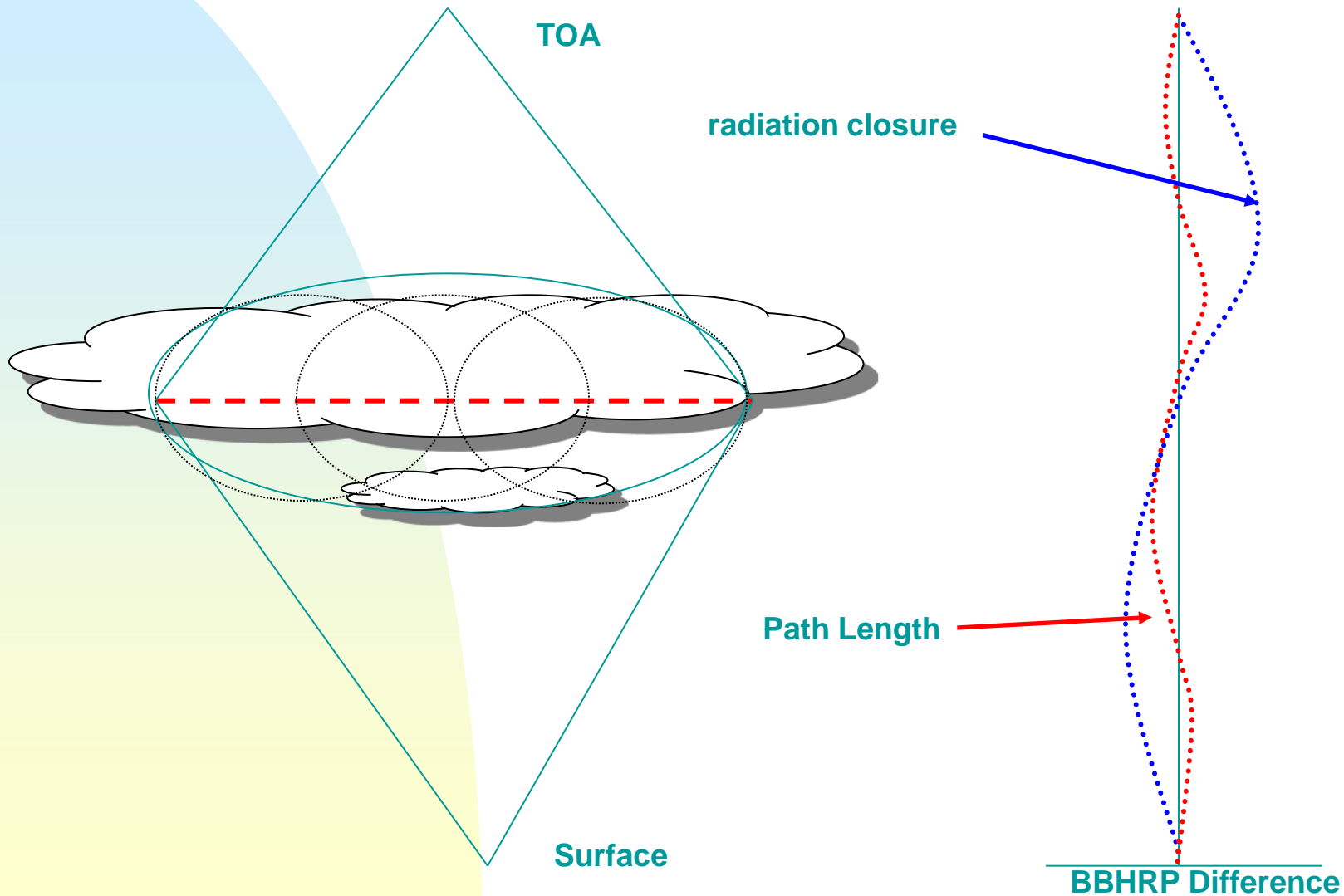


# High-resolution Oxygen A-band Spectrometer (HABS) and Photon path distribution

**Qilong Min**

**ASRC, State University of New York at Albany**

# Validate Broadband Heating Rate Profiles (BBHRP)

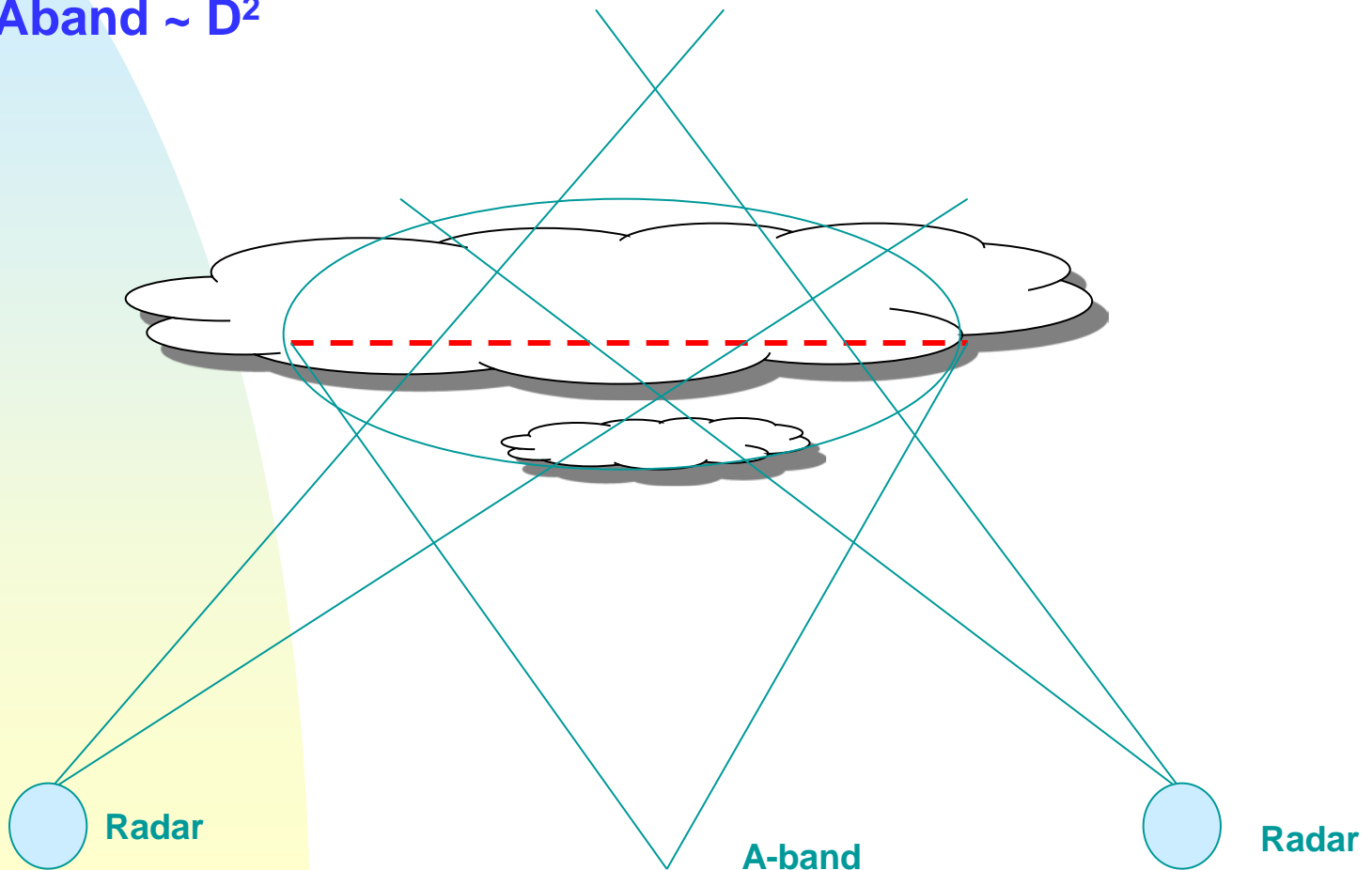


# AVA and A-band (Active and passive )

LWP  $\sim D^3$

R\_radar  $\sim D^6$

T\_Aband  $\sim D^2$



# Radiative Transfer and Photon Pathlength Distribution

Equivalence Theorem: [Irvine, 1964]

$$I_{\nu}(\mu, \phi; \mu_0, \phi_0) = I_0(\mu, \phi; \mu_0, \phi_0) \int_0^{\infty} p(l, \mu, \phi; \mu_0, \phi_0) e^{-\kappa_{\nu} l} dl$$

Where  $p(l)$  is photon path length distribution with path length  $l$

$\kappa_{\nu}$  is gaseous absorption coefficient

$I_0(\mu, \phi; \mu_0, \phi_0) \Rightarrow$  cloud optical properties

$p(l, \mu, \phi; \mu_0, \phi_0) \Rightarrow$  cloud geometry

$\Rightarrow$  Radiation field

- Mean path length to ensure the radiation closure at the surface and TOA
- Higher moments to ensure the accuracy of heating profiles

# Radiative Transfer and Photon Pathlength Distribution

## Equivalence Theorem: [Irvine, 1964]

$$I_\nu(\mu, \phi; \mu_0, \phi_0) = I_0(\mu, \phi; \mu_0, \phi_0) \int_0^\infty p(l, \mu, \phi; \mu_0, \phi_0) e^{-\kappa_\nu l} dl$$

