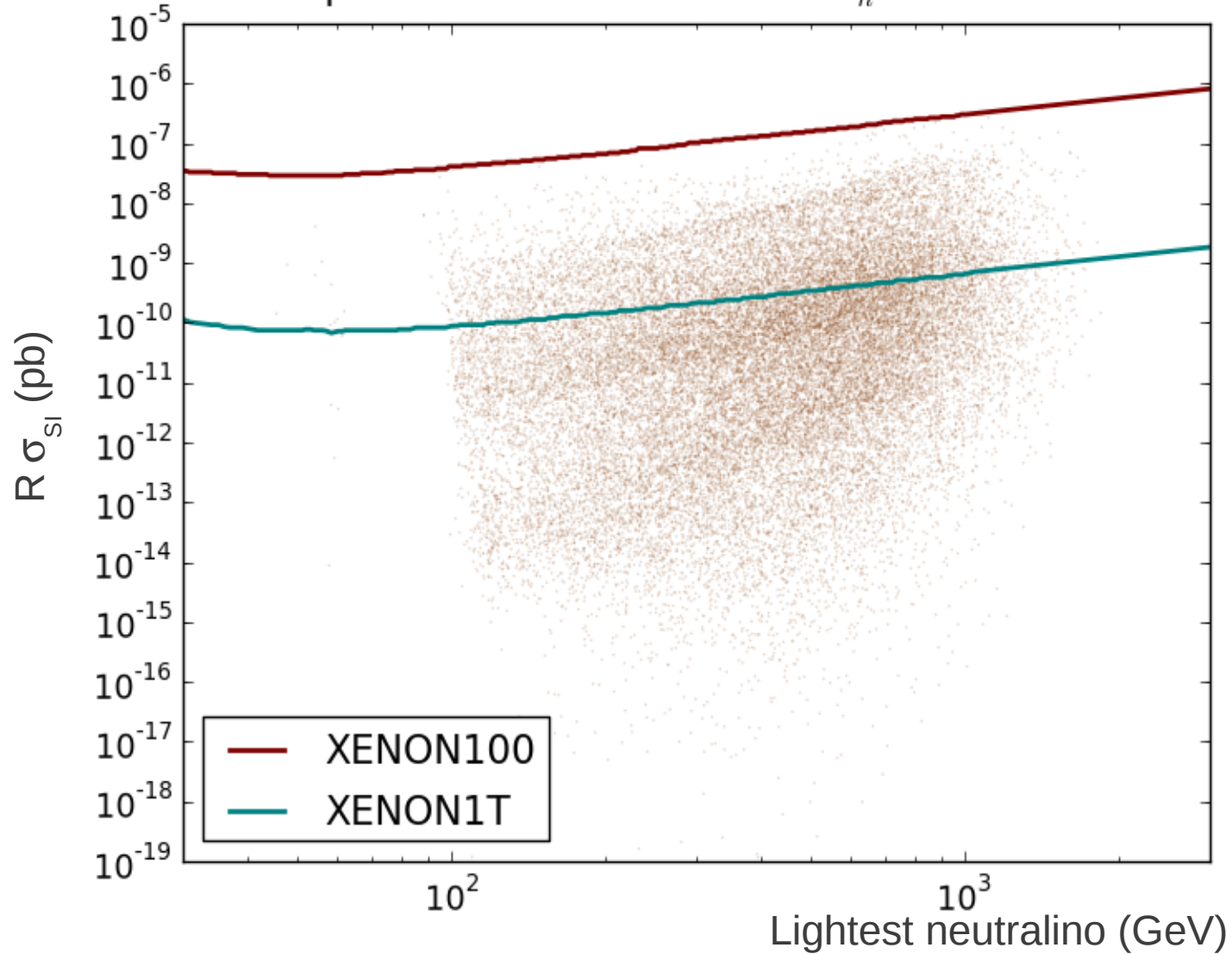


Neutralino LSP

Dark matter and LHC searches for SUSY  
complement each other

# LHC searches

All previous constraints with  $m_h = 126 \pm 3$  GeV



Higgs mass cut is approximately independent of LHC and DM searches

# LHC searches

Gravitino LSP

Analysis	7 TeV 4.7 fb <sup>-1</sup>
Jets + MET	18.8%
Many jets + MET	2.7%
1 $\ell$ + jets + MET	5.7%
Multi-leptons	14.2%
Stop/sbottom	12.0%
HSCP	16.3%
Disappearing tracks	0.9%
$B_s \rightarrow \mu\mu, \phi \rightarrow \tau\tau$	2.6%
Remaining models	60.1%

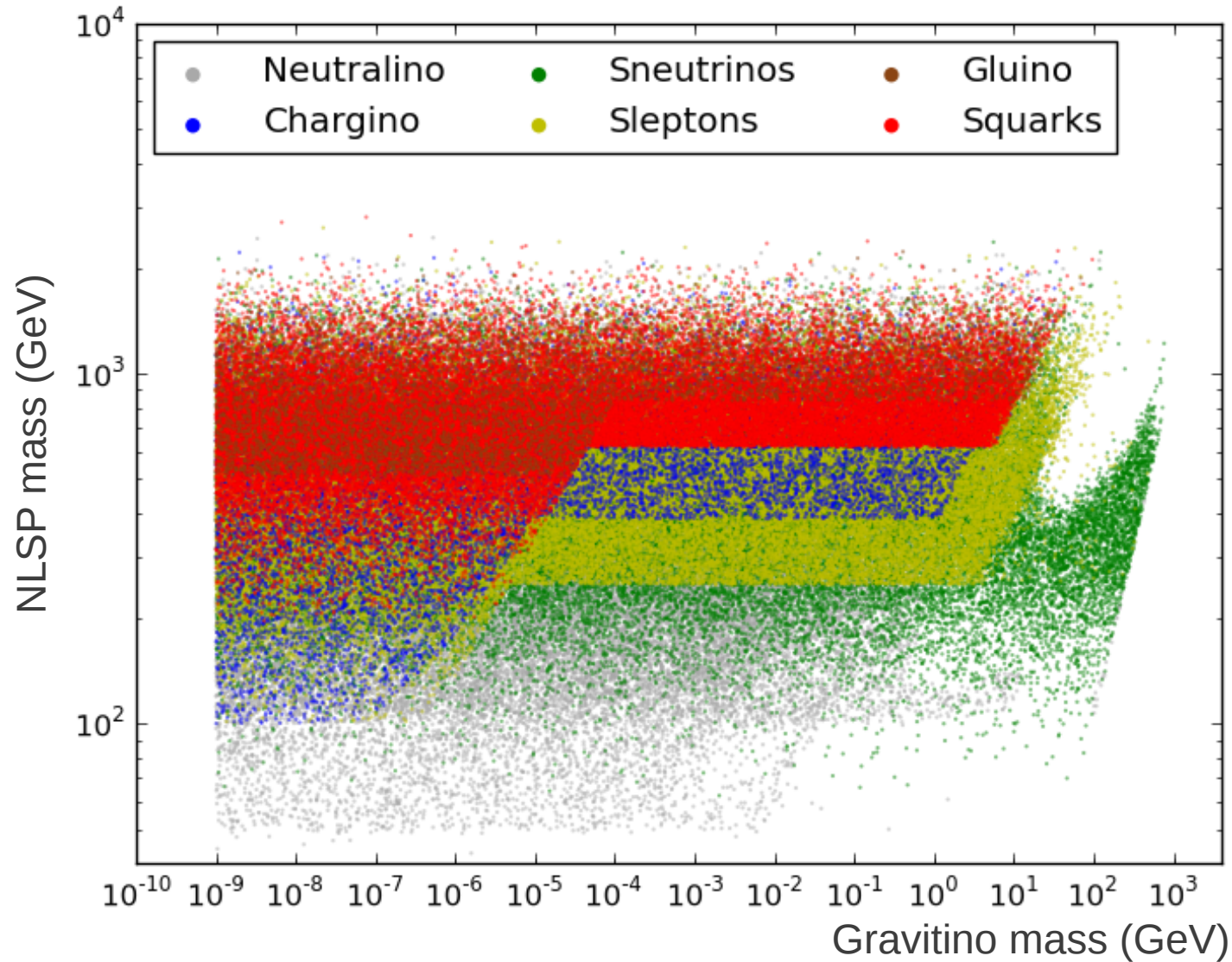
8 TeV  
TBD

- More MET since gravitino is nearly massless
- Very weakly interacting gravitino often leads to stable NLSP, so HSCP searches do much better than in neutralino LSP case



