

Modeling the Atmospheric Transport and Deposition of Mercury



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Materials assembled for
“Mercury in Maryland” Meeting, Appalachian Lab,
Univ. of Maryland Center for Environmental Science
301 Braddock Road, Frostburg MD, Nov 2-3, 2005

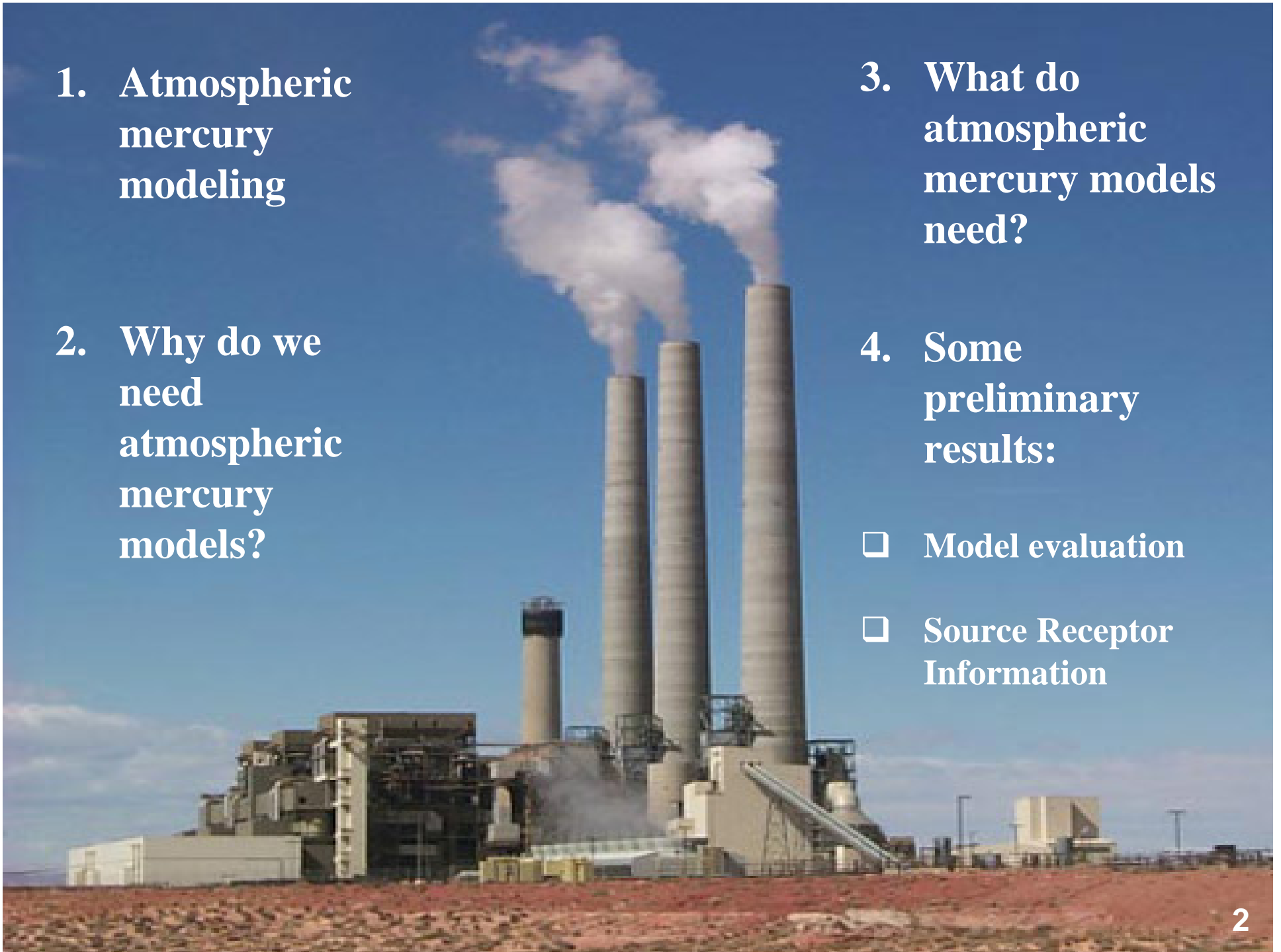
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2. Why do we need atmospheric mercury models?

3. What do atmospheric mercury models need?

4. Some preliminary results:

- Model evaluation
- Source Receptor Information



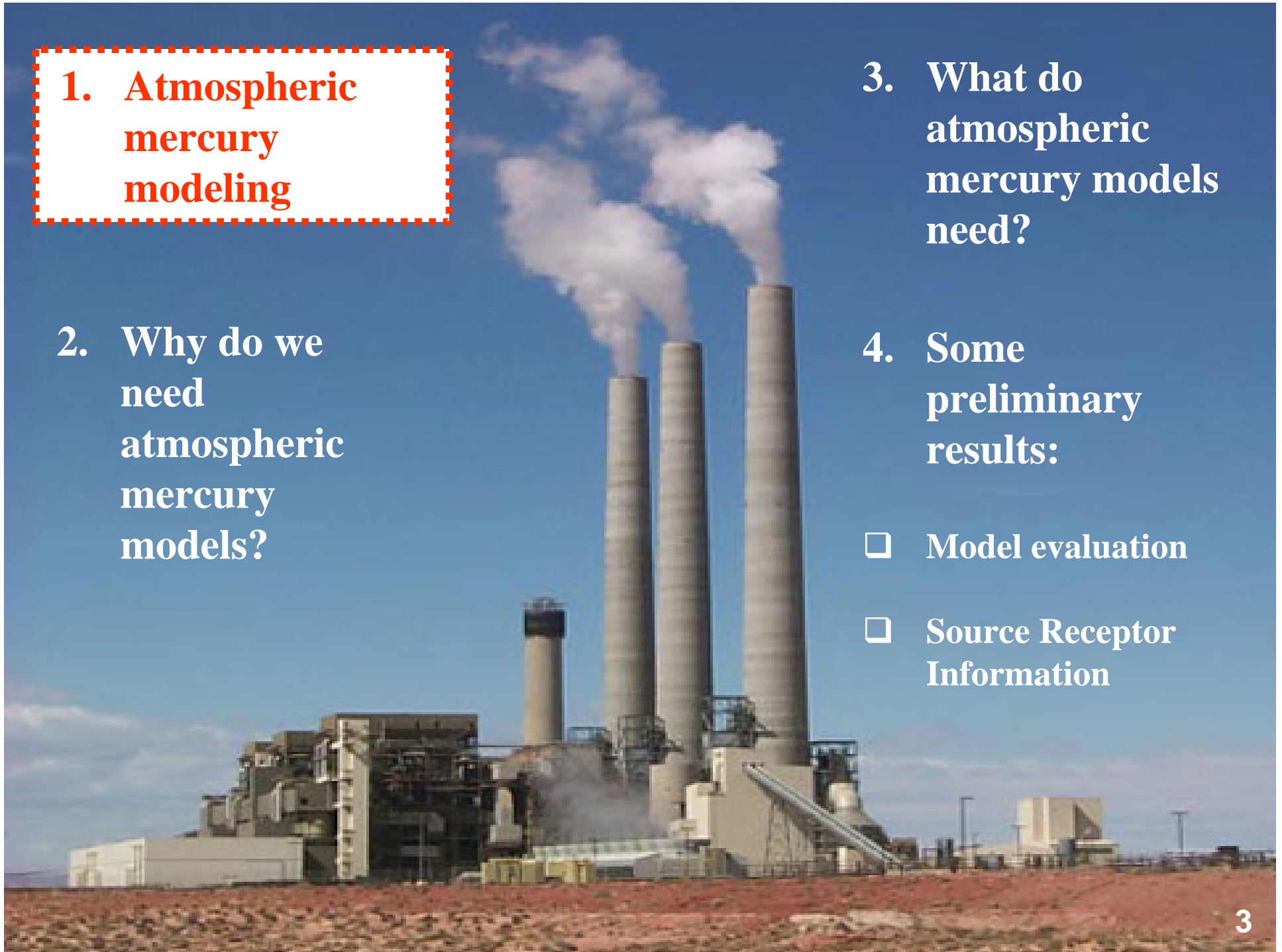
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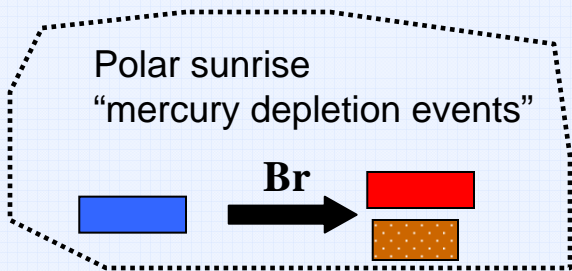
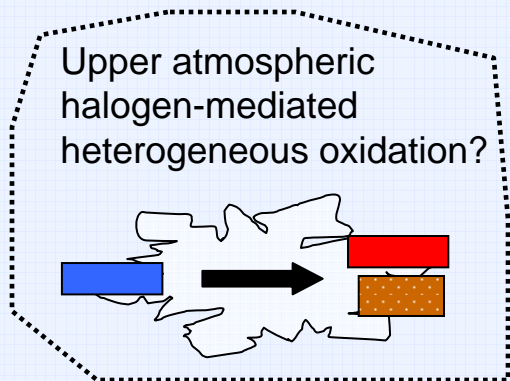
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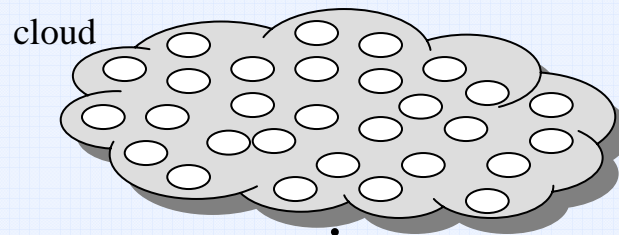
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Atmospheric Mercury Fate Processes



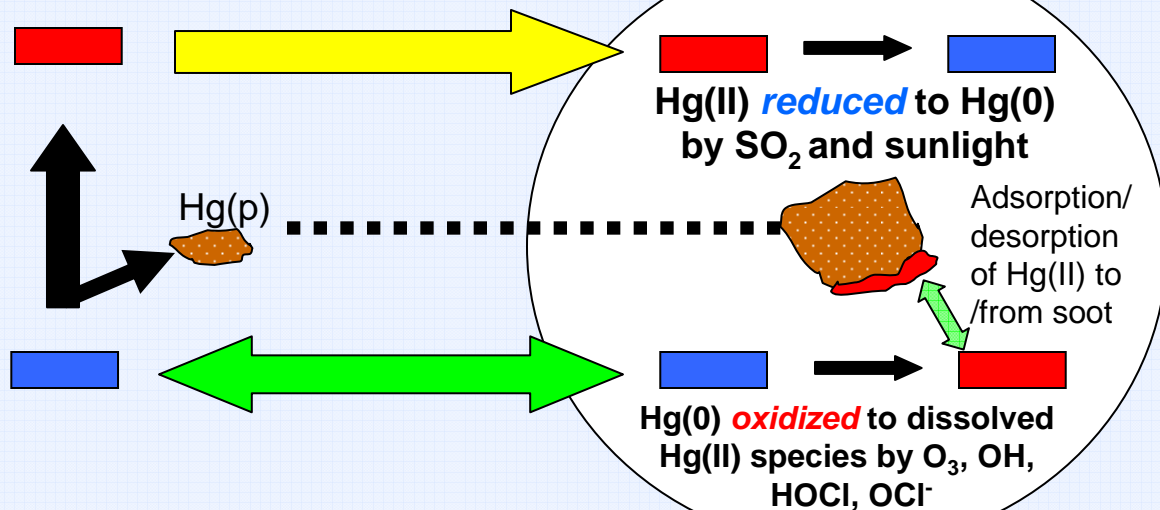
- Elemental Mercury [Hg(0)]
- Hg(II), ionic mercury, RGM
- Particulate Mercury [Hg(p)]



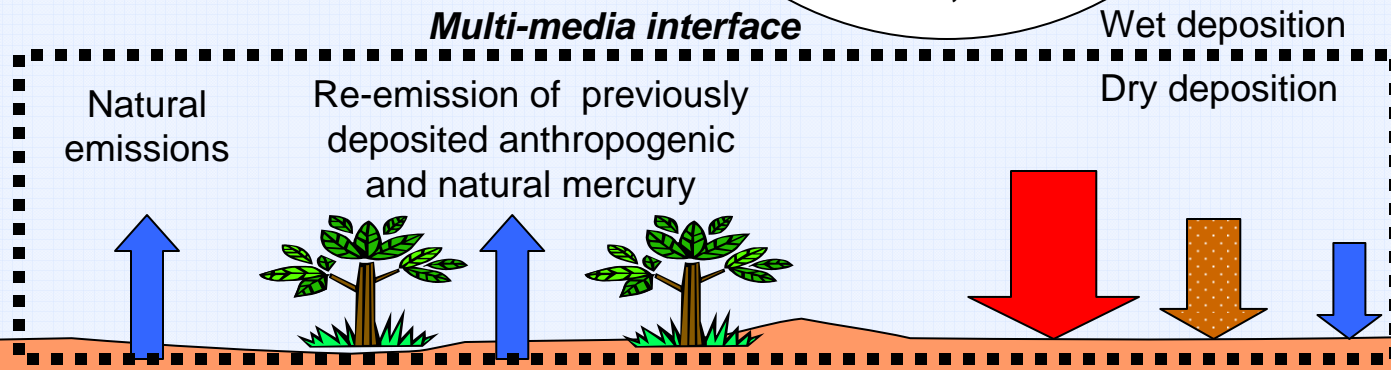
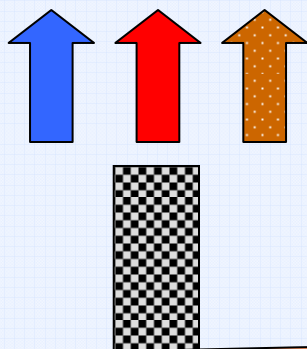
CLOUD DROPLET

Vapor phase:

Hg(0) oxidized to RGM and Hg(p) by O_3 , H_2O_2 , Cl_2 , OH, HCl

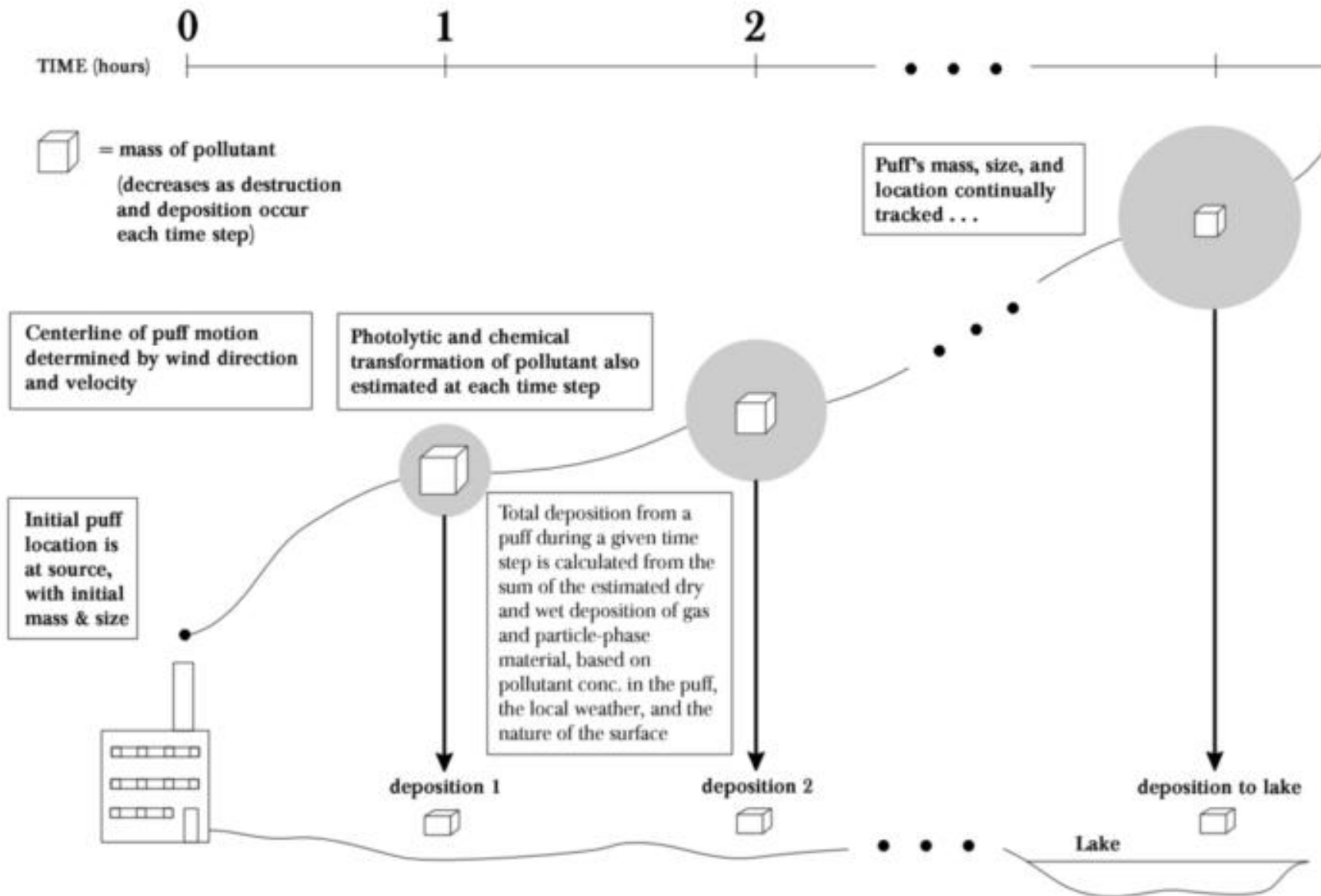


Primary Anthropogenic Emissions

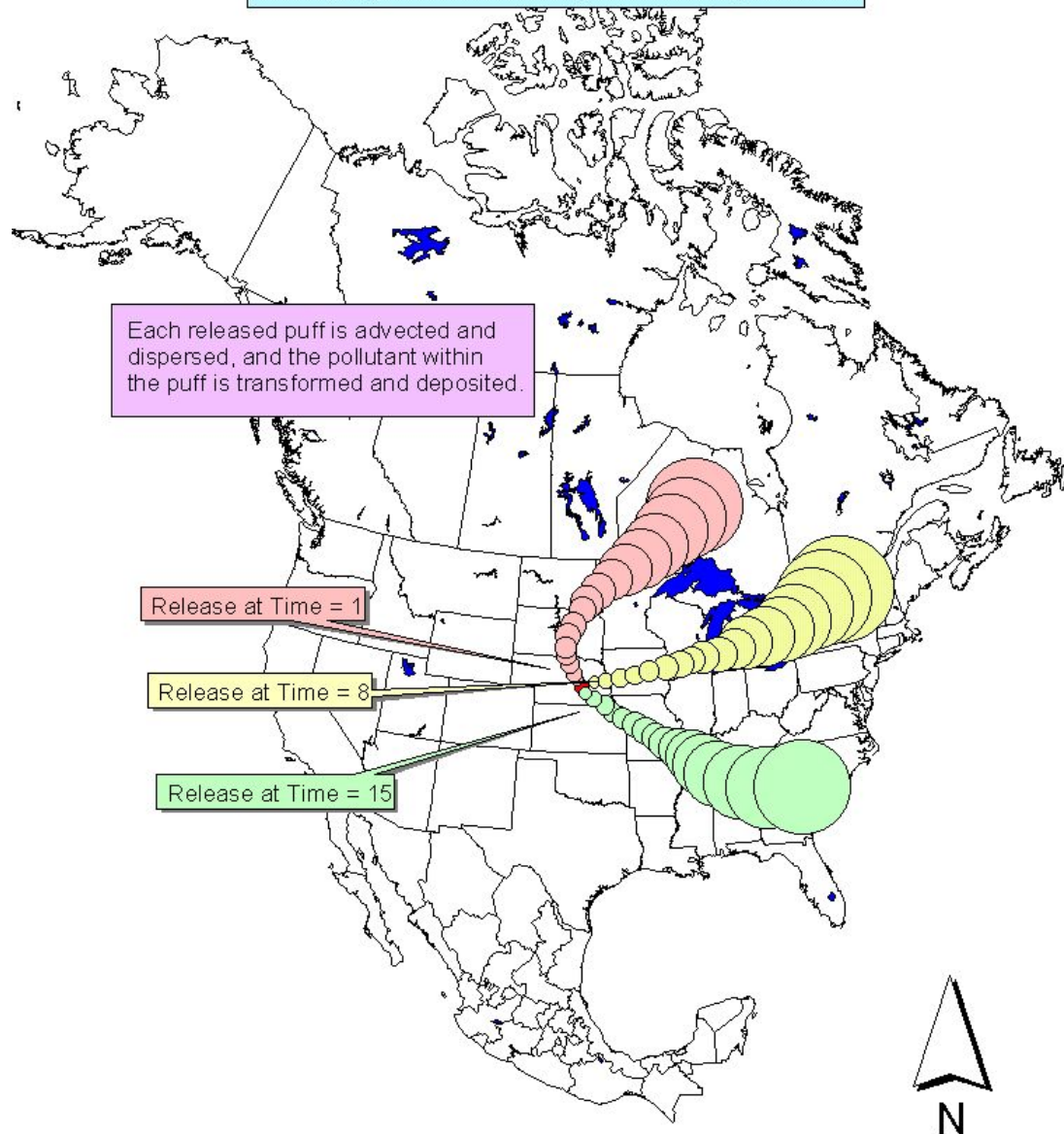


NOAA HYSPLIT MODEL

Lagrangian Puff Air Transport and Deposition Model



Over the entire modeling period (e.g., one year), puffs are released at periodic intervals (e.g., once every 7 hours).



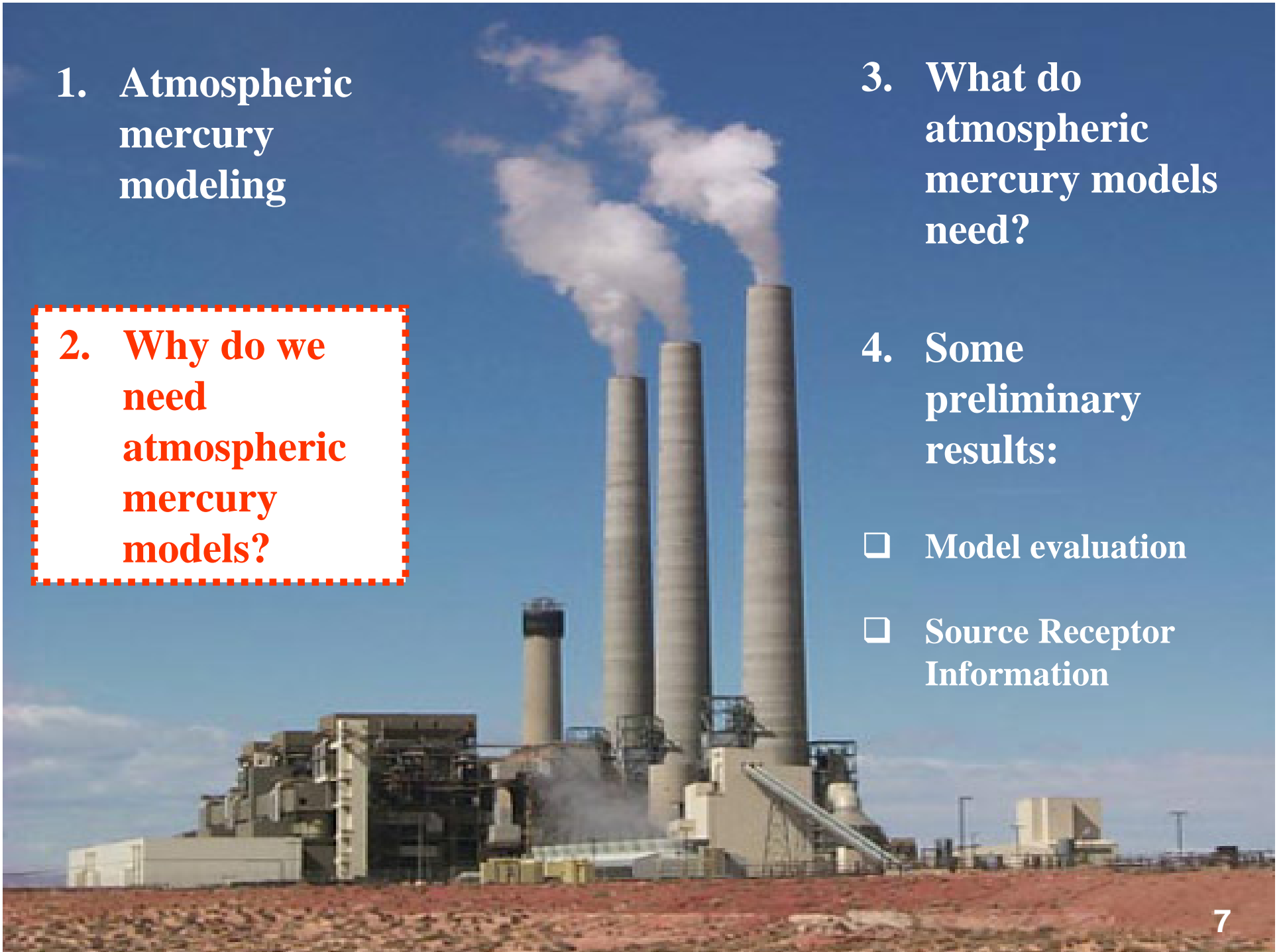
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Why do we need atmospheric mercury models?