Where  $p(l)$  is photon path length distribution with path length  $l$   
 $\kappa_\nu$  is gaseous absorption coefficient

Taking inverse Laplace transform:

$$p(l, \mu, \phi; \mu_0, \phi_0) = L^{-1} \left( \frac{I_\nu(k_\nu, \mu, \phi; \mu_0, \phi_0)}{I_0(\mu, \phi; \mu_0, \phi_0)} \right) = L^{-1} R(k_\nu, \mu, \phi; \mu_0, \phi_0)$$

Radiation measurements of  $I_0(\mu, \phi; \mu_0, \phi_0)$  and  $R(k_\nu, \mu, \phi; \mu_0, \phi_0)$

provide a complete set for understanding RT in the atmosphere

# Photon Pathlength Distribution and Oxygen A-band

## Oxygen A-band: 759 -770 nm

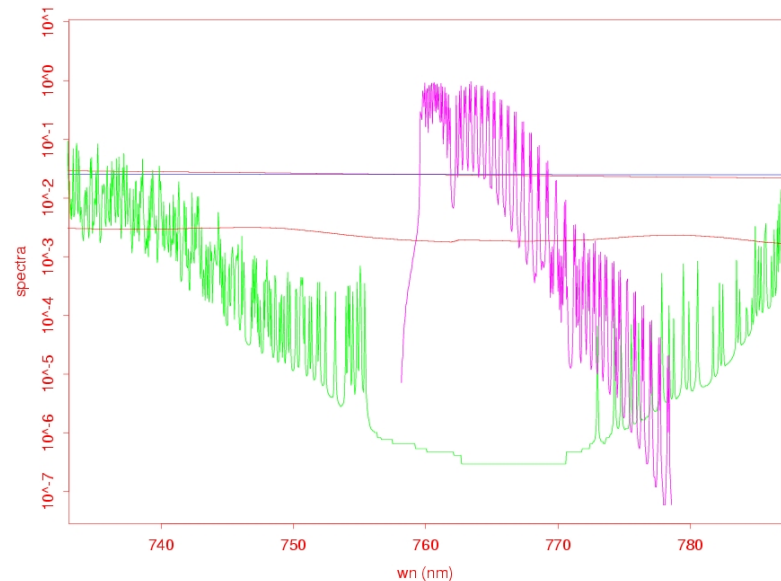
- Vertical profile of Oxygen is well known (uniform mixture)
- No other absorbers interfere within Oxygen A-band
- A number of absorption lines cover a suitable dynamic range
- The lines are regular and “looser” to be resolved by an instrument
- 760nm is the central wavelength to represent SW

**In the lower atmosphere, the line shape of Oxygen A-band**

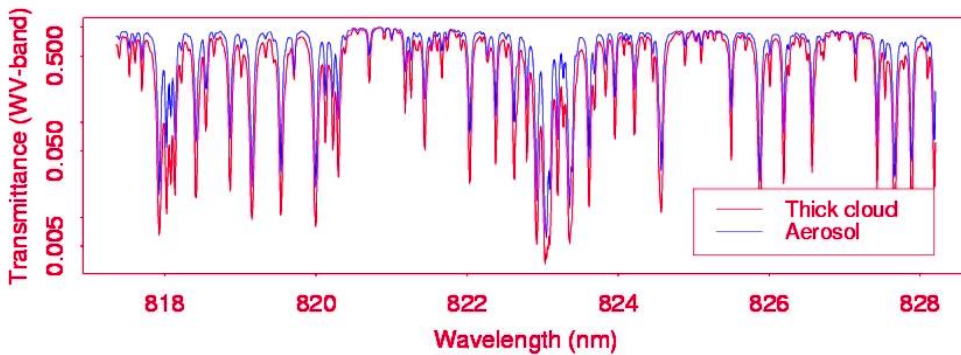
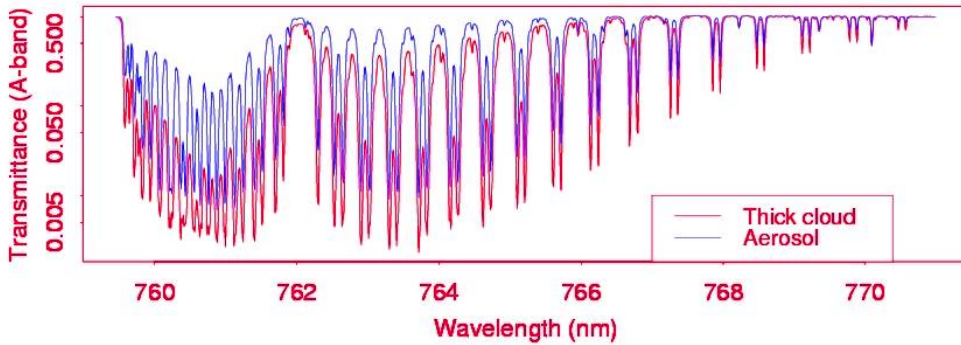
$$k_i = \frac{S_i}{\pi} \frac{\alpha_i}{(\nu - \nu_i)^2 + \alpha_i^2}$$

$$\alpha_i = \alpha_i^0 \frac{p}{p_0} \left(\frac{T_0}{T}\right)^{1/2}$$

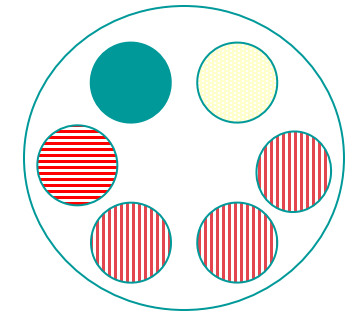
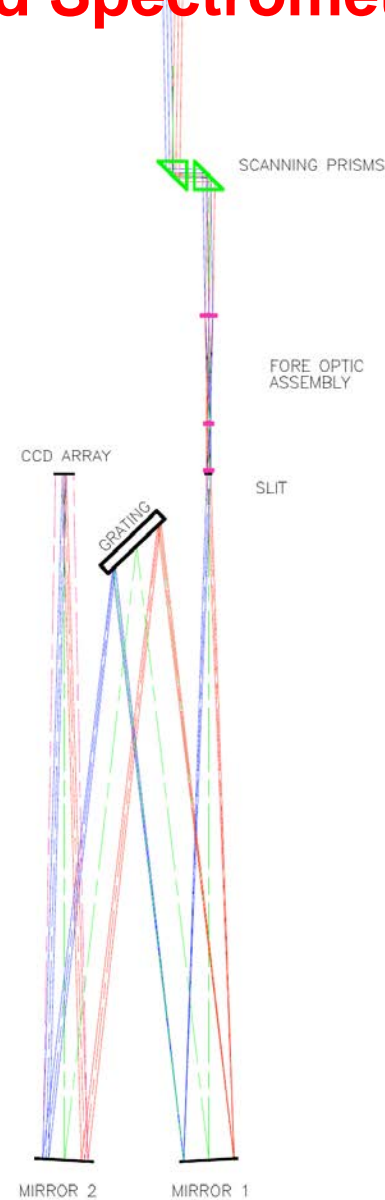
$$S_i = S(T_0) \frac{T_0}{T} \exp[1.439E'' \left(\frac{1}{T_0} - \frac{1}{T}\right)]$$



# High resolution A-band and Water vapor band Spectrometer (HAWS)



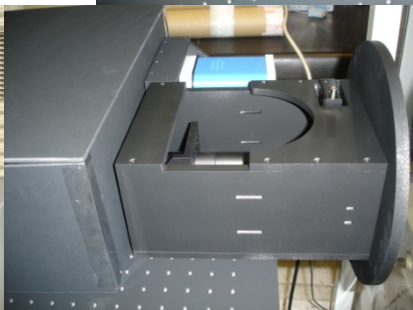
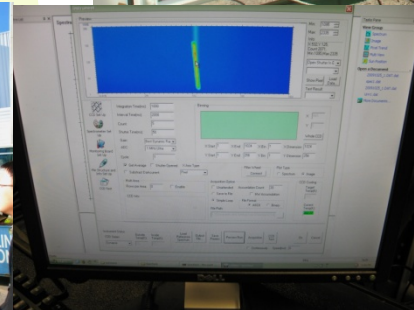
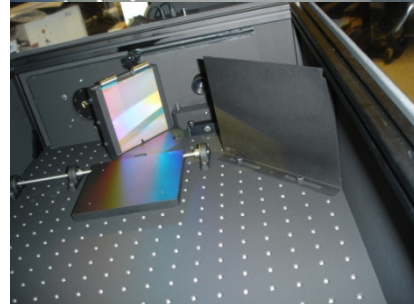
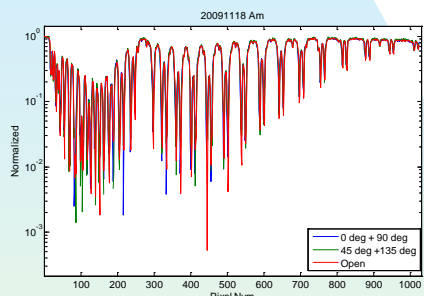
# High Resolution Oxygen A-band Spectrometer (HABS)