# LHC searches

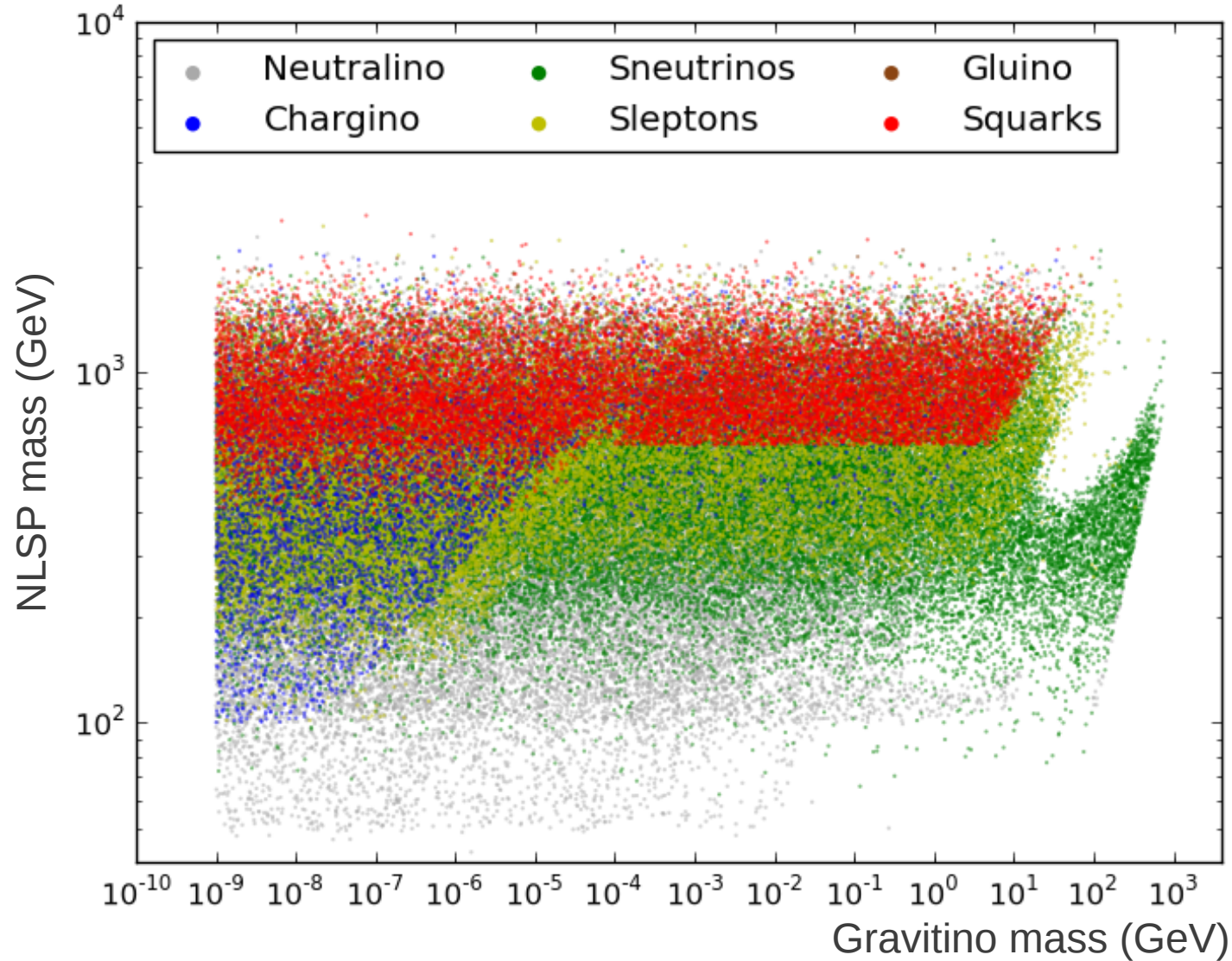
Full model set



Stable particles in gravitino LSP set are  
constrained by HSCP searches

# LHC searches

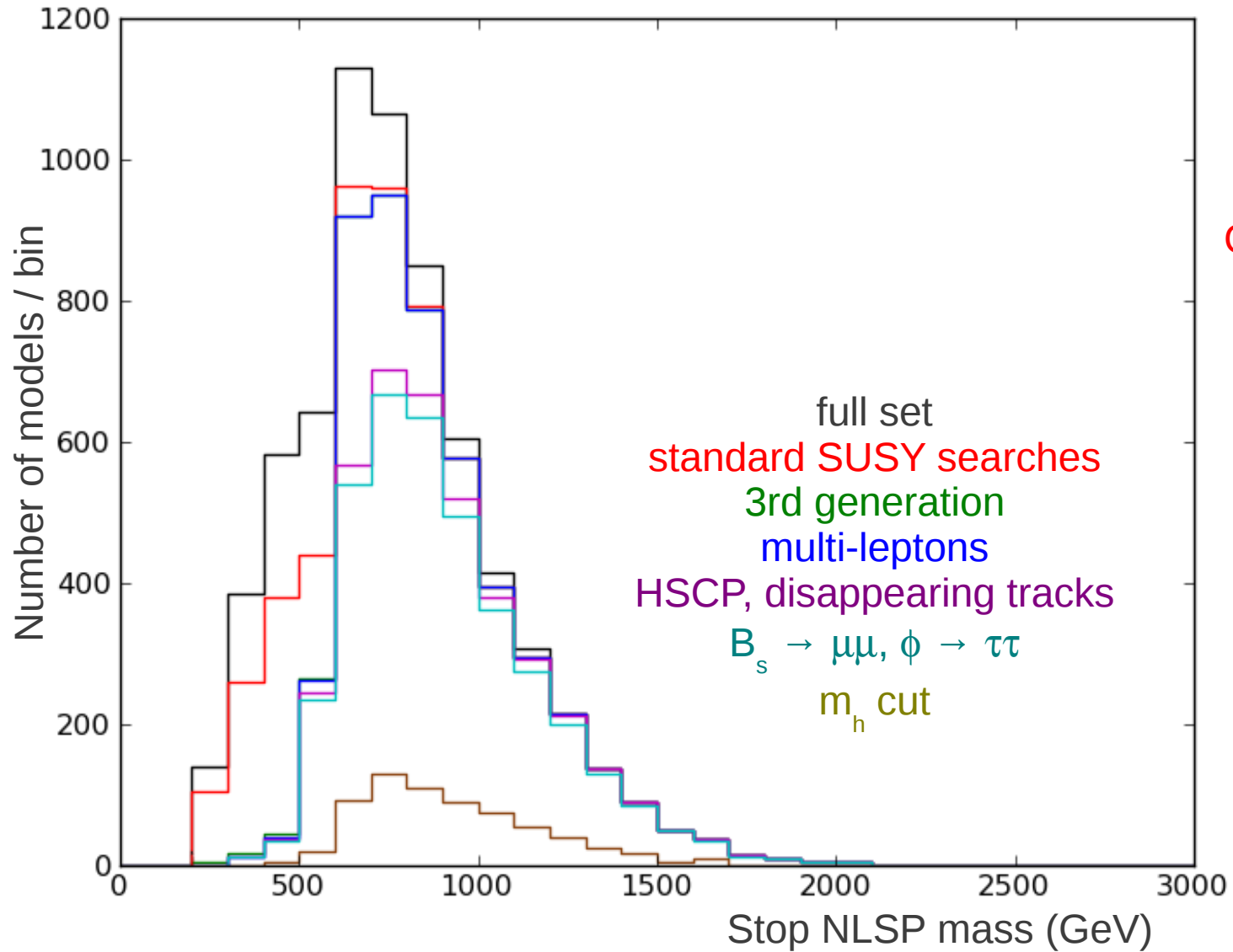
All searches



Stable particles in gravitino LSP set are  
constrained by HSCP searches

# LHC searches

After searches



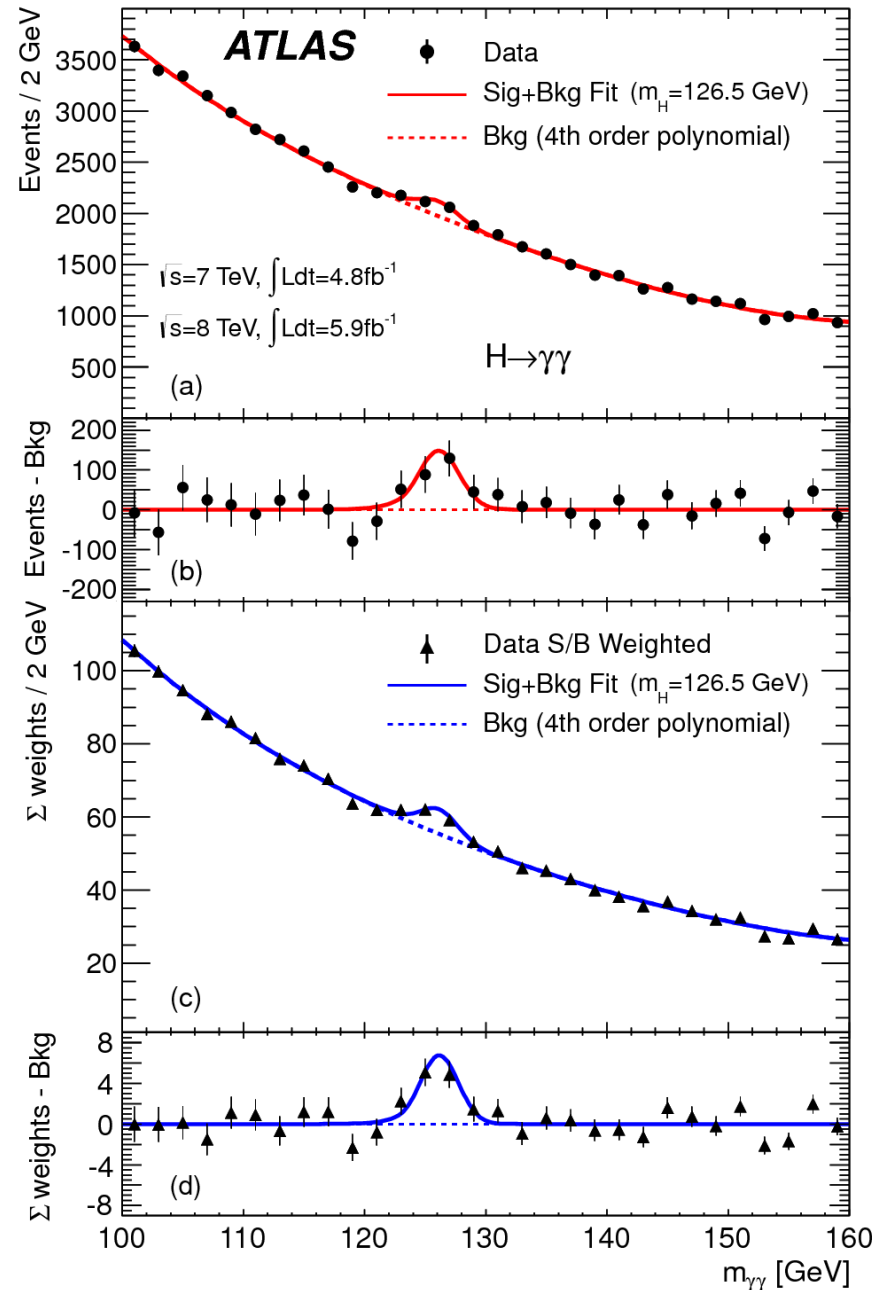
Light stops are very likely to be excluded if stop is NLSP

# Outline

- The phenomenological MSSM
- LHC search results
- Higgs discovery implications
- Fine-tuning
- Outlook

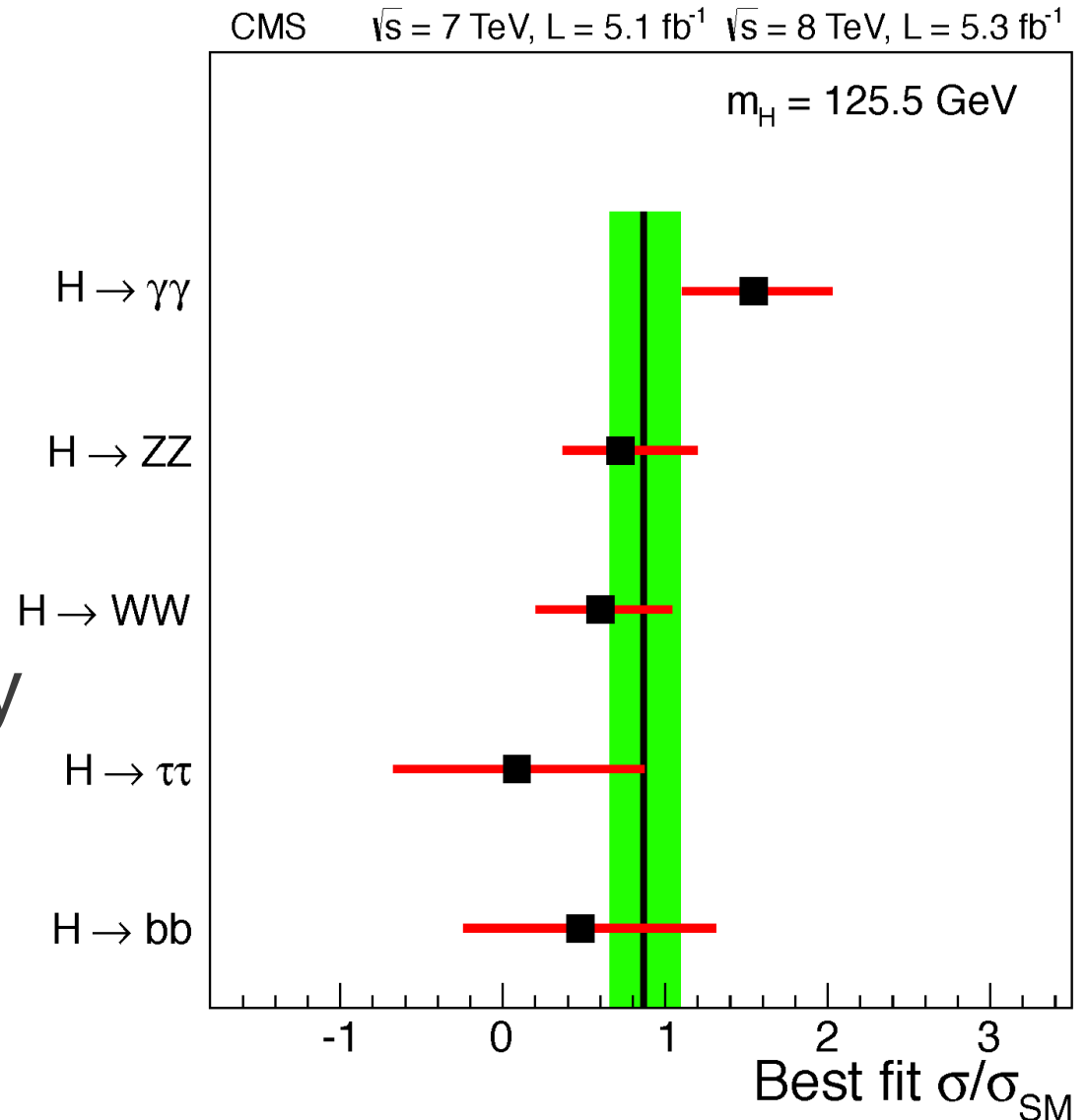
# Higgs discovery implications

- Excess events in both ATLAS and CMS Higgs searches near 126 GeV
- Greatest significance obtained from diphoton channel
- Can we easily obtain such a Higgs in the pMSSM?

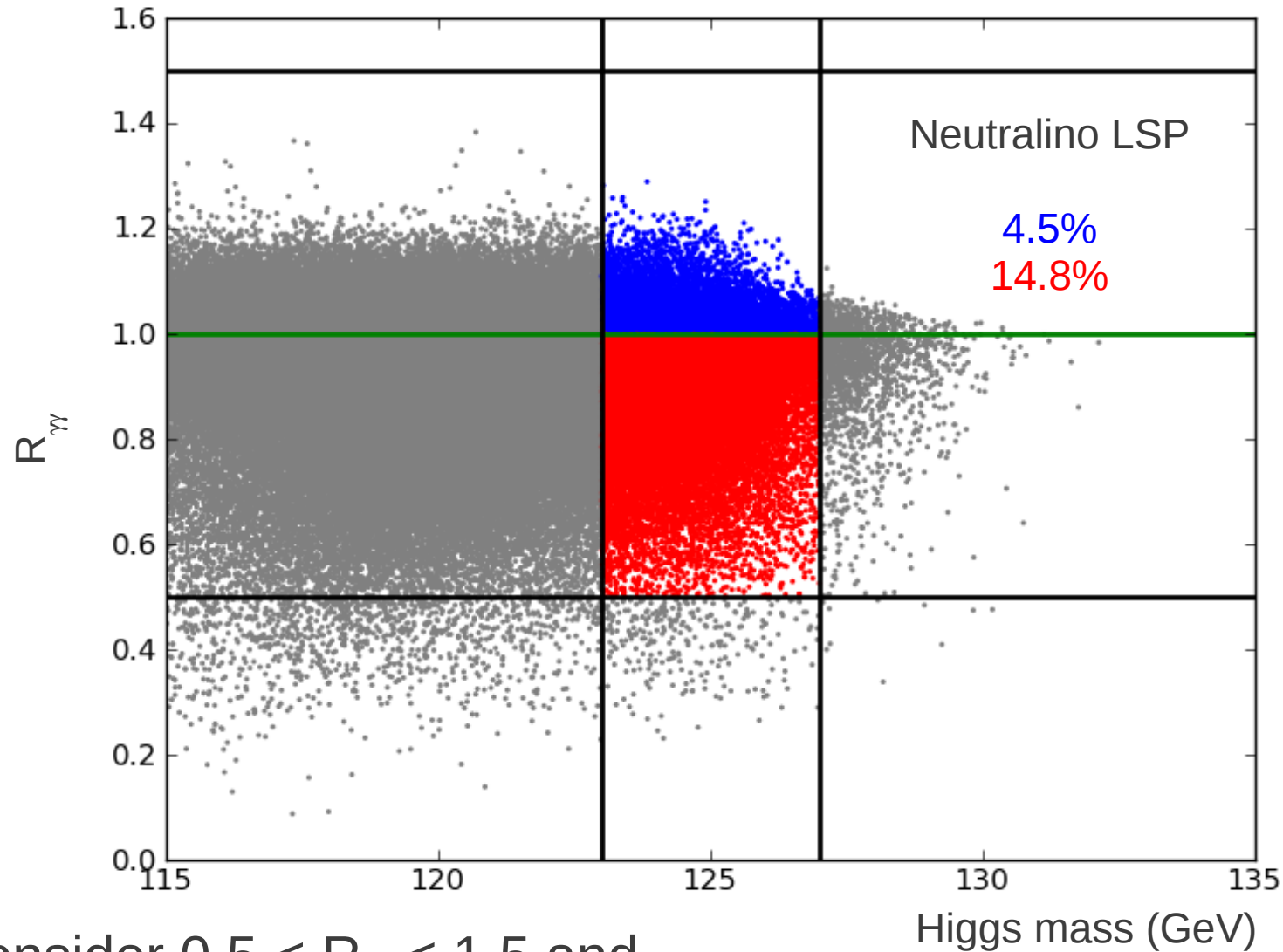


# Higgs discovery implications

- We calculate  $\sigma(gg \rightarrow h \rightarrow \gamma\gamma)_{\text{pMSSM/SM}}$  in the narrow width approximation and call the ratio  $R_{\gamma\gamma}$
- Can do same for any other mode  $h \rightarrow XX$  to obtain  $R_{XX}$