- to get *comprehensive source attribution* information ---
we don't just want to know how much is depositing at any given location, we also want to know where it came from...
- to estimate *deposition over large regions*,
...because deposition fields are highly spatially variable,
and one can't measure everywhere all the time...
- to estimate *dry deposition*
- to evaluate *potential consequences* of alternative future emissions scenarios

**But models
must have
measurements**



**Monitoring
required to
develop models
and to evaluate
their accuracy**

**Modeling
needed to help
interpret
measurements
and estimate
source-receptor
relationships**

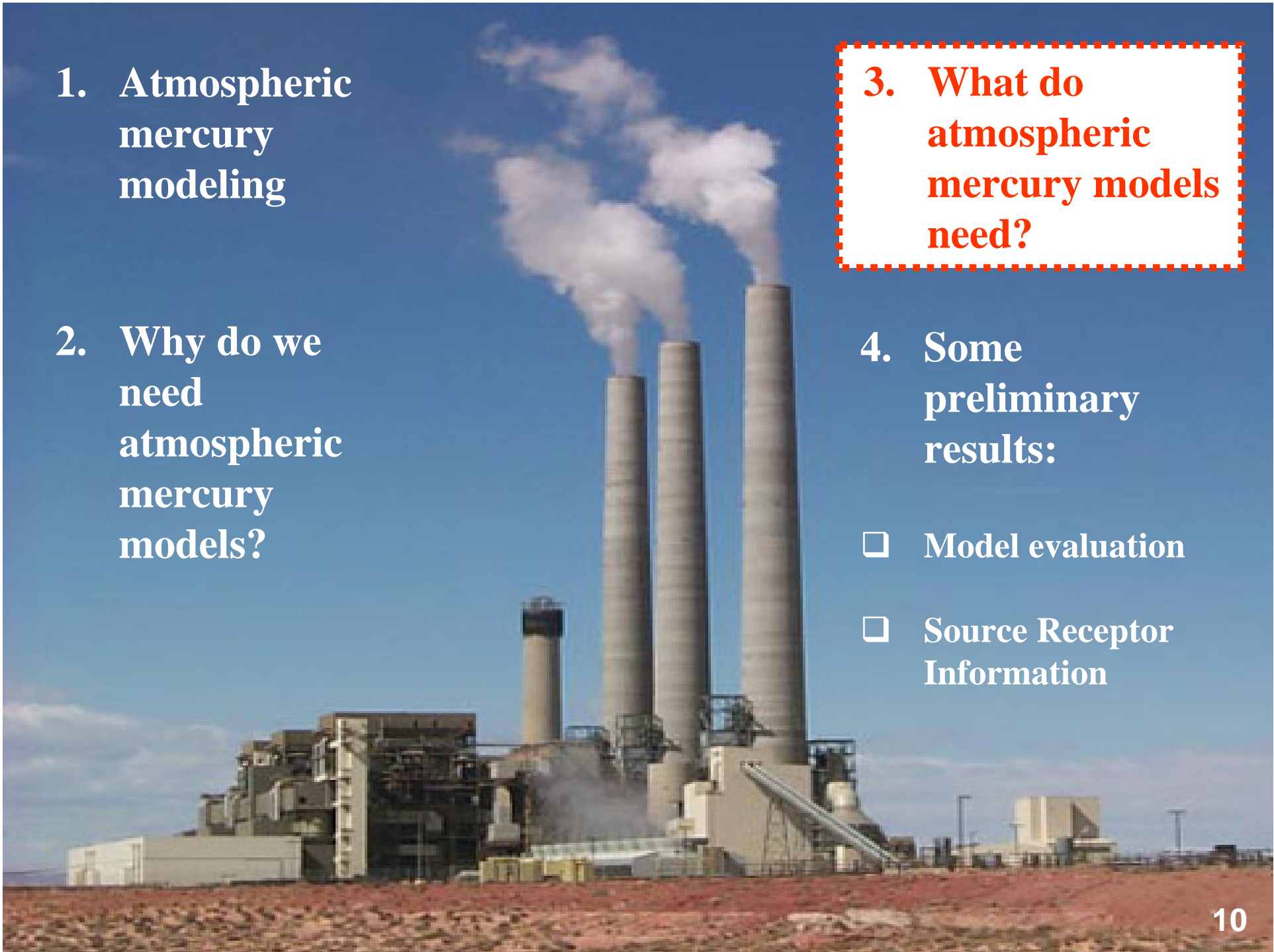
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What do atmospheric mercury models need?

**Emissions
Inventories**

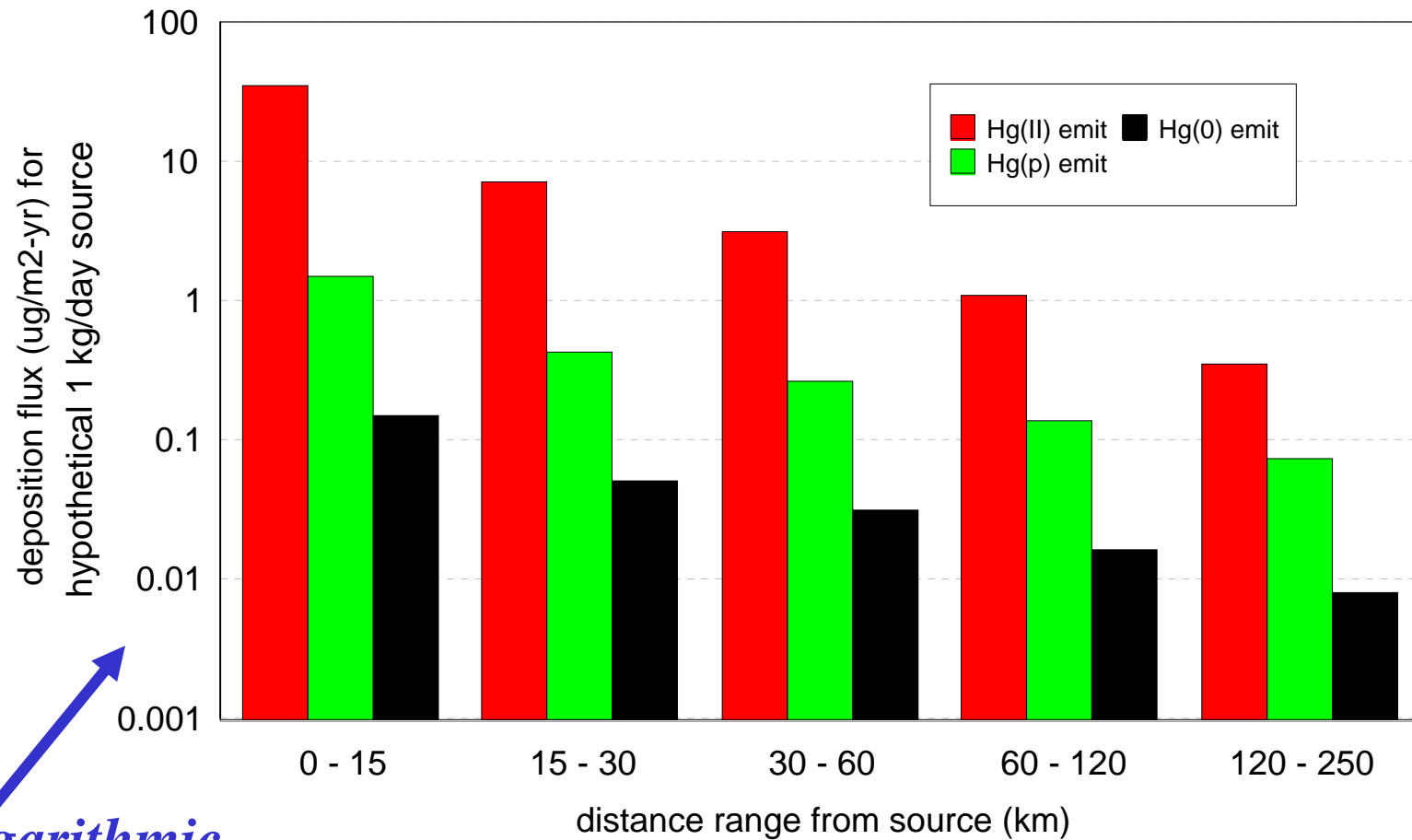
**Meteorological
Data**

**Scientific understanding of
phase partitioning,
atmospheric chemistry,
and deposition processes**

**Ambient data for comprehensive
model evaluation and improvement**

| | some challenges facing mercury modeling |
|-----------------------------------|--|
| emissions inventories | <ul style="list-style-type: none"> • need <i>all</i> sources • accurately divided into <i>different Hg forms</i> • U.S. 1996, 1999, 2003 / CAN 1995, 2000, 2005 • <i>temporal</i> variations (e.g. shut downs) |
| meteorological data | <ul style="list-style-type: none"> • precipitation not well characterized |
| scientific understanding | <ul style="list-style-type: none"> • what is RGM? what is Hg(p)? • accurate info for known reactions? • do we know all significant reactions? • natural emissions, re-emissions? |
| ambient data for model evaluation | <ul style="list-style-type: none"> • Mercury Deposition Network (MDN) is great, but: • also need RGM, Hg(p), and Hg(0) concentrations • also need data above the surface (e.g., from aircraft) • also need source-impacted sites (not just background) |

Why is emissions speciation information critical?



Logarithmic

*Hypothesized rapid reduction of Hg(II) in plumes?
If true, then dramatic impact on modeling results...*

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Atmospheric Chemical Reaction Scheme for Mercury

| Reaction | Rate | Units | Reference |
|---|--|-------------------------------|--|
| <i>GAS PHASE REACTIONS</i> | | | |
| $\text{Hg}^0 + \text{O}_3 \rightarrow \text{Hg(p)}$ | 3.0E-20 | cm ³ /molec-sec | Hall (1995) |
| $\text{Hg}^0 + \text{HCl} \rightarrow \text{HgCl}_2$ | 1.0E-19 | cm ³ /molec-sec | Hall and Bloom (1993) |
| $\text{Hg}^0 + \text{H}_2\text{O}_2 \rightarrow \text{Hg(p)}$ | 8.5E-19 | cm ³ /molec-sec | Tokos et al. (1998) (upper limit based on experiments) |
| $\text{Hg}^0 + \text{Cl}_2 \rightarrow \text{HgCl}_2$ | 4.0E-18 | cm ³ /molec-sec | Calhoun and Prestbo (2001) |
| $\text{Hg}^0 + \text{OHC} \rightarrow \text{Hg(p)}$ | 8.7E-14 | cm ³ /molec-sec | Sommar et al. (2001) |
| <i>AQUEOUS PHASE REACTIONS</i> | | | |
| $\text{Hg}^0 + \text{O}_3 \rightarrow \text{Hg}^{+2}$ | 4.7E+7 | (molar-sec) ⁻¹ | Munthe (1992) |
| $\text{Hg}^0 + \text{OHC} \rightarrow \text{Hg}^{+2}$ | 2.0E+9 | (molar-sec) ⁻¹ | Lin and Pehkonen(1997) |
| $\text{HgSO}_3 \rightarrow \text{Hg}^0$ | $T * e^{((31.971 * T) - 12595.0) / T} \text{ sec}^{-1}$ [T = temperature (K)] | | Van Loon et al. (2002) |
| $\text{Hg(II)} + \text{HO}_2\text{C} \rightarrow \text{Hg}^0$ | ~ 0 | (molar-sec) ⁻¹ | Gardfeldt & Jonnson (2003) |
| $\text{Hg}^0 + \text{HOCl} \rightarrow \text{Hg}^{+2}$ | 2.1E+6 | (molar-sec) ⁻¹ | Lin and Pehkonen(1998) |
| $\text{Hg}^0 + \text{OCl}^{-1} \rightarrow \text{Hg}^{+2}$ | 2.0E+6 | (molar-sec) ⁻¹ | Lin and Pehkonen(1998) |
| $\text{Hg(II)} \leftrightarrow \text{Hg(II)}_{(\text{soot})}$ | 9.0E+2 | liters/gram; t = 1/hour | eqnbrm: Seigneur et al. (1998) rate: Bullock & Brehme (2002). |
| $\text{Hg}^{+2} + \text{h} \leftrightarrow \text{Hg}^0$ | 6.0E-7 | (sec) ⁻¹ (maximum) | Xiao et al. (1994); Bullock and Brehme (2002) |

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Some Additional Measurement Issues (from a modeler's perspective)

- **Data availability**
- **Simple vs. Complex Measurements**

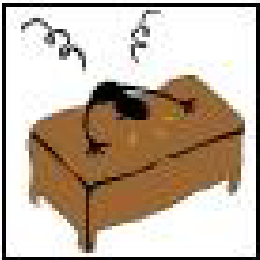
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Data availability



A major impediment to evaluating and improving atmospheric Hg models has been the lack of speciated Hg air concentration data



There have been very few measurements to date, and these data are rarely made available in a practical way (timely, complete, etc.)



The data being collected at Piney Reservoir could be extremely helpful!

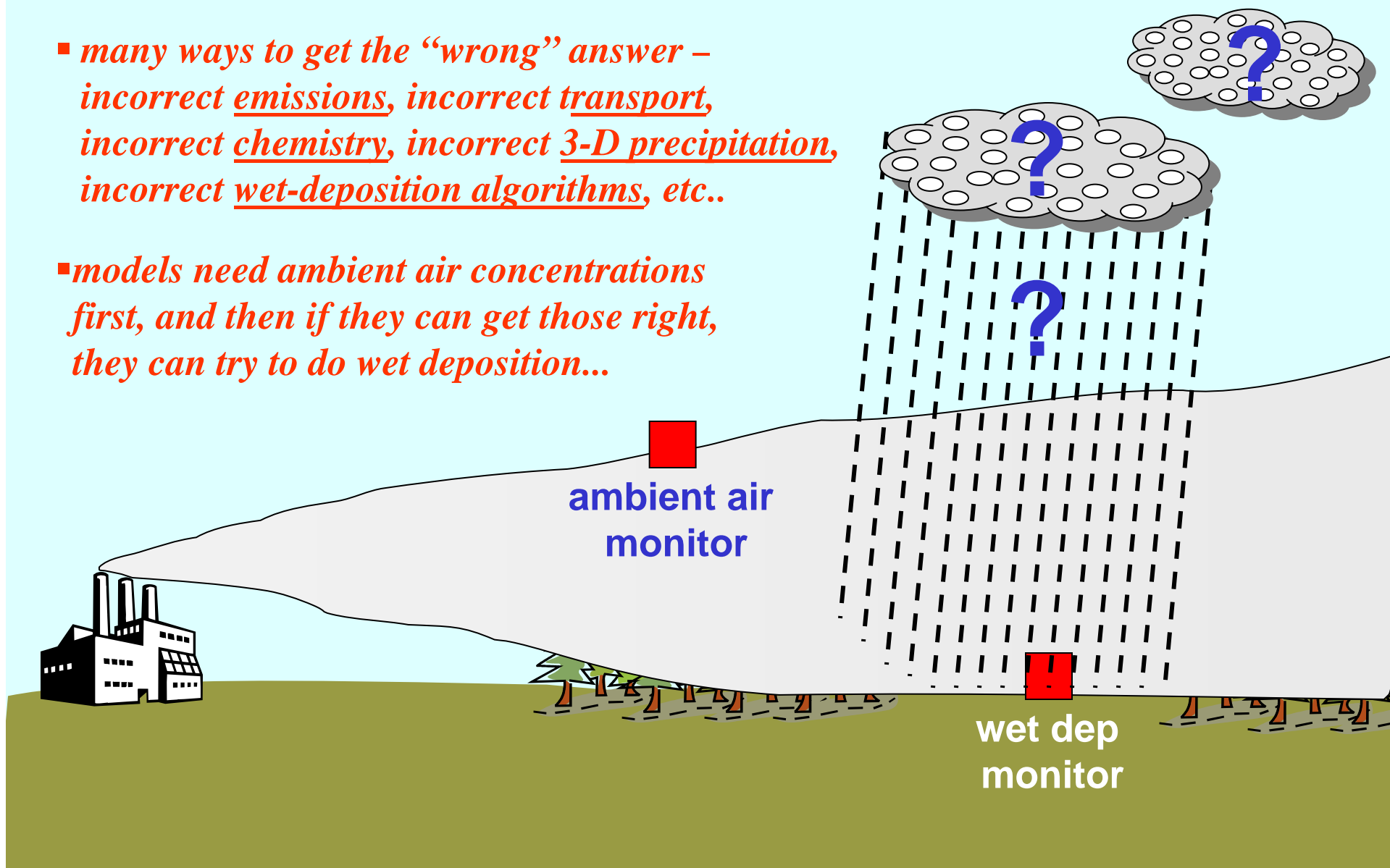
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Simple vs. Complex Measurements:

1. Wet deposition is a very complicated phenomena...

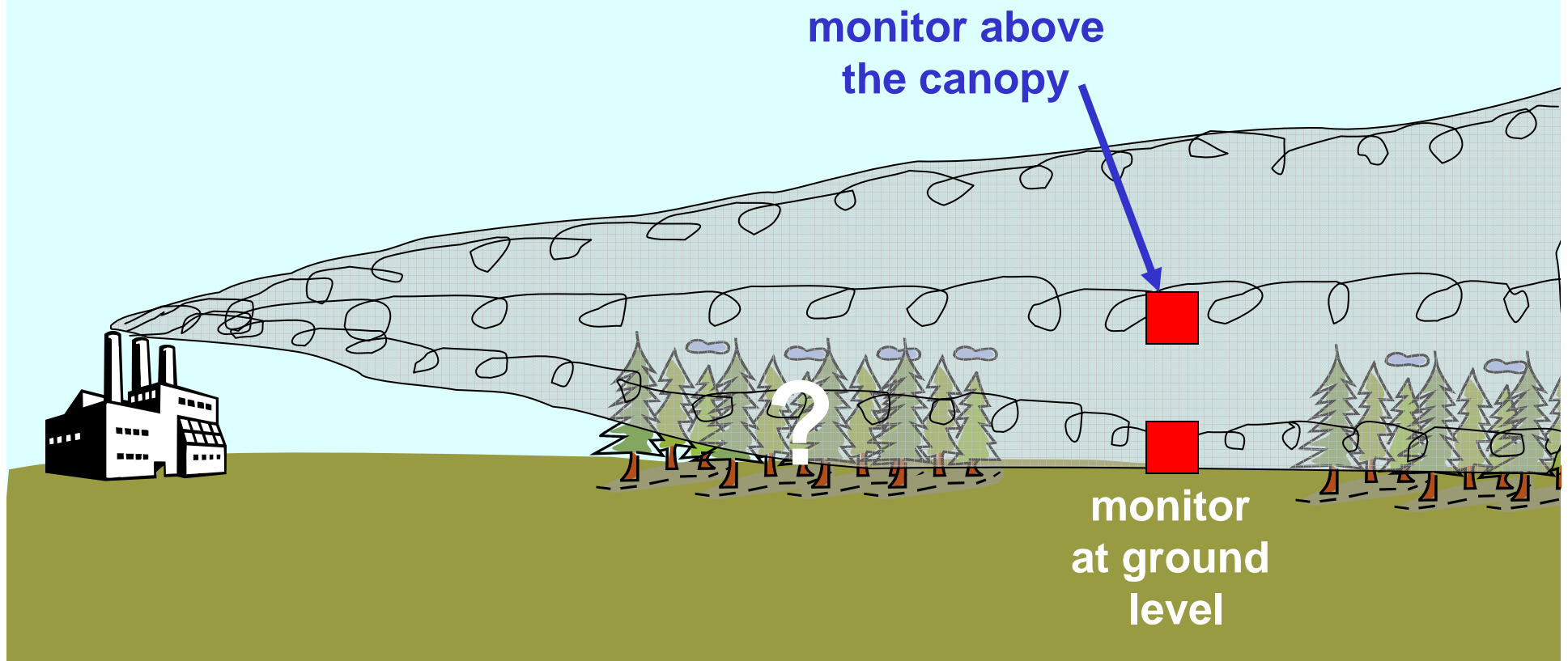
- *many ways to get the “wrong” answer – incorrect emissions, incorrect transport, incorrect chemistry, incorrect 3-D precipitation, incorrect wet-deposition algorithms, etc..*
- *models need ambient air concentrations first, and then if they can get those right, they can try to do wet deposition...*



Simple vs. Complex Measurements:

2. Potential complication with ground-level monitors... (“fumigation”, “filtration”, etc.)...

- *atmospheric phenomena are complex and not well understood;*
- *models need “simple” measurements for diagnostic evaluations;*
- *ground-level data for rapidly depositing substances (e.g., RGM) hard to interpret*
- *elevated platforms might be more useful (at present level of understanding)*



Simple vs. Complex measurements - 3. Urban areas:

- a. Emissions inventory poorly known**
- b. Meteorology very complex (flow around buildings)**
- c. So, measurements in urban areas not particularly useful for current large-scale model evaluations**



Simple vs. Complex Measurements – 4: extreme near-field measurements



- Sampling near intense sources?
- Must get the fine-scale met “perfect”

Ok, if one wants to develop hypotheses regarding *whether or not this is actually a source of the pollutant (and you can't do a stack test for some reason!)*.

Complex vs. Simple Measurements – 5: Need some source impacted measurements

- **Major questions regarding plume chemistry and near-field impacts (are there “hot spots”?)**
- **Most monitoring sites are designed to be “regional background” sites (e.g., most Mercury Deposition Network sites).**
- **We need some source-impacted sites as well to help resolve near-field questions**
- *But not too close – maybe 20-30 km is ideal (?)*

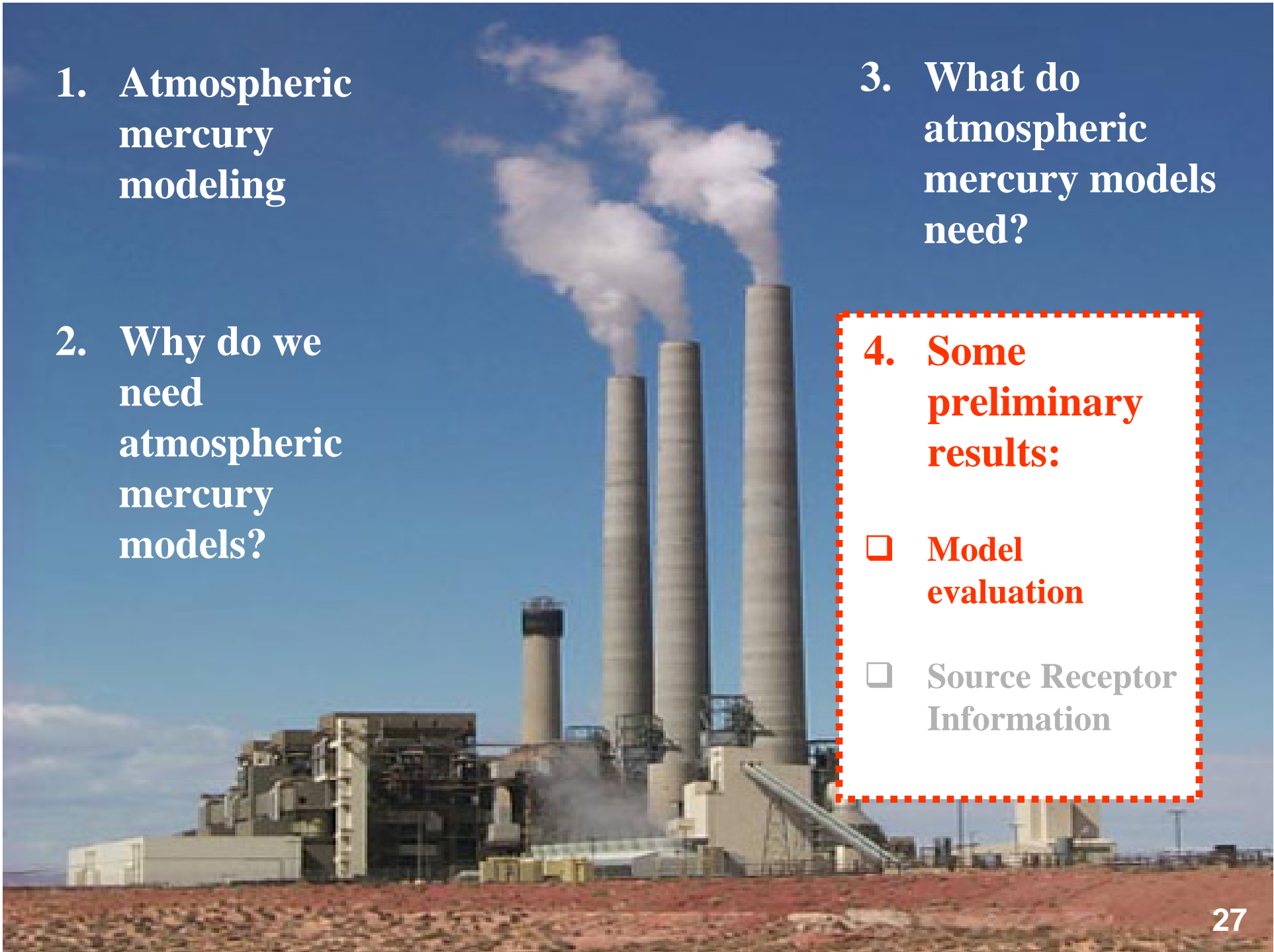
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|-------------------|-----------|-----------------|-------|-----|-----------|---------|---------|------------------|
| | Chemistry | Hg ⁰ | Hg(p) | RGM | Wet Dep | Dry Dep | Budgets | |

Participants

D. Syrakov Bulgaria....NIMH
 A. Dastoor, D. Davignon Canada..... MSC-Can
 J. Christensen Denmark...NERI
 G. Petersen, R. Ebinghaus Germany...GKSS
 J. Pacyna Norway.....NILU
 J. Munthe, I. Wängberg Sweden..... IVL
 R. BullockUSA.....EPA
 M. Cohen, R. Artz, R. DraxlerUSA.....NOAA
 C. Seigneur, K. LohmanUSA..... AER/EPRI
 A. Ryaboshapko, I. Ilyin, O.Travnikov...EMEP..... MSC-E

| | | | | | | | | |
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Intercomparison Conducted in 3 Stages

- I. Comparison of chemical schemes for a cloud environment
- II. Air Concentrations in Short Term Episodes
- III. Long-Term Deposition and Source-Receptor Budgets