1. Open
2. 0-Polarization
3. 45-Polarization
4. 90-Polarization
5. 135-polarization
6. Diffusor

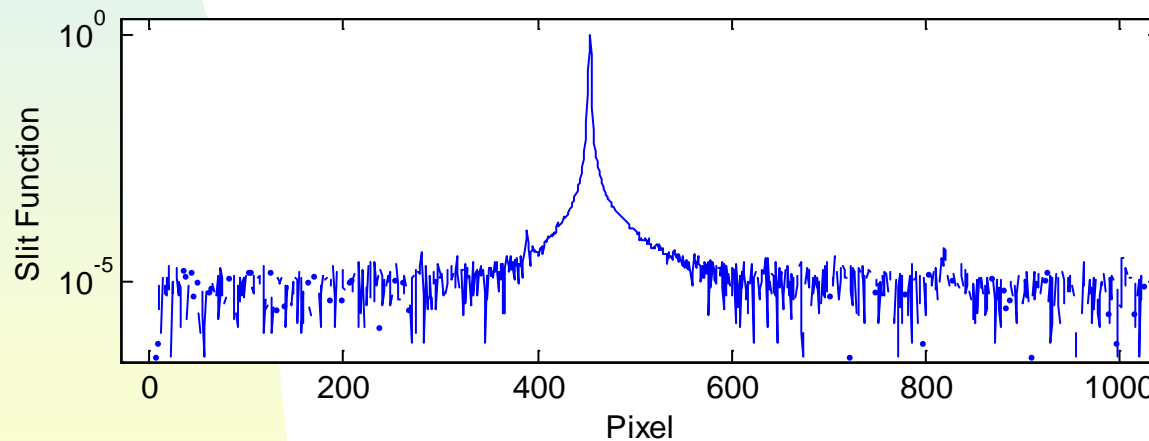
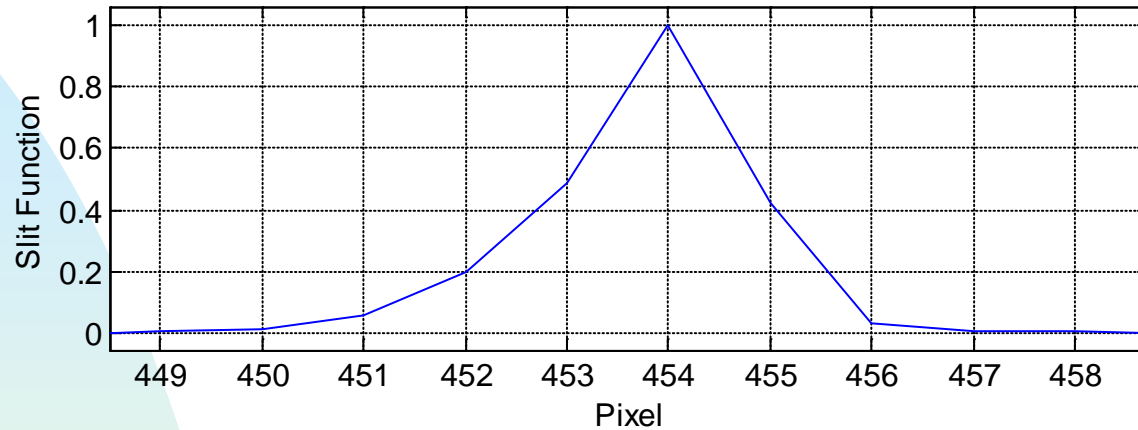


# High Resolution Oxygen A-band Spectrometer (HABS)



# High Resolution Oxygen A-band Spectrometer (HABS)

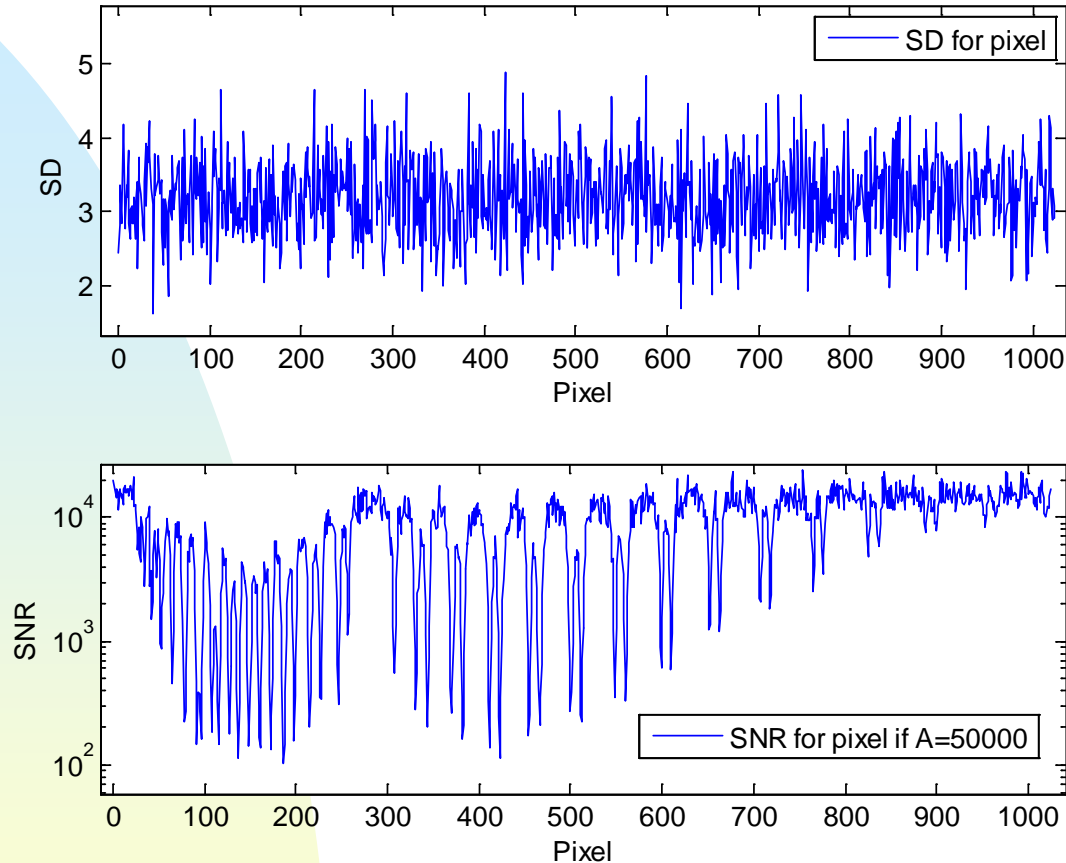
## Slit Function:



- the FWHM is about 1.85 pixels or 0.019 nm, which is better than 1/3 of wavenumber.
- The out of band rejection of the HABS is  $10^{-5}$ , better than the requirement.

# High Resolution Oxygen A-band Spectrometer (HABS)

## Signal-to-Noise Ratio (SNR):

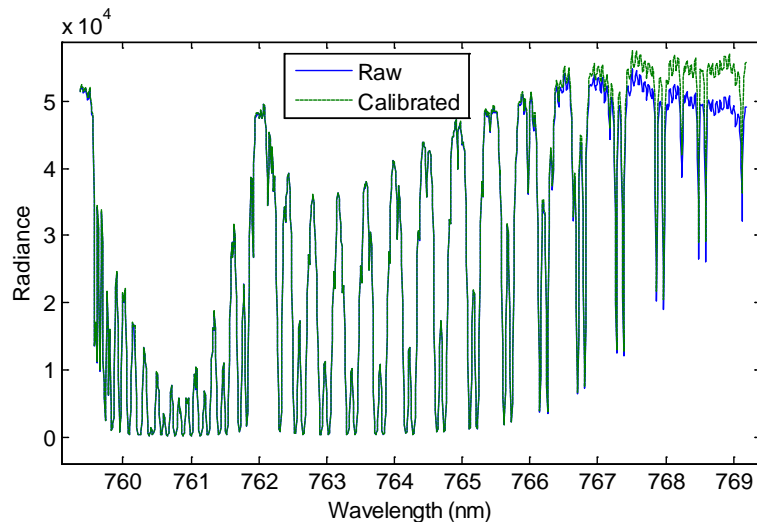
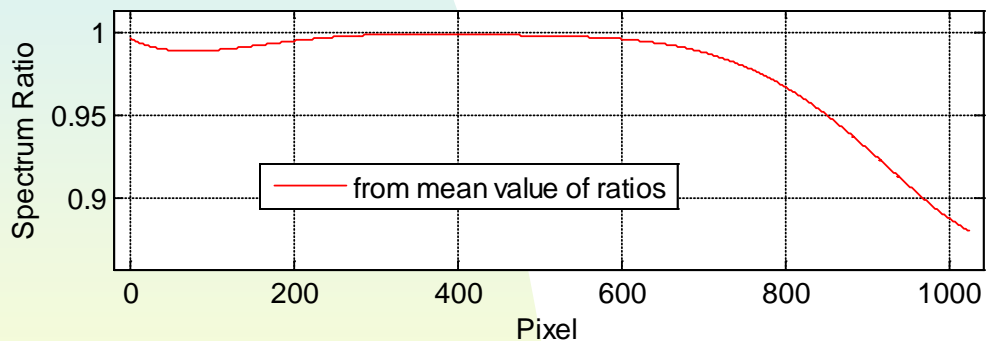
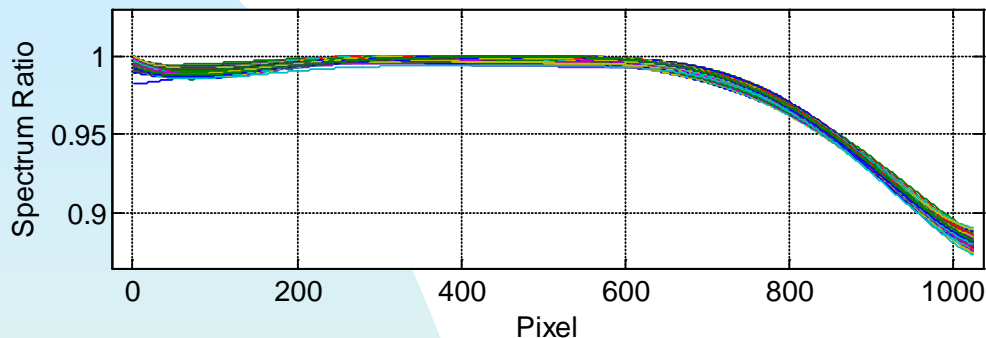


Our test shows that the maximum readout noise is 8.35 e, slightly larger than manufactory specification (6 e).