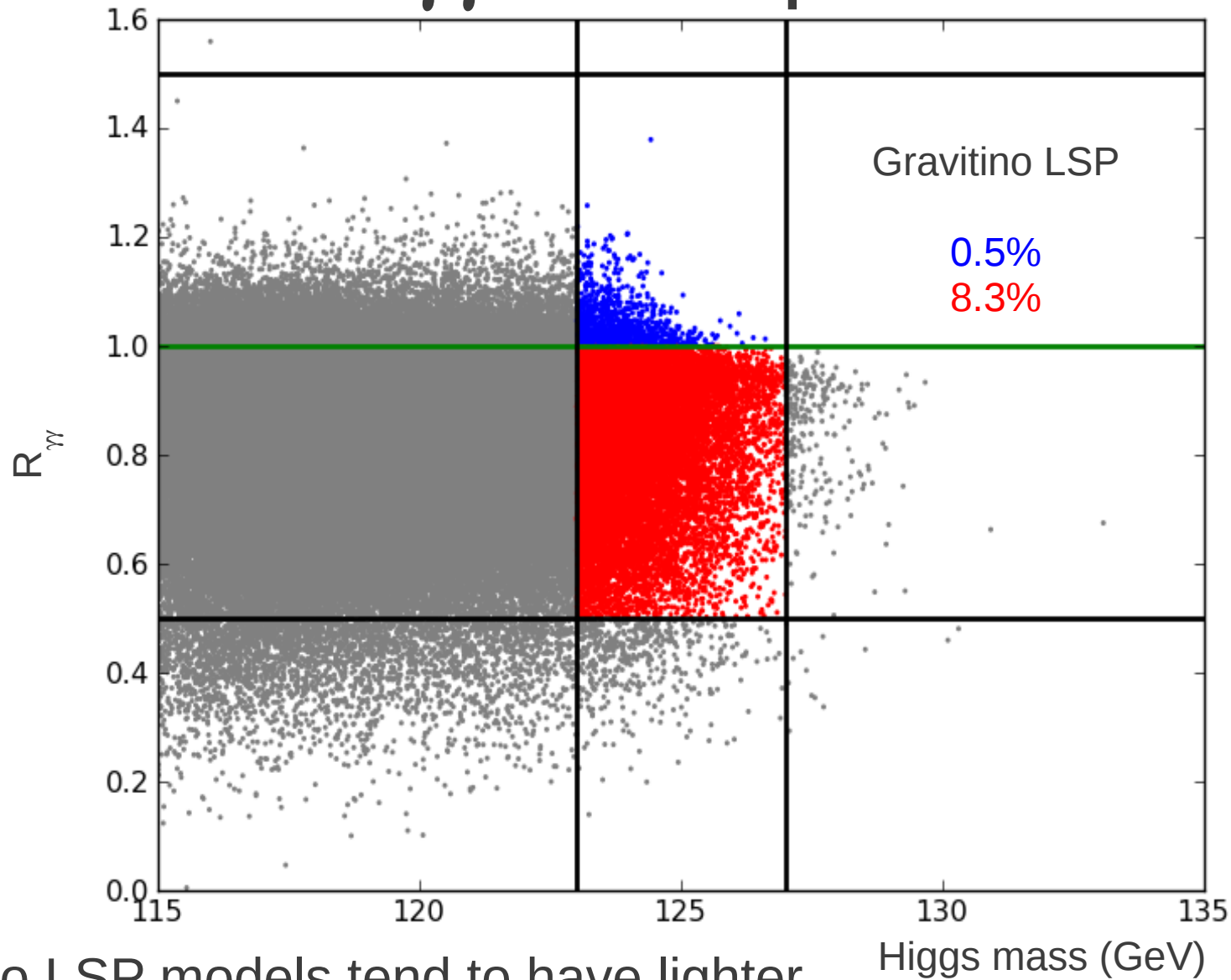


# $h \rightarrow \gamma\gamma$ in the pMSSM



We consider  $0.5 < R_{\gamma\gamma} < 1.5$  and  
 $123 \text{ GeV} < m_h < 127 \text{ GeV}$

# $h \rightarrow \gamma\gamma$ in the pMSSM

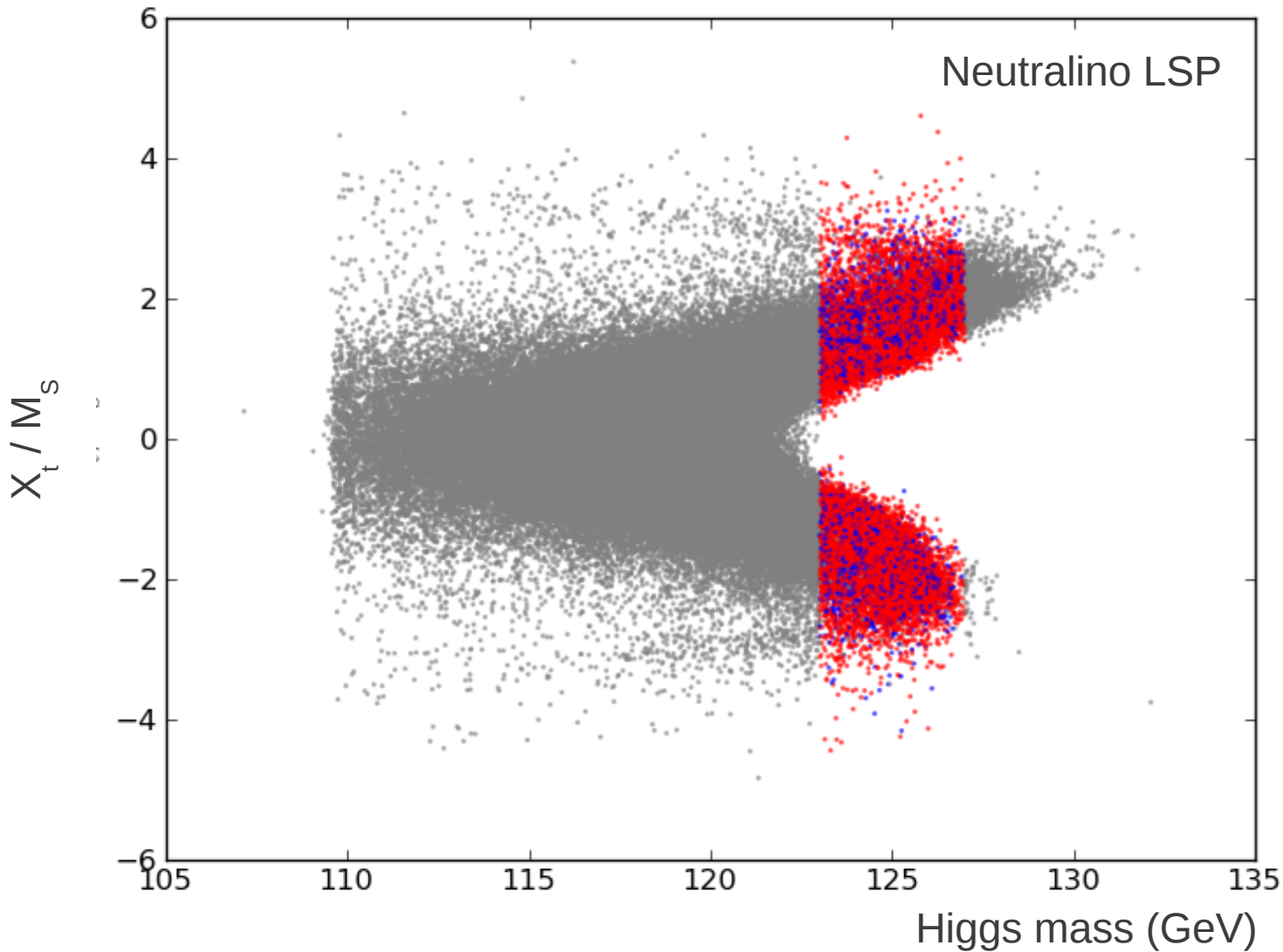


Gravitino LSP models tend to have lighter Higgses, but 125 GeV still quite viable



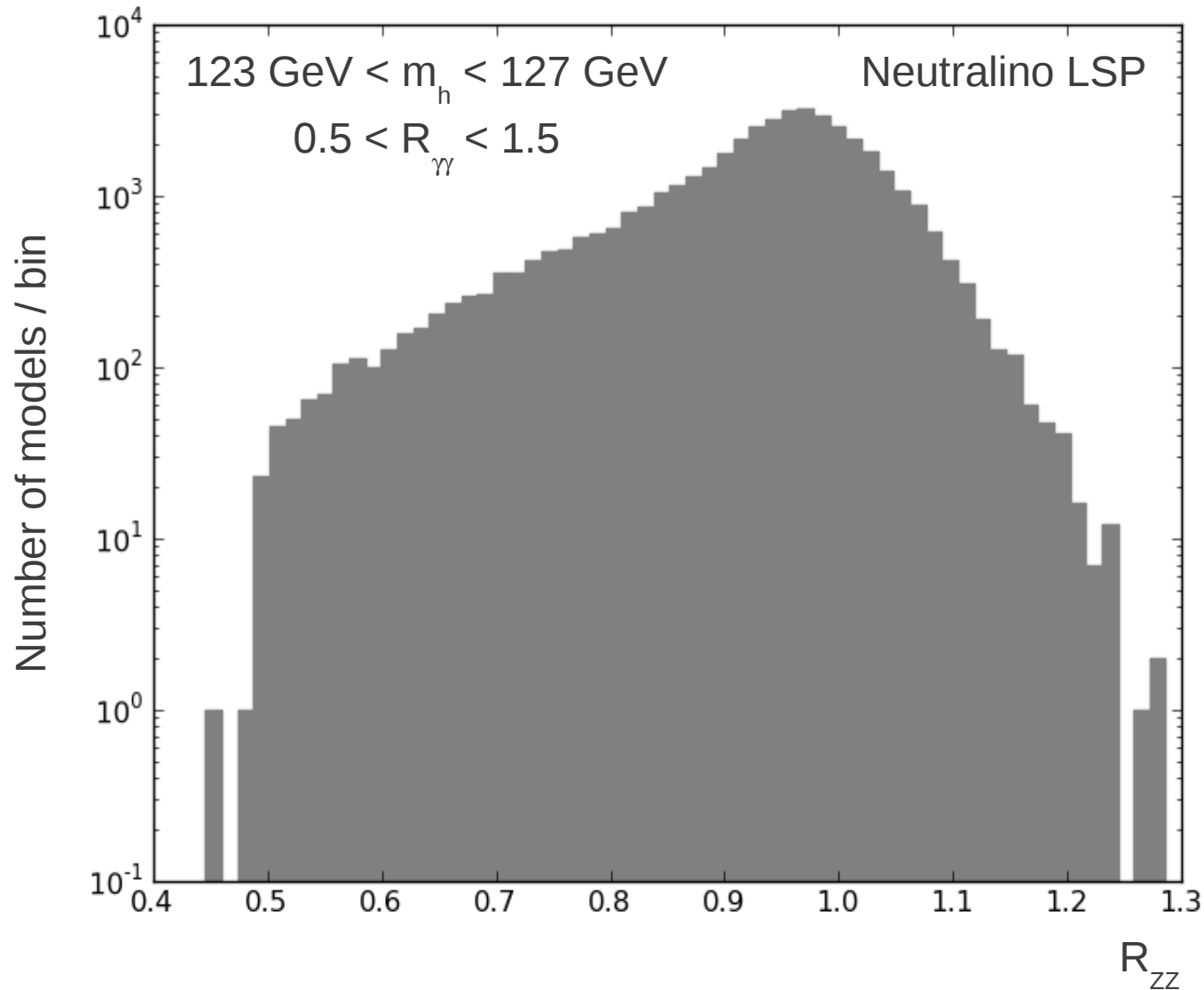


# Getting a heavy Higgs



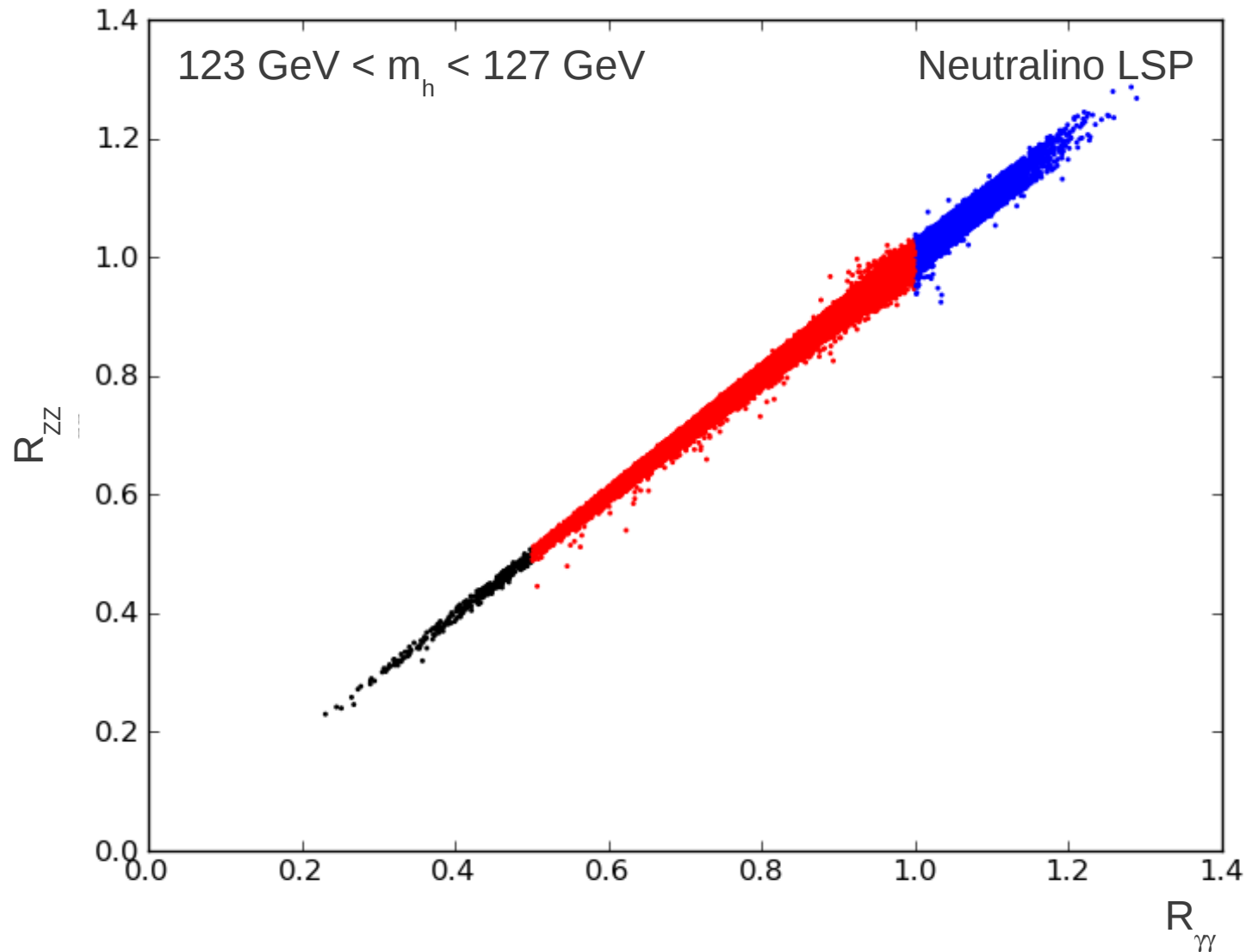
To get heavy Higgs, need large stop mixing  $X_t = A_t - \mu \cot \beta$

# Higgs LHC phenomenology



Wide spread of  $gg \rightarrow h \rightarrow ZZ$  cross section....