EMEP Intercomparison Study of Numerical Models for Long-Range Atmospheric Transport of Mercury

| | | | | | | | | |
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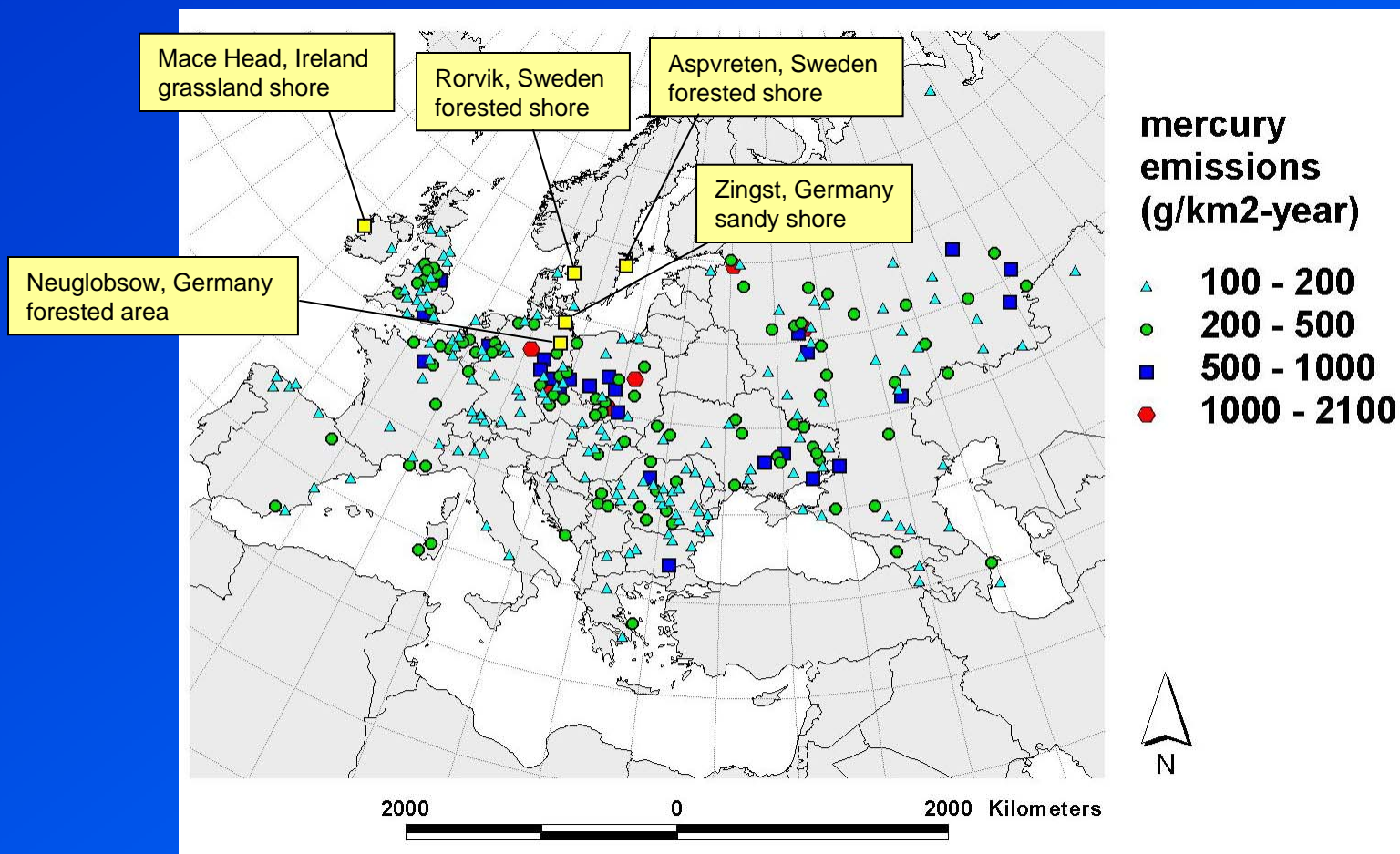
Participating Models

| Model Acronym | Model Name and Institution | Stage | | |
|---------------|--|-------|----|-----|
| | | I | II | III |
| CAM | <i>Chemistry of Atmos. Mercury model</i> , Environmental Institute, Sweden | | | |
| MCM | <i>Mercury Chemistry Model</i> , Atmos. & Environmental Research, USA | | | |
| CMAQ | <i>Community Multi-Scale Air Quality model</i> , US EPA | | | |
| ADOM | <i>Acid Deposition and Oxidants Model</i> , GKSS Research Center, Germany | | | |
| MSCE-HM | <i>MSC-E heavy metal regional model</i> , EMEP MSC-E | | | |
| GRAHM | <i>Global/Regional Atmospheric Heavy Metal model</i> , Environment Canada | | | |
| EMAP | <i>Eulerian Model for Air Pollution</i> , Bulgarian Meteo-service | | | |
| DEHM | <i>Danish Eulerian Hemispheric Model</i> , National Environmental Institute | | | |
| HYSPLIT | <i>Hybrid Single Particle Lagrangian Integrated Trajectory model</i> , US NOAA | | | |
| MSCE-HM-Hem | <i>MSC-E heavy metal hemispheric model</i> , EMEP MSC-E | | | |

| | | | | | | | | |
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Anthropogenic Mercury Emissions Inventory and Monitoring Sites for Phase II

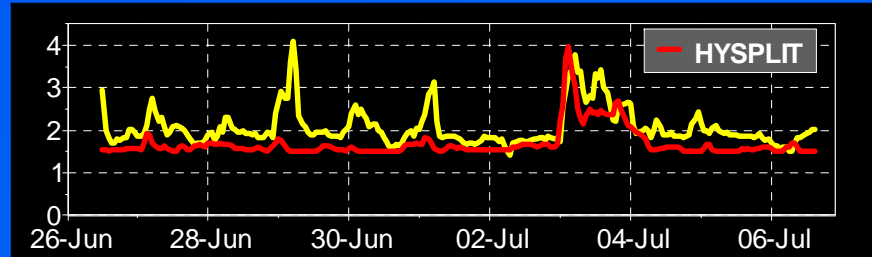
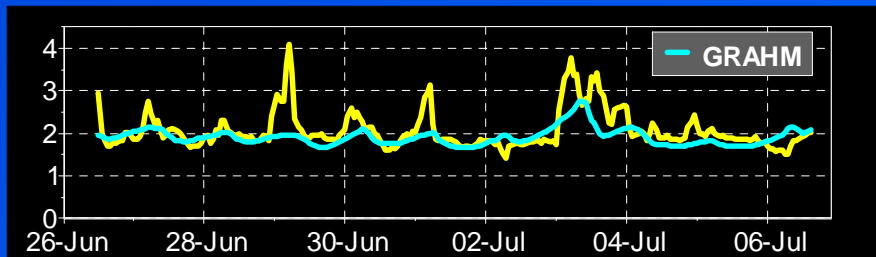
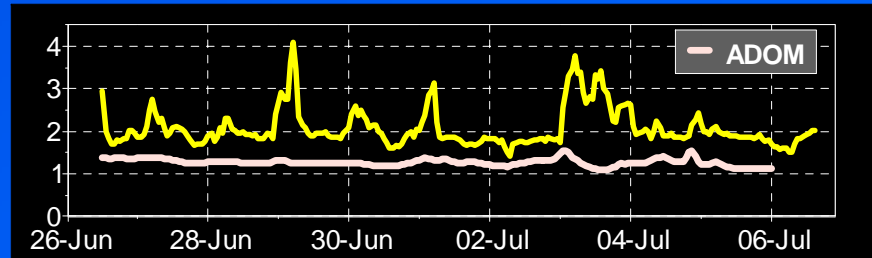
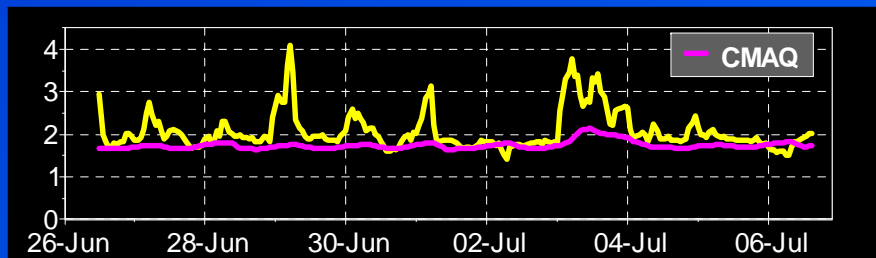
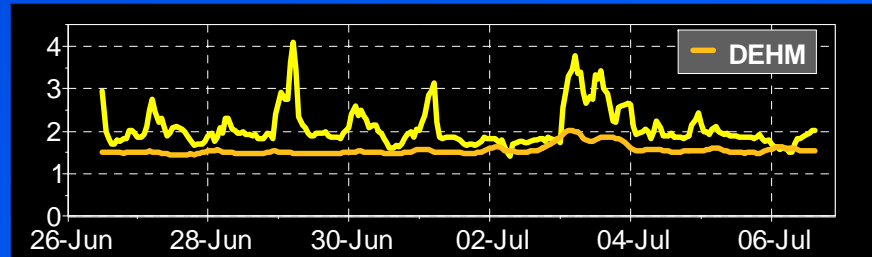
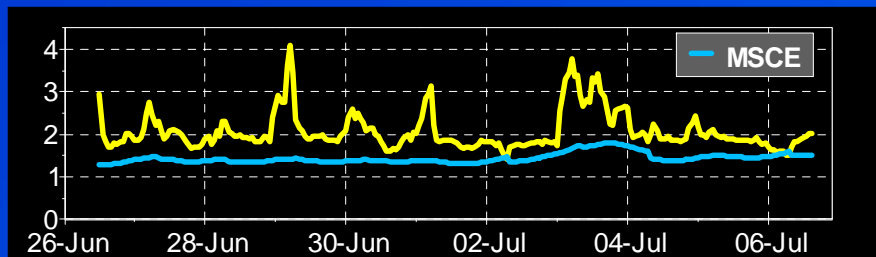
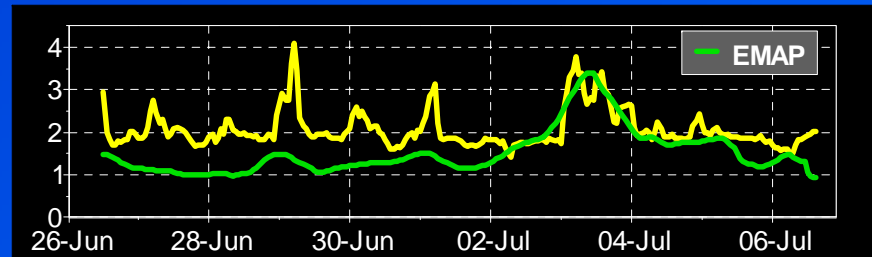
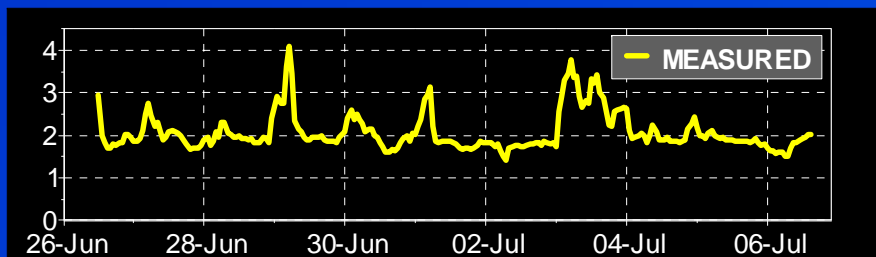
(note: only showing largest emitting grid cells)



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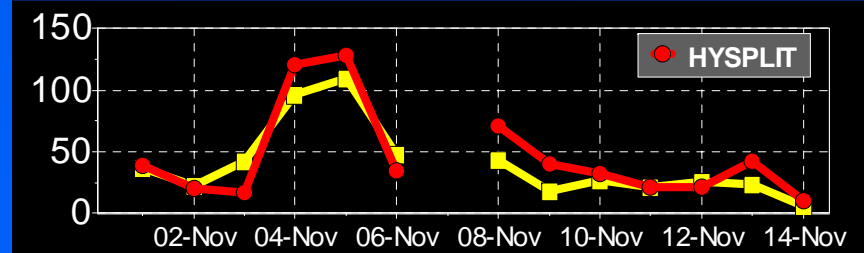
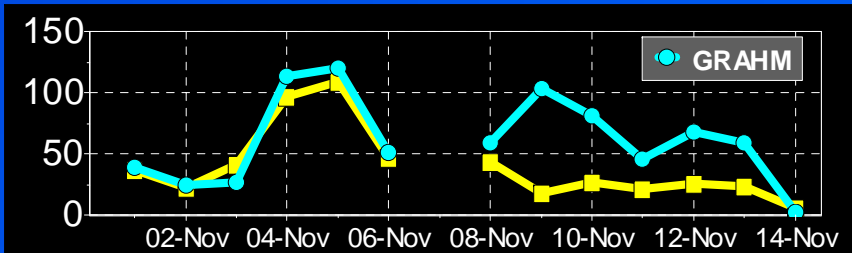
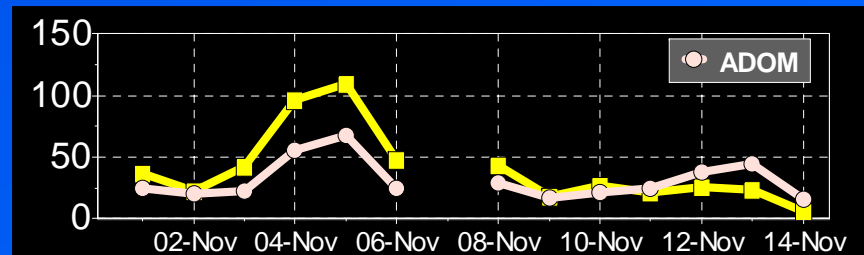
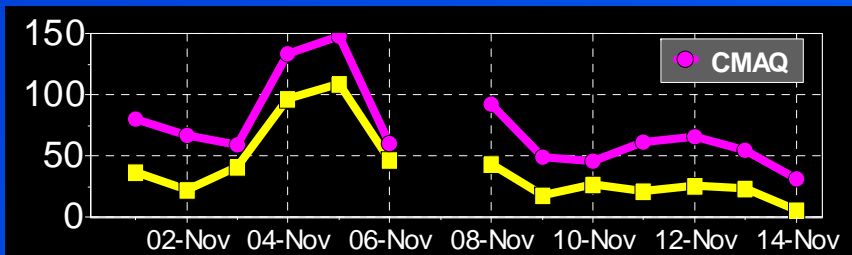
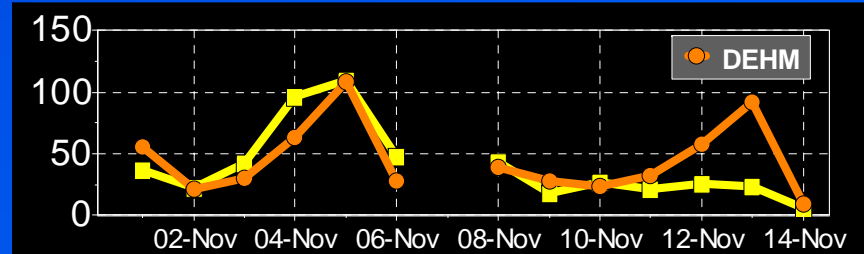
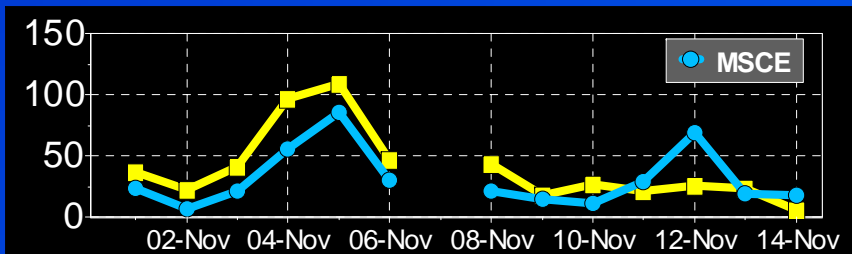
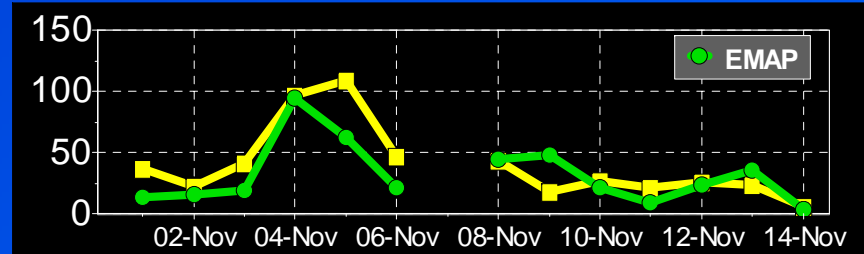
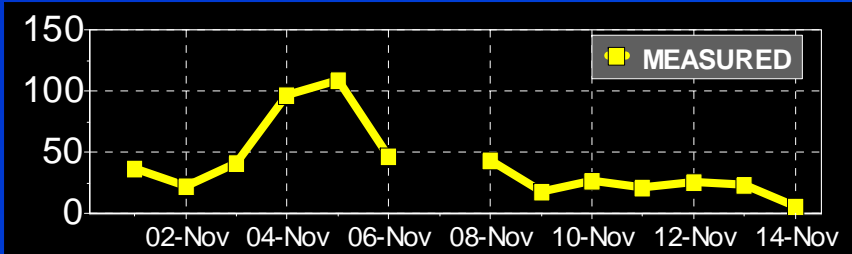
Total Gaseous Mercury (ng/m³) at Neuglobsow: June 26 – July 6, 1995



EMEP Intercomparison Study of Numerical Models for Long-Range Atmospheric Transport of Mercury

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Total *Particulate* Mercury (pg/m³) at Neuglobsow, Nov 1-14, 1999



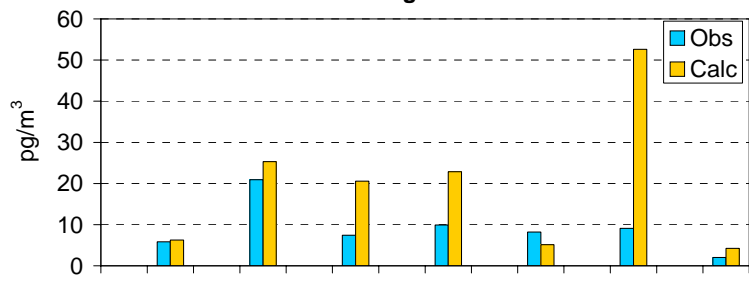
EMEP Intercomparison Study of Numerical Models for

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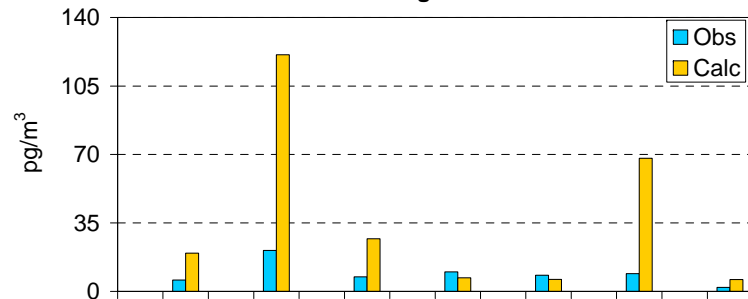
conclu-
ions

Reactive Gaseous Mercury at Neuglobsow, Nov 1-14, 1999

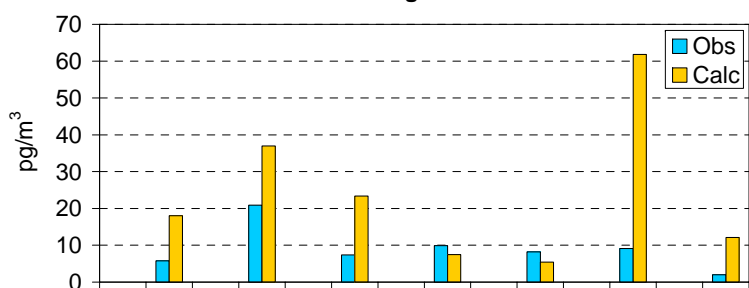
MSCE Neuglobsow RGM



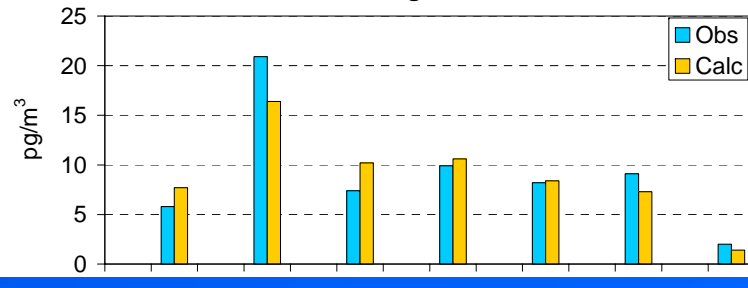
GRAHM Neuglobsow RGM



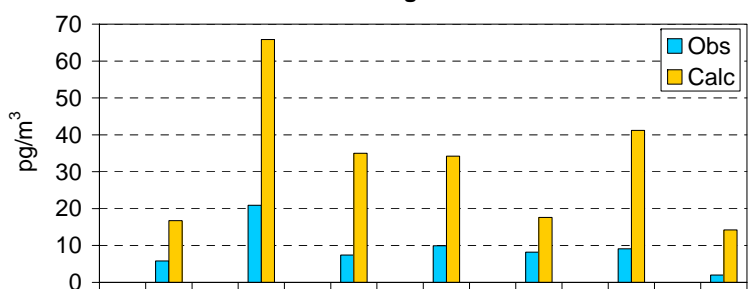
ADOM Neuglobsow RGM



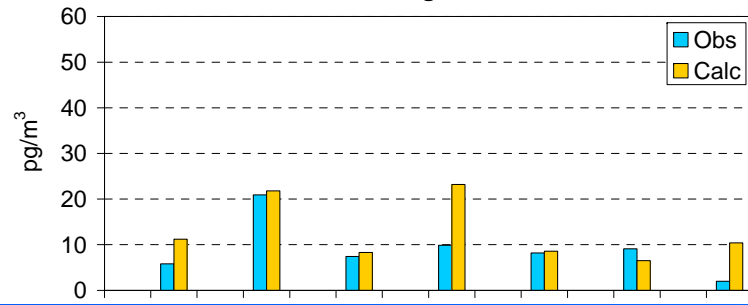
EMAP Neuglobsow RGM



CMAQ Neuglobsow RGM



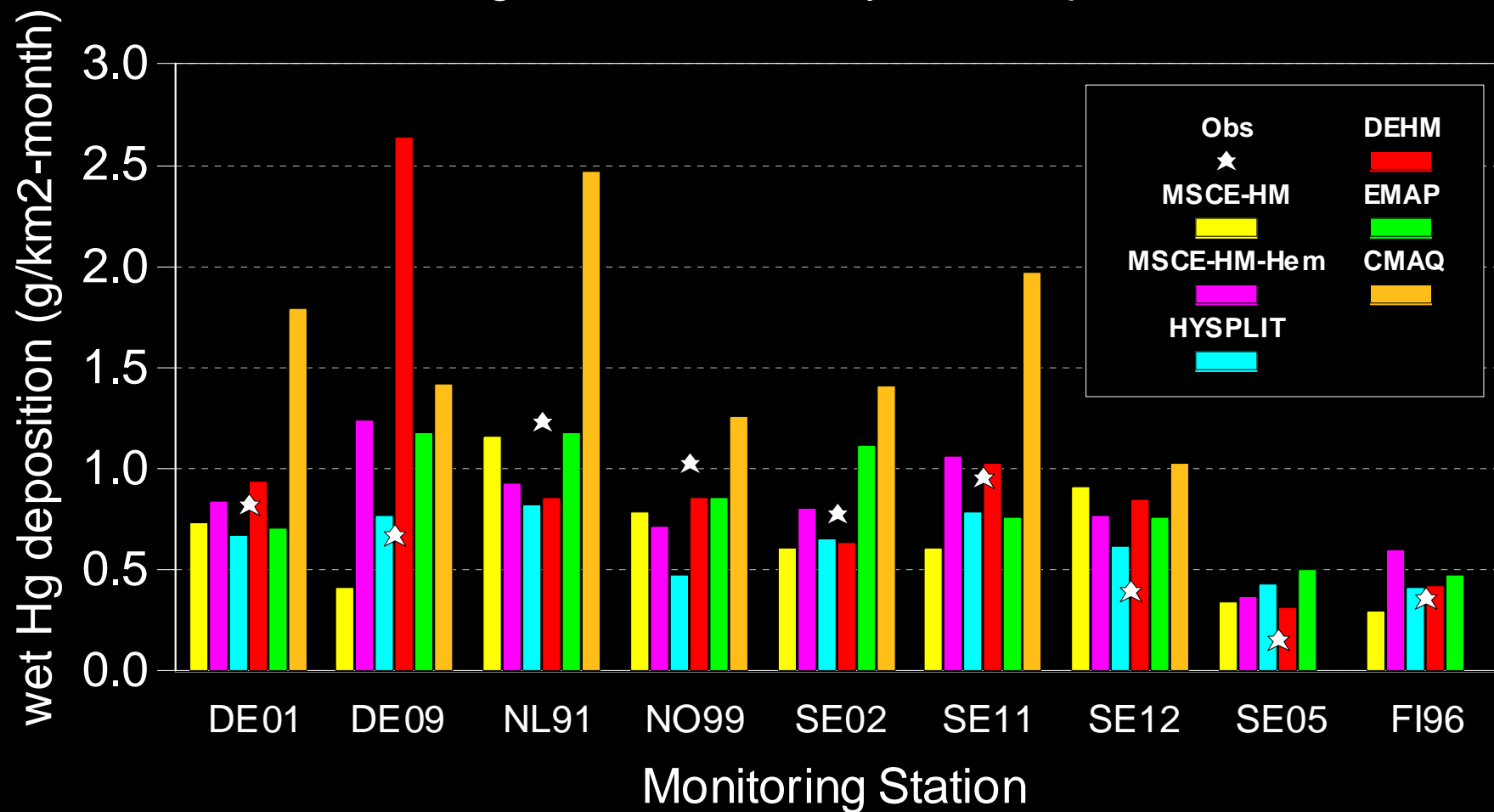
HYSPLIT Neuglobsow RGM



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| | Chemistry | Hg ⁰ | Hg(p) | RGM | Wet Dep | Dry Dep | Budgets | |