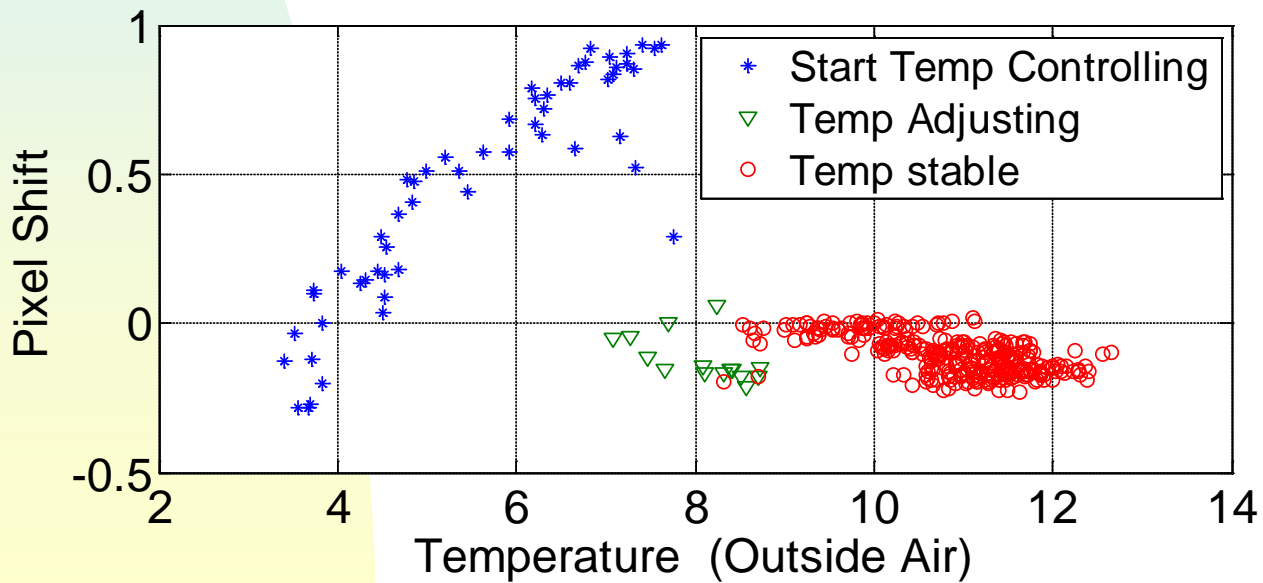
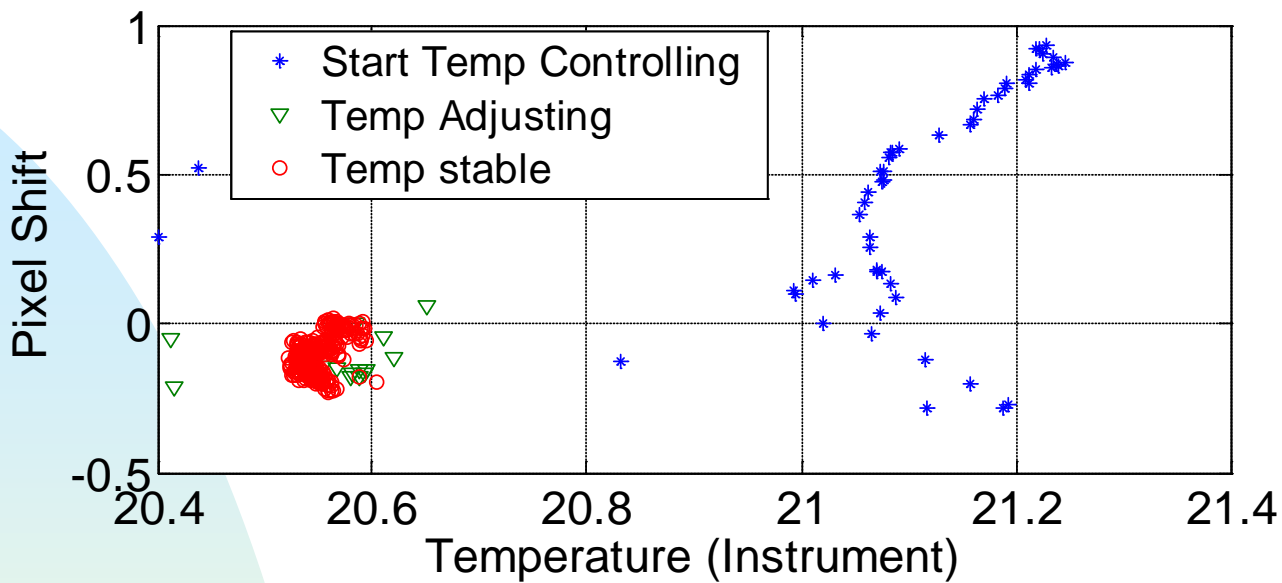
- With an optimized exposure for zenith measurements, the SNR is above 100 even for the darkest pixel.
- It demonstrates the SNR is better than our design requirement.

# High Resolution Oxygen A-band Spectrometer (HABS)

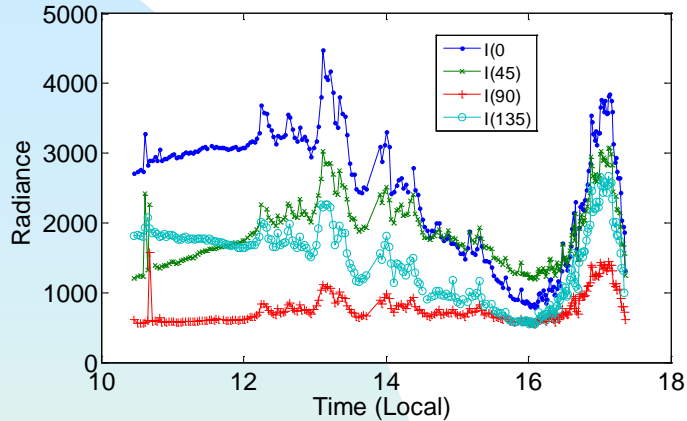
## Spectral calibration



The response of overall filter function has been calibrated using GS0937 lamp .  
The tail at the long wavelength is calibrated up and consists with radiance at the short wavelength.



# HABS polarization (and polarization Spectrum)

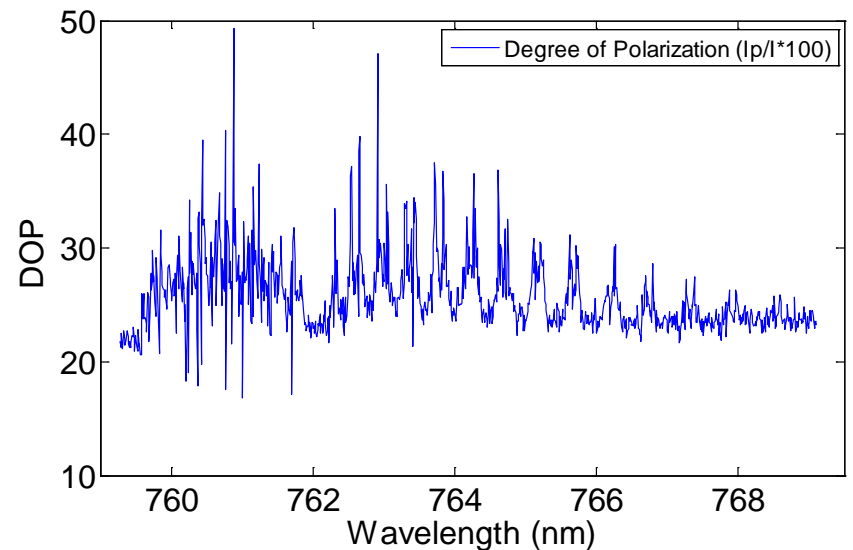
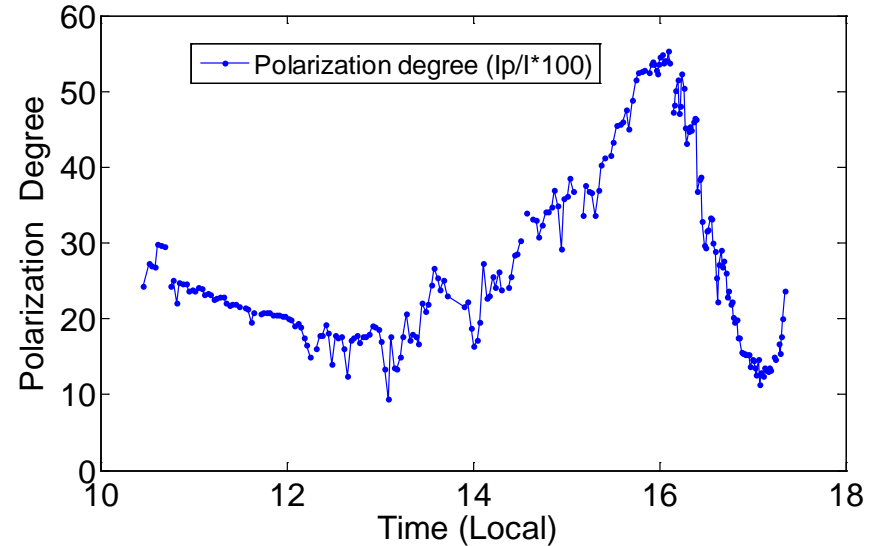
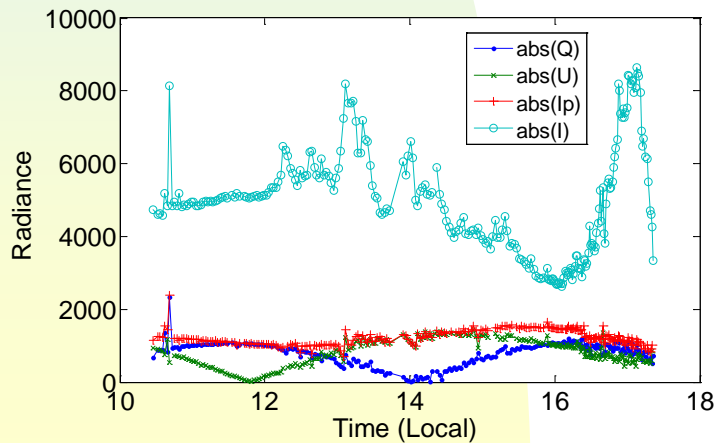


