

The $F_L(x, Q^2)$ Structure Function.

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The neutral current inclusive e-nucleon cross section can be written as a combination of two terms:

$$\sigma_T(x, Q^2, y) = \frac{d^2\sigma}{dx dQ^2} \cdot \frac{Q^4 x}{2\pi\alpha^2 Y_+} = F_2(x, Q^2) - \frac{y^2}{Y_+} \cdot F_L(x, Q^2),$$

“reduced cross section”

$F_2 \propto (\sigma_T + \sigma_L)$ $F_L \sim \sigma_L$ where σ_T and σ_L are the γ^*P cross sections for transverse and longitudinally polarized virtual photons. The value of F_2 is determined by the sum of quark and anti-quark distributions, whereas F_L depends on quark + gluon distributions. Above some value of Q^2 , quark PDFs are much smaller than gluon PDFs hence F_L is mainly “driven” by gluon PDFs

The measurement has experimental difficulties:

The highest sensitivity is obtained at high values of $y = Q^2/sx$, at low values of Q^2 high y corresponds to low energy of the scattered e that can be obscured by hadronic background. (One possible reason to have redundant electron PID).

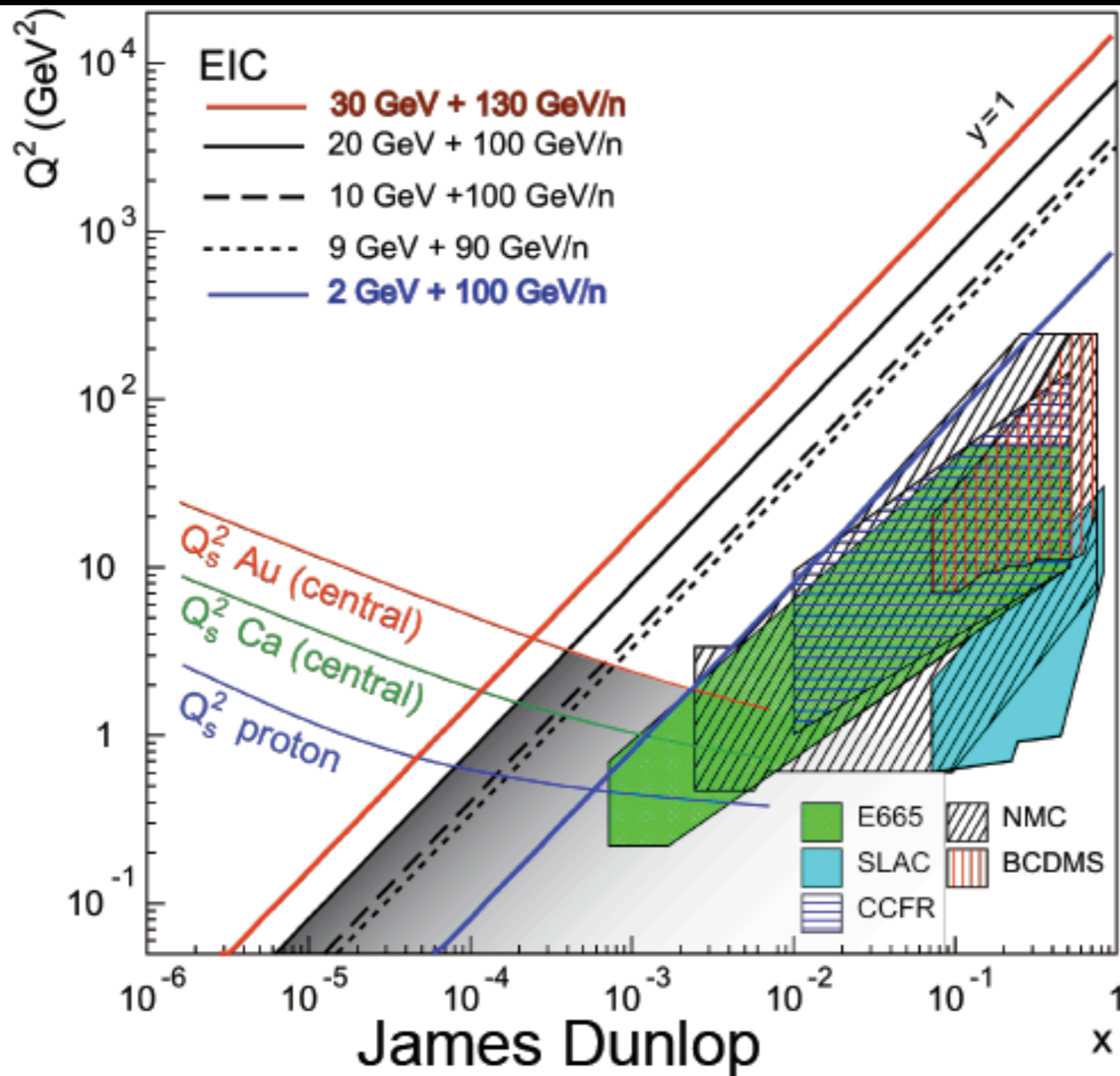
The FL measurements is done at different values of y for fixed x and Q^2 . (at HERA they fix the e^+ beam energy and change the energy of the proton beam.