August 1999 Mercury Wet Deposition



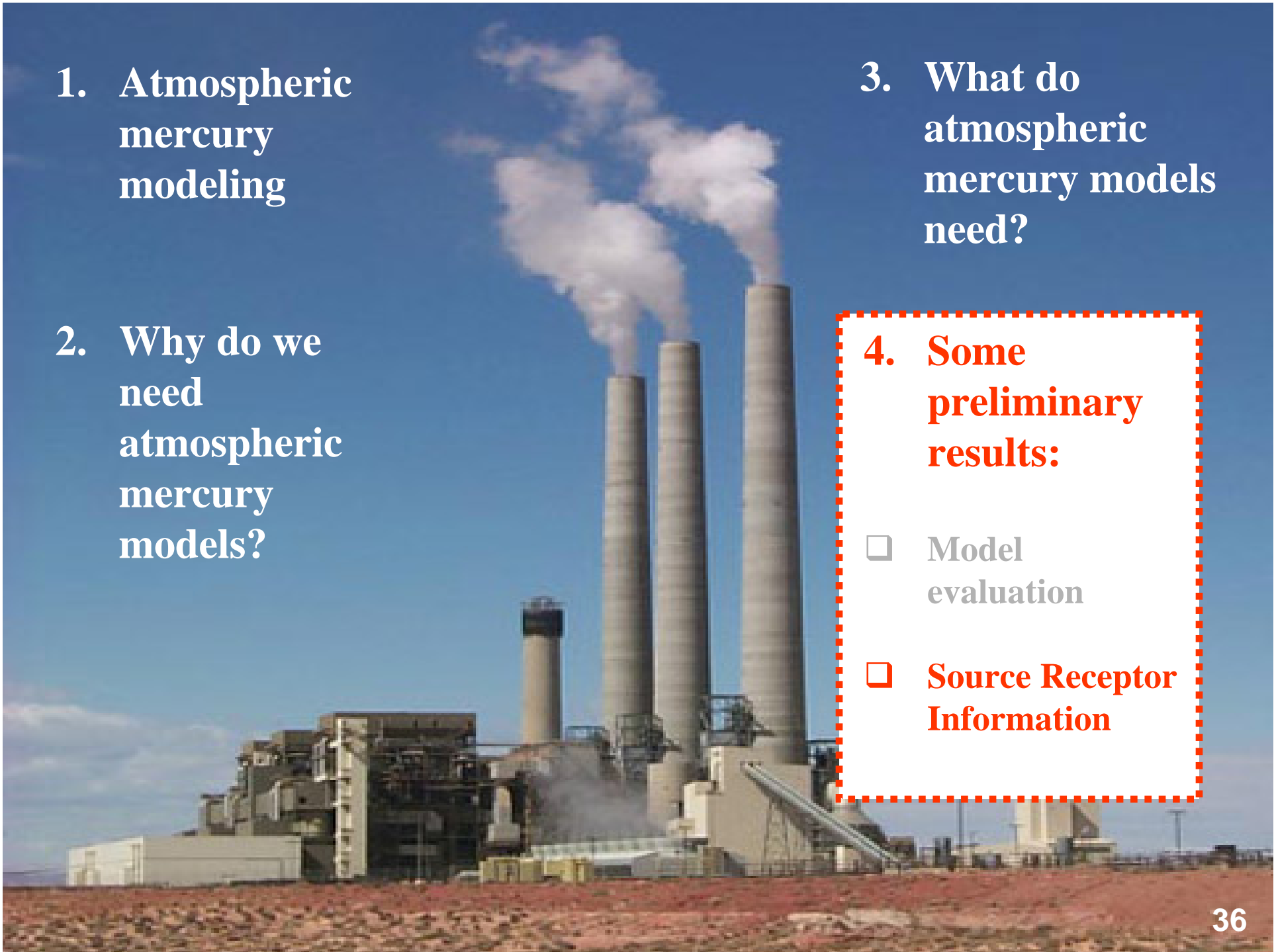
1. Atmospheric mercury modeling

2. Why do we need atmospheric mercury models?

3. What do atmospheric mercury models need?

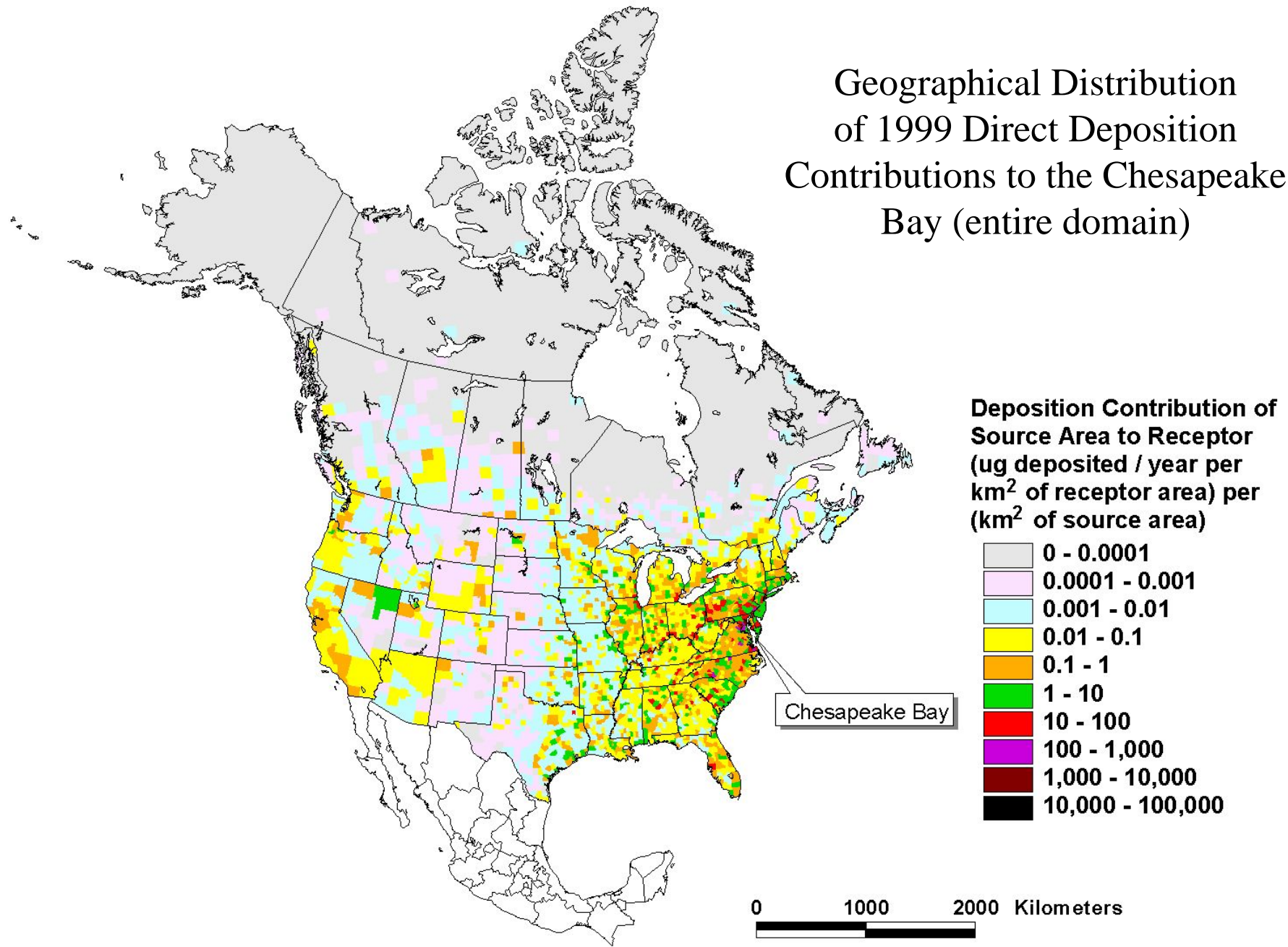
4. **Some preliminary results:**

- Model evaluation
- Source Receptor Information**

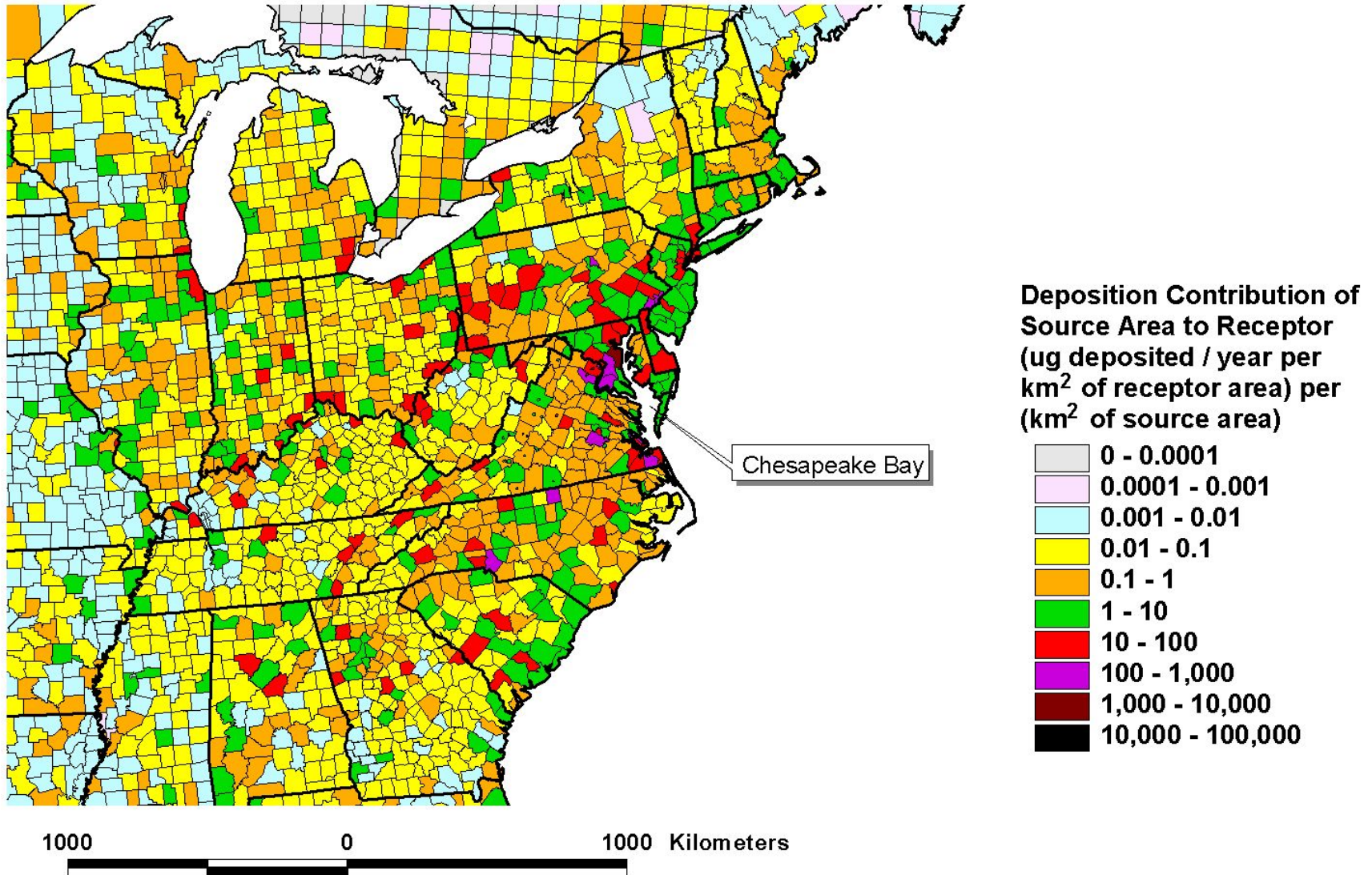


**Example of
Detailed Results:
1999 Results for
Chesapeake Bay**

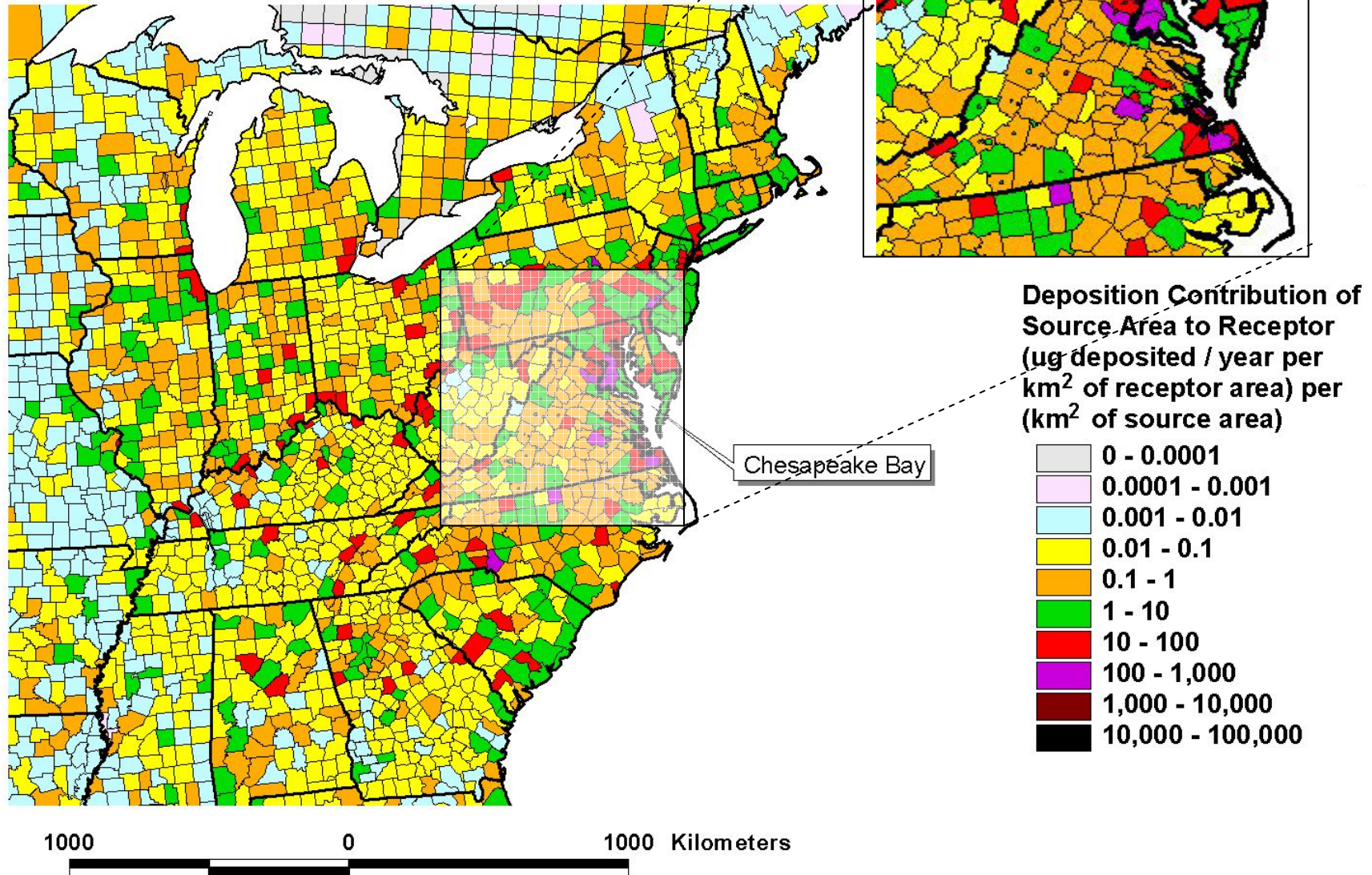
Geographical Distribution of 1999 Direct Deposition Contributions to the Chesapeake Bay (entire domain)



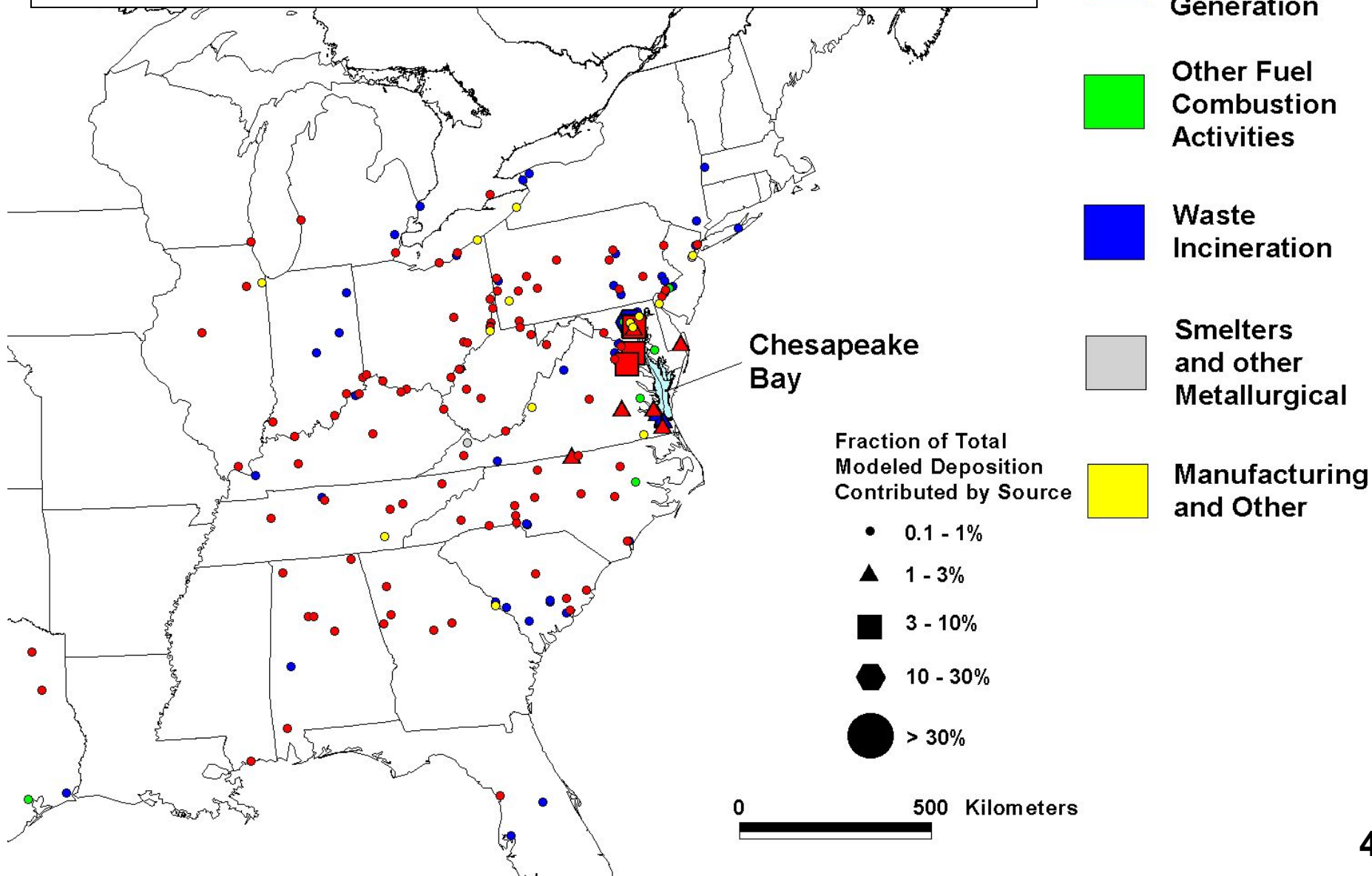
Geographical Distribution of 1999 Direct Deposition Contributions to the Chesapeake Bay (regional close-up)



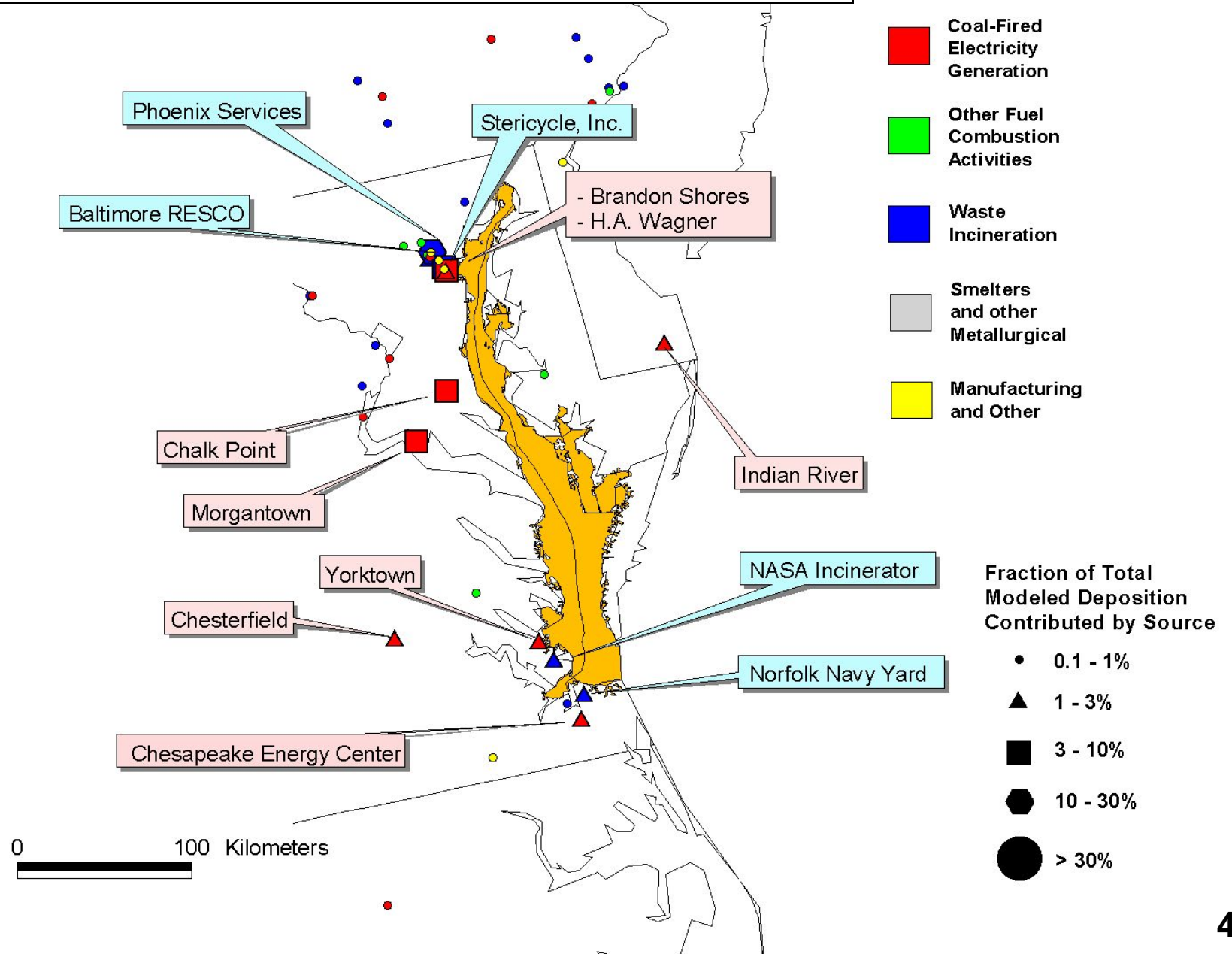
Geographical Distribution of 1999 Direct Deposition Contributions to the Chesapeake Bay (local close-up)



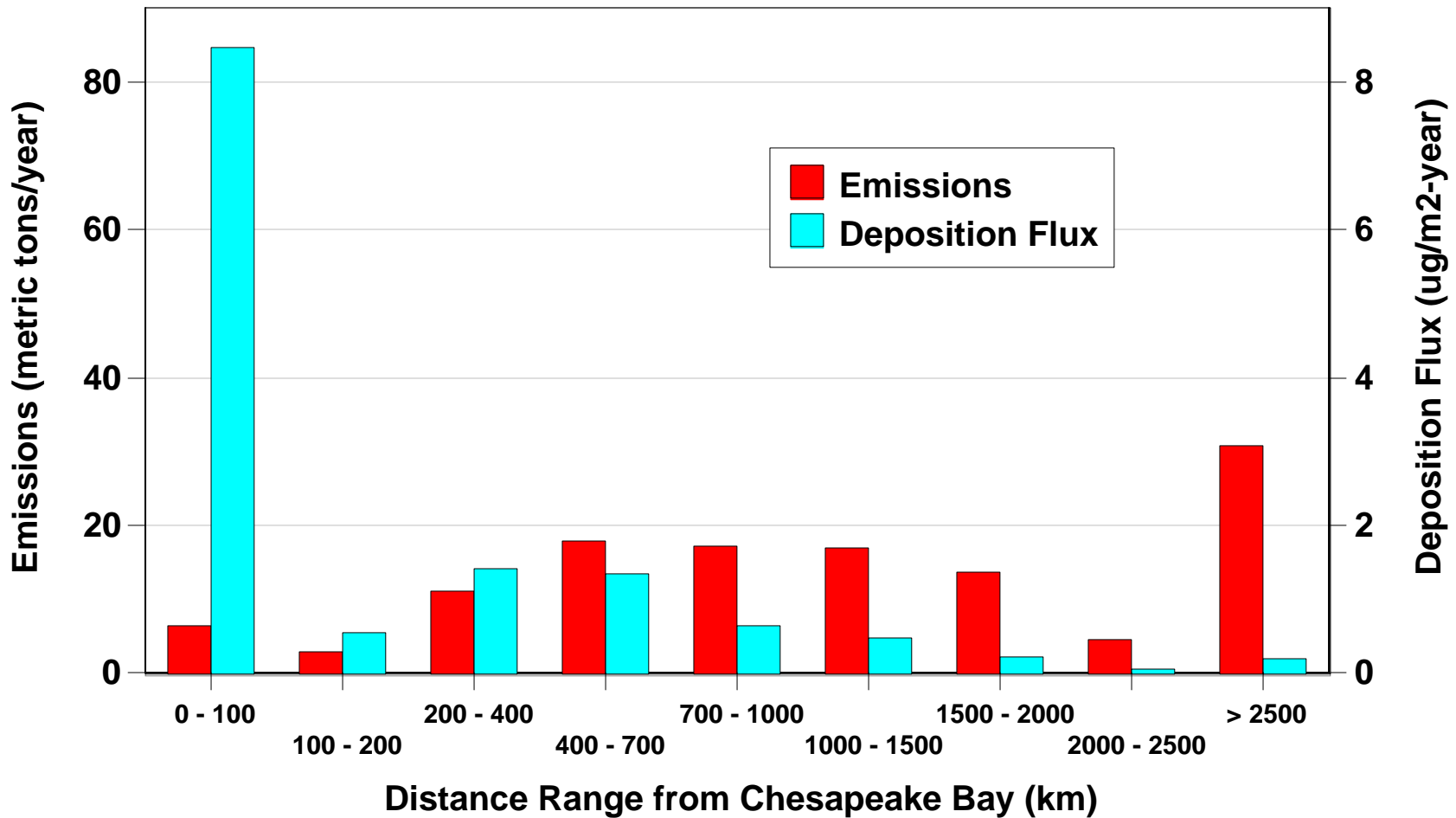
Largest Regional Individual Sources Contributing to 1999 Mercury Deposition Directly to the Chesapeake Bay



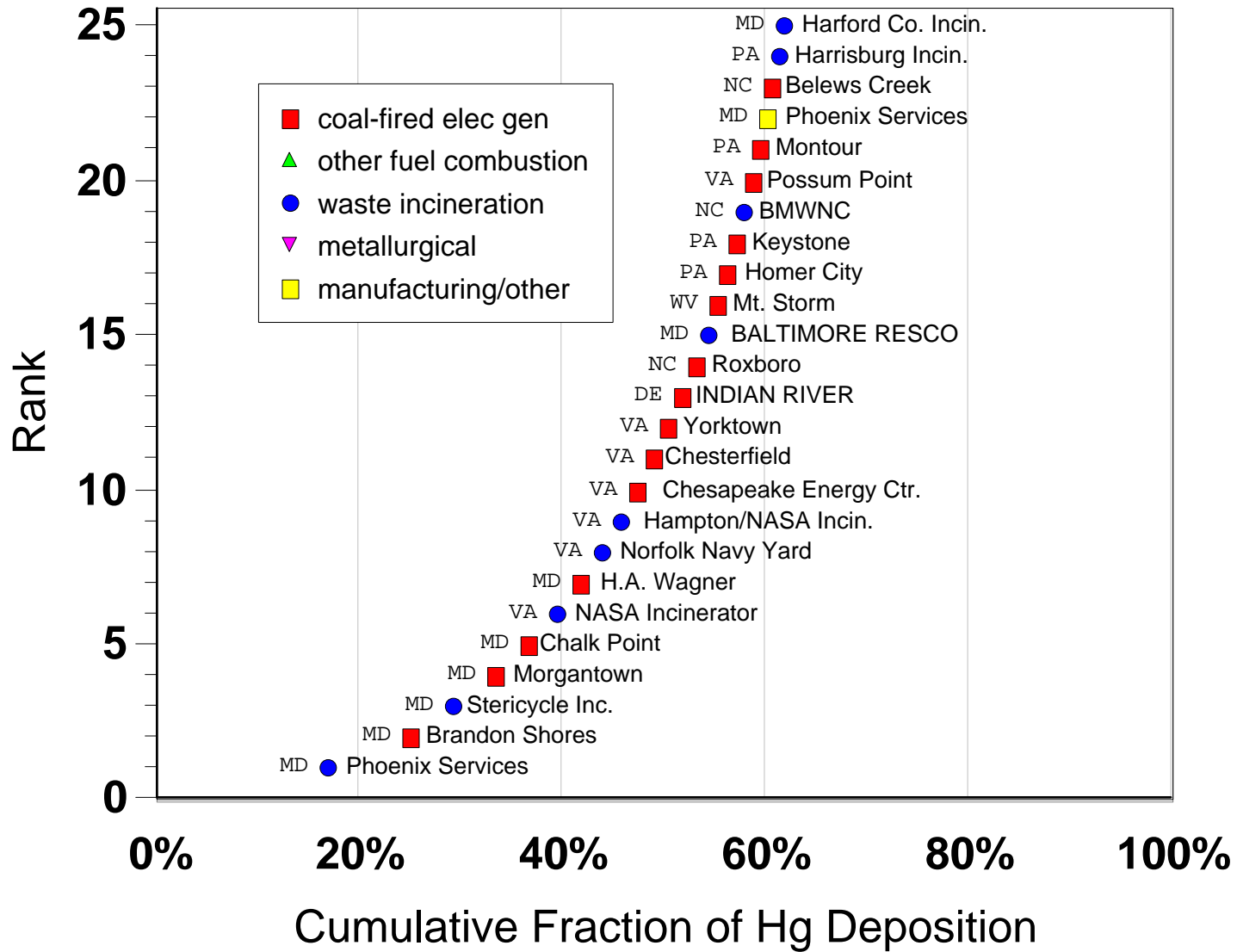
Largest Local Individual Sources Contributing to 1999 Mercury Deposition Directly to the Chesapeake Bay



