

**V_{ud} , V_{us} , THE CABIBBO ANGLE,
AND CKM UNITARITY**

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The Cabibbo-Kobayashi-Maskawa (CKM) [1,2] three-generation quark mixing matrix written in terms of the Wolfenstein parameters (λ, A, ρ, η) [3] nicely illustrates the orthonormality constraint of unitarity and central role played by λ .

$$V_{\text{CKM}} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} = \begin{pmatrix} 1 - \lambda^2/2 & \lambda & A\lambda^3(\rho - i\eta) \\ -\lambda & 1 - \lambda^2/2 & A\lambda^2 \\ A\lambda^3(1 - \rho - i\eta) & -A\lambda^2 & 1 \end{pmatrix} + \mathcal{O}(\lambda^4). \quad (1)$$

That cornerstone is a carryover from the two-generation Cabibbo angle, $\lambda = \sin(\theta_{\text{Cabibbo}}) = V_{us}$. Its value is a critical ingredient in determinations of the other parameters and in tests of CKM unitarity.

Unfortunately, the precise value of λ has been somewhat controversial in the past, with kaon decays suggesting [4] $\lambda \simeq 0.220$, while hyperon decays [5] and indirect determinations via nuclear β -decays imply a somewhat larger $\lambda \simeq 0.225 - 0.230$. That discrepancy is often discussed in terms of a deviation from the unitarity requirement

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1. \quad (2)$$

For many years, using a value of V_{us} derived from $K \rightarrow \pi e \nu$ (K_{e3}) decays, that sum was consistently 2–2.5 sigma below unity, a potential signal [6] for new physics effects. Below, we discuss the current status of V_{ud} , V_{us} , and their associated unitarity test in Eq. (2). (Since $|V_{ub}|^2 \simeq 1 \times 10^{-5}$ is negligibly small, it is ignored in this discussion.)

V_{ud}

The value of V_{ud} has been obtained from superallowed nuclear, neutron, and pion decays. Currently, the most precise

determination of V_{ud} comes from superallowed nuclear beta-decays [6] ($0^+ \rightarrow 0^+$ transitions). Measuring their half-lives, t , and Q values which give the decay rate factor, f , leads to a precise determination of V_{ud} via the master formula [7–9]

$$|V_{ud}|^2 = \frac{2984.48(5) \text{ sec}}{ft(1 + \text{RC})} \quad (3)$$

where RC denotes the entire effect of electroweak radiative corrections, nuclear structure, and isospin violating nuclear effects. RC is nucleus-dependent, ranging from about +3.0% to +3.6% for the nine best measured superallowed decays. In Table 1, we give updated [10] ft values along with their implied V_{ud} for the nine best measured superallowed decays [6, 10]. They collectively give a weighted average (with errors combined in quadrature) of

$$V_{ud} = 0.97418(27) \text{ (superallowed)} , \quad (4)$$

which, assuming unitarity, corresponds to $\lambda = 0.226(1)$. We note that the new average value of V_{ud} is shifted upward compared to our 2005 value of 0.97377(27) primarily because of a recent reevaluation of the isospin breaking Coulomb corrections by Towner and Hardy [10].

Combined measurements of the neutron lifetime, τ_n , and the ratio of axial-vector/vector couplings, $g_A \equiv G_A/G_V$, via neutron decay asymmetries can also be used to determine V_{ud} :

$$|V_{ud}|^2 = \frac{4908.7(1.9) \text{ sec}}{\tau_n(1 + 3g_A^2)} , \quad (5)$$

where the error stems from uncertainties in the electroweak radiative corrections [8] due to hadronic loop effects. Those effects have been recently updated and their error was reduced by about a factor of 2 [9], leading to a ± 0.0002 theoretical uncertainty in V_{ud} (common to all V_{ud} extractions). Using the world averages from this *Review*

$$\begin{aligned} \tau_n^{\text{ave}} &= 885.7(8) \text{ sec} \\ g_A^{\text{ave}} &= 1.2695(29) \end{aligned} \quad (6)$$

Table 1: Values of V_{ud} implied by various precisely measured superallowed nuclear beta decays. The ft values and Coulomb isospin breaking corrections are taken from Towner and Hardy [10]. Uncertainties in V_{ud} correspond to 1) nuclear structure and $Z^2\alpha^3$ uncertainties [6, 11] added in quadrature with the ft error; 2) a common error assigned to nuclear Coulomb distortion effects [11]; and 3) a common uncertainty in the radiative corrections from quantum loop effects [9]. Only the first error is used to obtain the weighted average.