$$I_0' = I_2(0^\circ) + I_2(90^\circ) / R_G$$

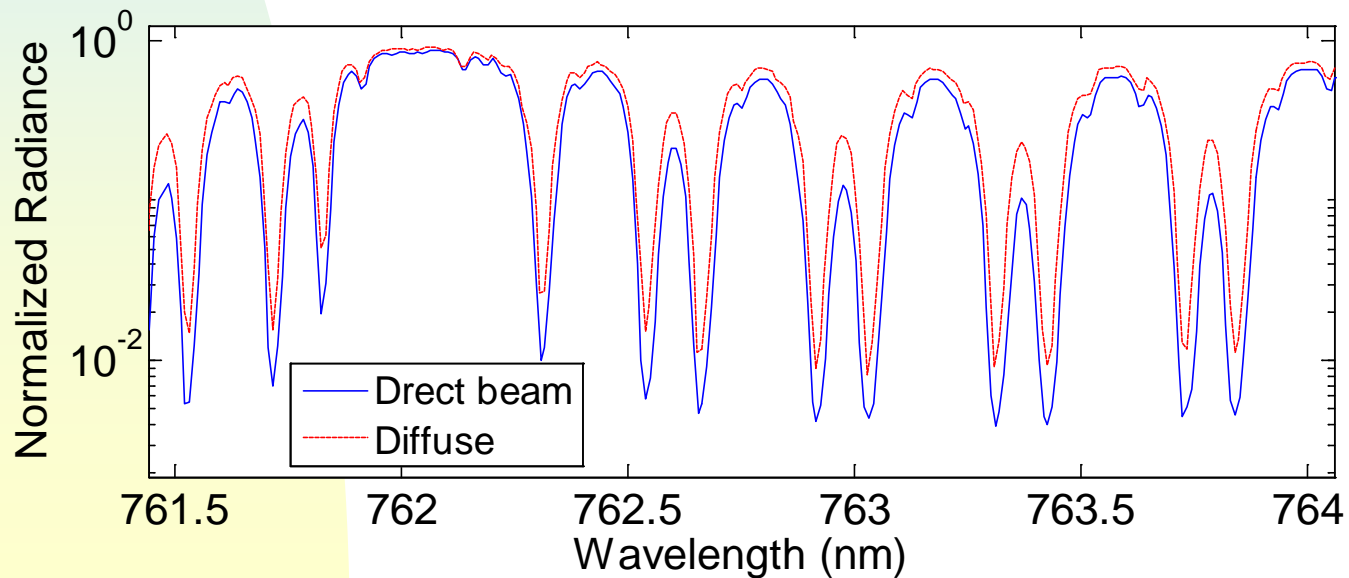
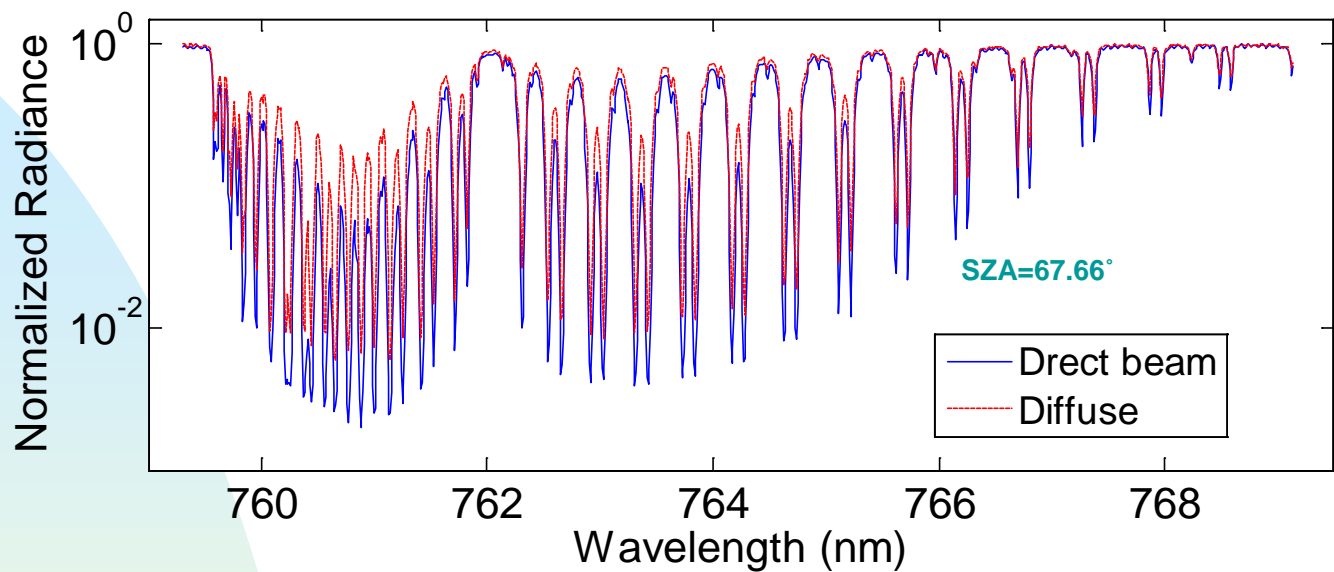
$$Q = I_2(0^\circ) - I_2(90^\circ) / R_G$$