Emissions and Direct Deposition Contributions from Different Distance Ranges Away From the Chesapeake Bay

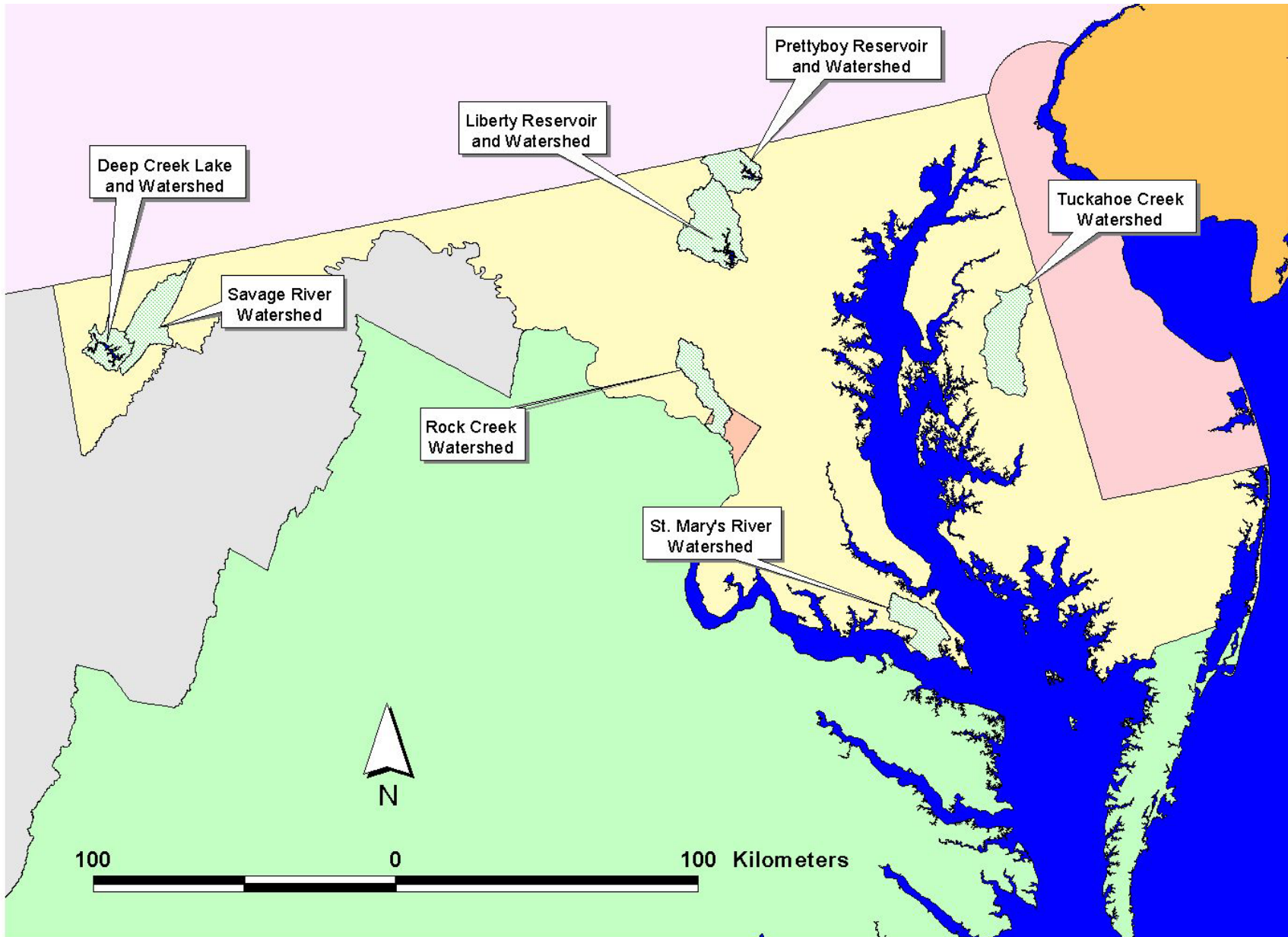


Top 25 Contributors to 1999 Hg Deposition Directly to the Chesapeake Bay

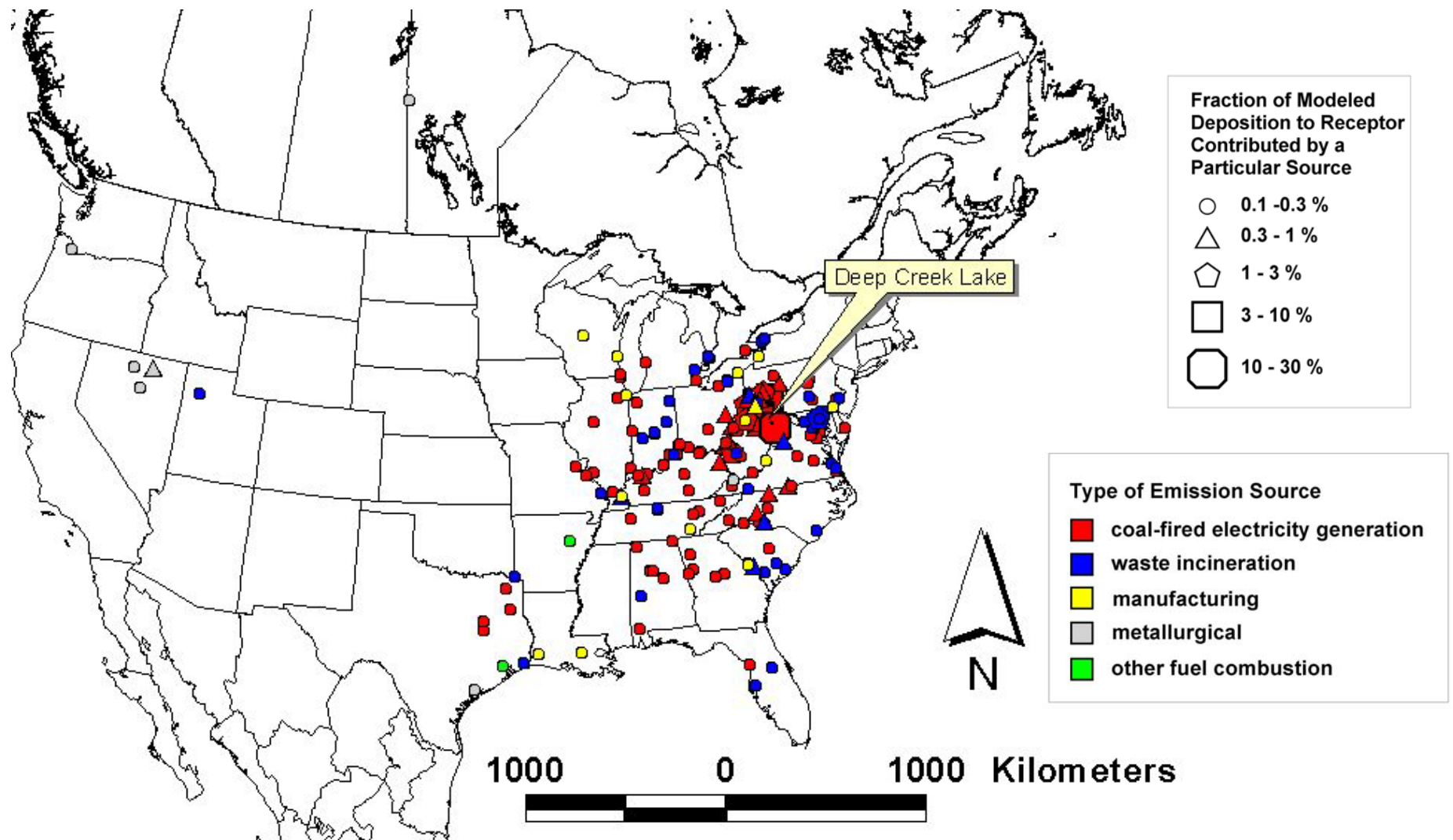


Preliminary Results for other Maryland Receptors

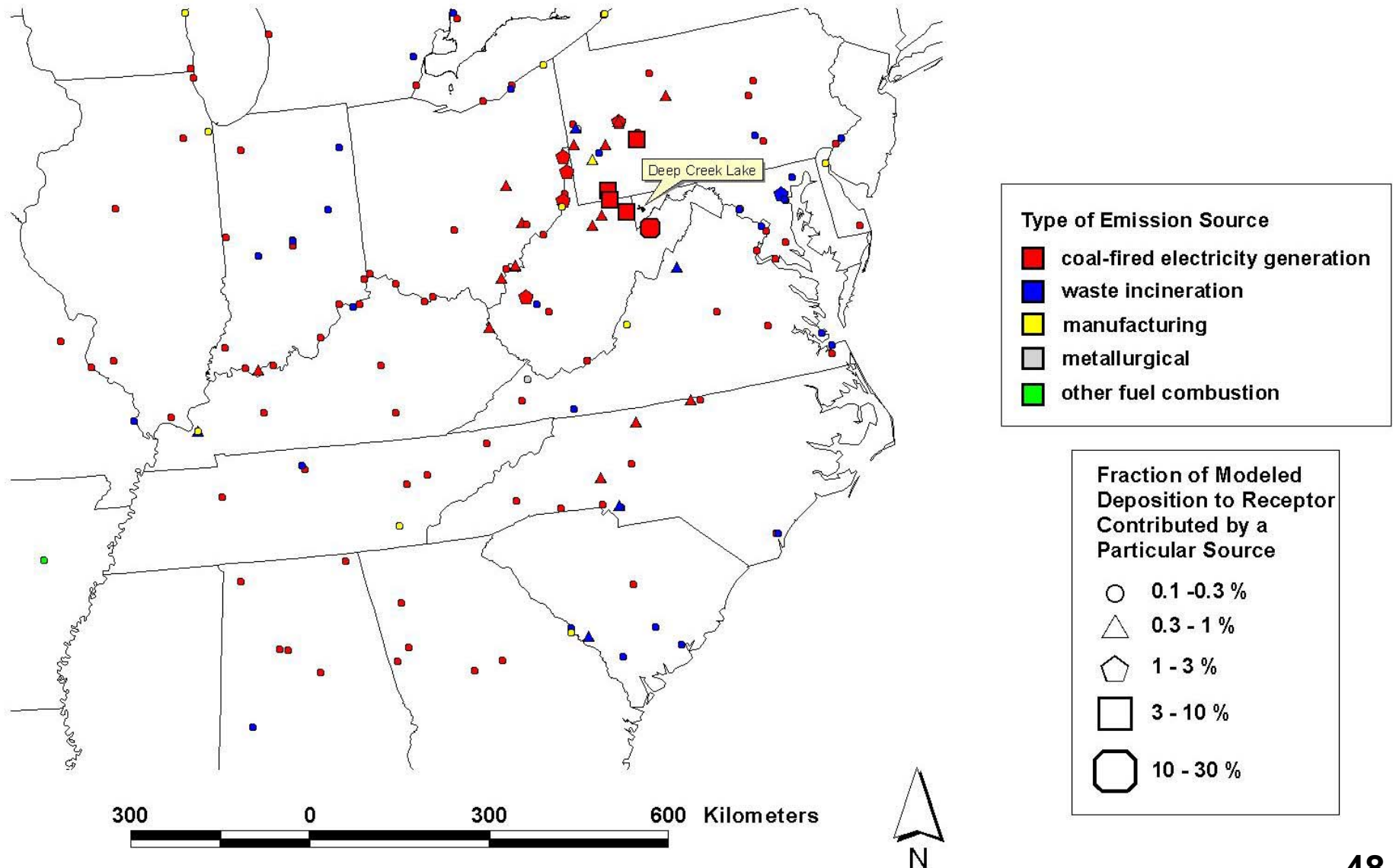
Maryland Receptors Included in Recent Preliminary HYSPLIT-Hg modeling (*but modeling was not optimized for these receptors!*)



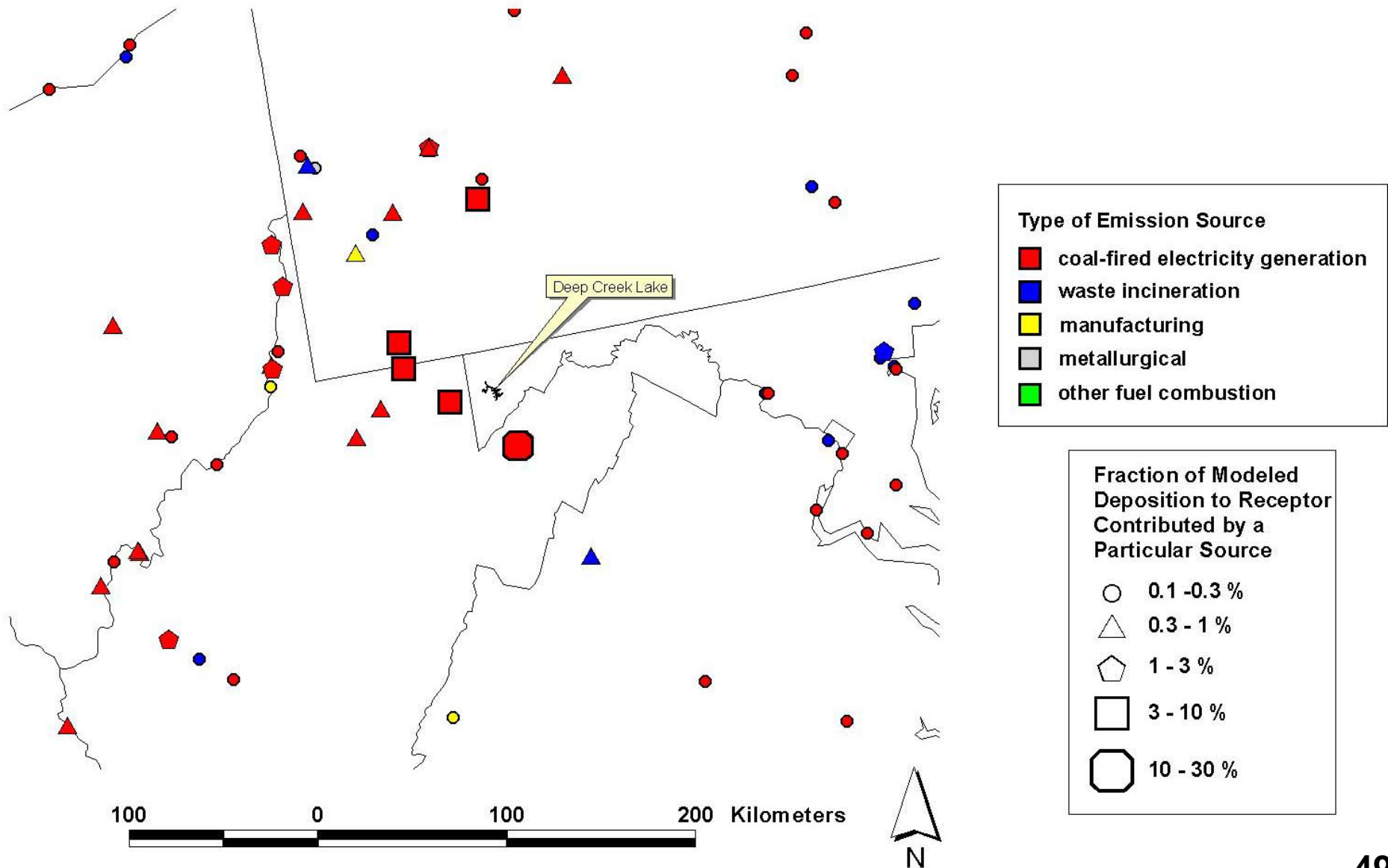
Largest Modeled Atmospheric Deposition Contributors Directly to
Deep Creek Lake based on 1999 USEPA Emissions Inventory
(national view)



Largest Modeled Atmospheric Deposition Contributors Directly to Deep Creek Lake based on 1999 USEPA Emissions Inventory (regional view)









Largest Modeled Atmospheric Deposition Contributors Directly to Deep Creek Lake based on 1999 USEPA Emissions Inventory (close-up view)



Some Next Steps

- Use more highly resolved meteorological data grids
- Expand model domain to include global sources
- Simulate natural emissions and re-emissions of previously deposited Hg
- Additional model evaluation exercises ... more sites, more time periods, more variables
- Sensitivity analyses and examination of atmospheric Hg chemistry (e.g. marine boundary layer, upper atmosphere)
- Dynamic linkage with ecosystem cycling models

Conclusions

-  **Models needed for source-receptor and other info**
-  **At present, many model uncertainties & data limitations**
-  **Monitoring data required to evaluate and improve models**
-  **For this, simple may be better than complex measurements**
-  **Some useful model results appear to be emerging**
-  **Future is much brighter because of this coordination!**

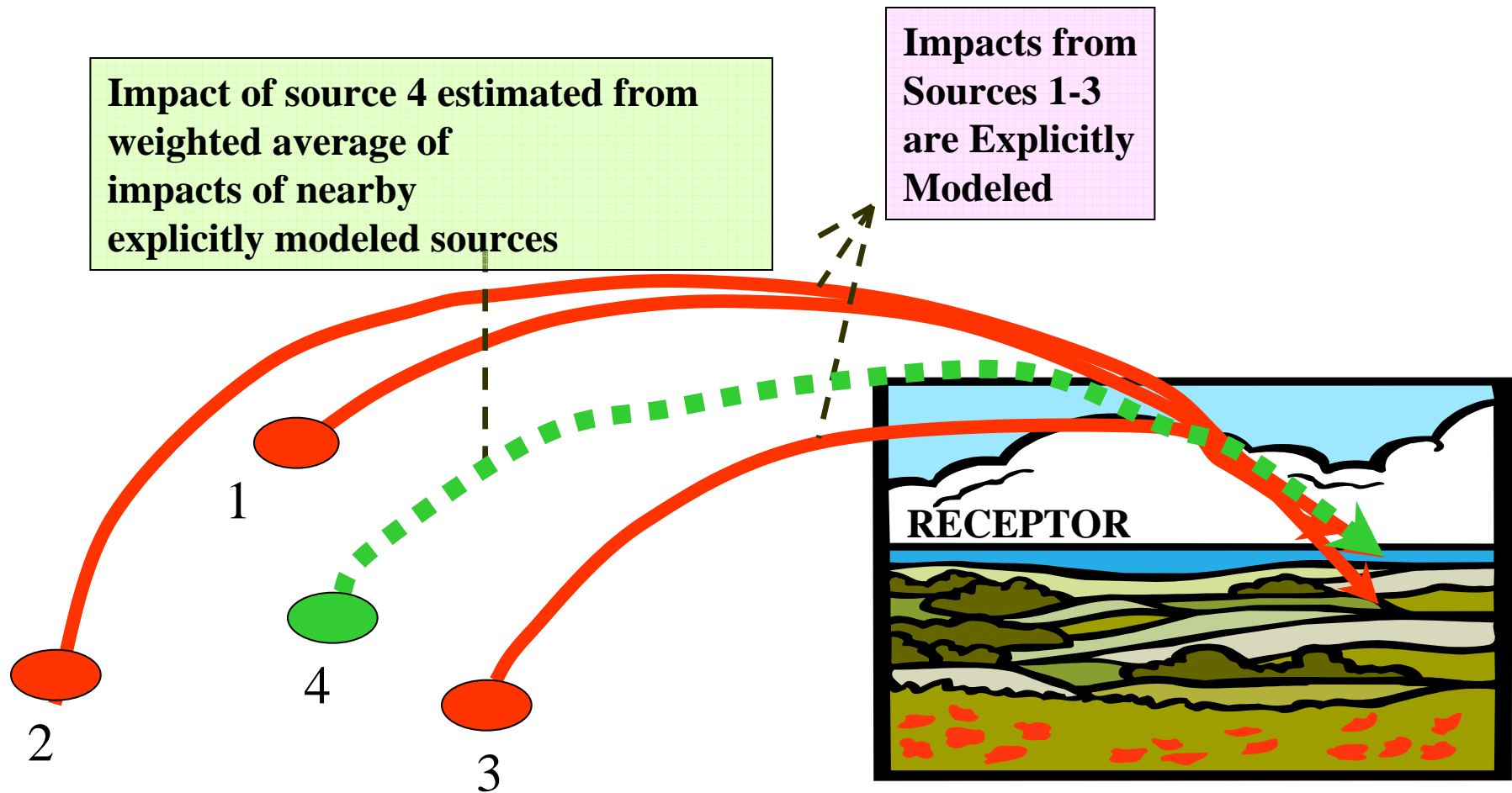
Thanks

EXTRA SLIDES

Why might the atmospheric fate of mercury emissions be essentially linearly independent?

- Hg is present at extremely trace levels in the atmosphere
- Hg won't affect meteorology (can simulate meteorology independently, and provide results to drive model)
- Most species that complex or react with Hg are generally present at *much* higher concentrations than Hg
- Other species (e.g. OH) generally react with many other compounds than Hg, so while present in trace quantities, their concentrations cannot be strongly influenced by Hg
- Wet and dry deposition processes are generally 1st order with respect to Hg
- The current “consensus” chemical mechanism (equilibrium + reactions) does not contain any equations that are not 1st order in Hg

Spatial interpolation

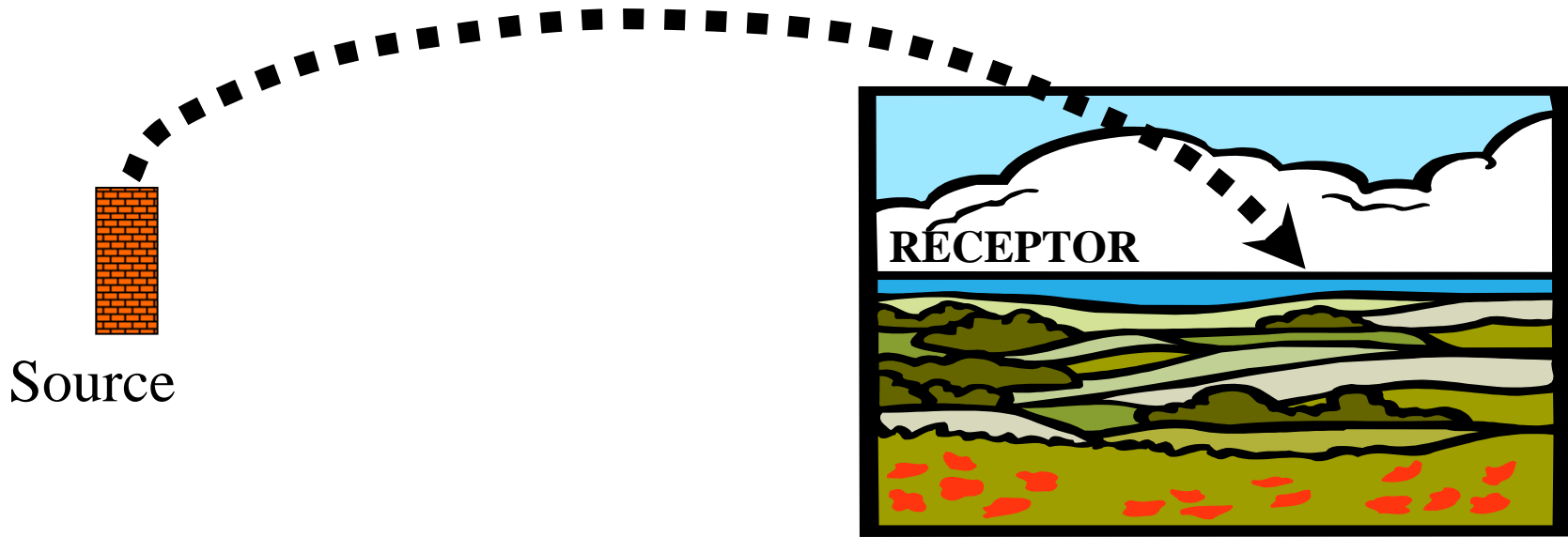


- **Perform separate simulations at each location for emissions of pure Hg(0), Hg(II) and Hg(p)**

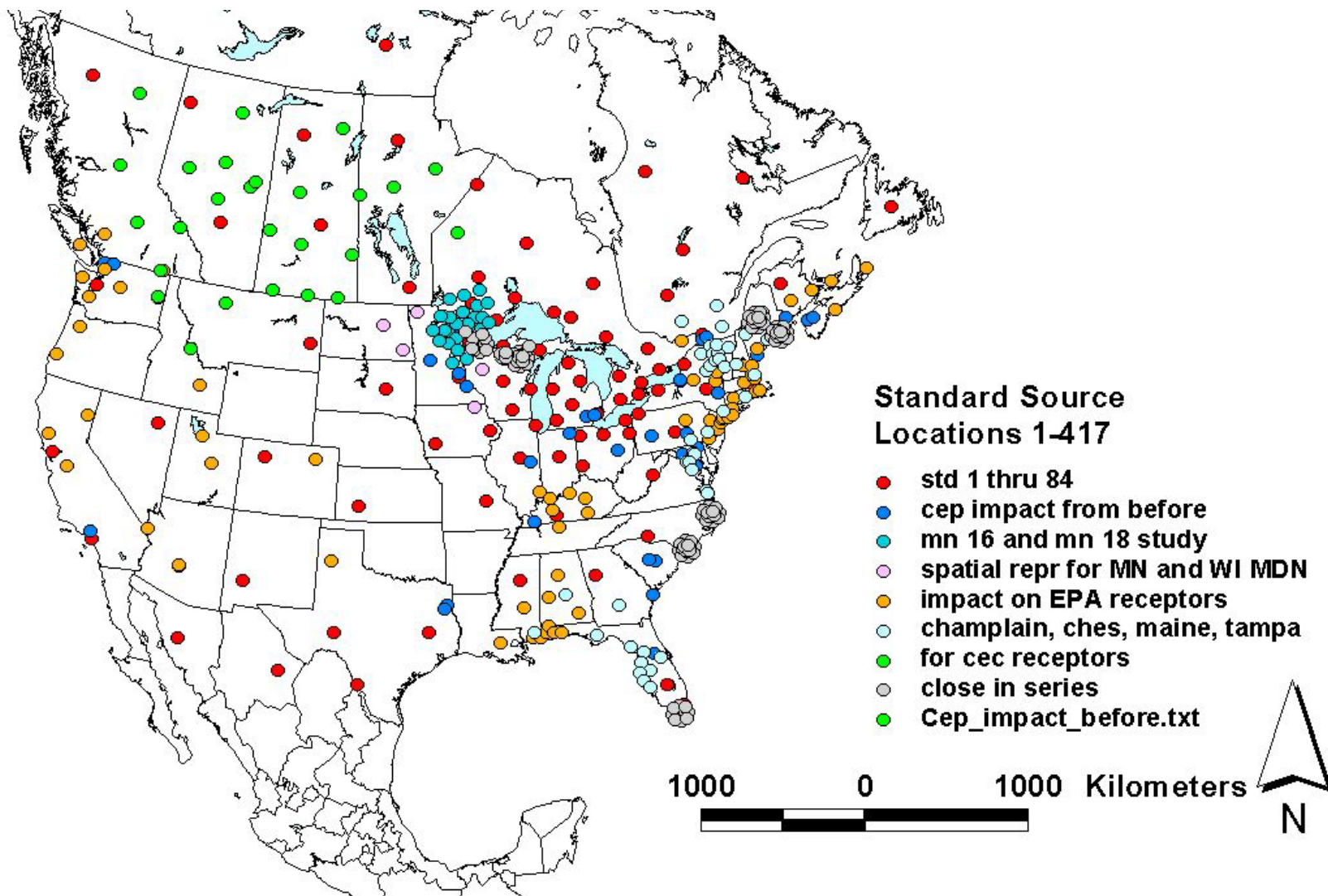
[after emission, simulate transformations between Hg forms]