Nucleus	ft (sec)	V_{ud}
^{10}C	3039.5(47)	0.97370(80)(14)(19)
^{14}O	3042.5(27)	0.97411(51)(14)(19)
^{26}Al	3037.0(11)	0.97400(24)(14)(19)
^{34}Cl	3050.0(11)	0.97417(34)(14)(19)
^{38}K	3051.1(10)	0.97413(39)(14)(19)
^{42}Sc	3046.4(14)	0.97423(44)(14)(19)
^{46}V	3049.6(16)	0.97386(49)(14)(19)
^{50}Mn	3044.4(12)	0.97487(45)(14)(19)
^{54}Co	3047.6(15)	0.97490(54)(14)(19)
Weighted Ave.		0.97418(13)(14)(19)

leads to

$$V_{ud} = 0.9746(4)_{\tau_n(18)}g_A(2)_{\text{RC}} \quad (7)$$

with the error dominated by g_A uncertainties (which have been expanded due to experimental inconsistencies). We note that a recent precise measurement [12] of $\tau_n = 878.5(7)(3)$ sec is also inconsistent with the world average from this *Review* and would lead to a considerably larger $V_{ud} = 0.9786(4)(18)(2)$. Future neutron studies are expected to resolve these inconsistencies and significantly reduce the uncertainties in g_A and τ_n , potentially making them the best way to determine V_{ud} .

The recently completed PIBETA experiment at PSI measured the very small ($\mathcal{O}(10^{-8})$) branching ratio for $\pi^+ \rightarrow \pi^0 e^+ \nu_e$ with about $\pm 1/2\%$ precision. Their result gives [13]

$$V_{ud} = 0.9749(26) \left[\frac{BR(\pi^+ \rightarrow e^+ \nu_e(\gamma))}{1.2352 \times 10^{-4}} \right]^{\frac{1}{2}} \quad (8)$$

which is normalized using the very precisely determined theoretical prediction for $BR(\pi^+ \rightarrow e^+ \nu_e(\gamma)) = 1.2352(5) \times 10^{-4}$ [7], rather than the experimental branching ratio from this *Review* of $1.230(4) \times 10^{-4}$ which would lower the value to $V_{ud} = 0.9728(30)$. Theoretical uncertainties in that determination are very small; however, much higher statistics would be required to make this approach competitive with others.

V_{us}

$|V_{us}|$ may be determined from kaon decays, hyperon decays, and tau decays. Previous determinations have most often used $K\ell 3$ decays:

$$\Gamma_{K\ell 3} = \frac{G_F^2 M_K^5}{192\pi^3} S_{EW} (1 + \delta_K^\ell + \delta_{SU2}) C^2 |V_{us}|^2 f_+^2(0) I_K^\ell. \quad (9)$$

Here, ℓ refers to either e or μ , G_F is the Fermi constant, M_K is the kaon mass, S_{EW} is the short-distance radiative correction, δ_K^ℓ is the mode-dependent long-distance radiative correction, $f_+(0)$ is the calculated form factor at zero momentum transfer for the $\ell\nu$ system, and I_K^ℓ is the phase-space integral, which depends on measured semileptonic form factors. For charged kaon decays, δ_{SU2} is the deviation from one of the ratio of $f_+(0)$ for the charged to neutral kaon decay; it is zero for the neutral kaon. C^2 is 1 (1/2) for neutral (charged) kaon decays. Most determinations of $|V_{us}|$ have been based only on $K \rightarrow \pi e \nu$ decays; $K \rightarrow \pi \mu \nu$ decays have not been used because of large uncertainties in I_K^μ . The experimental measurements are the semileptonic decay widths (based on the semileptonic branching fractions and lifetime) and form factors (allowing calculation of the phase space integrals). Theory is needed for S_{EW} , δ_K^ℓ , δ_{SU2} , and $f_+(0)$.