$$U = (I_2(45^\circ) - I_2(135^\circ)) R_{12}$$

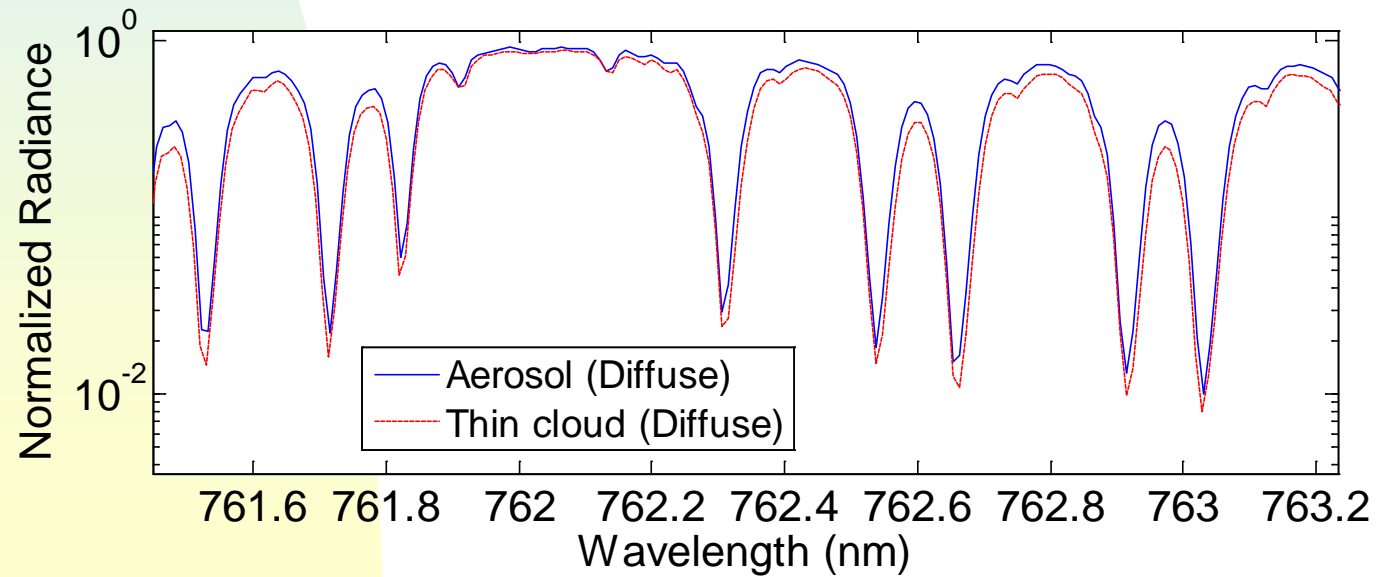
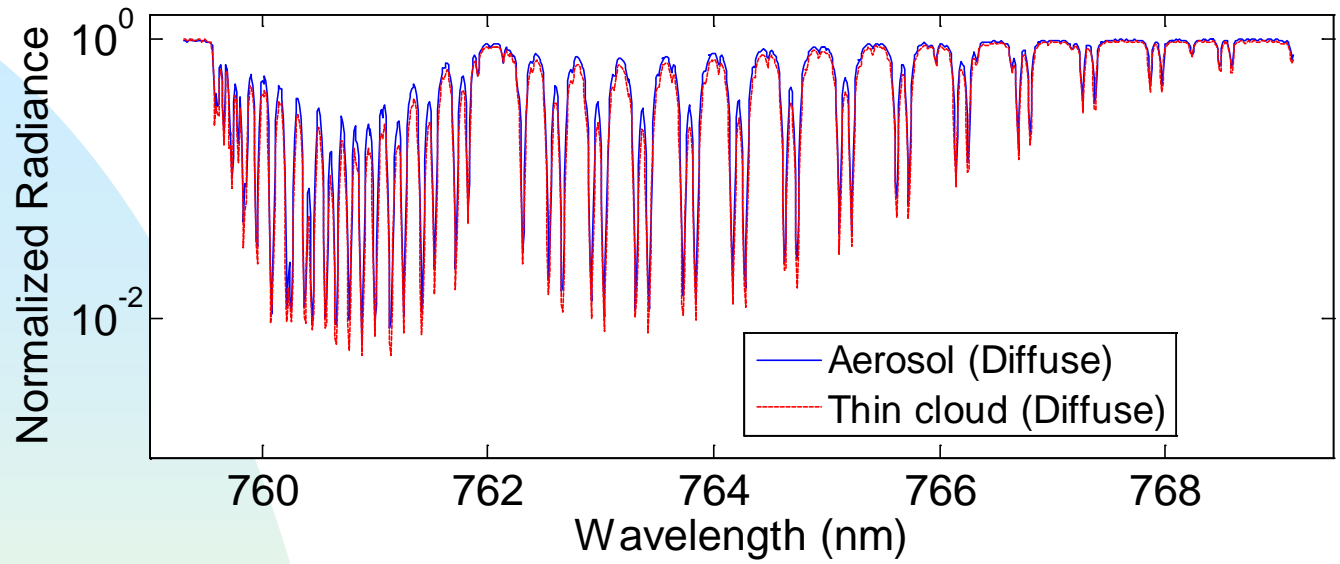
$$I_{0P}' = \sqrt{Q^2 + U^2}$$