- **Impact of emissions mixture taken as a linear combination of impacts of pure component runs on any given receptor**

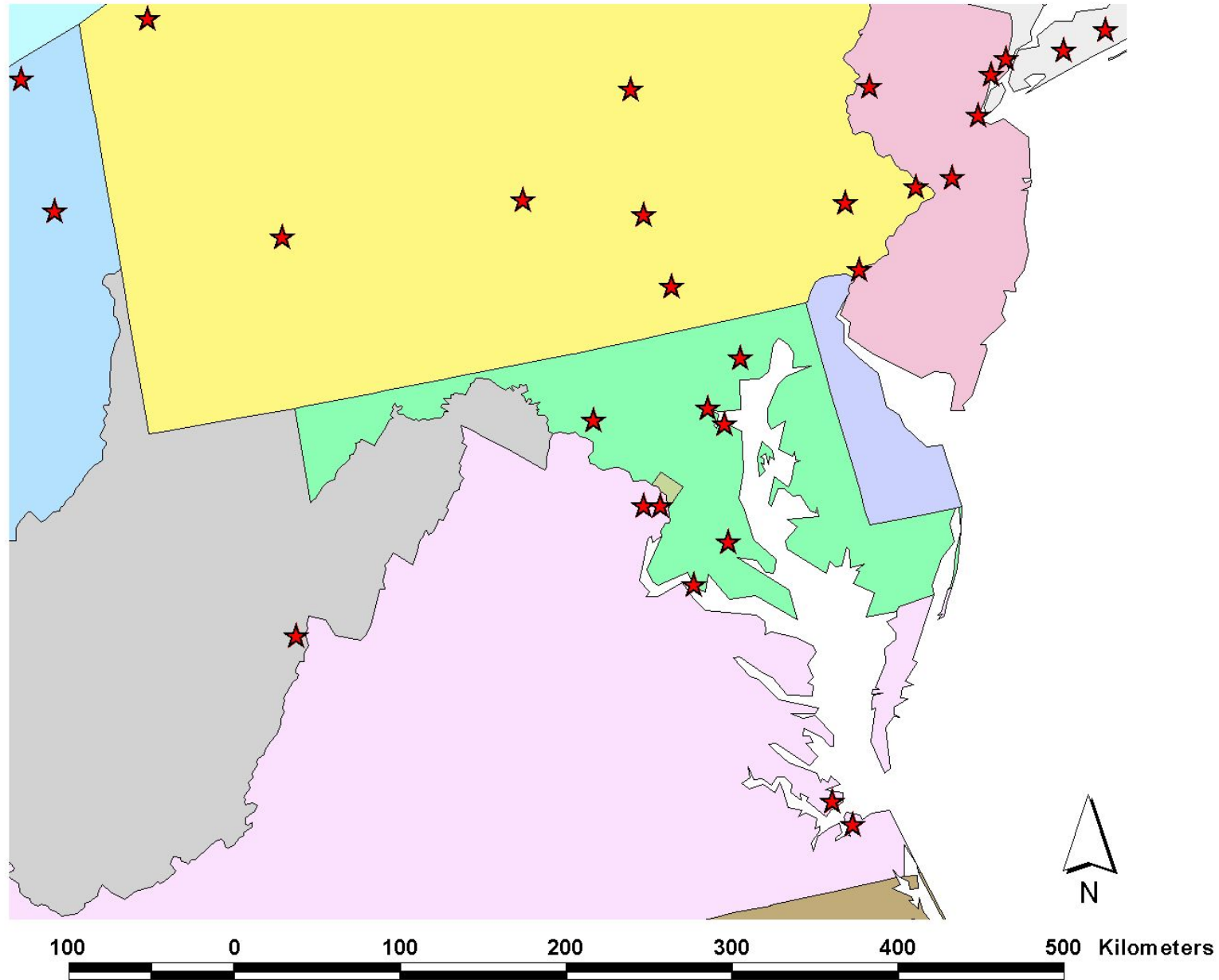
“Chemical Interpolation”



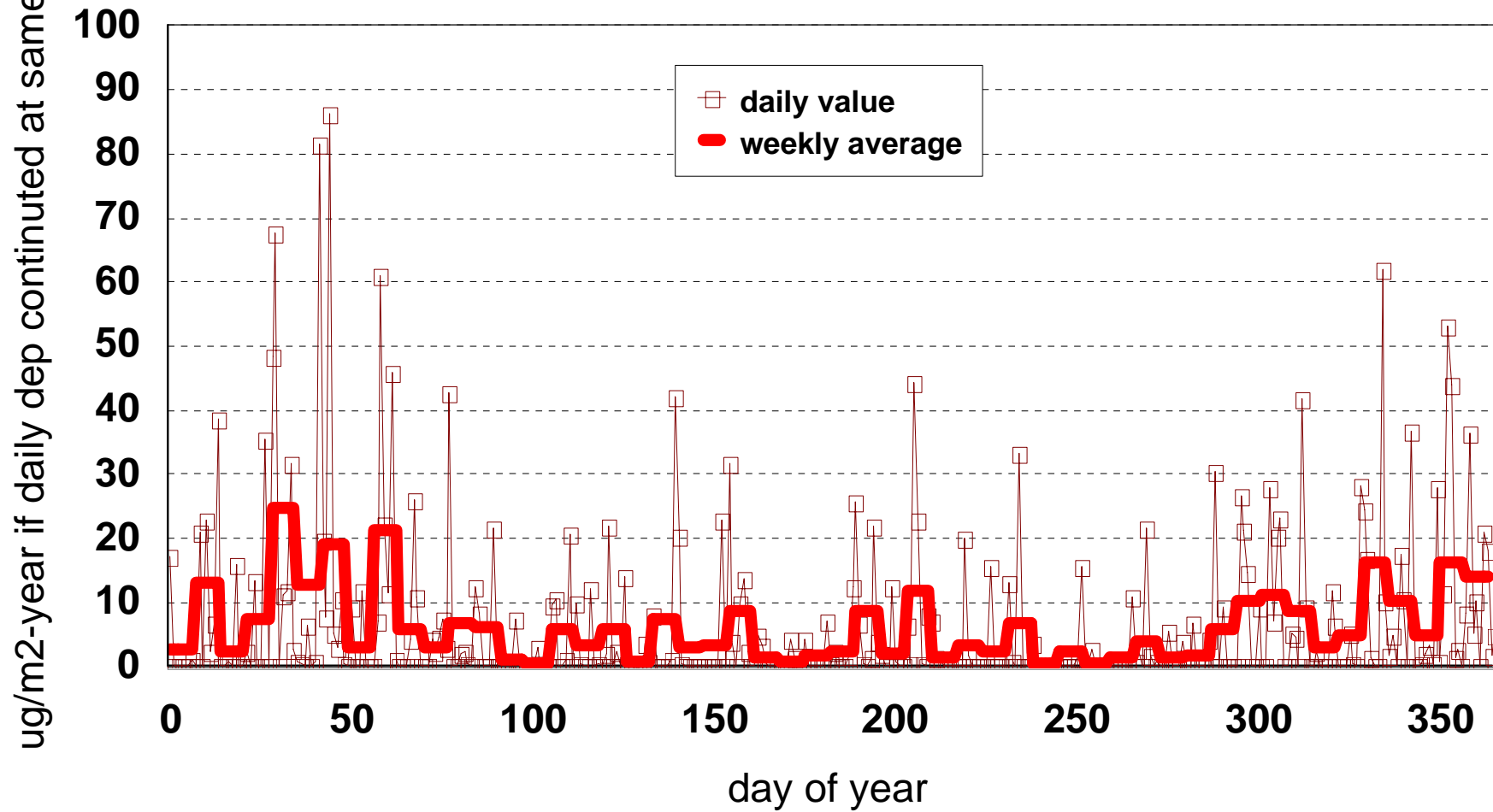
| | | | |
|---|----------|--------------|--|
| Impact of Source Emitting 30% Hg(0) 50% Hg(II) 20% Hg(p) | = | 0.3 x | Impact of Source Emitting Pure Hg(0) |
| | | + | |
| | | 0.5 x | Impact of Source Emitting Pure Hg(II) |
| | | + | |
| | | 0.2 x | Impact of Source Emitting Pure Hg(p) |

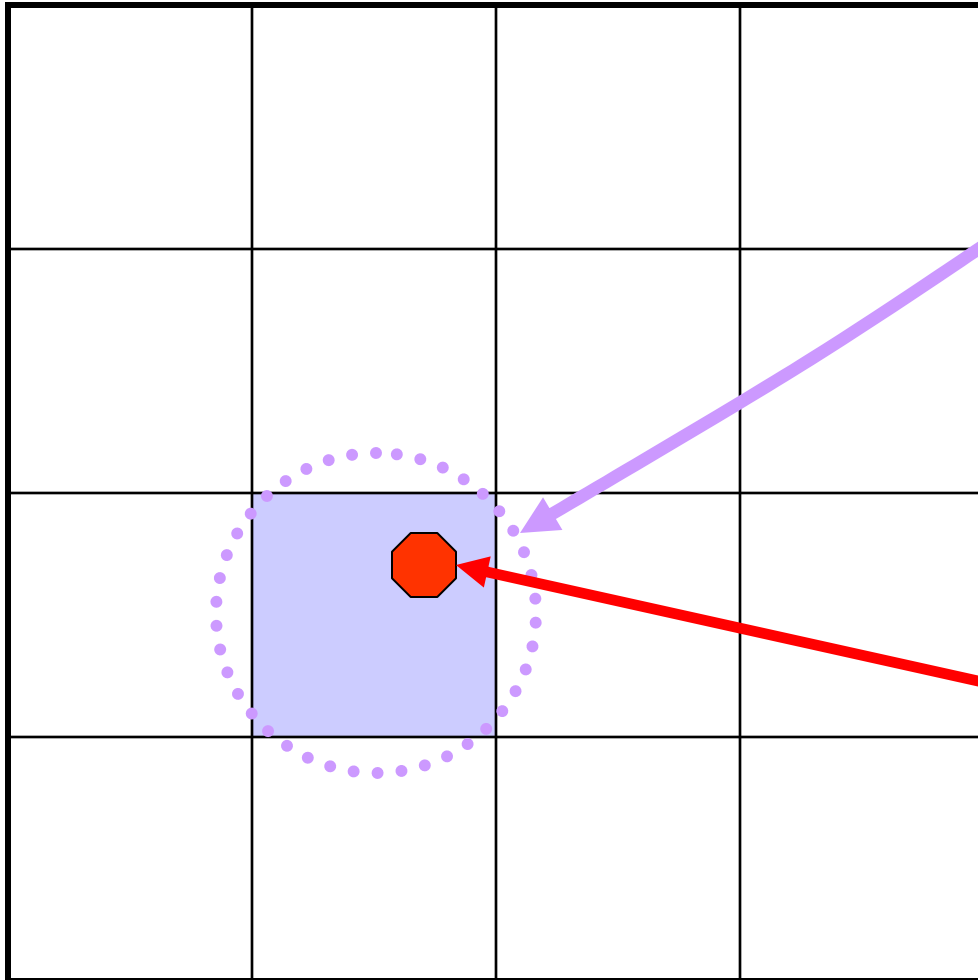


Standard Source Locations in Maryland region during recent simulation



Illustrative example of total deposition at a location
~40 km "downwind" of a 1 kg/day RGM source

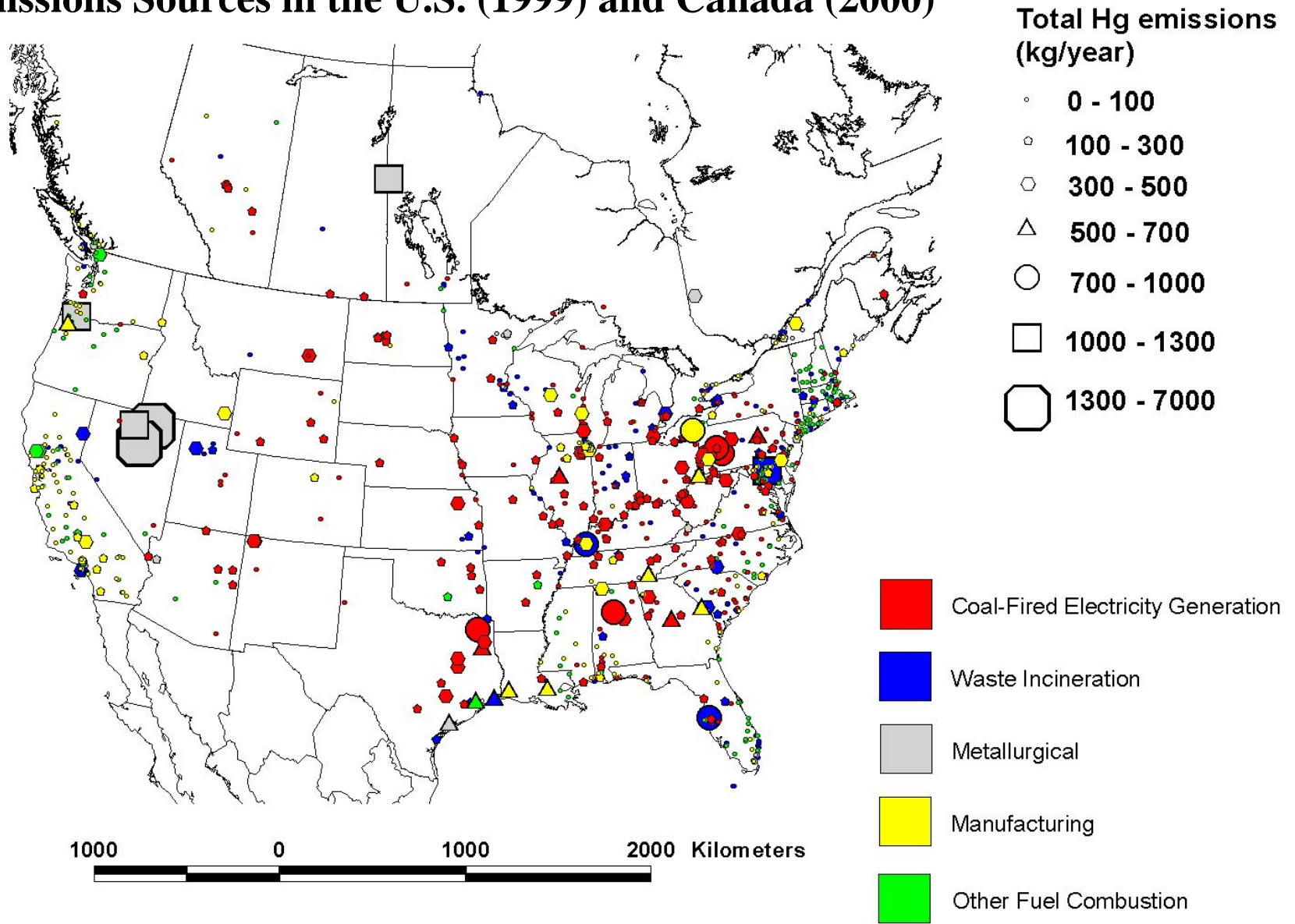


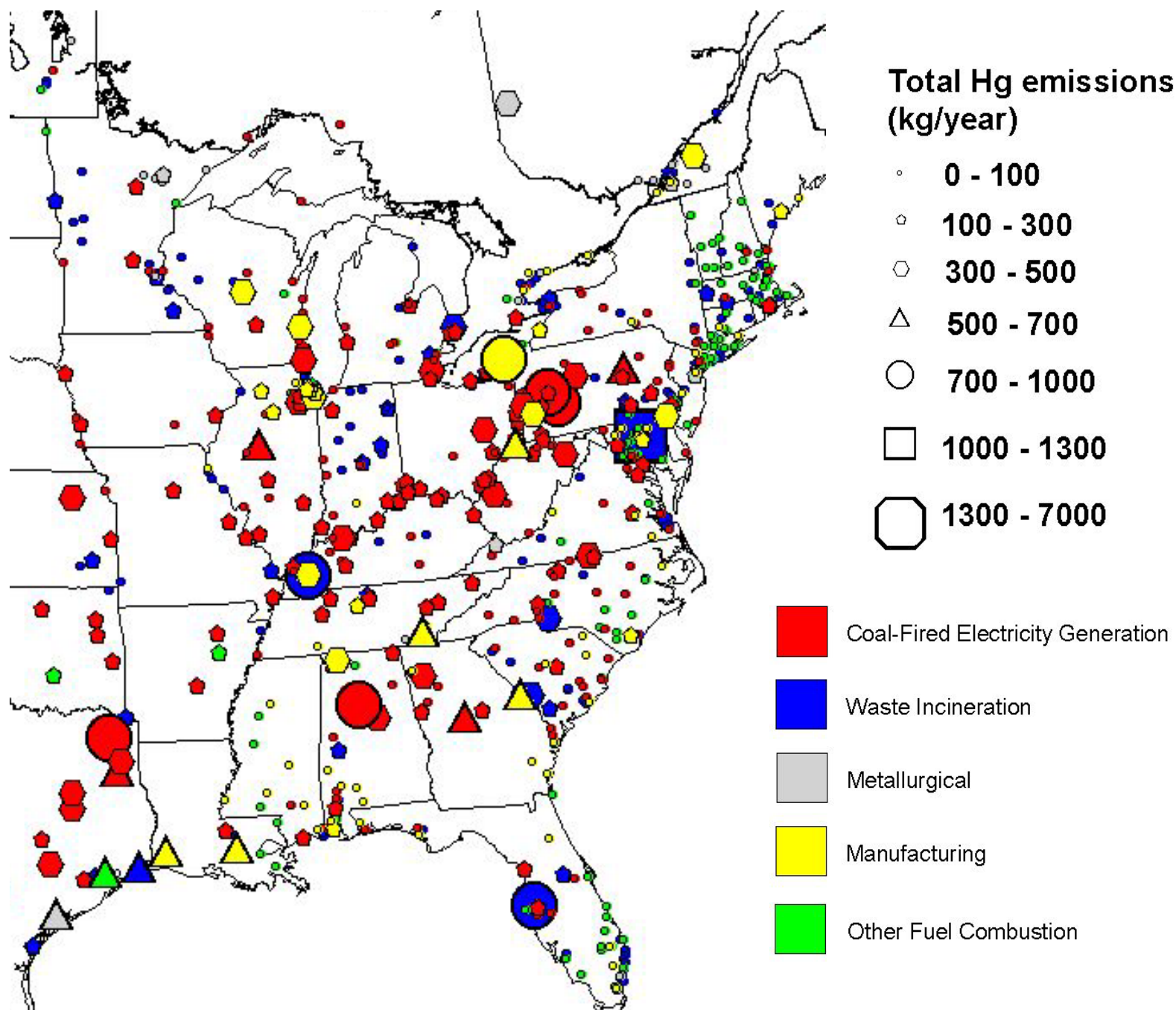


**Eulerian grid
models give
grid-averaged
estimates –**

**...difficult to
compare against
measurement at a
single location**

Geographic Distribution of Largest Anthropogenic Mercury Emissions Sources in the U.S. (1999) and Canada (2000)





- **In principle, we need do this for each source in the inventory**
- **But, since there are more than 100,000 sources in the U.S. and Canadian inventory, we need shortcuts...**
- **Shortcuts described in Cohen *et al* *Environmental Research* 95(3), 247-265, 2004**



Modeling the atmospheric transport and deposition of mercury to the Great Lakes[☆]

Mark Cohen,^{a,*} Richard Artz,^a Roland Draxler,^a Paul Miller,^b Laurier Poissant,^c David Niemi,^d Dominique Ratté,^d Marc Deslauriers,^d Roch Duval,^e Rachelle Laurin,^{e,d} Jennifer Slotnick,^f Todd Nettesheim,^g and John McDonald^h

^a NOAA Air Resources Laboratory, 1315 East West Highway, BART, Room 316, Silver Spring, MD 20910, USA

^b Commission for Environmental Cooperation, Montreal, Que., Canada

^c Atmospheric Toxic

^d Bio

Cohen, M., Artz, R., Draxler, R., Miller, P., Poissant, L., Niemi, D., Ratté, D., Deslauriers, M., Duval, R., Laurin, R., Slotnick, J., Nettesheim, T., McDonald, J.
 “Modeling the Atmospheric Transport and Deposition of Mercury to the Great Lakes.” *Environmental Research* 95(3), 247-265, 2004.

Abstract

A special version of mercury in a North American region and provide estimates of atmospheric mercury available for model evaluation in the Great Lakes region from the Great Lakes. Significant contribution to atmospheric mercury is published by Elsevier.

Keywords: Mercury; Atm

Mercury contamination of other ecosystems is a serious environmental problem. Human exposure to mercury, and significant mercury levels are believed to be of concern (Cohen et al., 2000). Historical mercury production using the Great Lakes region is believed to have caused in-

[☆] Supplementary data associated with this article can be found in the online version, at [doi:10.1016/j.envres.2003.11.007](http://doi.org/10.1016/j.envres.2003.11.007).

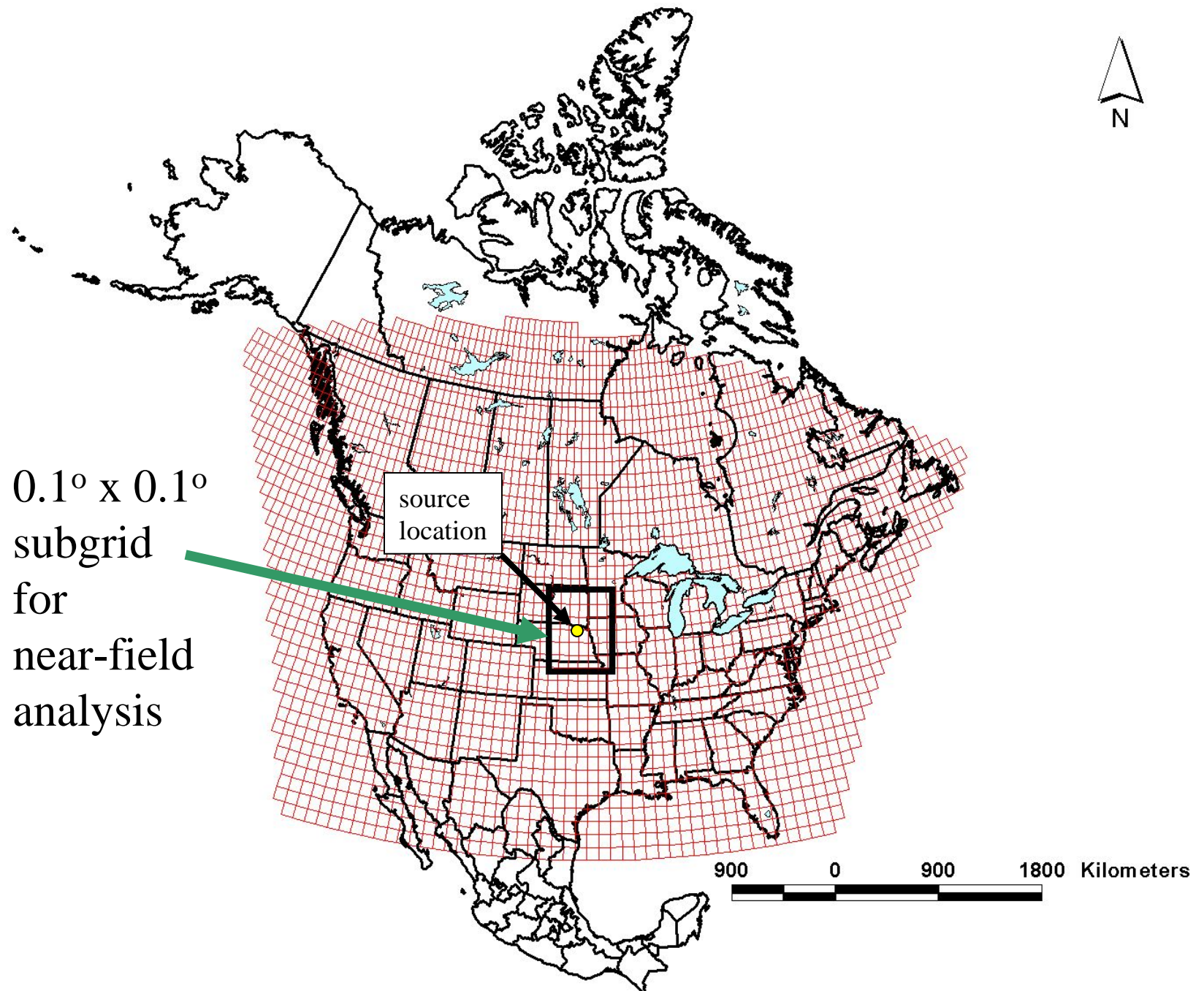
* Corresponding author.

E-mail address: mark.cohen@noaa.gov (M. Cohen).