Jamie does something similar with two calculators of e - p scattering GRB98 and MRST.

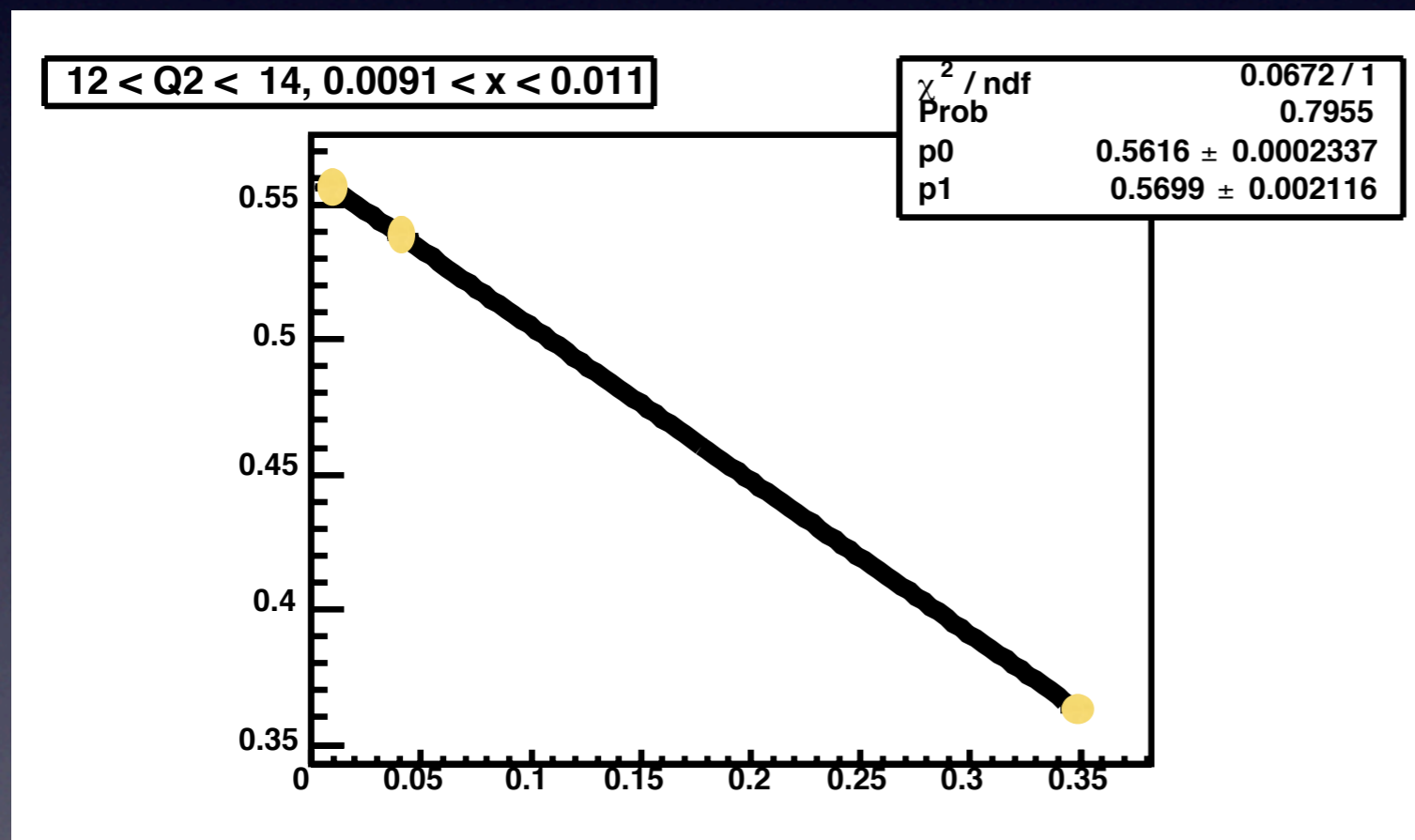
Both calculations produce cross sections in bins of x and Q^2 at six beam energy combinations:

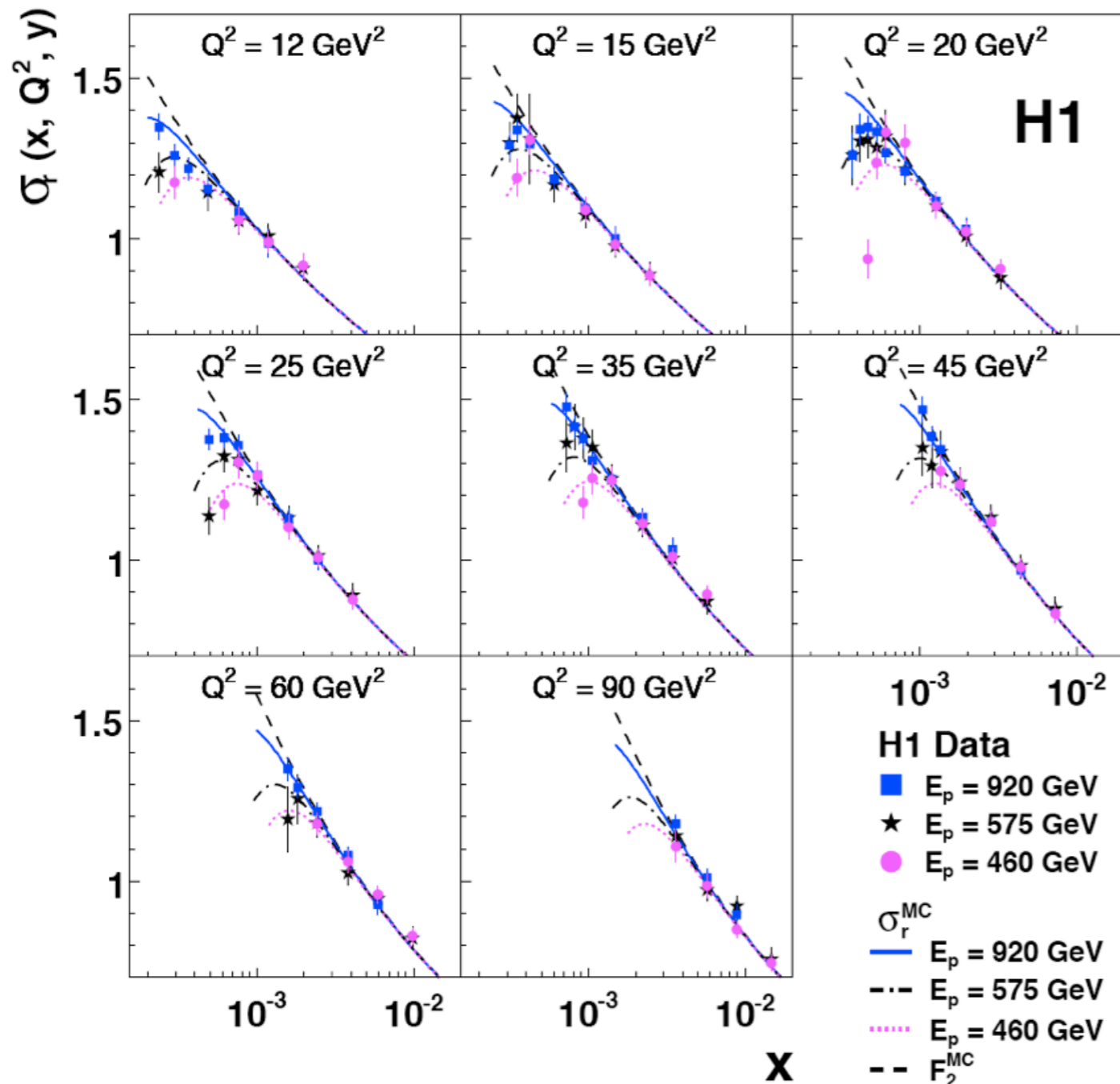
data ela / 5., 10., 5., 10., 10., 5./

data epa / 50., 50., 250., 250., 100., 100./



For Q^2 and x bins the data is plotted as counts versus y^2/Y_+ and each plot is fitted to a line: the intercept is equal to F_2 and the slope is F_L .