Many new measurements during the last few years have resulted in a significant shift in V_{us} . Most importantly, recent measurements of the $K \rightarrow \pi e \nu$ branching fractions are significantly different than earlier PDG averages, probably as

a result of inadequate treatment of radiation in older experiments. This effect was first observed by BNL E865 [14] in the charged kaon system and then by KTeV [15,16] in the neutral kaon system; subsequent measurements were made by KLOE [17–20], NA48 [21–23], and ISTRA+ [24]. Current averages (*e.g.*, by the PDG [25] or Flavianet [26]) of the semileptonic branching fractions are based only on recent, high-statistics experiments where the treatment of radiation is clear. In addition to measurements of branching fractions, new measurements of lifetimes [27] and form factors [28–32], have resulted in improved precision for all of the experimental inputs to V_{us} . Precise measurements of form factors for $K_{\mu 3}$ decay now make it possible to use both semileptonic decay modes to extract V_{us} .

Following the analysis of the Flavianet group [26], one finds the values of $|V_{us}|f_+(0)$ in Table 2. The average of these measurements gives

$$f_+(0)|V_{us}| = 0.21668(45). \quad (10)$$

Figure 1 shows a comparison of these results with the PDG evaluation from 2002 [33], as well as $f_+(0)(1-|V_{ud}|^2-|V_{ub}|^2)^{1/2}$, the expectation for $f_+(0)|V_{us}|$ assuming unitarity, based on $|V_{ud}| = 0.9742 \pm 0.0003$, $|V_{ub}| = (3.6 \pm 0.7) \times 10^{-3}$, and the widely used Leutwyler-Roos calculation of $f_+(0) = 0.961 \pm 0.008$ [34]. Using the result in Eq. (10) with the Leutwyler-Roos calculation of $f_+(0)$ gives

$$|V_{us}| = \lambda = 0.2255 \pm 0.0019. \quad (11)$$

Similar results for $f_+(0)$ were recently obtained from lattice gauge theory calculations [35,36]. For example, and recent 2+1 fermion dynamical wall calculation [36] gave $f_+(0) = 0.9609(51)$. Other calculations of $f_+(0)$ result in $|V_{us}|$ values that differ by as much as 2% from the result in Eq. (11). For example, a recent chiral perturbation theory calculation [37, 38] gives $f_+(0) = 0.974 \pm 0.012$, which implies a lower value of $|V_{us}| = 0.2225 \pm 0.0028$ [39].

Table 2: $|V_{us}|f_+(0)$ from $K\ell 3$.

Decay Mode	$ V_{us} f_+(0)$
$K^\pm e 3$	0.21746 ± 0.00085
$K^\pm \mu 3$	0.21810 ± 0.00114
$K_L e 3$	0.21638 ± 0.00055
$K_L \mu 3$	0.21678 ± 0.00067
$K_S e 3$	0.21554 ± 0.00142
Average	0.21668 ± 0.00045