# Higgs LHC phenomenology

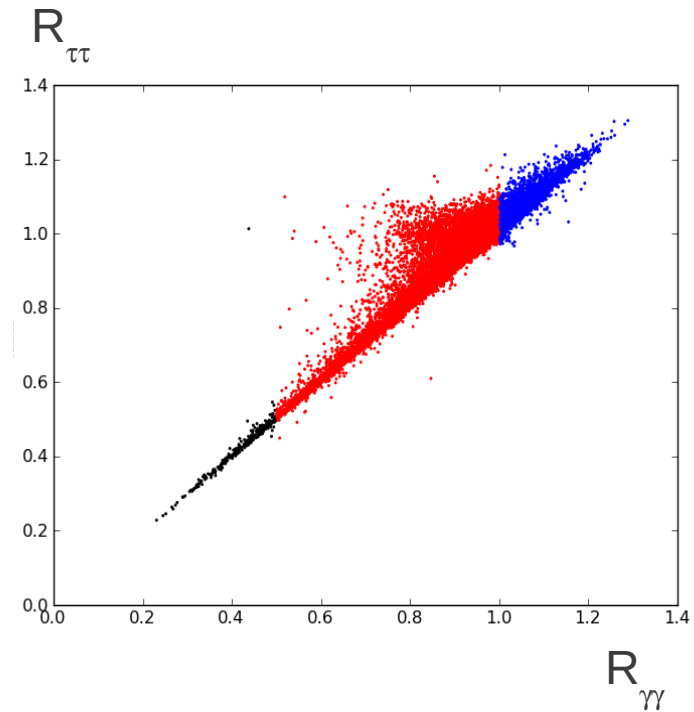
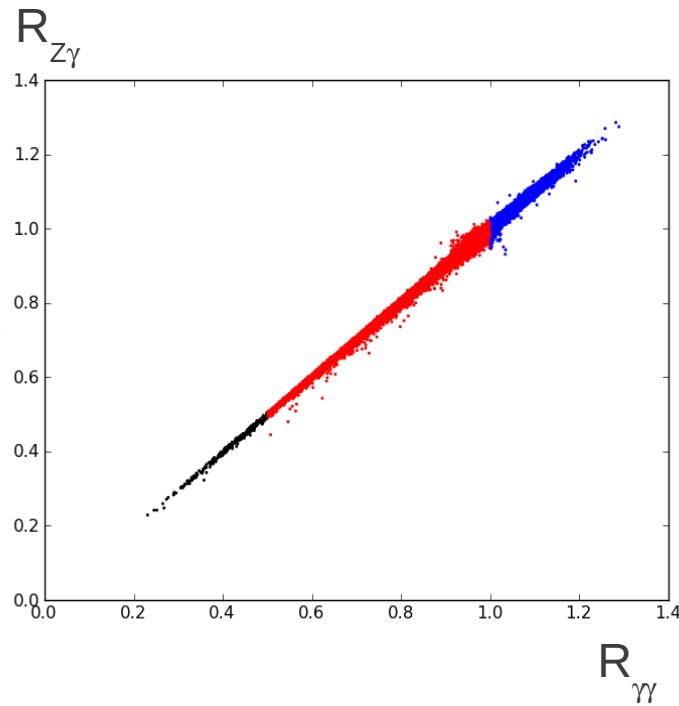
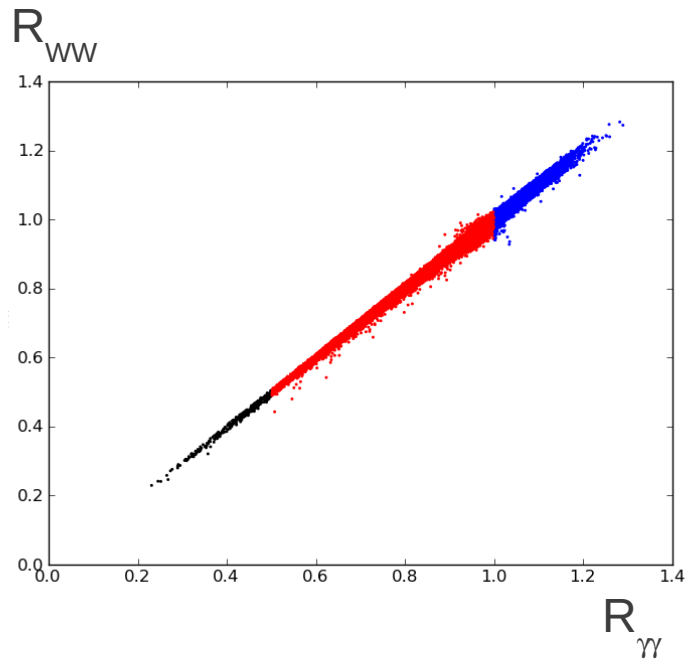


...but strongly correlated with number of  $h \rightarrow \gamma\gamma$  events!  
happens because most Higgs couplings are  $\sim$ SM

# Higgs LHC phenomenology

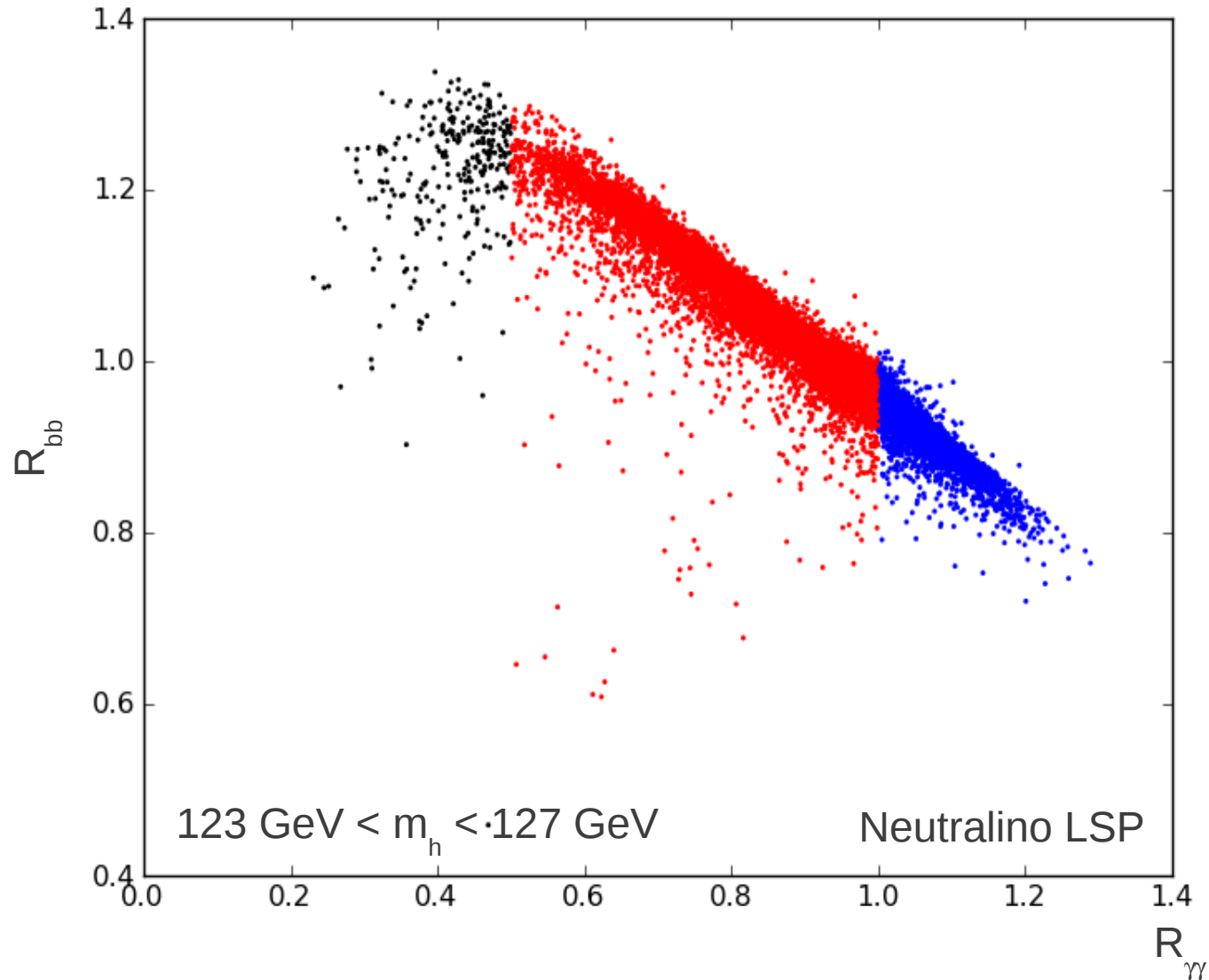
Neutralino LSP

$123 \text{ GeV} < m_h < 127 \text{ GeV}$



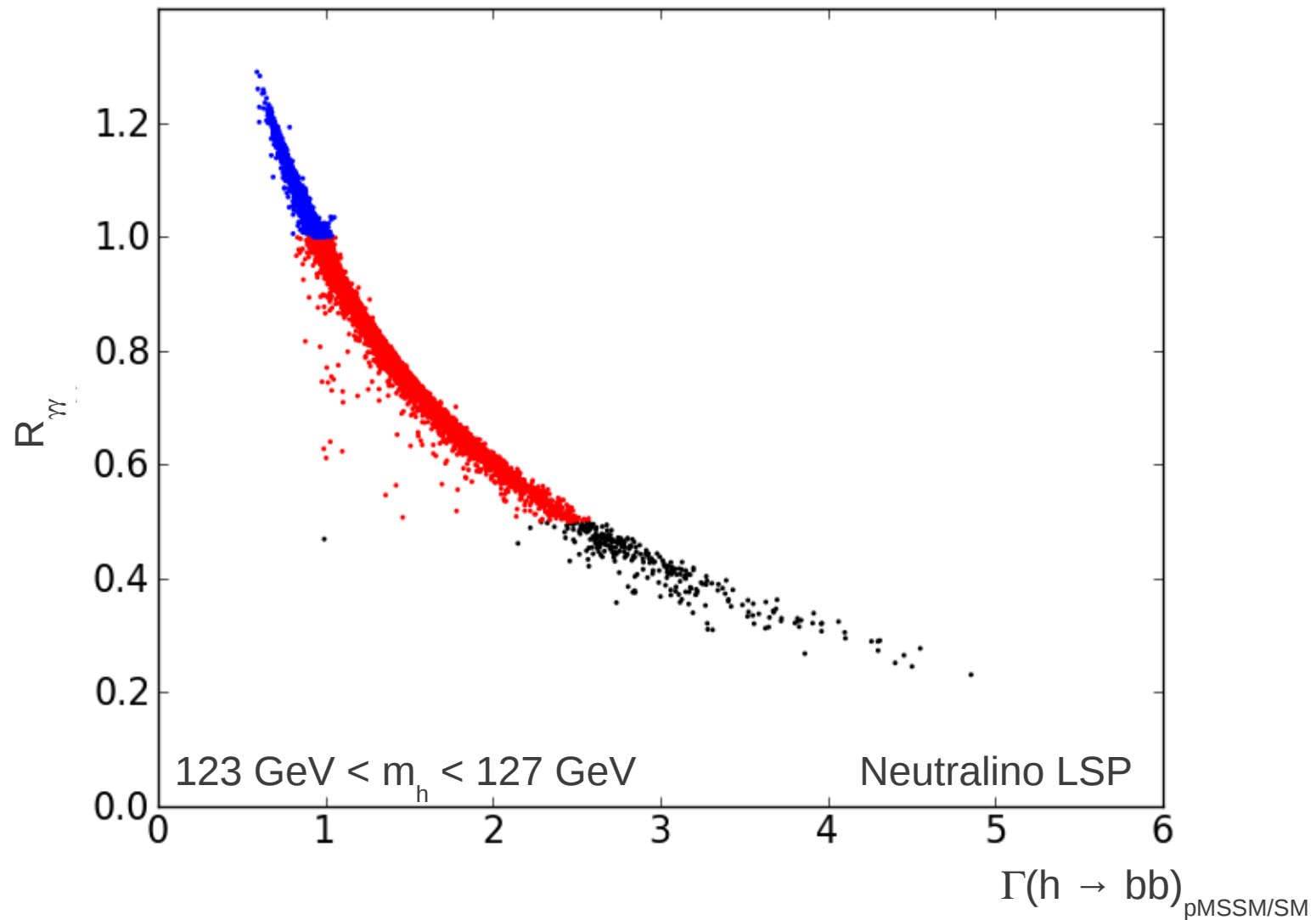
Other modes behave the same way because of decoupling