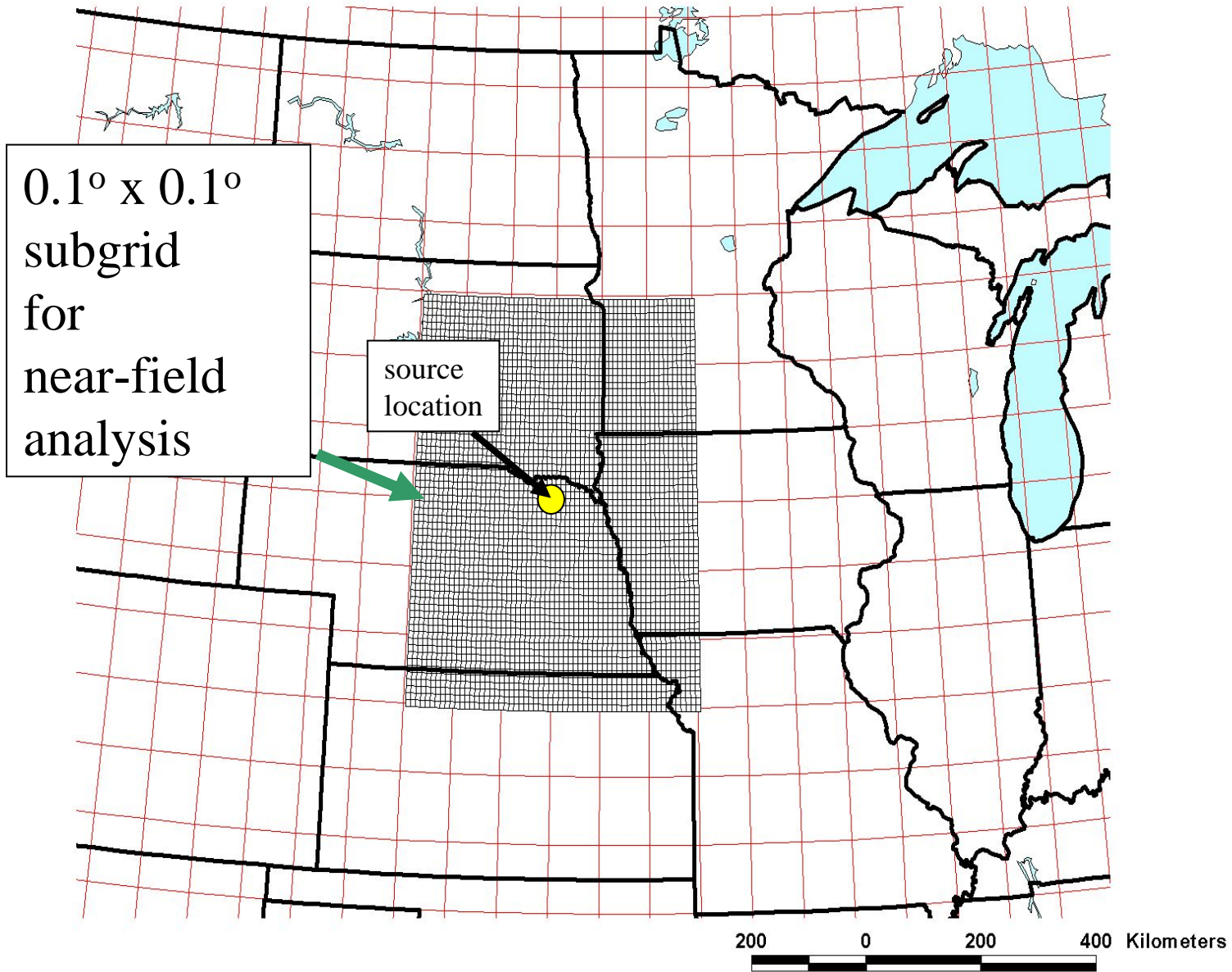
^c Current address: IPRAC, Canada, The Institute of Environmental Research, Concord, Ontario, Canada.

has developed detailed source-receptor relationships for the Great Lakes, as advocated in Annex 15 of the Great

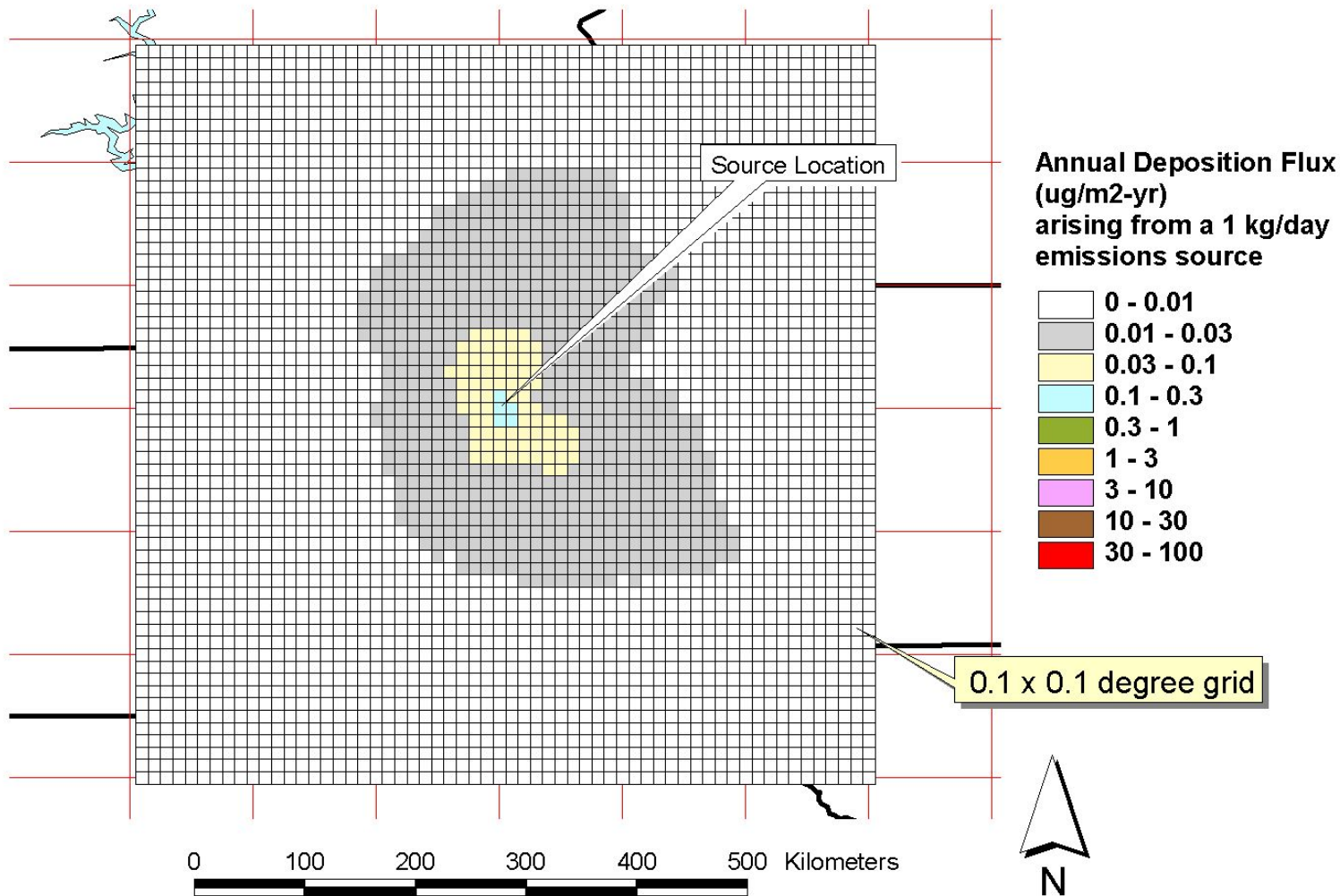
- **For each run, simulate fate and transport *everywhere*, but only keep track of impacts on each selected receptor (e.g., Great Lakes, Chesapeake Bay, etc.)**
- **Only run model for a limited number (~100) of hypothetical, individual unit-emissions sources throughout the domain**
- **Use spatial interpolation to estimate impacts from sources at locations not explicitly modeled**



0.1° x 0.1°
subgrid
for
near-field
analysis

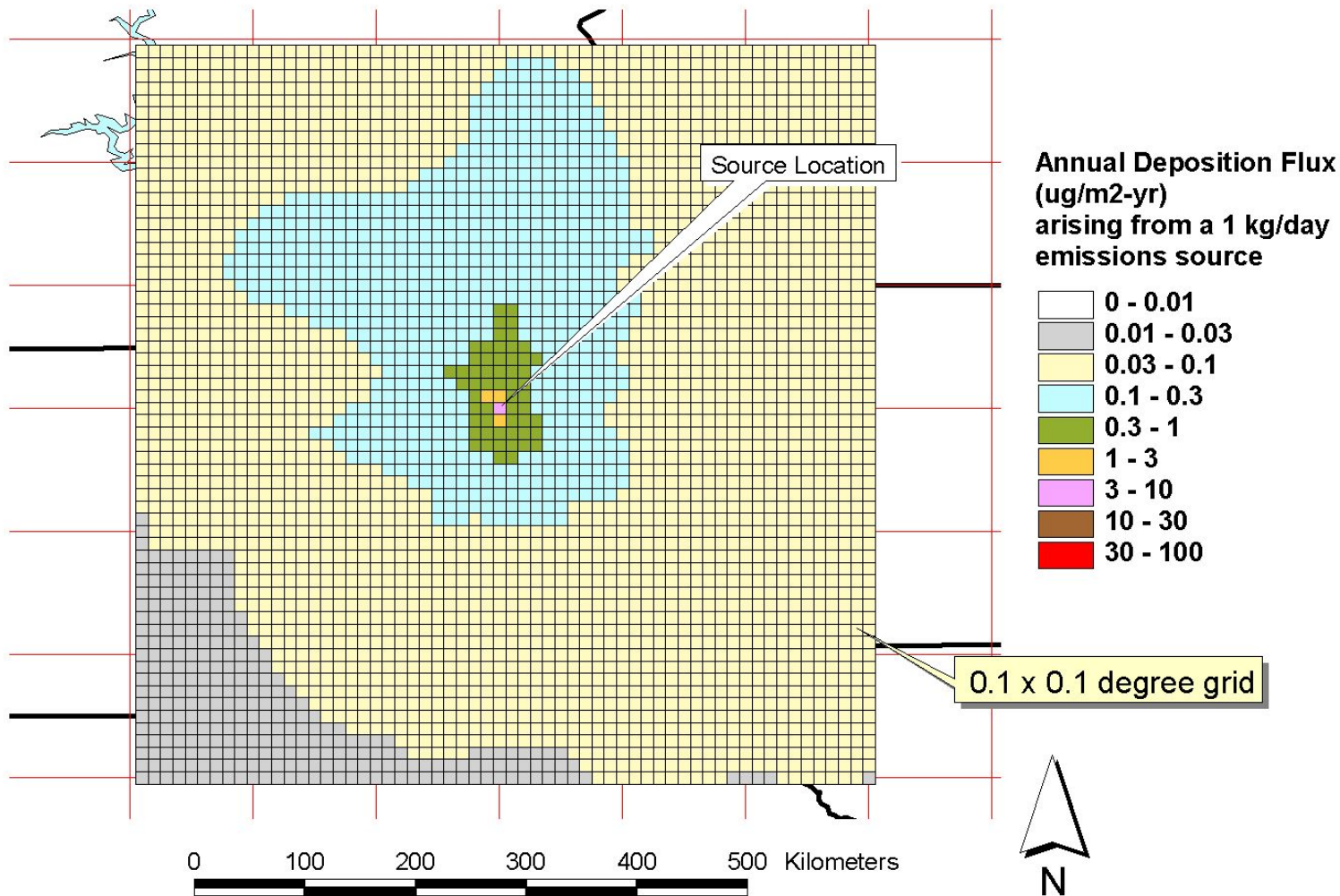


Annual deposition summary for emissions of elemental Hg from a 250 meter high source



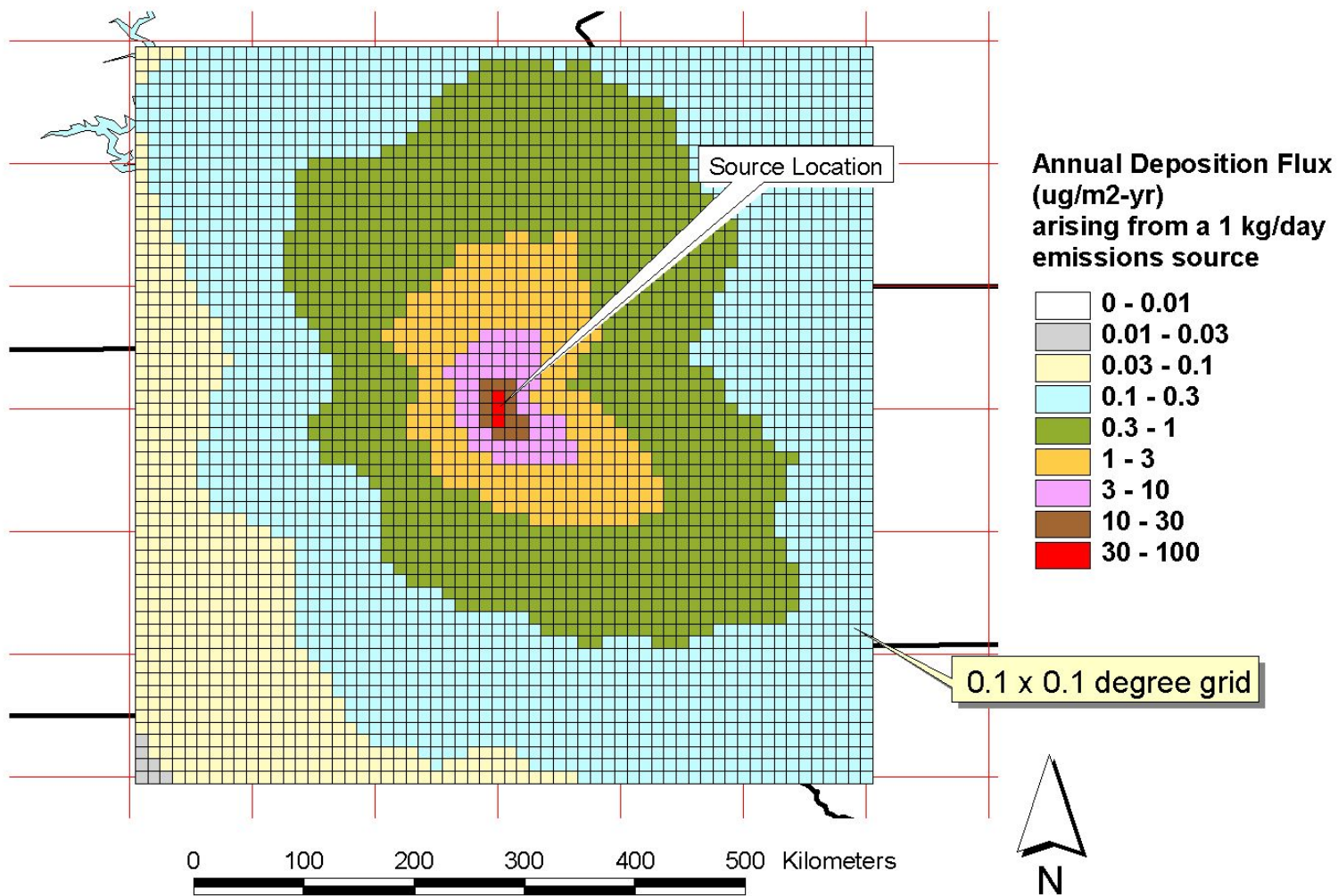
Hypothetical emissions source at lat = 42.5, long = -97.5;
simulation for entire year 1996 using archived NGM meteorology (180 km resolution)

Annual deposition summary for emissions of particulate Hg from a 250 meter high source



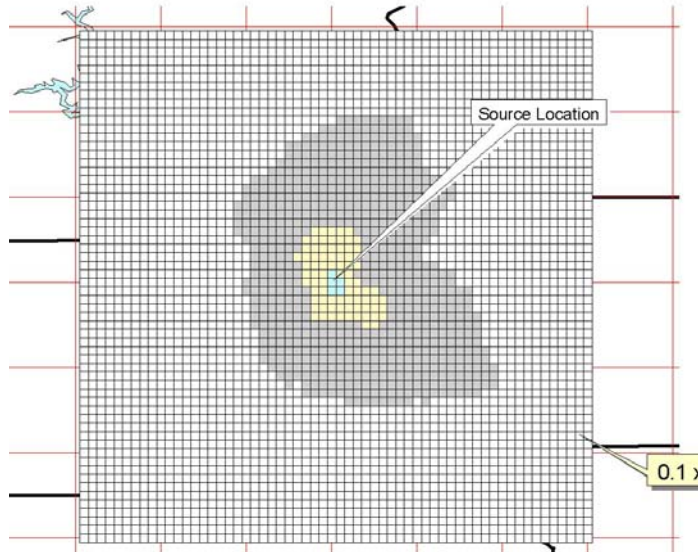
Hypothetical emissions source at lat = 42.5, long = -97.5;
simulation for entire year 1996 using archived NGM meteorology (180 km resolution)

Annual deposition summary for emissions of ionic Hg from a 250 meter high source

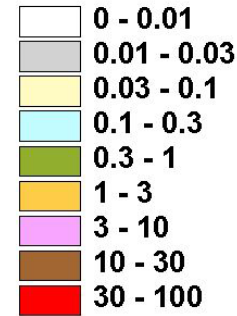


Hypothetical emissions source at lat = 42.5, long = -97.5;
simulation for entire year 1996 using archived NGM meteorology (180 km resolution)

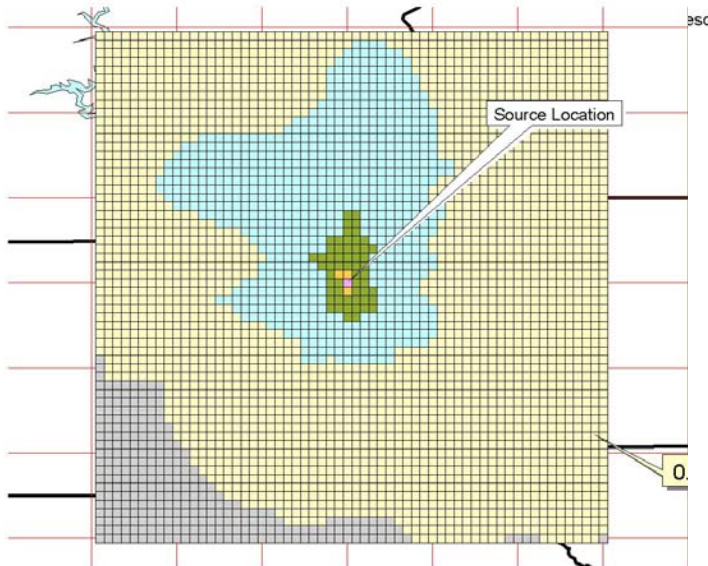
Annual deposition summary for emissions of elemental Hg from a 250 meter high source



Annual Deposition Flux (ug/m2-yr) arising from a 1 kg/day emissions source



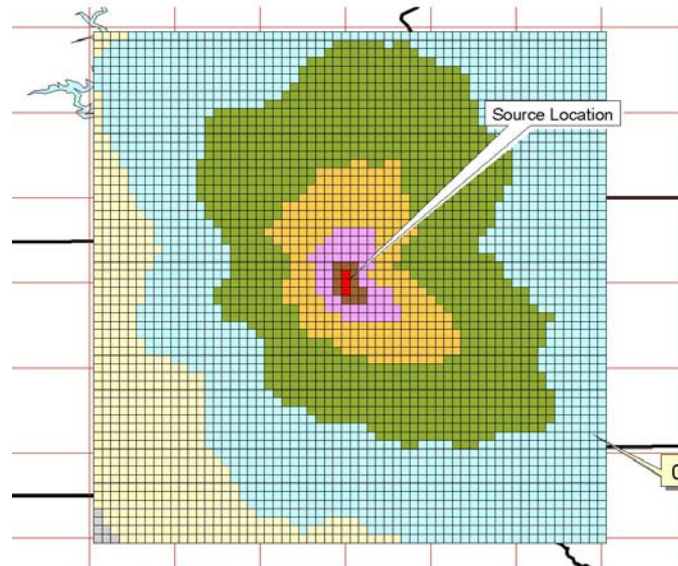
Annual deposition summary for emissions of particulate Hg from a 250 meter high source



0 100 200 300 400 500 Kilometers

Hypothetical emissions source at lat = 42.5, long = -97.5; simulation for entire year 1996 using archived NGM meteorology (180 km r

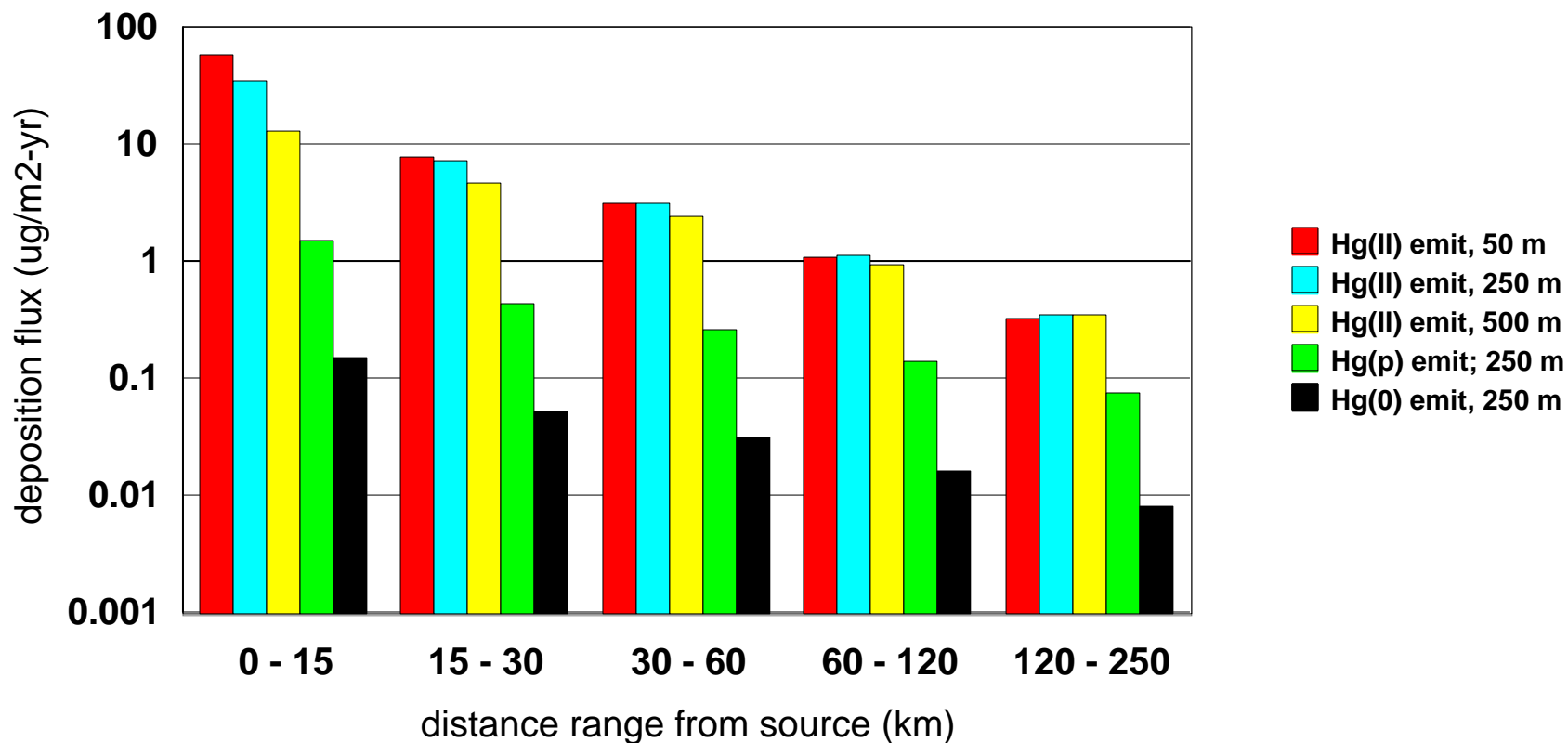
Annual deposition summary for emissions of ionic Hg from a 250 meter high source



0 100 200 300 400 500 Kilometers

Hypothetical emissions source at lat = 42.5, long = -97.5; simulation for entire year 1996 using archived NGM meteorology (180 km r

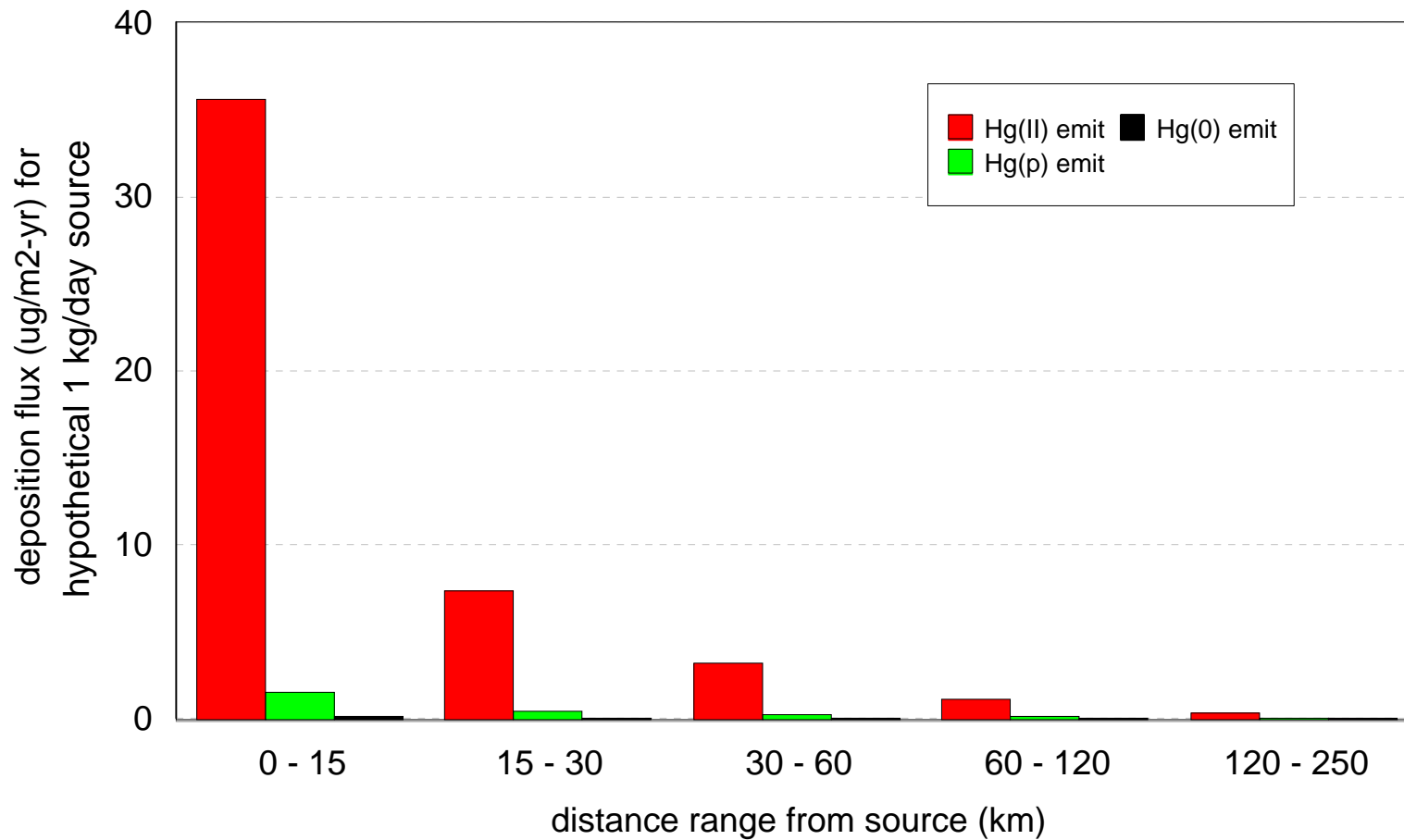
Deposition flux within different distance ranges from a hypothetical 1 kg/day source



Source at Lat = 42.5, Long = -97.5; simulation for entire year 1996 using archived NGM meteorological data

*Hypothesized rapid reduction of Hg(II) in plumes?
If true, then dramatic impact on modeling results...*

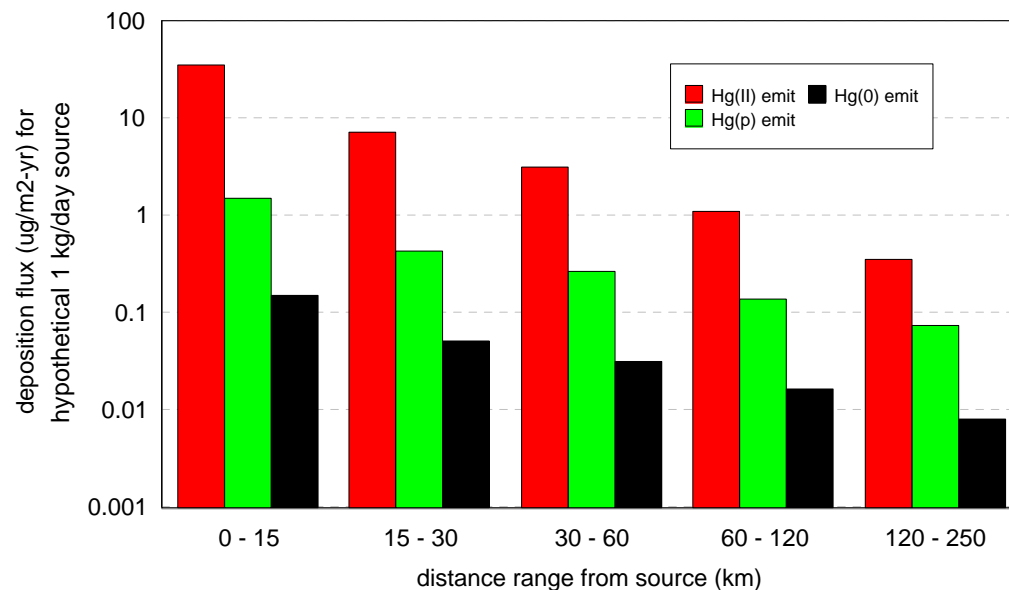
Why is emissions speciation information critical?