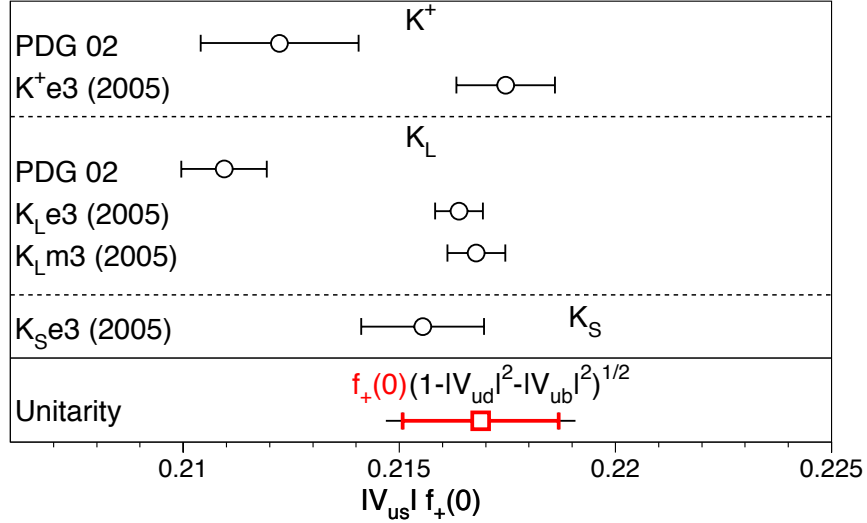


Figure 1: Comparison of determinations of $|V_{us}|f_+(0)$ from this review (labeled 2005), from the PDG 2002, and with the prediction from unitarity using $|V_{ud}|$ and the Leutwyler-Roos calculation of $f_+(0)$ [34]. For $f_+(0)(1-|V_{ud}|^2-|V_{ub}|^2)^{1/2}$, the inner error bars are from the quoted uncertainty in $f_+(0)$; the total uncertainties include the $|V_{ud}|$ and $|V_{ub}|$ errors.

A value of V_{us} can also be obtained from a comparison of the radiative inclusive decay rates for $K \rightarrow \mu\nu(\gamma)$ and $\pi \rightarrow \mu\nu(\gamma)$ combined with a lattice gauge theory calculation of f_K/f_π via [40]

$$\frac{|V_{us}|f_K}{|V_{ud}|f_\pi} = 0.2387(4) \left[\frac{\Gamma(K \rightarrow \mu\nu(\gamma))}{\Gamma(\pi \rightarrow \mu\nu(\gamma))} \right]^{\frac{1}{2}} \quad (12)$$

with the small error coming from electroweak radiative corrections. Employing

$$\frac{\Gamma(K \rightarrow \mu\nu(\gamma))}{\Gamma(\pi \rightarrow \mu\nu(\gamma))} = 1.3337(46), \quad (13)$$

which averages in the KLOE result [41], $B(K \rightarrow \mu\nu(\gamma)) = 63.66(9)(15)\%$ and [42, 43]

$$f_K/f_\pi = 1.208(2)(+7/-14) \quad (14)$$

along with the value of V_{ud} in Eq. (4) leads to

$$|V_{us}| = 0.2223(5)(1.208f_\pi/f_K). \quad (15)$$

It should be mentioned that hyperon decay fits suggest [5]

$$|V_{us}| = 0.2250(27) \text{ Hyperon Decays} \quad (16)$$

modulo SU(3) breaking effects that could shift that value up or down. We note that a recent representative effort [44] that incorporates SU(3) breaking found $V_{us} = 0.226(5)$. Similarly, strangeness changing tau decays give [45]

$$|V_{us}| = 0.2208(34) \text{ Tau Decays} \quad (17)$$

where the central value depends on the strange quark mass.

Employing the value of V_{ud} in Eq. (4) and V_{us} in Eq. (11) leads to the unitarity consistency check

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 0.9999(5)(9), \quad (18)$$

where the first error is the uncertainty from $|V_{ud}|^2$ and the second error is the uncertainty from $|V_{us}|^2$. The result is in good agreement with unitarity. Averaging the direct determination of λ (V_{us}) with the determination derived from unitarity and V_{ud} gives $\lambda = 0.226(1)$. Although unitarity now seems well established, issues regarding the Q values in superallowed nuclear β -decays, τ_n , g_A , $f_+(0)$ and f_K/f_π must still be resolved before a definitive confirmation is possible.

CKM Unitarity Constraints

The current good experimental agreement with unitarity, $|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 0.9999(10)$ provides strong confirmation of Standard Model radiative corrections (which range between 3-4% depending on the nucleus used) at better than the 30 sigma level [46]. In addition, it implies constraints on “New Physics” effects at both the tree and quantum loop levels. Those effects could be in the form of contributions to nuclear beta decays, $K_{\ell 3}$ decays and/or muon decays, with the last of these providing normalization via the muon lifetime [47], which is used to obtain the Fermi constant, $G_\mu = 1.166371(6) \times 10^{-5} \text{GeV}^{-2}$.

We illustrate the implications of CKM unitarity for: 1) exotic muon decays [48] (beyond ordinary muon decay $\mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu$); and 2) new heavy quark mixing V_{uD} [49]. Other examples in the literature [50,51] include Z_χ boson quantum loop effects, supersymmetry, leptoquarks, compositeness etc.