# Higgs LHC phenomenology



bb production is anti-correlated with other decay modes

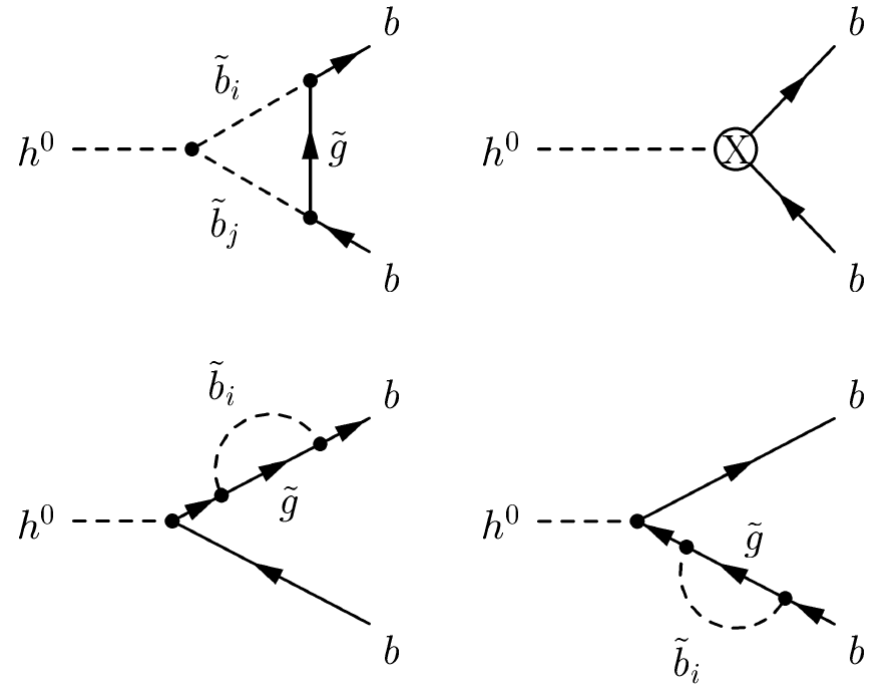
# Higgs LHC phenomenology



SUSY corrections to bb width reduce other branching ratios!

# $h \rightarrow bb$ decoupling

- $\Gamma = \Gamma_0 (1 + 2 \delta g^{\text{QCD}} / g + 2 \delta g^{\text{SQCD}} / g)$
- $\delta g^{\text{SQCD}}$  receives contributions from vertex correction, b wave function renormalization, and hbb counterterm



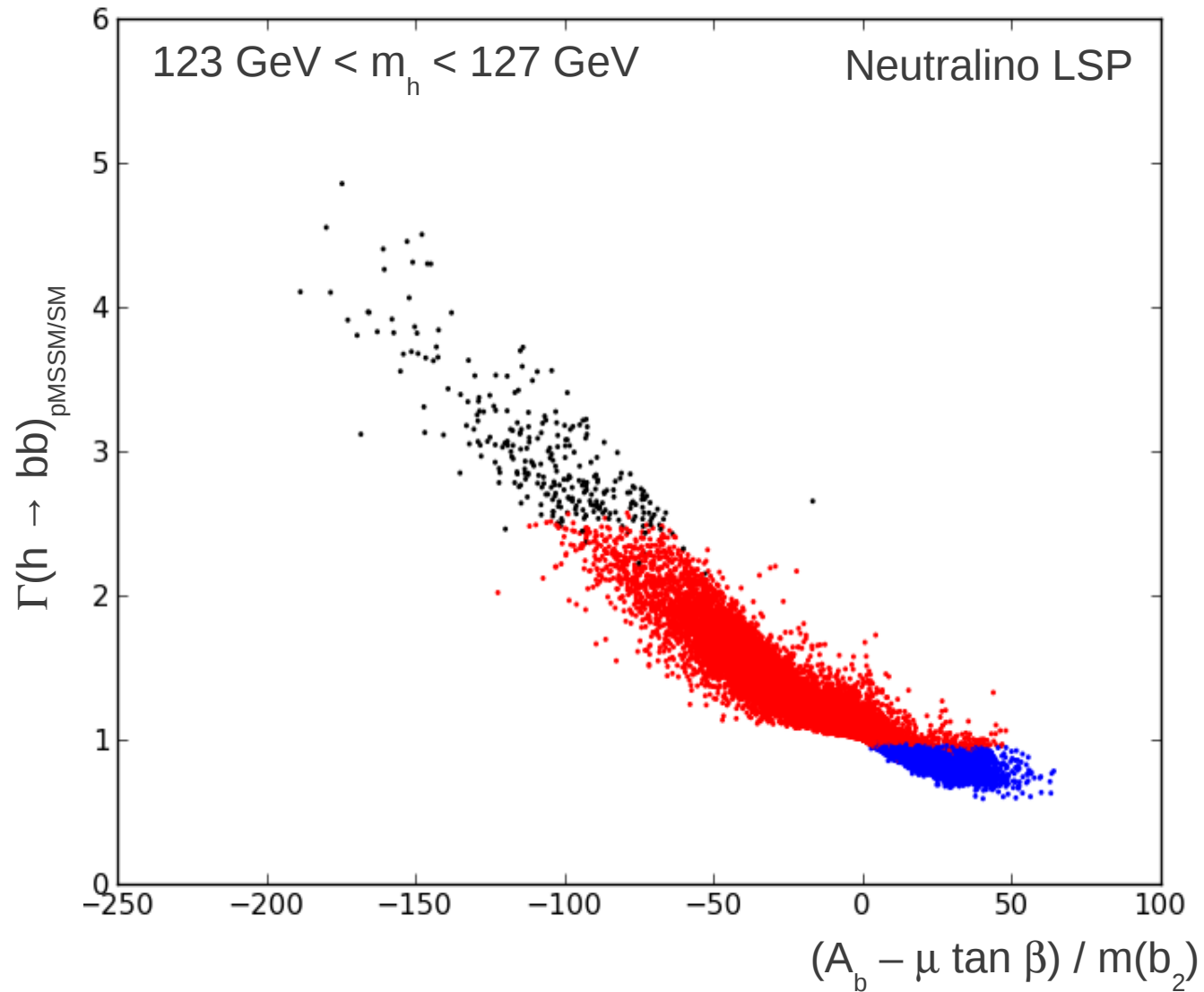
Haber et al., hep-ph/0007006



# $h \rightarrow bb$ decoupling

- As sparticles get heavier, SUSY corrections to  $h \rightarrow bb$  width usually decouple quickly
- However, in certain limits, e.g. near-maximal sbottom mixing with large  $\tan \beta$ , the decoupling happens very slowly, and corrections can be large
- Large resulting corrections push up  $bb$  width, decreasing all other branching ratios accordingly

# $h \rightarrow bb$ decoupling



# Outline

- The phenomenological MSSM
- LHC search results
- Higgs discovery implications
- **Fine-tuning**
- Outlook

# Fine-tuning

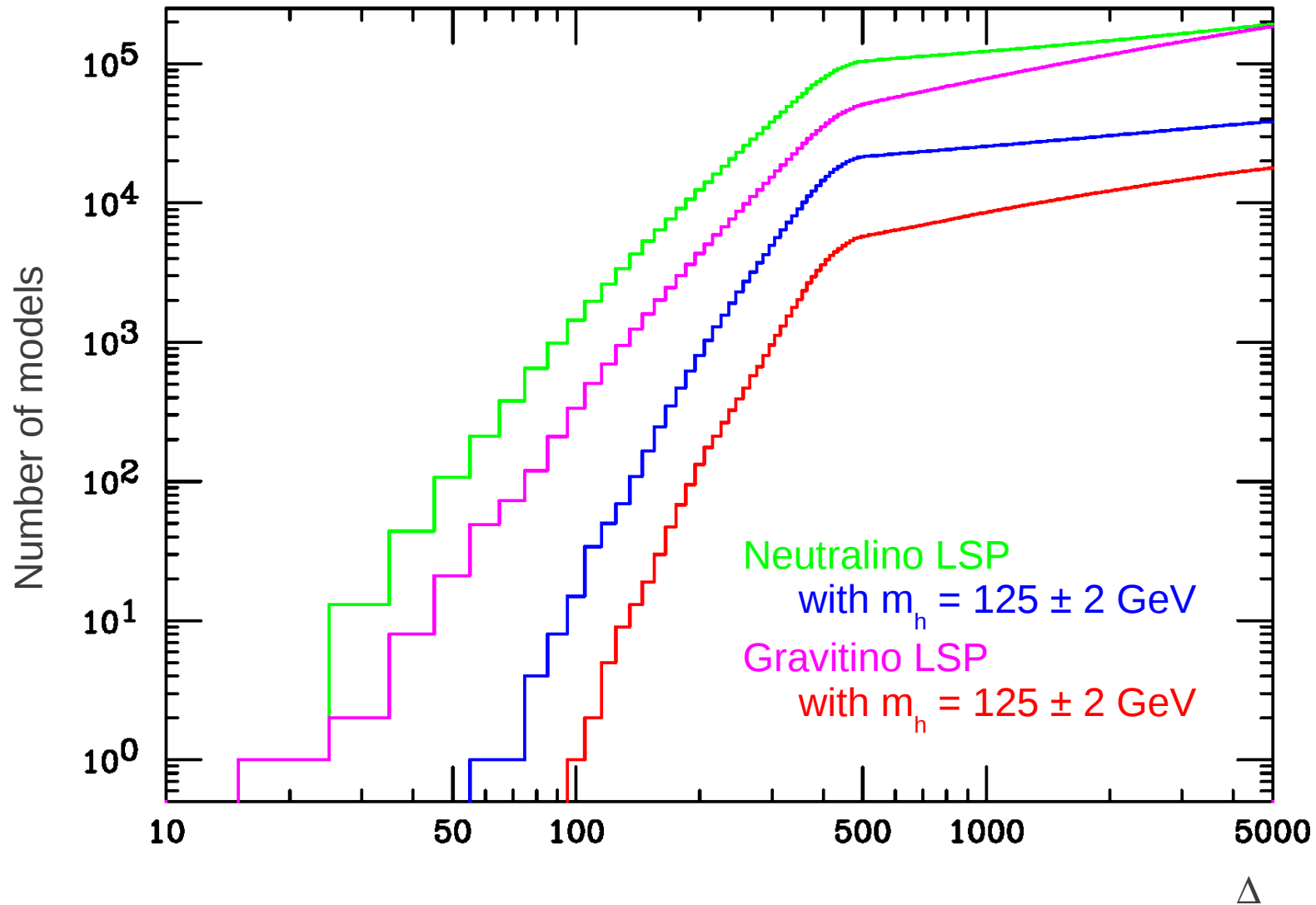
- Measure sensitivity of electroweak symmetry breaking scale to each pMSSM parameter  $p_i$

Barbieri and Giudice, Nucl.Phys. B306 (1988) 63

$$M_Z^2 = -2\mu^2 + 2 \frac{m_{H_d}^2 - t_\beta^2 m_{H_u}^2}{t_\beta^2 - 1}$$

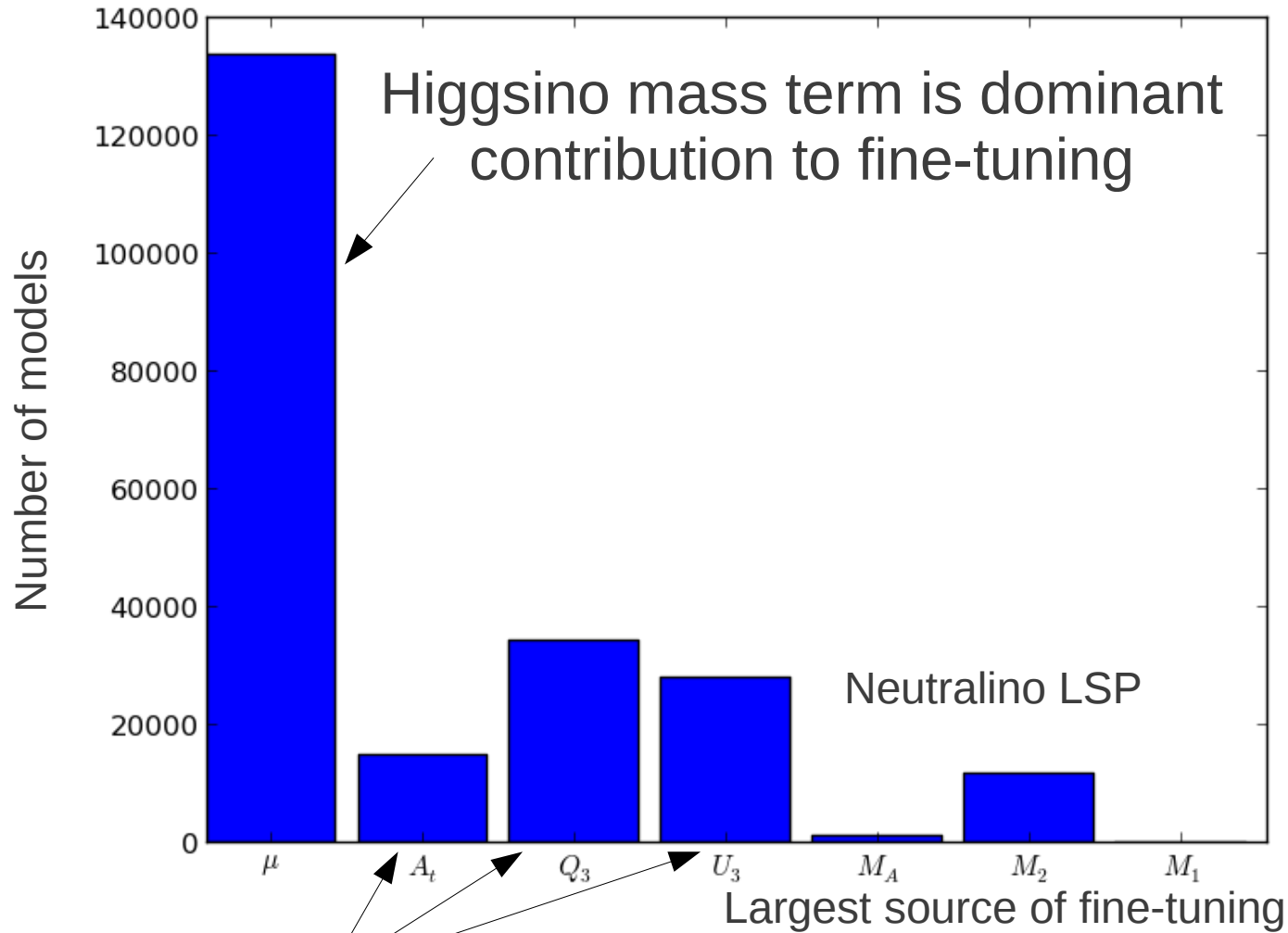
- $A_i = \partial(\log M_Z^2) / \partial(\log p_i)$ ,  $1 \leq i \leq 19$
- Most sensitive to  $\mu$  and stop mass parameters, but gluino mass enters at higher order
- Take maximum of all  $A_i$  to get fine-tuning  $\Delta$