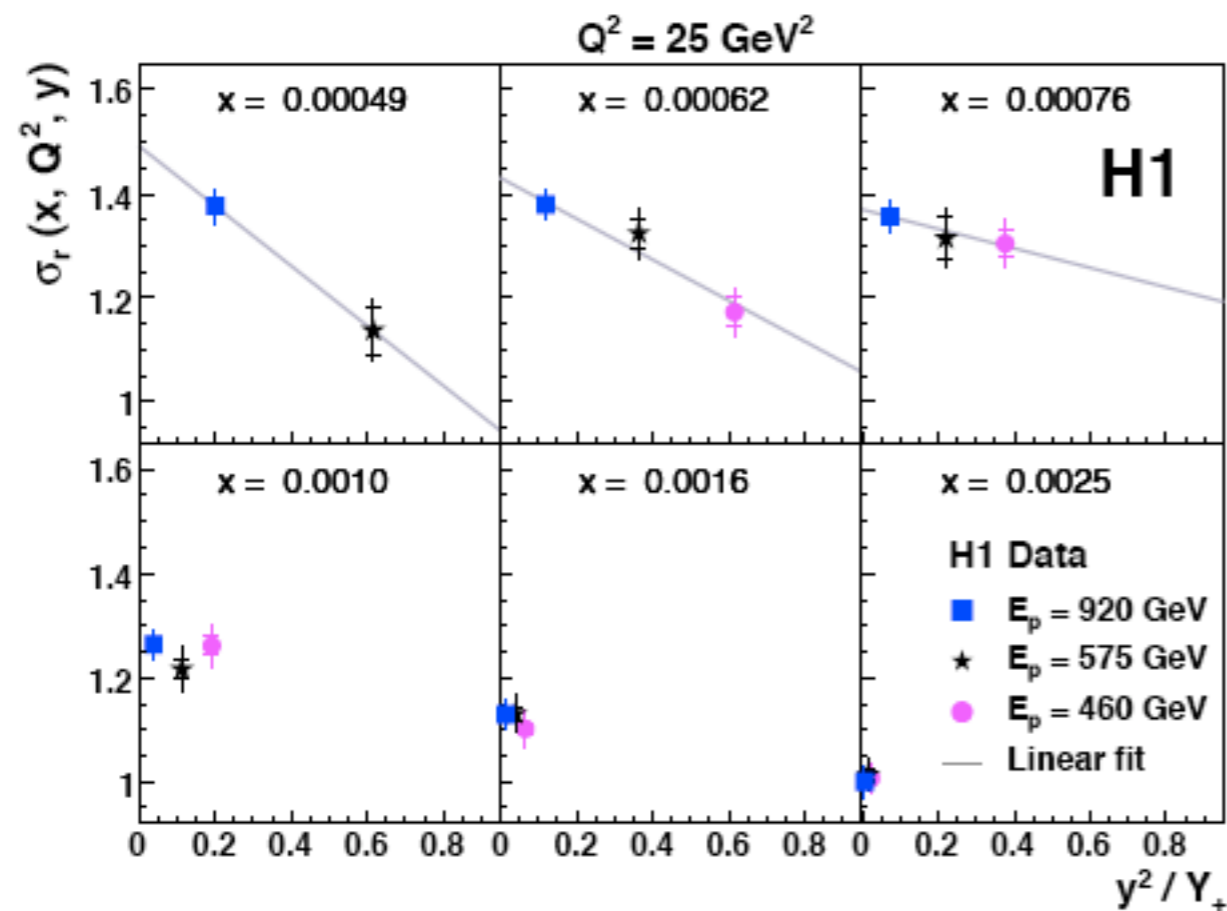




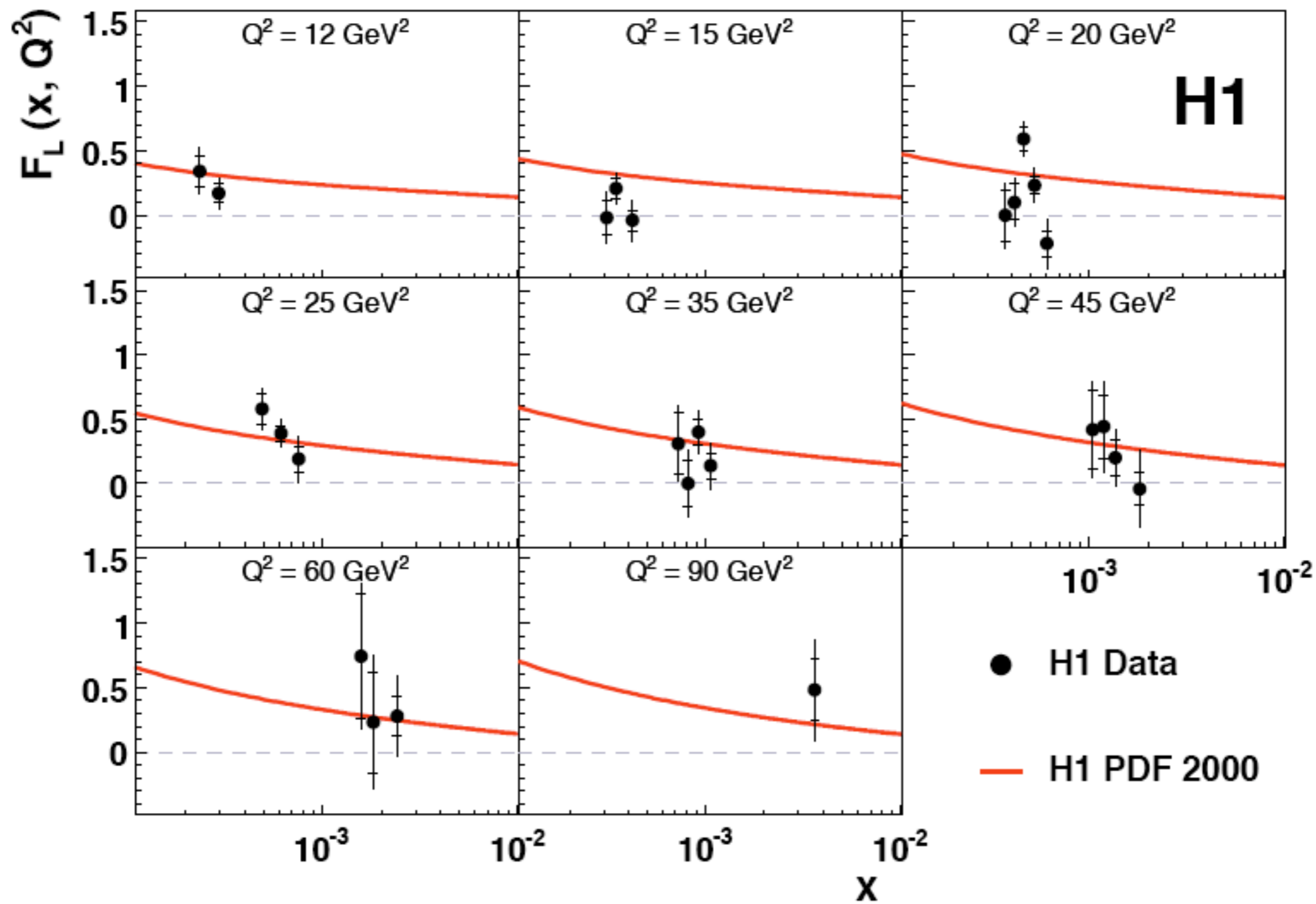
The first F_L measurement in a collider: H1 and Zeus.