Exotic Muon Decays

If additional lepton flavor violating decays such as $\mu^+ \rightarrow e^+ \bar{\nu}_e \nu_\mu$ (wrong neutrinos) occur, they would cause confusion in searches for neutrino oscillations at, for example, muon storage rings/neutrino factories or other neutrino sources from muon decays. Calling the rate for all such decays $\Gamma(\text{exotic } \mu \text{ decays})$, they should be subtracted before the extraction of G_μ and normalization of the CKM matrix. Since that is not done and unitarity works, one has (at one-sided 95% CL)

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1 - BR(\text{exotic } \mu \text{ decays}) \geq 0.9982 \quad (19)$$

or

$$BR(\text{exotic } \mu \text{ decays}) < 0.0018 . \quad (20)$$

That bound is a factor of 6–7 better than the direct experimental bound on $\mu^+ \rightarrow e^+ \bar{\nu}_e \nu_\mu$.

New Heavy Quark Mixing

Heavy D quarks naturally occur in fourth quark generation models and some heavy quark “new physics” scenarios such as E_6 grand unification. Their mixing with ordinary quarks gives rise to V_{ud} which is constrained by unitarity (one sided 95% CL)

$$\begin{aligned} |V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 &= 1 - |V_{uD}|^2 > 0.9982 \\ |V_{uD}| &< 0.04 . \end{aligned} \quad (21)$$

A similar constraint applies to heavy neutrino mixing and the couplings $V_{\mu N}$ and V_{eN} .

References

1. N. Cabibbo, Phys. Rev. Lett. **10**, 531 (1963).
2. M. Kobayashi and T. Maskawa, Prog. Theor. Phys. **49**, 652 (1973).
3. L. Wolfenstein, Phys. Rev. Lett. **51**, 1945 (1983).
4. S. Eidelman *et al.*, [Particle Data Group], Phys. Lett. **B592**, 1 (2004).
5. N. Cabibbo, E.C. Swallow, and R. Winston, Phys. Rev. Lett. **92**, 251803 (2004) [hep-ph/0307214].
6. J.C. Hardy and I.S. Towner, Phys. Rev. Lett. **94**, 092502 (2005) [nucl-th/0412050].
7. W.J. Marciano and A. Sirlin, Phys. Rev. Lett. **71**, 3629 (1993).
8. A. Czarnecki, W.J. Marciano, and A. Sirlin, Phys. Rev. **D70**, 093006 (2004) [hep-ph/0406324].
9. W.J. Marciano and A. Sirlin, Phys. Rev. Lett. **96**, 032002 (2006) [hep-ph/0510099].
10. I.S. Towner and J.C. Hardy, arXiv:0710.3181 [nucl-th].
11. I.S. Towner and J.C. Hardy, [nucl-th/0209014].
12. A. Serebrov *et al.*, Phys. Lett. **B605**, 72 (2005) [nucl-ex/0408009].
13. D. Pocanic *et al.*, Phys. Rev. Lett. **93**, 181803 (2004) [hep-ex/0312030].
14. A. Sher *et al.*, Phys. Rev. Lett. **91**, 261802 (2003).
15. T. Alexopoulos *et al.*, [KTeV Collab.], Phys. Rev. Lett. **93**, 181802 (2004) [hep-ex/0406001].
16. T. Alexopoulos *et al.*, [KTeV Collab.], Phys. Rev. **D70**, 092006 (2004) [hep-ex/0406002].