# Fine-tuning



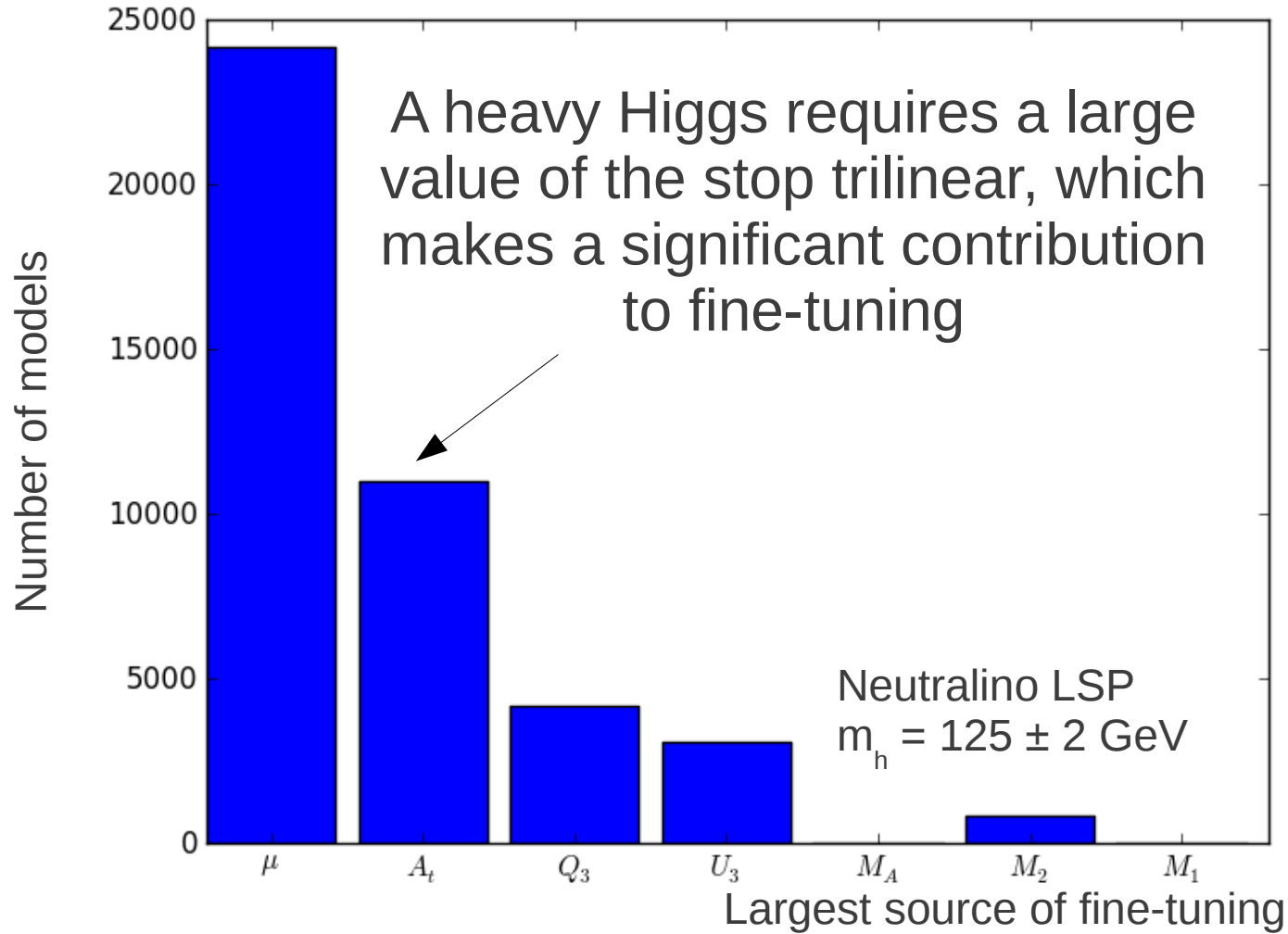
models with Higgs near 125 GeV  
are more fine-tuned

# Sources of fine-tuning



Stop mass terms also important, but even with strong coupling, loop-induced gluino contribution is less than wino FT

# Sources of fine-tuning

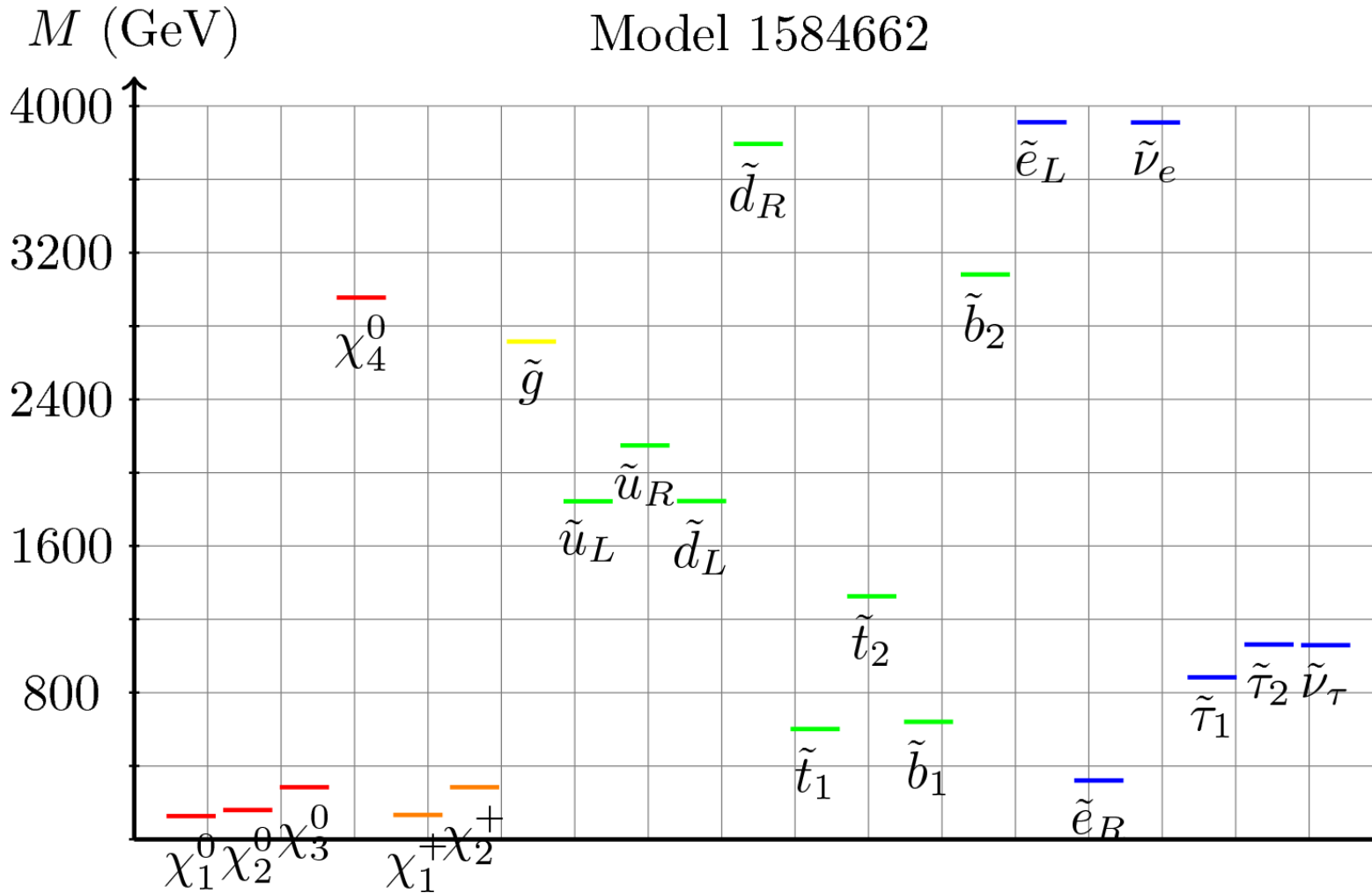


# Features of models with low FT

- Look at models with  $\Delta < 100$ , Higgs near 125 GeV, and passing all existing constraints
- 9 (0) such models in neutralino (gravitino) LSP model set
- Light higgsinos, usually light winos
- Moderately light 3rd generation squarks, heavy 1st/2nd generation squarks
- Gluino is constrained by LHC searches, but not naturalness at this level of fine-tuning