H1 has a publication in the arXiv: [hep-ex:0805.2809v2](https://arxiv.org/abs/hep-ex/0805.2809v2)

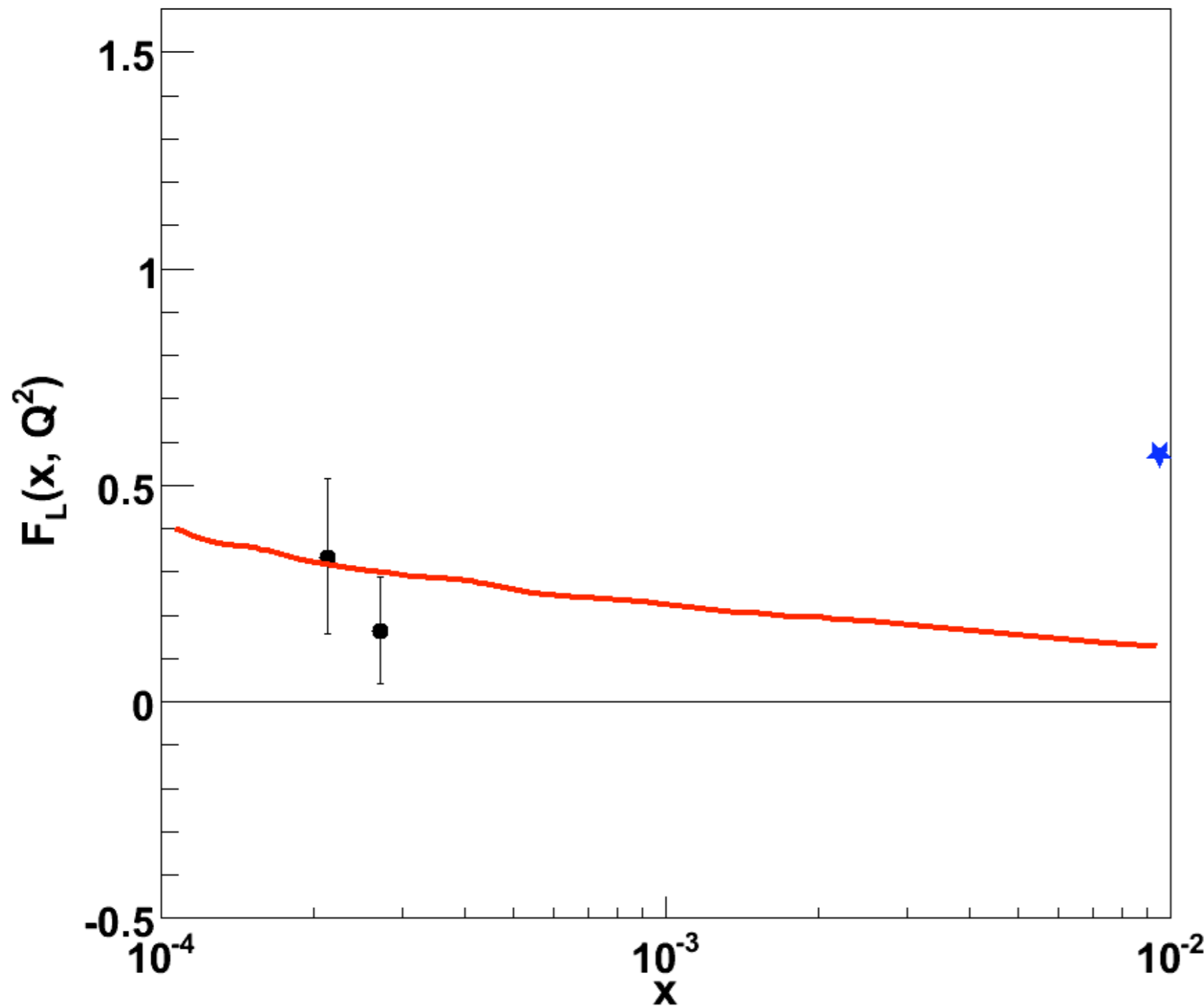
Zeus results are still a collaboration document.



Linear fits to measurements at different y at fixed x and Q^2 .



HI measurement at $Q^2=12 \text{ GeV}^2$



The **blue star** is a value obtained by Jamie with the MRST calculation. (Statistical errors smaller than the marker)

HI at $Q^2=12 \text{ GeV}^2$

Things to do:

- Extend the range of y .
- Produce data in the HI range to see how good are these calculations.
- Event generator with a simple detector geometry.
- Try other models.