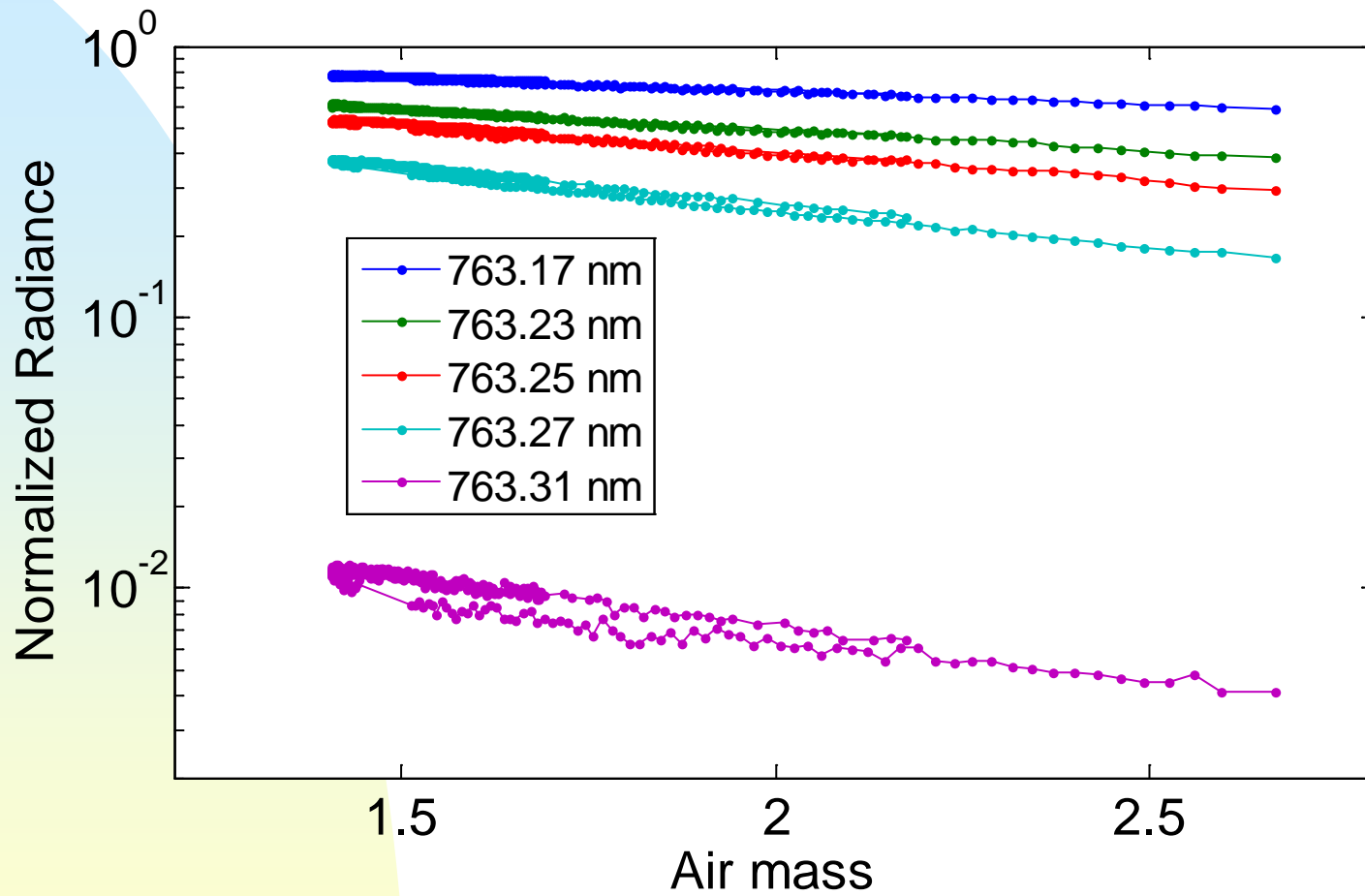
## Direct beam vs. diffuse



# Aerosol vs. Thin cloud



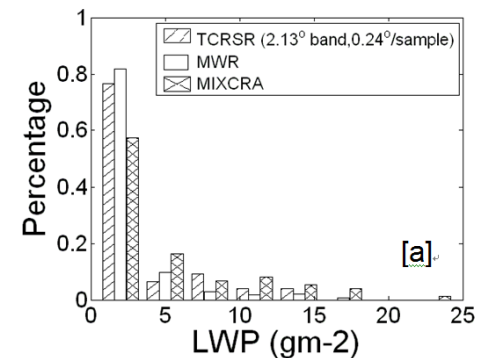
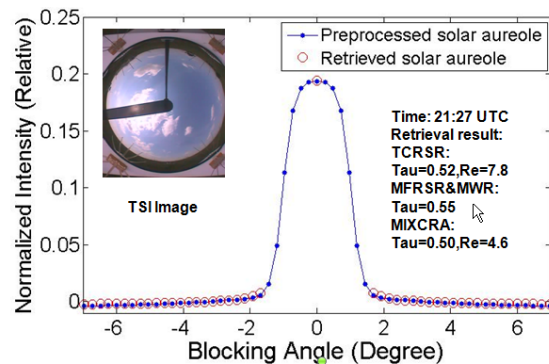
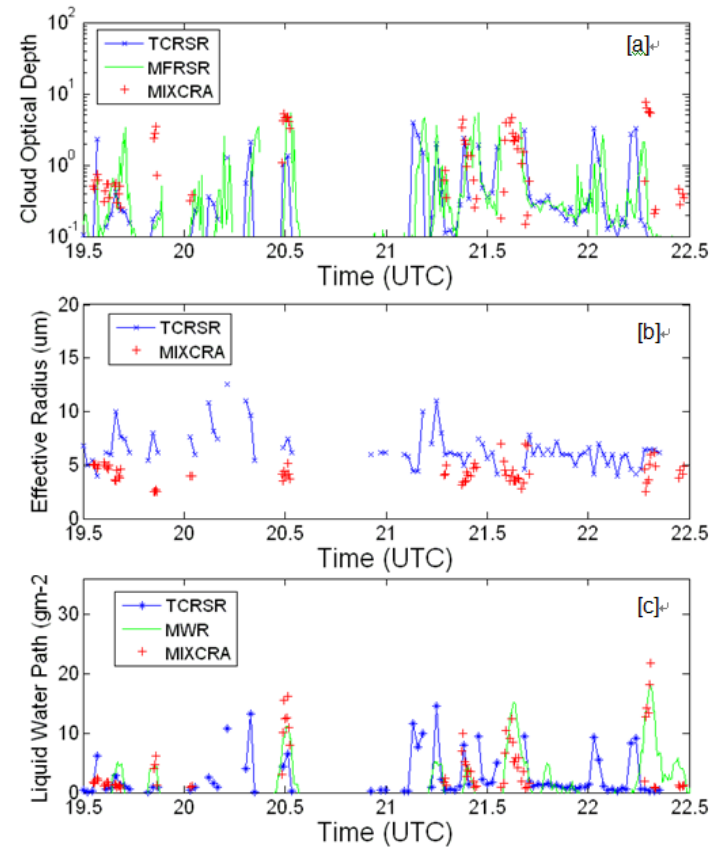
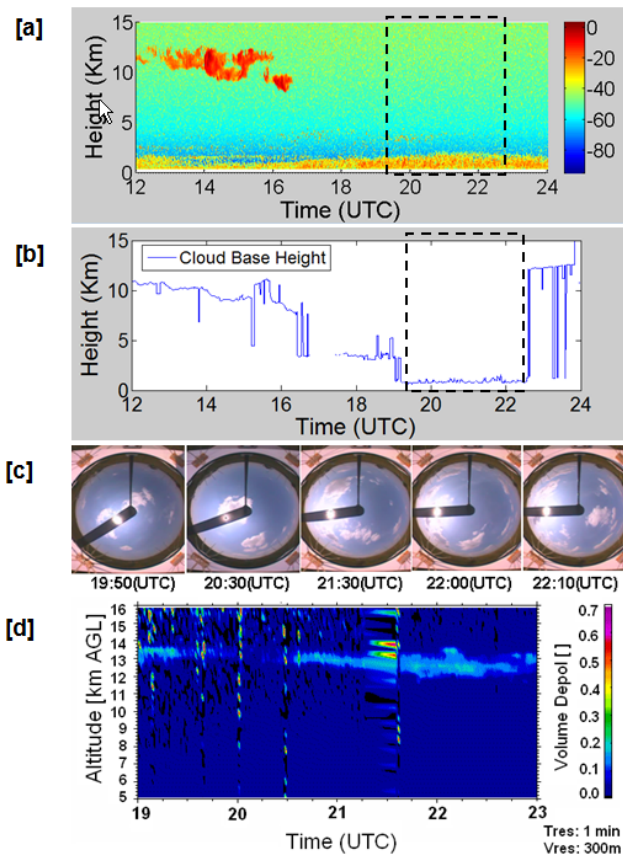




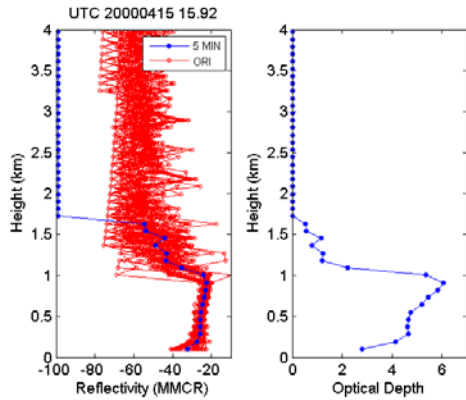
## High Resolution Oxygen A-band Spectrometer (HABS)

- Has the capability to measure both zenith and direct-beam radiances with a field of view of 2.7 degrees. The direct-beam measurements can be used to calibrate the spectrometer and construct retrieval kernels for zenith measurements.
- Measures polarizations of A-band spectra with four polarizers, which substantially enhances the retrieval ability for aerosols and ice clouds.
- Achieves an out-of-band rejection of  $10^{-5}$ , a resolution of better than  $0.3 \text{ cm}^{-1}$ , and a high signal-to-noise ratio

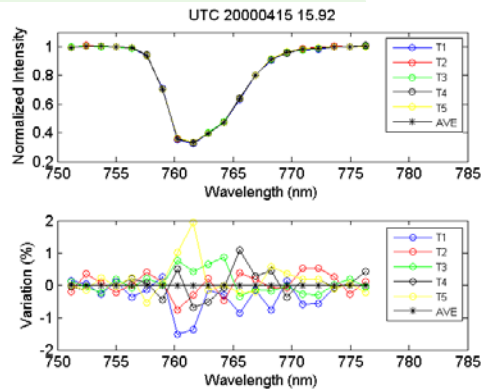
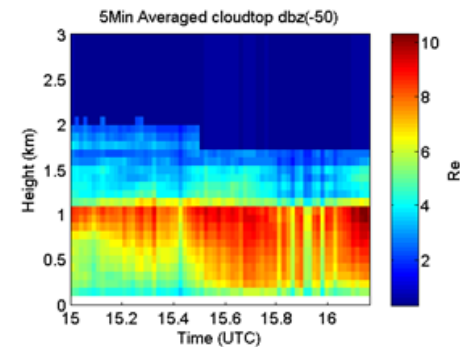
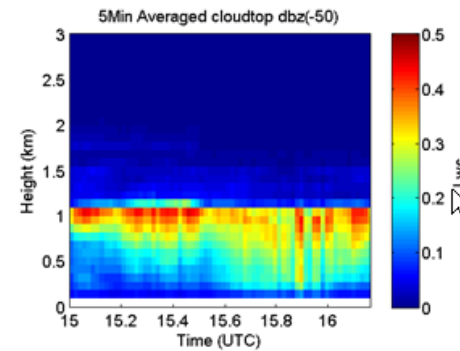
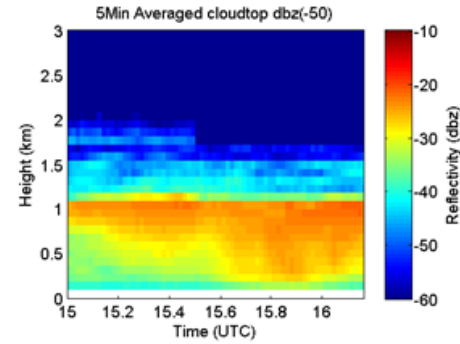
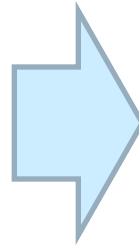
# TCRSR retrievals for optically thin clouds: poster by Yin et al.



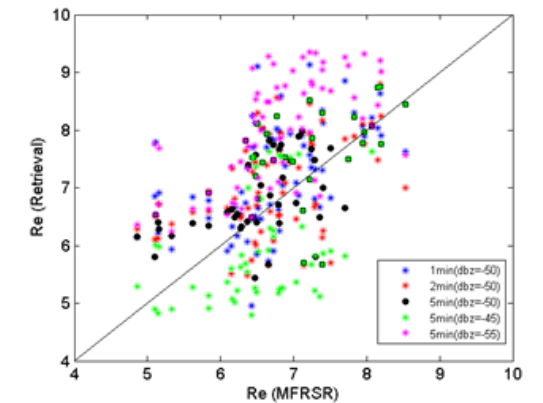
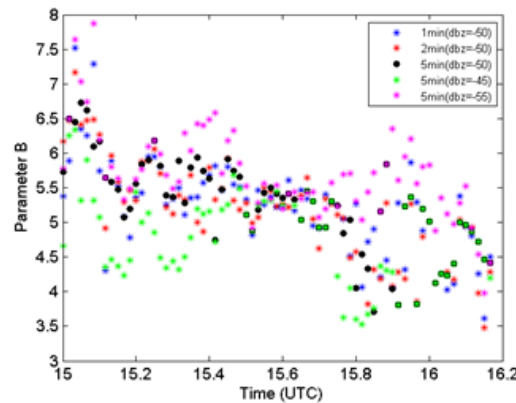
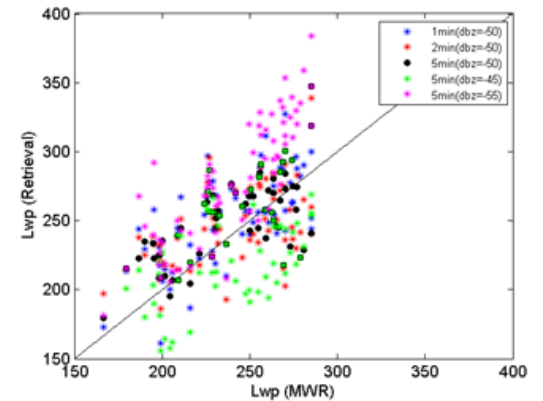
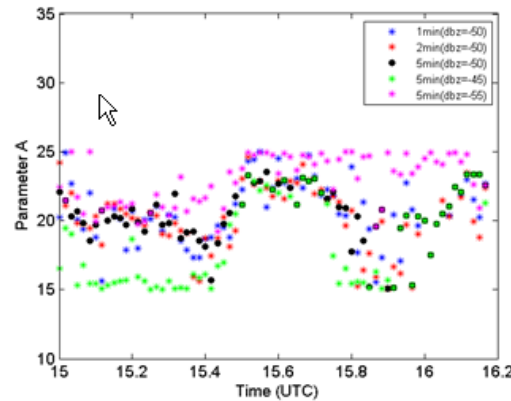
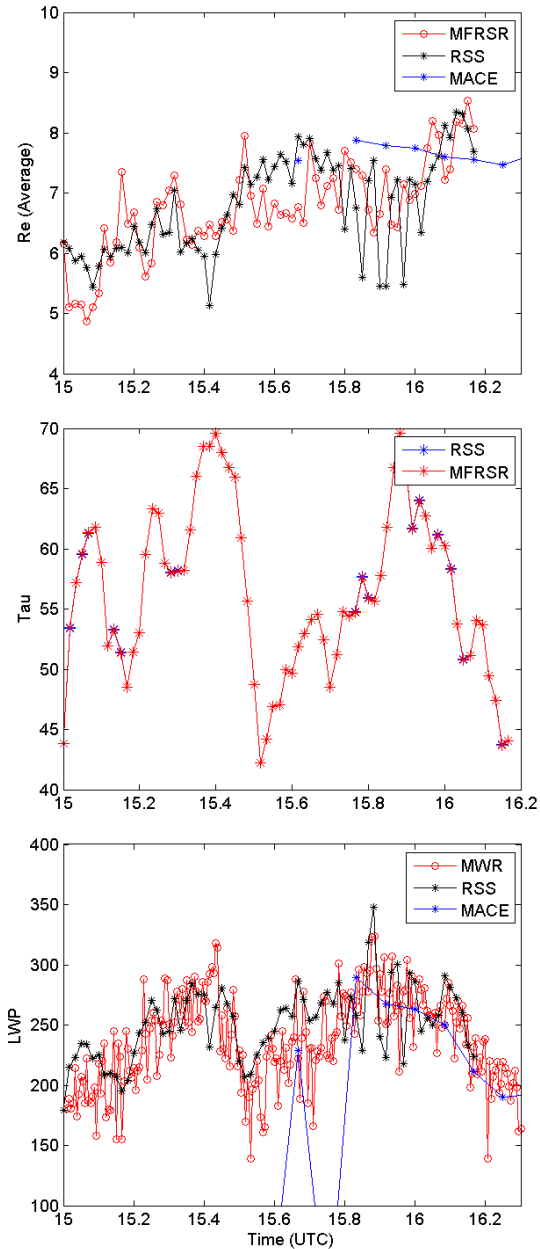
# Combining Radar and A-band to retrieve cloud optical property profiles:



MMCR reflectivity:  $Re = a f(Z)$  &  $LWC = b f(Z)$   
 RSS A-band and MFRSR Tau



# Combining Radar and A-band to retrieve cloud optical property profiles:

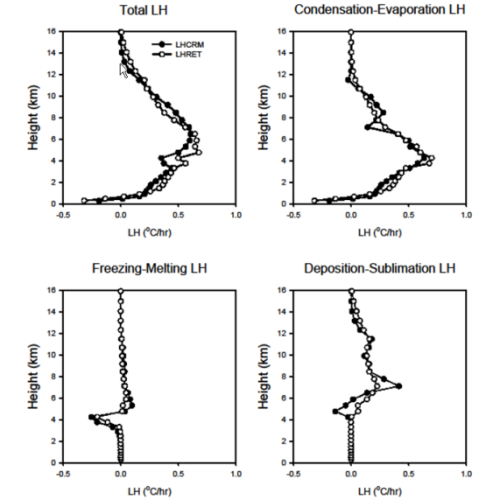
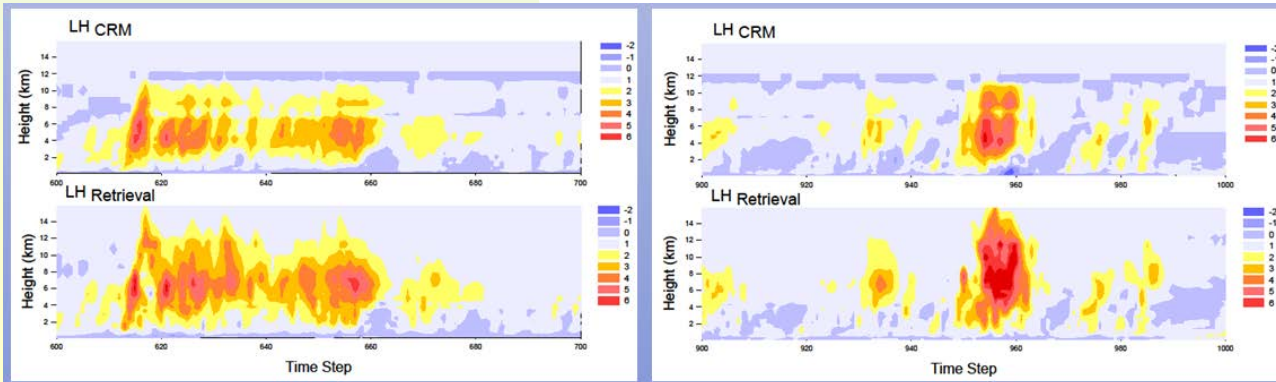
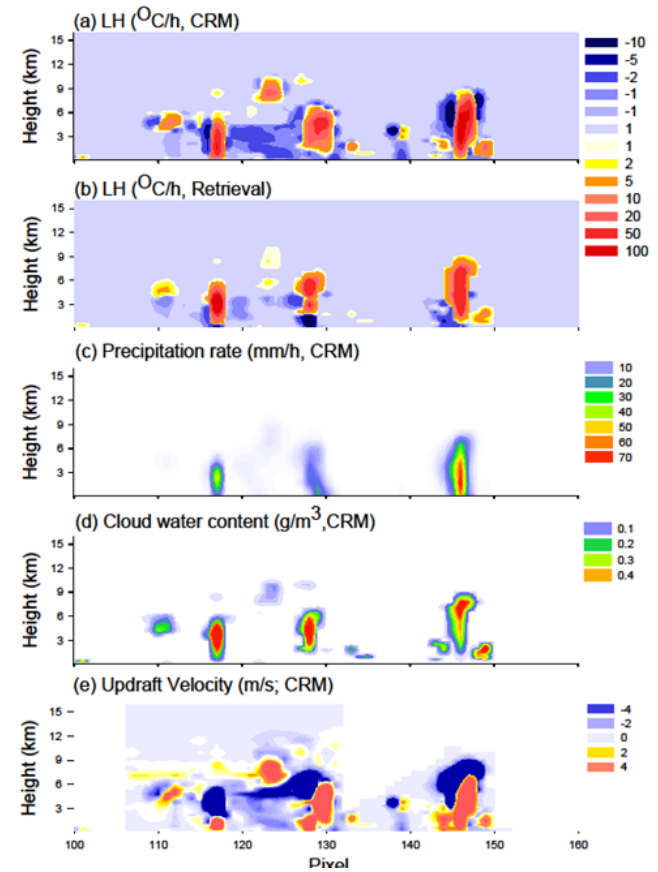
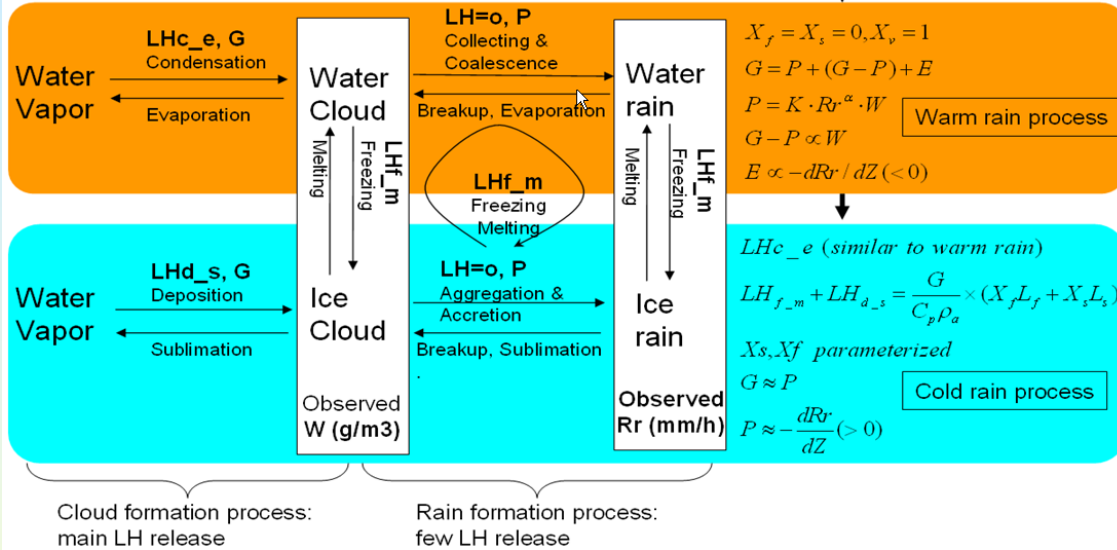


Retrievals are sensitive to the top portion of clouds and temporal variation

# Retrieving Latent Heat Vertical Structure Using Precipitation and Cloud Profiles and Cloud Resolving Model Simulations

## Poster by Li, Min, and Wu

$$LH = LH_{c\_e} + LH_{f\_m} + LH_{d\_s} = \frac{G}{C_p \rho_a} \times (X_v L_v + X_f L_f + X_s L_s)$$



# Sketch map of the O2 A-Band Spectrometer