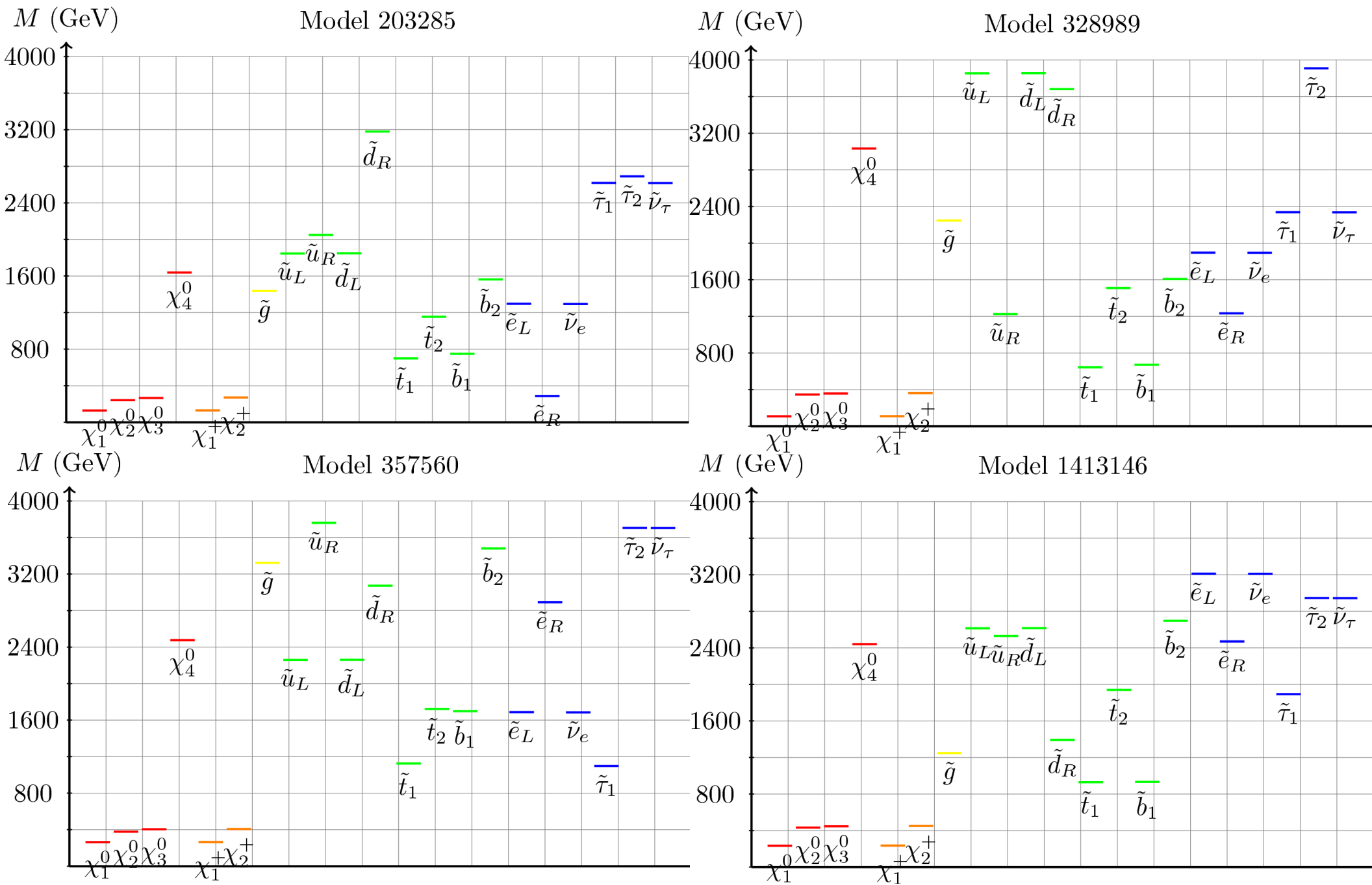


# Sample spectrum



Many possible cascades for light stops and sbottoms

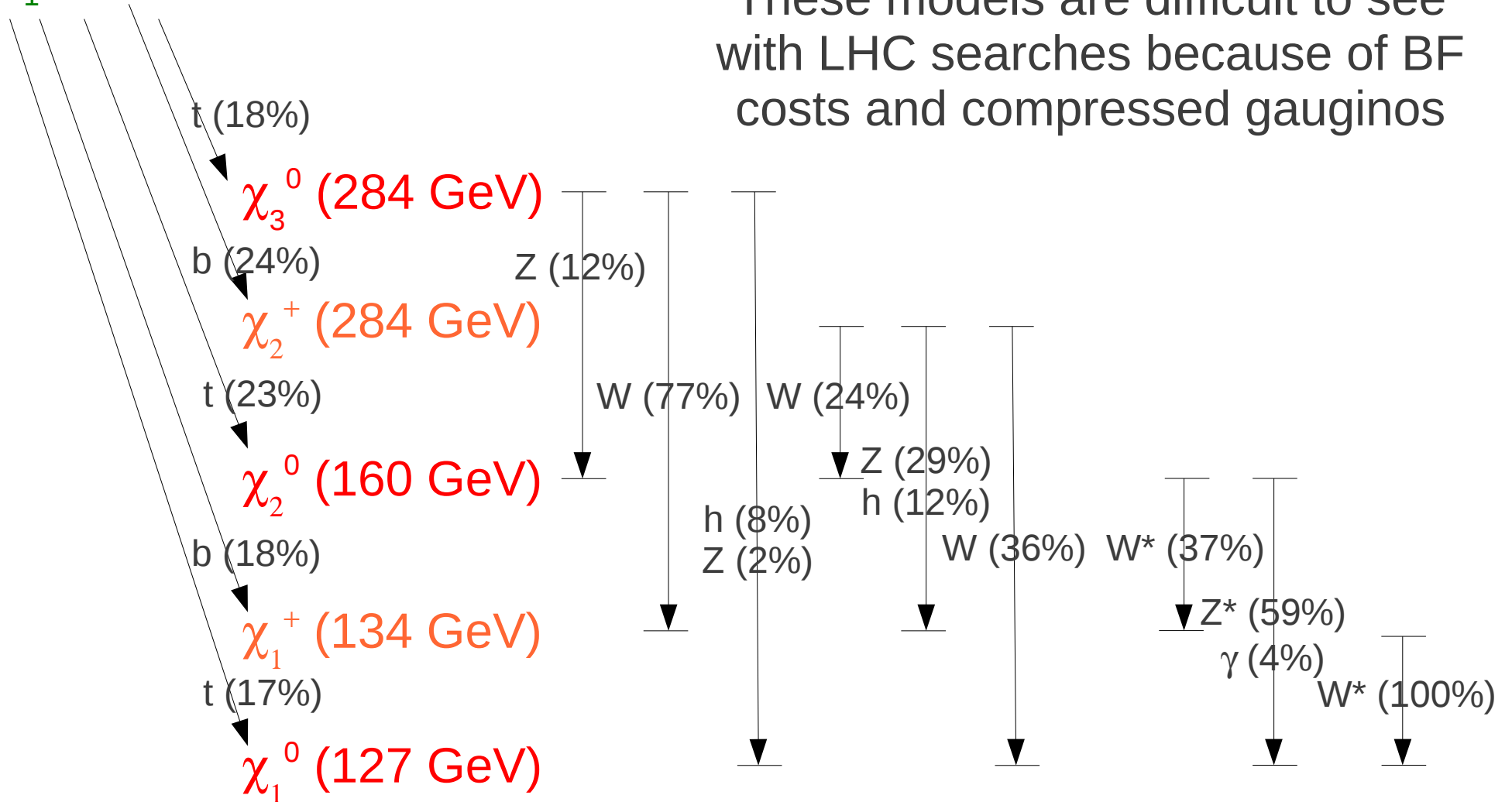
# More sample spectra



# Sample spectrum

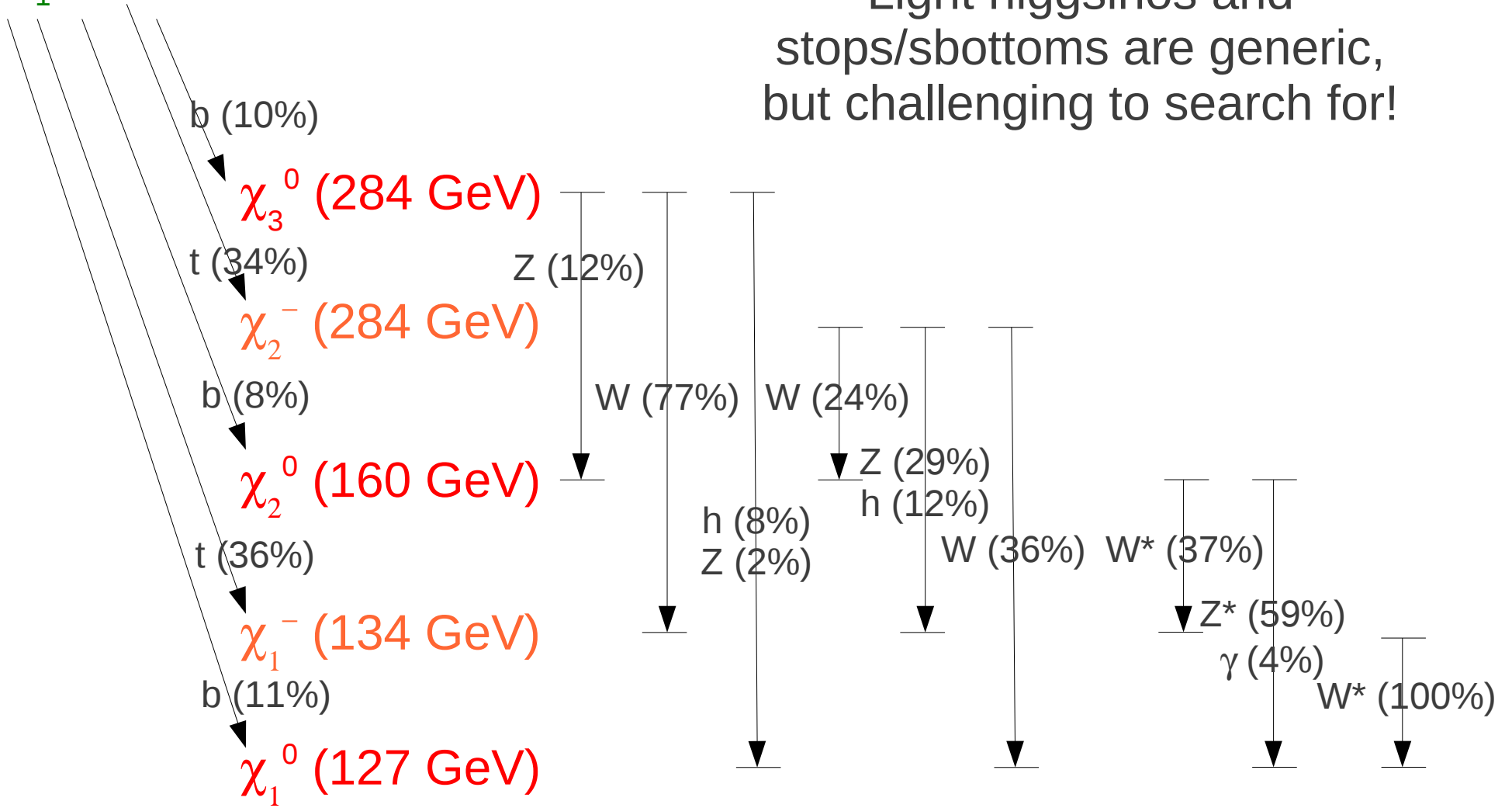
$t_1$  (601 GeV)

These models are difficult to see with LHC searches because of BF costs and compressed gauginos

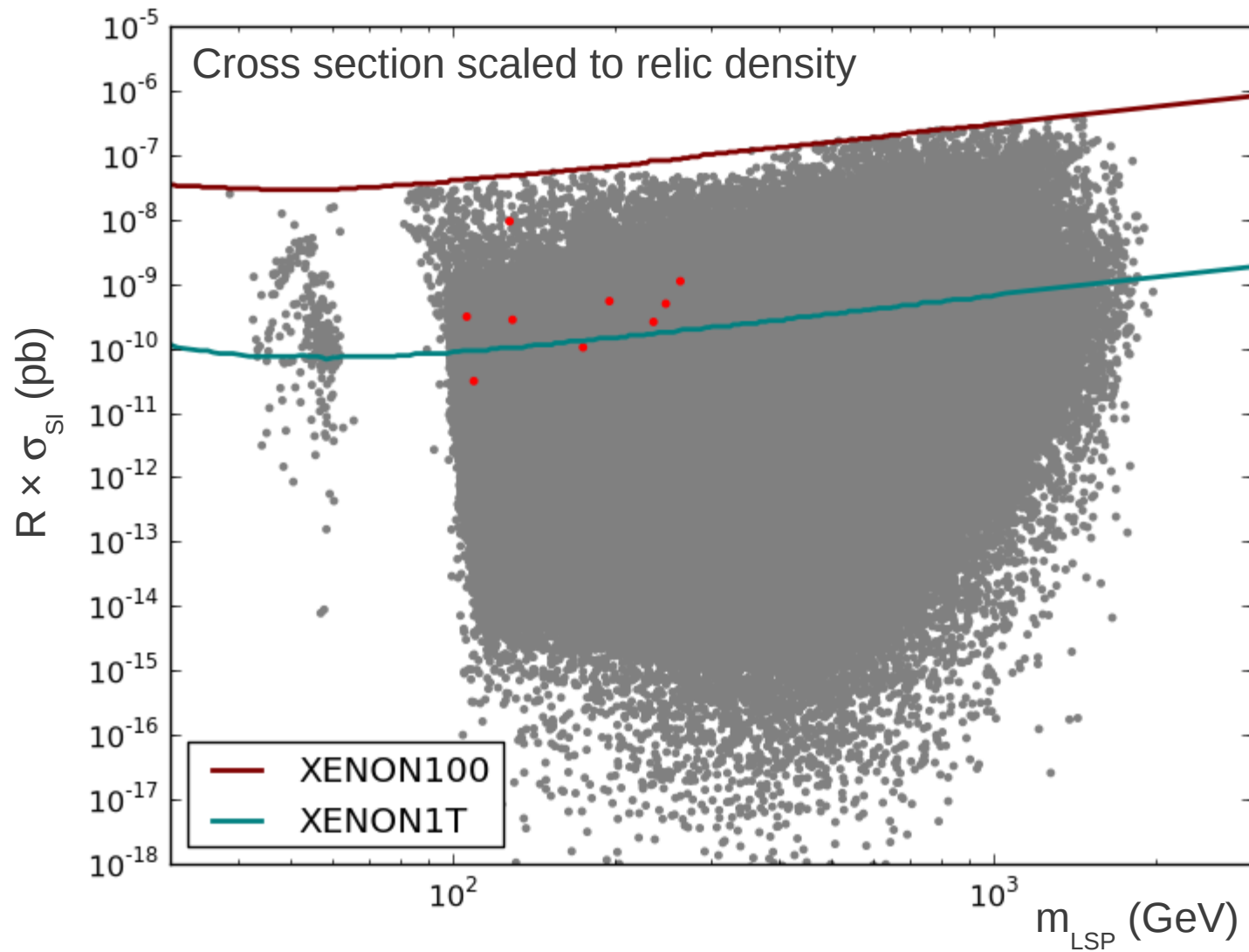


# Sample spectrum

$b_1$  (641 GeV)



# Features of models with low FT



Low FT models will be probed by upcoming XENON results

# Outline

- The phenomenological MSSM
- Model set generation
- LHC search results
- A 125 GeV Higgs
- Fine-tuning
- Outlook

# Outlook

- The pMSSM allows us to investigate complete, realistic supersymmetric spectra at the LHC and beyond
- Phenomenology different between neutralino and gravitino LSP model sets, though both have collider-stable particles
- LHC is already excluding models in our sets, through both MET and non-MET searches
- 8 TeV searches improve coverage overall, but tighter cuts means models can get missed