17. F. Ambrosino *et al.*, [KLOE Collab.], Phys. Lett. **B632**, 43 (2006) [[hep-ex/0508027](#)].
18. F. Ambrosino *et al.*, [KLOE Collab.], Phys. Lett. **B638**, 140 (2006) [[hep-ex/0603041](#)].
19. F. Ambrosino *et al.*, [KLOE Collab.], Phys. Lett. **B636**, 173 (2006) [[hep-ex/0601026](#)].
20. F. Ambrosino *et al.*, [KLOE Collab.], PoS **HEP2005**, 287 (2006) [Frascati Phys. Ser. **41**, 69 (2006)] [[hep-ex/0510028](#)].
21. A. Lai *et al.*, [NA48 Collab.], Phys. Lett. **B602**, 41 (2004) [[hep-ex/0410059](#)].
22. A. Lai *et al.*, [NA48 Collab.], Phys. Lett. **B645**, 26 (2007) [[hep-ex/0611052](#)].
23. J.R. Batley *et al.*, [NA48/2 Collab.], Eur. Phys. J. C **50**, 329 (2007) [[hep-ex/0702015](#)].
24. V.I. Romanovsky *et al.*, [[hep-ex/0704.2052](#)].
25. W.M. Yao *et al.*, [Particle Data Group], J. Phys. G **33**, 1 (2006).
26. Flavianet Working Group on Precise SM Tests in K Decays, <http://www.lnf.infn.it/wg/vus>.
27. F. Ambrosino *et al.*, [KLOE Collab.], Phys. Lett. **B626**, 15 (2005) [[hep-ex/0507088](#)].
28. T. Alexopoulos *et al.*, [KTeV Collab.], Phys. Rev. **D70**, 092007 (2004) [[hep-ex/0406003](#)].
29. E. Abouzaid *et al.*, [KTeV Collab.], Phys. Rev. **D74**, 097101 (2006) [[hep-ex/0608058](#)].
30. F. Ambrosino *et al.*, [KLOE Collab.], Phys. Lett. **B636**, 166 (2006) [[hep-ex/0601038](#)].
31. A. Lai *et al.*, [NA48 Collab.], Phys. Lett. **B604**, 1 (2004) [[hep-ex/0410065](#)].
32. O.P. Yushchenko *et al.*, Phys. Lett. **B589**, 111 (2004) [[hep-ex/0404030](#)].
33. K. Hagiwara *et al.*, [Particle Data Group], Phys. Rev. **D66**, 1 (2002).
34. H. Leutwyler and M. Roos, Z. Phys. **C25**, 91 (1984).
35. D. Becirevic *et al.*, Nucl. Phys. **B705**, 339 (2005) [[hep-ph/0403217](#)].
36. D.J. Antonio *et al.*, [[hep-lat/0702026](#)].
37. J. Bijnens and P. Talavera, Nucl. Phys. **B669**, 341 (2003).
38. V. Cirigliano *et al.*, JHEP **0504**, 006 (2005) [[hep-ph/0503108](#)].
39. M. Jamin, J.A. Oller, and A. Pich, JHEP **02**, 047 (2004).

40. W.J. Marciano, Phys. Rev. Lett. **93**, 231803 (2004) [[hep-ph/0402299](#)].
41. F. Ambrosino *et al.*, [KLOE Collab.], Phys. Lett. **B632**, 76 (2006) [[hep-ex/0509045](#)].
42. C. Aubin *et al.*, [MILC Collab.], Phys. Rev. **D70**, 114501 (2004) [[hep-lat/0407028](#)].
43. C. Bernard *et al.*, [MILC Collab.], PoS LAT 2005, 025 (2005) [[hep-lat/0509137](#)]; C. Bernard *et al.*, [hep-lat/0611024](#).
44. V. Mateu and A. Pich, JHEP **0510**, 041 (2005) [[hep-ph/0509045](#)].
45. E. Gamiz *et al.*, Phys. Rev. Lett. **94**, 011803 (2005) [[hep-ph/0408044](#)]; E. Gamiz *et al.*, [arXiv:0709.0282](#) [[hep-ph](#)].
46. A. Sirlin, Rev. Mod. Phys. **50**, 573 (1978).
47. D.B. Chitwood *et al.*, Phys. Rev. Lett. **99**, 032001 (2007).
48. K.S. Babu and S. Pakvasa, [hep-ph/0204236](#).
49. W. Marciano and A. Sirlin, Phys. Rev. Lett. **56**, 22 (1986); P. Langacker and D. London, Phys. Rev. **D38**, 886 (1988).
50. W. Marciano and A. Sirlin, Phys. Rev. **D35**, 1672 (1987).
51. R. Barbieri *et al.*, Phys. Lett. **156B**, 348 (1985); K. Hagiwara *et al.*, Phys. Rev. Lett. **75**, 3605 (1995); A. Kurylov and M. Ramsey-Musolf, Phys. Rev. Lett. **88**, 071804 (2000).