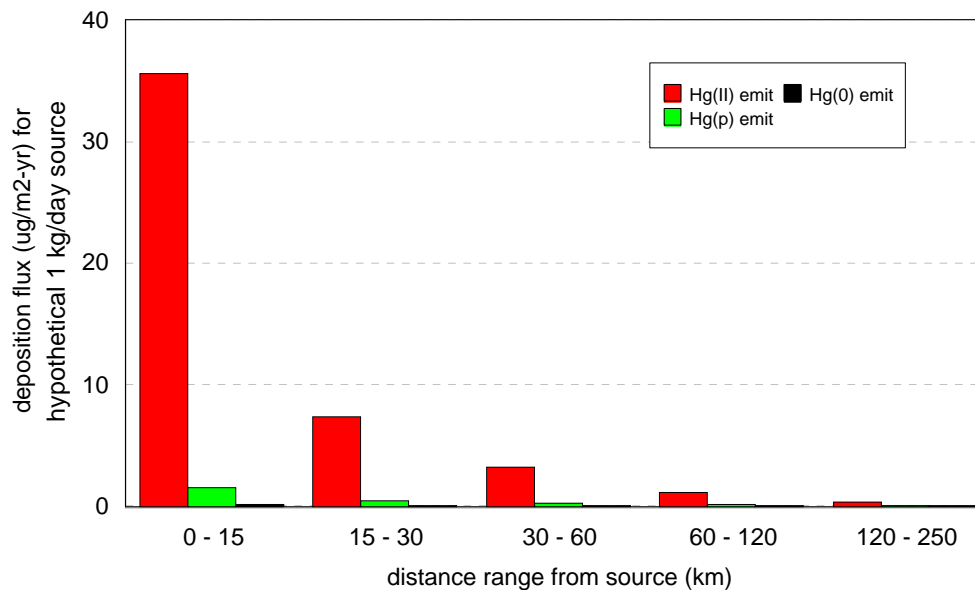
Linear

Why is emissions speciation information critical?

Logarithmic



Linear

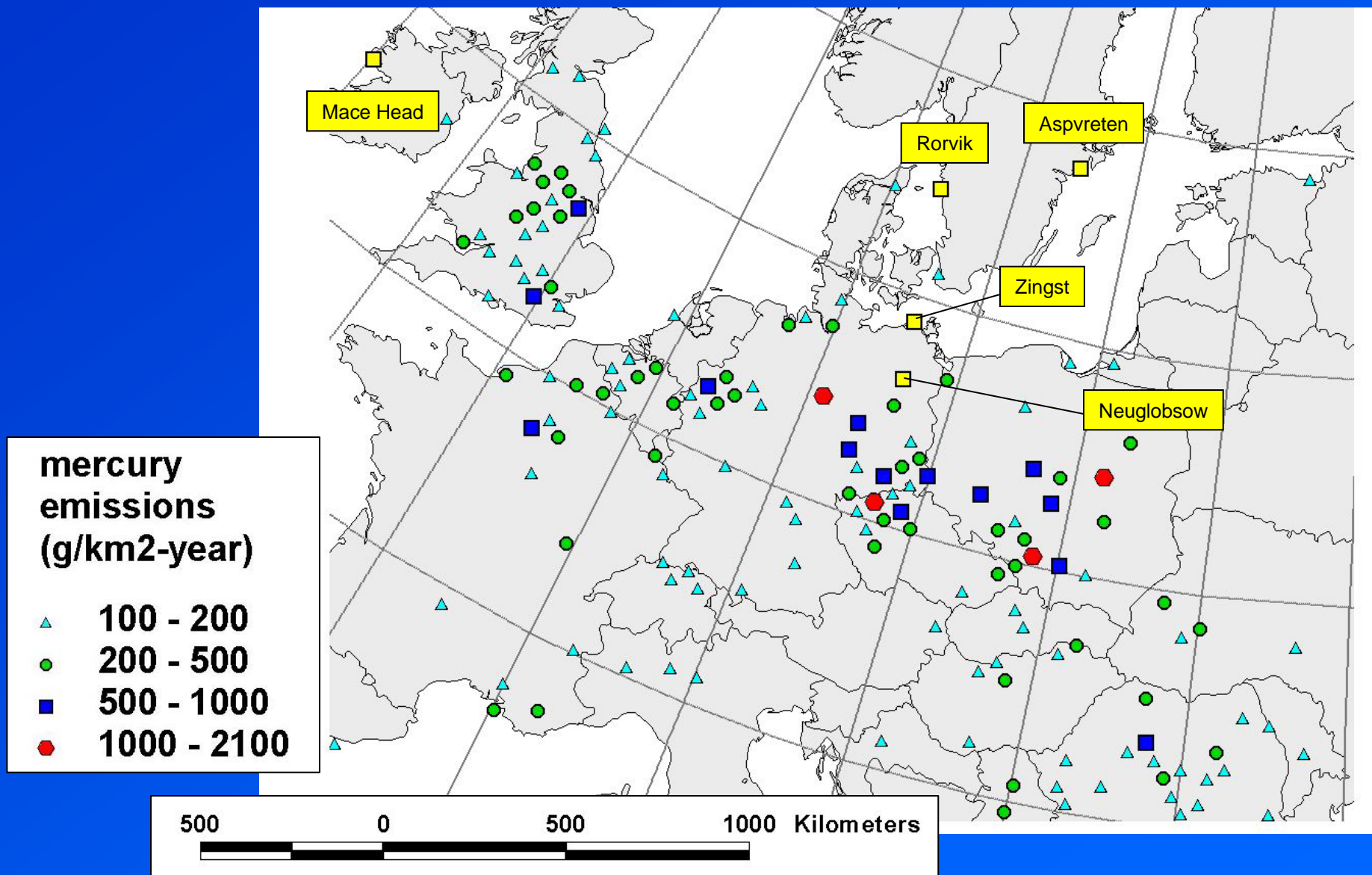


Emissions and Chemistry

- ❑ *The form of mercury emissions (elemental, ionic, particulate) is often very poorly known, but is a dominant factor in estimating deposition (and associated source-receptor relationships)*
- ❑ Questions regarding atmospheric chemistry of mercury may also be very significant
- ❑ *The above may contribute more to the overall uncertainties in atmospheric mercury models than uncertainties in dry and wet deposition algorithms*

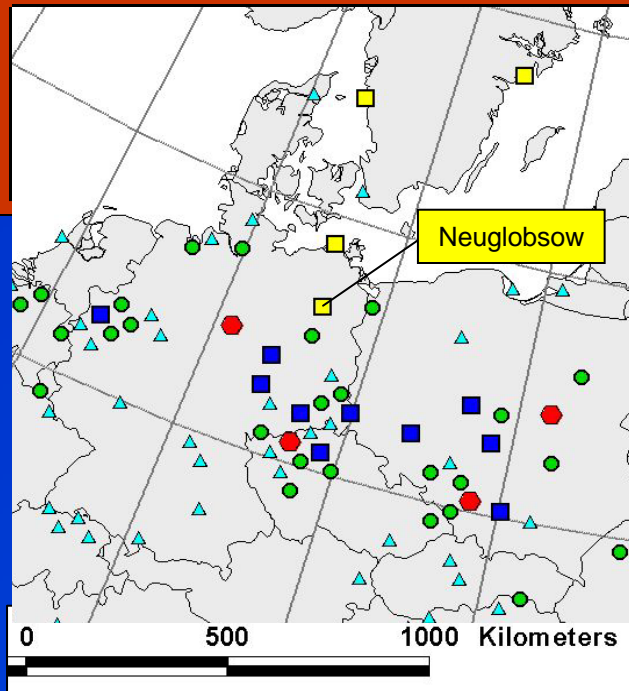
EMEP Intercomparison Study of Numerical Models for Long-Range Atmospheric Transport of Mercury

| | | | | | | | | |
|-------------------|-----------|-----------------|-------|-----|-----------|---------|---------|------------------|
| Intro- duction | Stage I | Stage II | | | Stage III | | | Conclu- sions |
| | Chemistry | Hg ⁰ | Hg(p) | RGM | Wet Dep | Dry Dep | Budgets | |

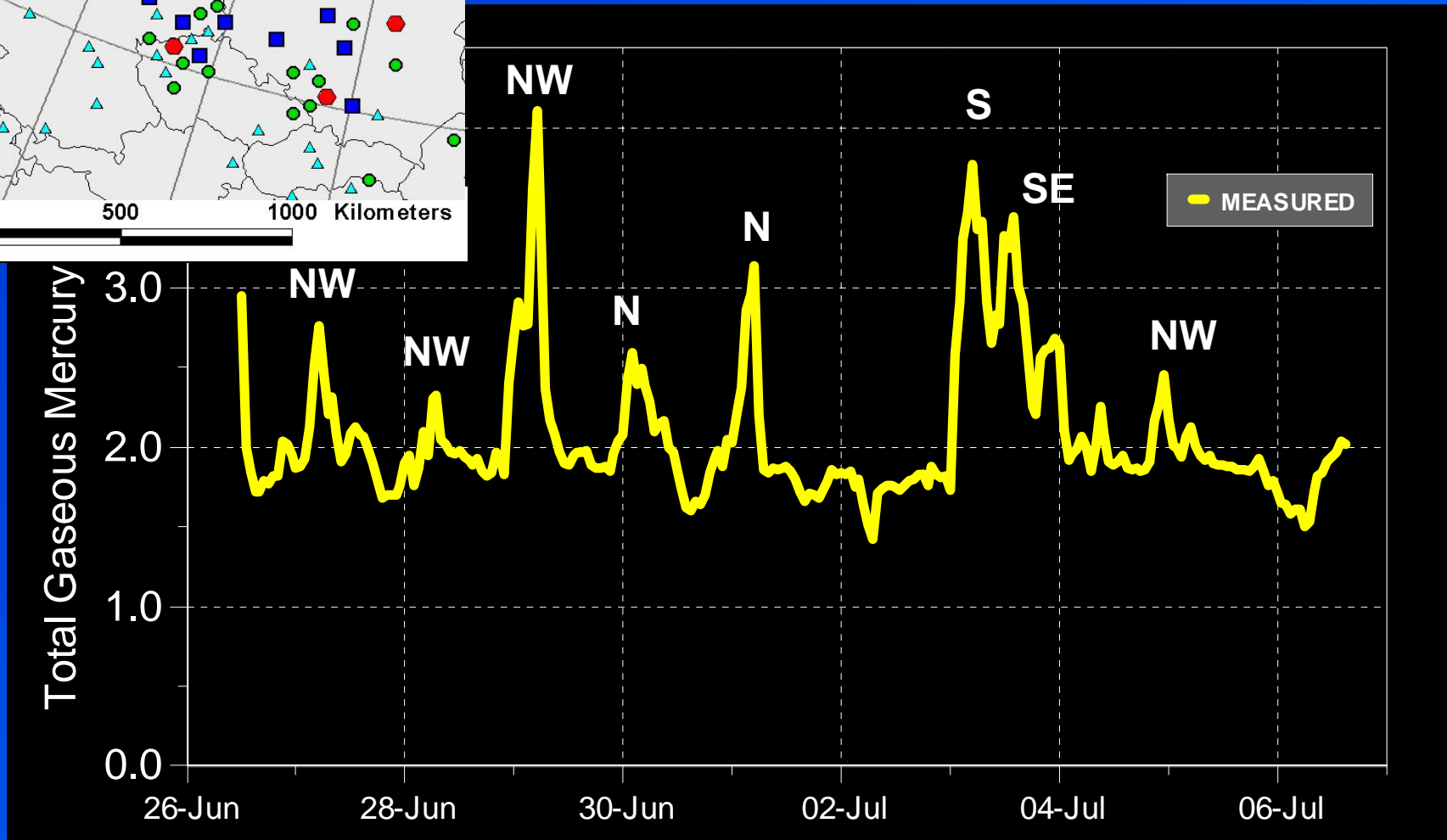


of Numerical Models for Long-Range Atmospheric Transport of Mercury

| Stage II | | Stage III | | | Conclu- sions |
|----------|-----|-----------|---------|---------|------------------|
| Hg(p) | RGM | Wet Dep | Dry Dep | Budgets | |



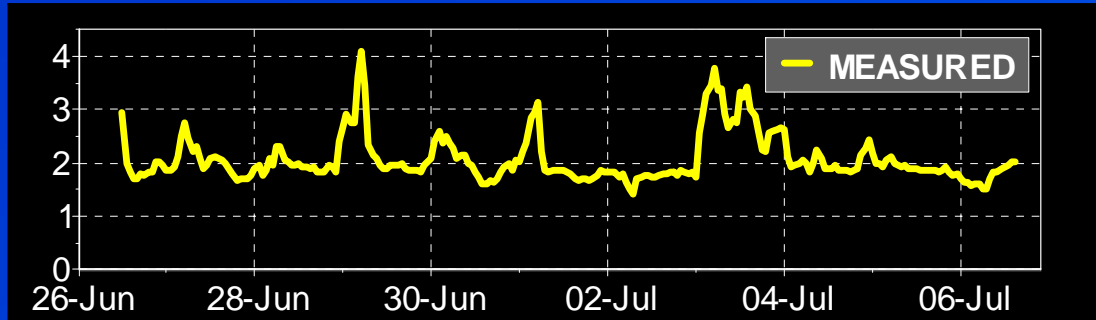
Mercury at Neuglobsow: June 26 – July 6, 1995



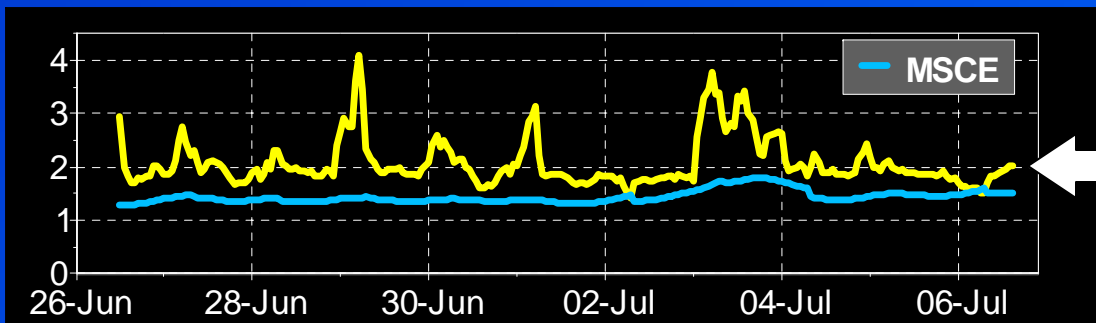
EMEP Intercomparison Study of Numerical Models for Long-Range Atmospheric Transport of Mercury

| | | | | | | | | |
|-------------------|-----------|-----------------|-------|-----|-----------|---------|---------|------------------|
| Intro- duction | Stage I | Stage II | | | Stage III | | | Conclu- sions |
| | Chemistry | Hg ⁰ | Hg(p) | RGM | Wet Dep | Dry Dep | Budgets | |

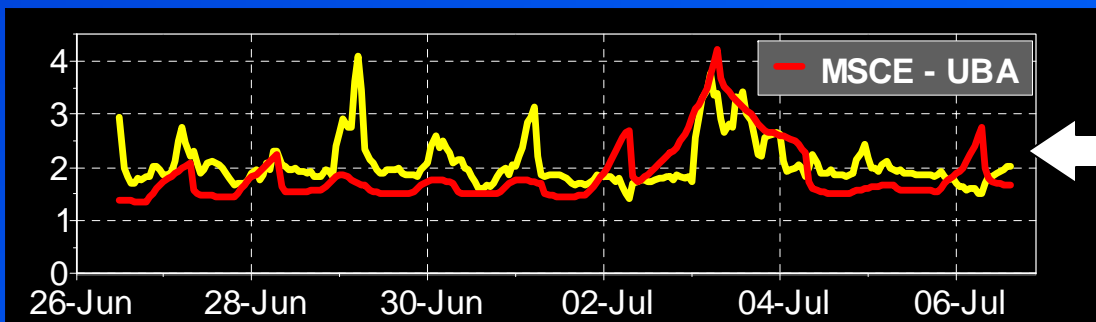
Total Gaseous Mercury (ng/m³) at Neuglobsow: June 26 – July 6, 1995



The emissions inventory is a critical input to the models...



Using default emissions inventory



Using alternative emissions inventory

Some Additional Measurement Issues (from a modeler's perspective)

- **Data availability**
- **Simple vs. Complex Measurements**
- **Process Information**

Process Information:

1. Dry Deposition - Resistance Formulation

$$V_d = \frac{1}{R_a + R_b + R_c + R_a R_b V_g} + V_g$$

in which

- R_a = aerodynamic resistance to mass transfer;
- R_b = resistance of the quasi-laminar sublayer;
- R_c = overall resistance of the canopy/surface (zero for particles)
- V_g = the gravitational settling velocity (zero for gases).

Dry Deposition

- ❑ depends intimately on vapor/particle partitioning and particle size distribution information
- ❑ resistance formulation [R_a , R_b , R_c ...]
- ❑ for gases, key uncertainty often R_c (e.g., “reactivity factor” f_0)
- ❑ for particles, key uncertainty often R_b
- ❑ How to evaluate algorithms when phenomena hard to measure?

Particle dry deposition phenomena

Ra

Atmosphere above the quasi-laminar sublayer

Rb

Quasi-laminar Sublayer (~ 1 mm thick)

Very small particles can *diffuse* through the layer like a gas

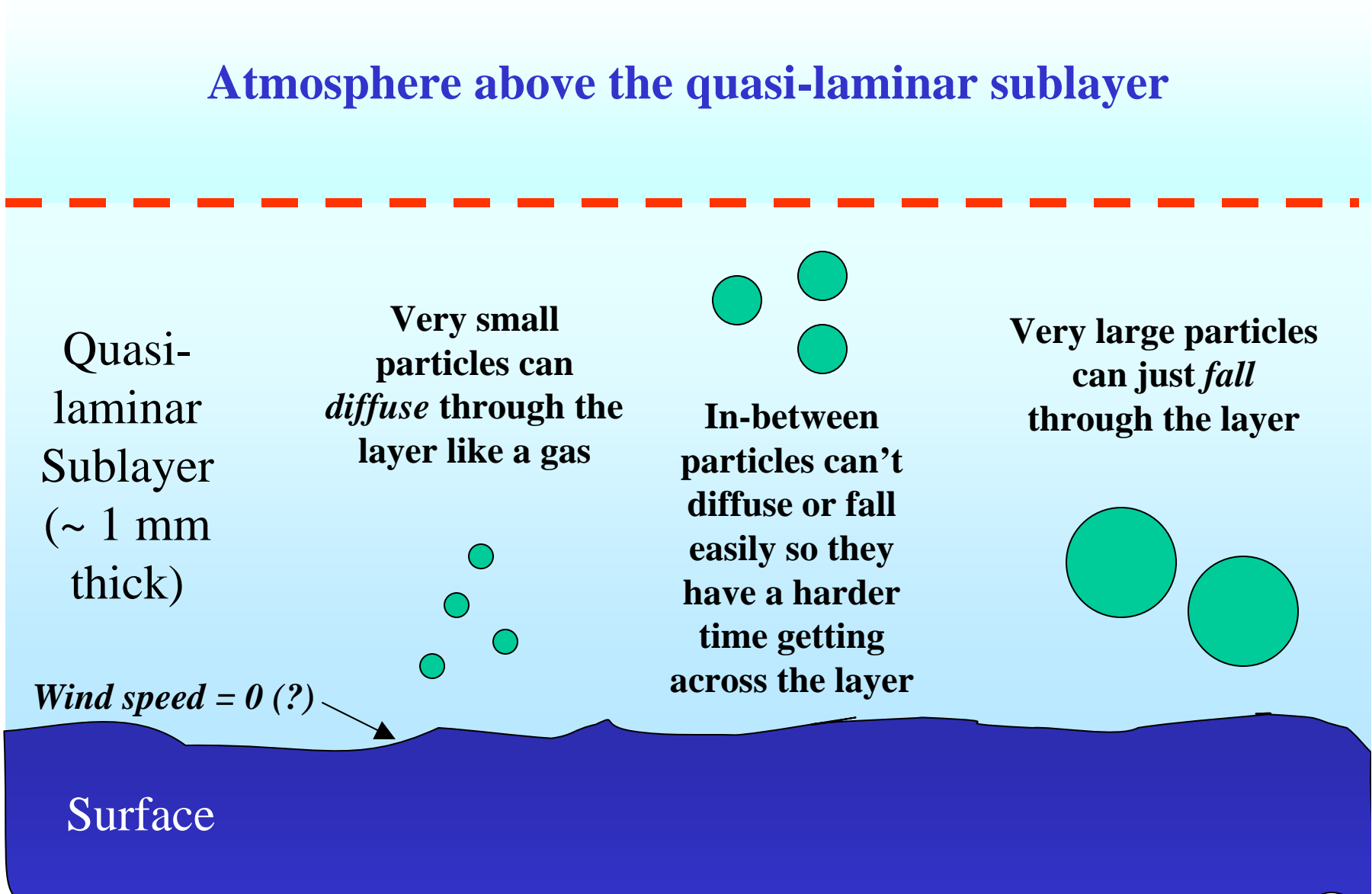
In-between particles can't diffuse or fall easily so they have a harder time getting across the layer

Very large particles can just *fall* through the layer

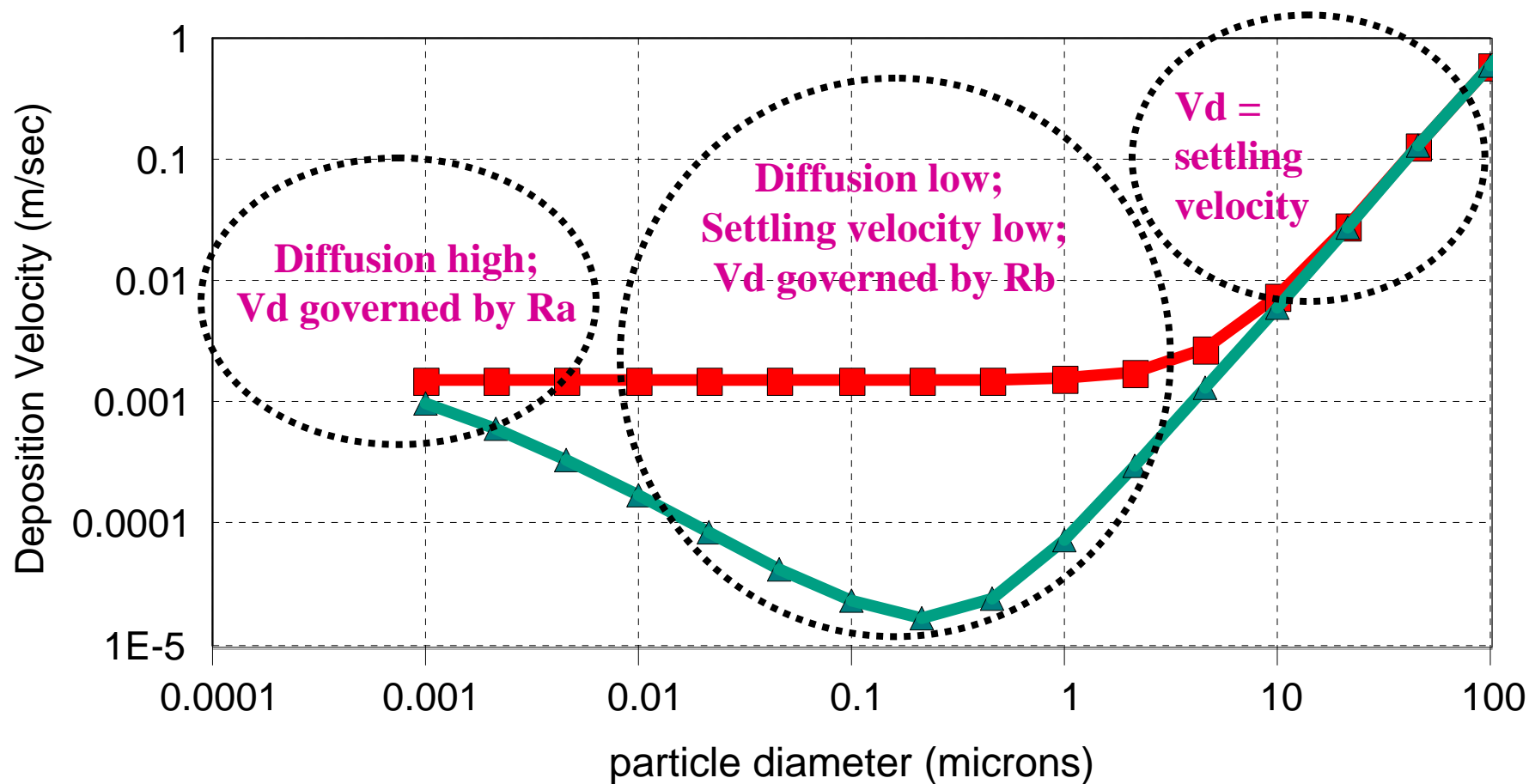
Rc

Wind speed = 0 (?)

Surface



Typical Deposition Velocities Over Water with Different Rb Formulations



■ Rb assumed small (= 10 sec/m) ▲ Slinn and Slinn

Process information needed:

1. For particle dry deposition, must have particle size distributions!

ATMOSPHERE

Gas-Phase
Pollutant

Particle-Phase
Pollutant

PROCESS INFORMATION:

2. The gas-exchange flux at a water surface depends on the concentration of pollutant in the gas-phase and the truly-dissolved phase (but these are rarely measured...)

Pollutant
Truly
Dissolved
in Water

Pollutant on
Suspended
Sediment

LAKE