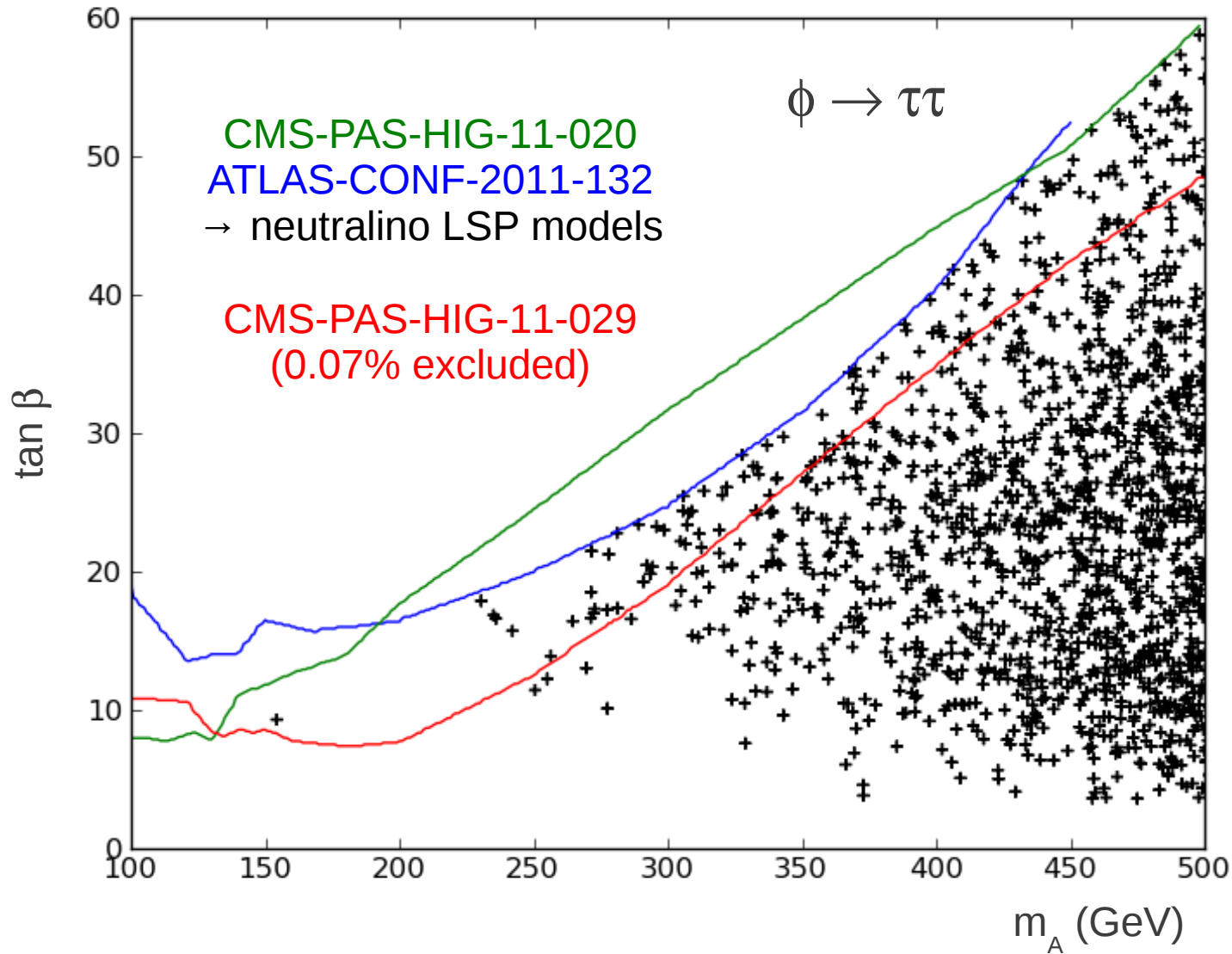
# Outlook

- Higgs production cross sections vary depending on final state, and are sensitive to  $hbb$  coupling
- Low FT models similar to “natural” spectra with light stops are still allowed despite 3rd generation searches due to cascade decays which are generic with light higgsinos
- Stay tuned for more!

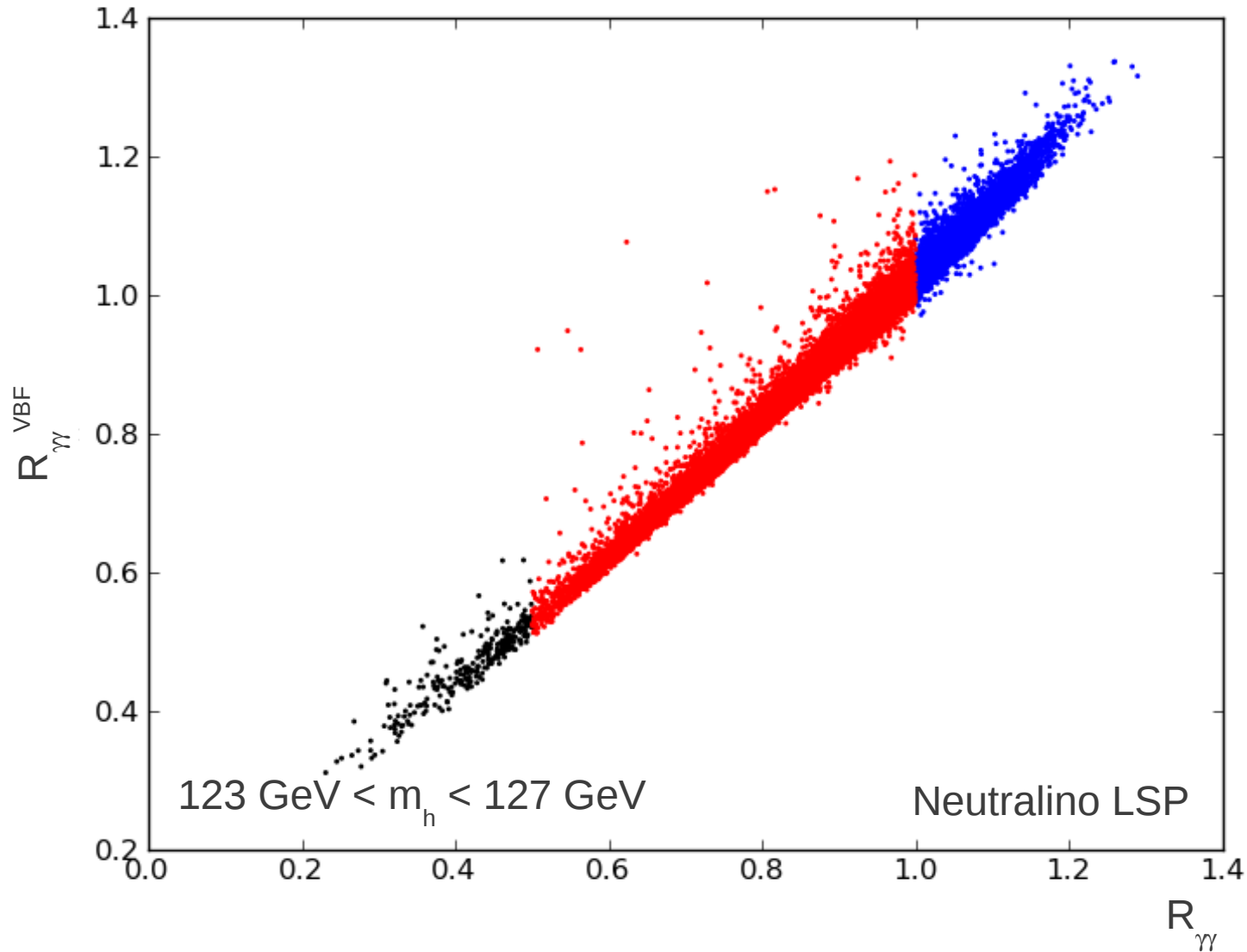


Backup

# Non-MET searches

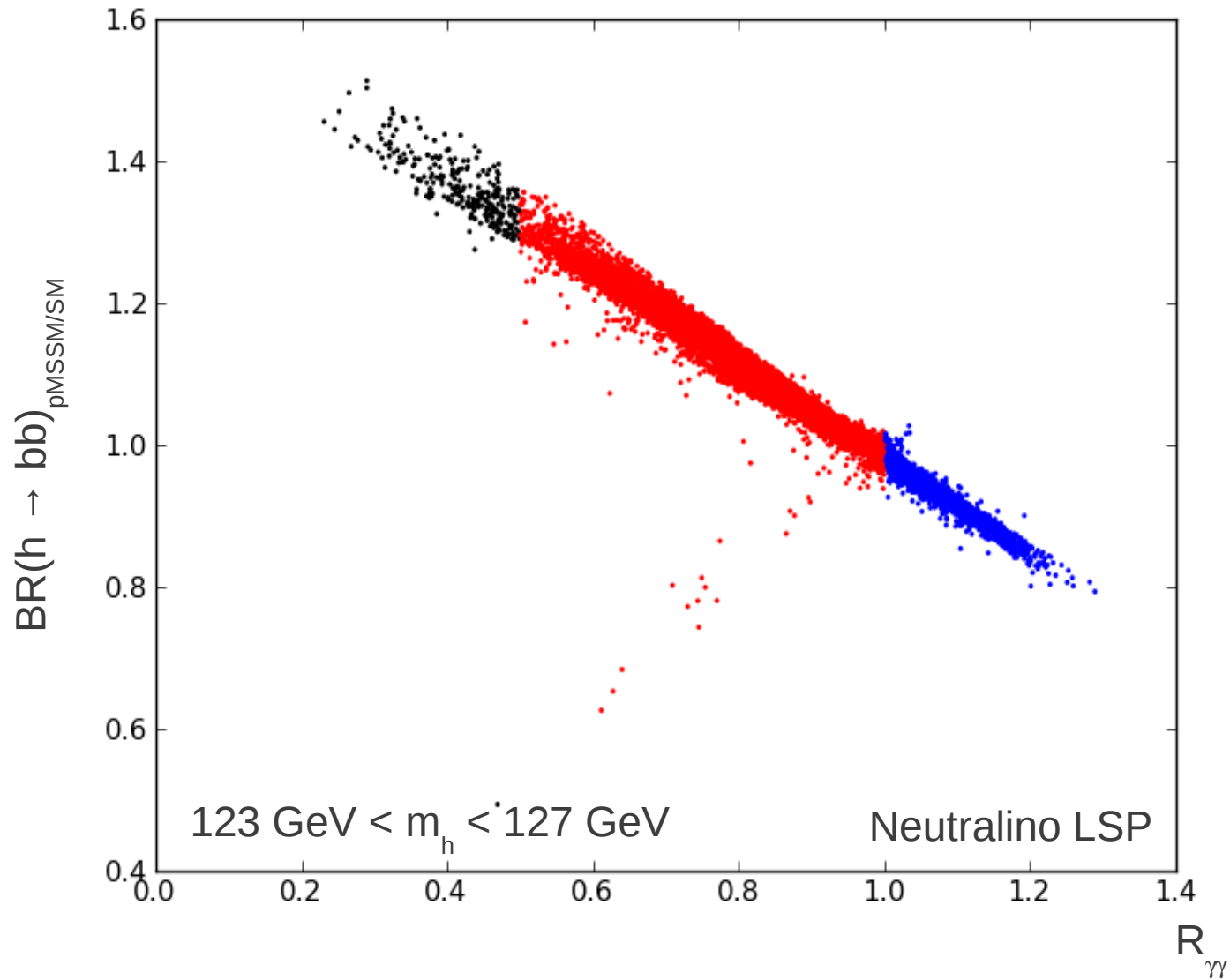


# Higgs LHC phenomenology



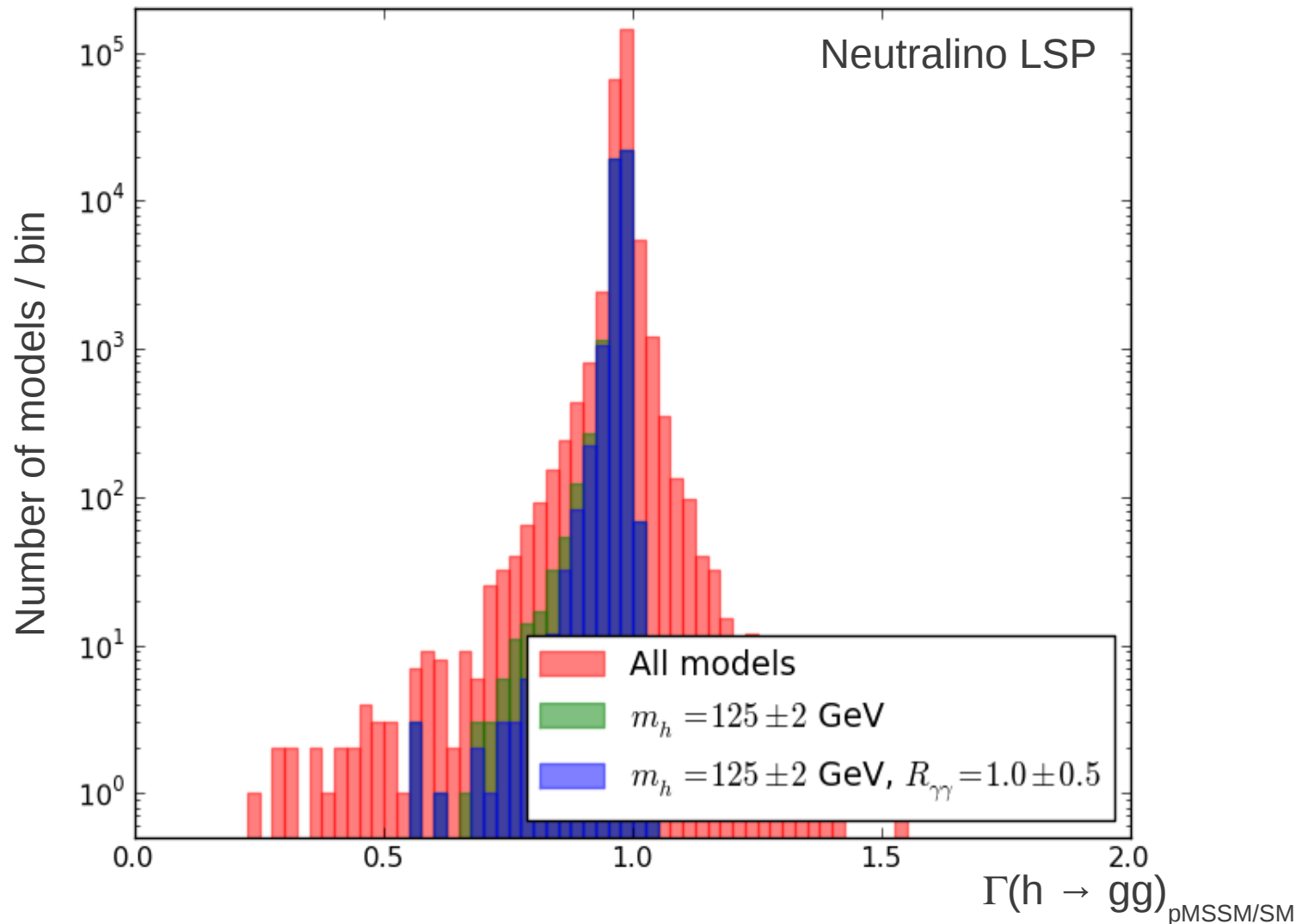
Can also look at vector boson fusion production  $WW \rightarrow h \rightarrow \gamma\gamma$

# Higgs LHC phenomenology



BR for  $h \rightarrow bb$  is anticorrelated with expected  $\gamma\gamma$  production

# Higgs LHC phenomenology



gg width is affected by less than  $\sim 25\%$  for models with Higgs near 125 GeV