

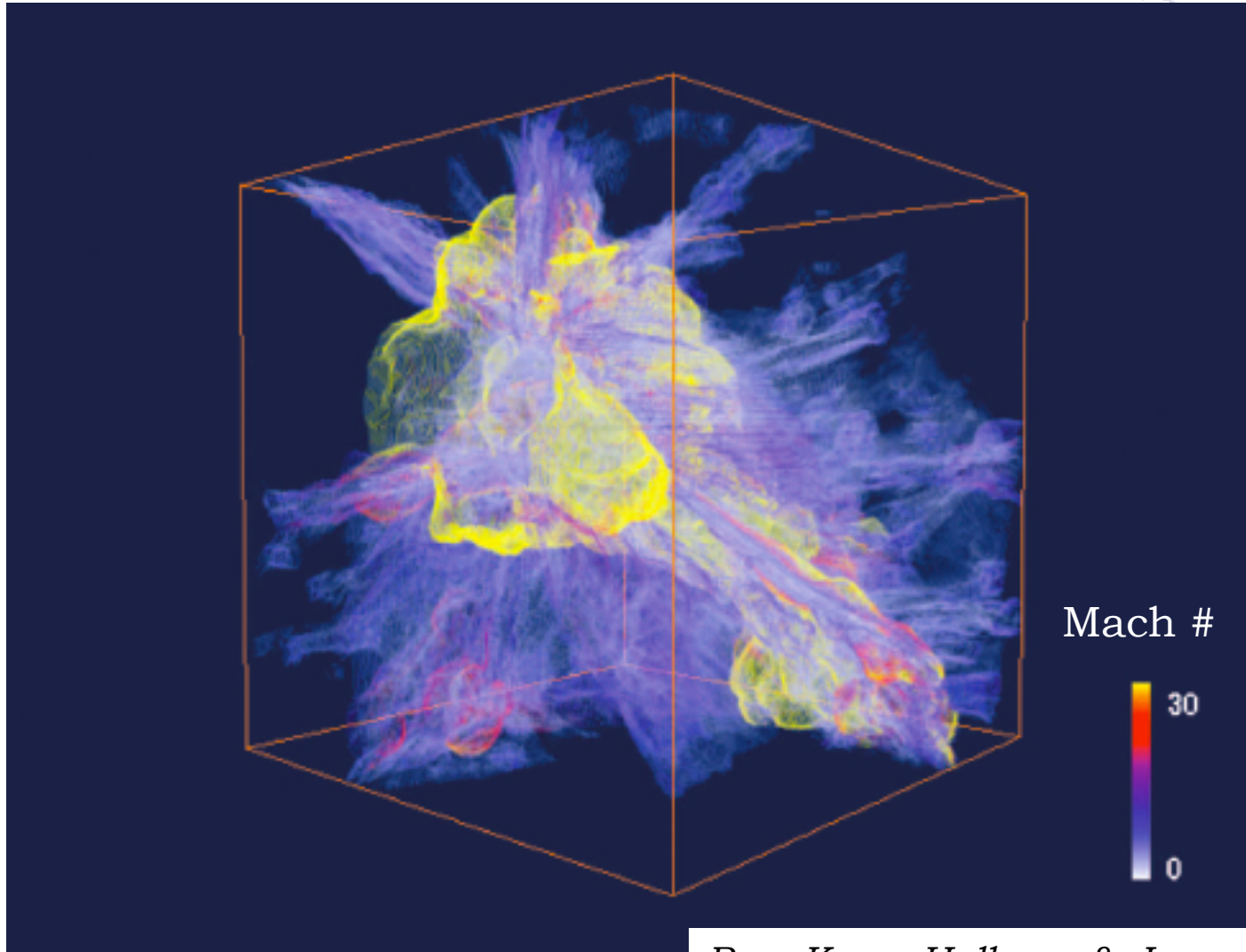
# Analytical Modeling of Cosmic Accretion Shocks: The Role Of Environment

Vasiliki Pavlidou  
Brian D. Fields

University of Illinois at Urbana-Champaign

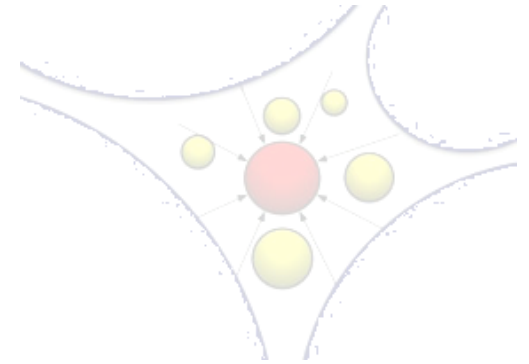
*TeV Particle Astrophysics Workshop*

*15 Jul 2005*



*Ryu, Kang, Hallman & Jones 2003*

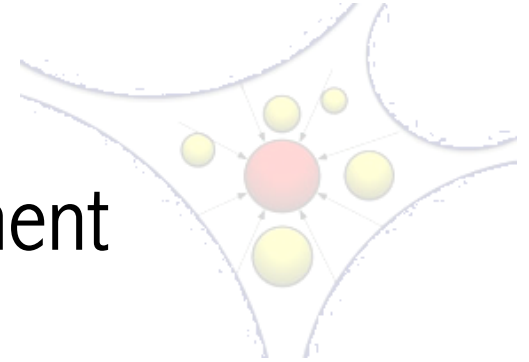
# Cosmic Shock Taxonomy



Classification according to driving mechanism

shock type	<b>Accretion</b>	<b>Merger</b>	<b>Filament</b>
driving mechanism	gravity of accretor	mutual gravitational attraction	expansion of void
artist's impression			

# Accretion Shocks and Environment



Properties of single shock:

- **Mach number**

$$= f(\text{accretion velocity}, \text{external sound speed})$$

- **Mass current**

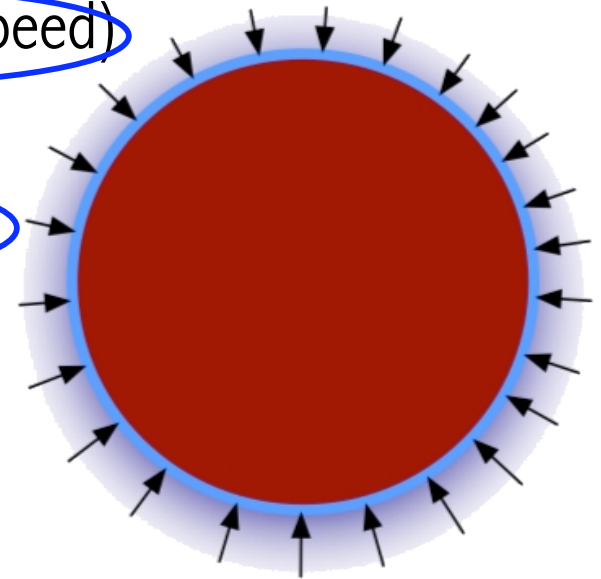
$$= f(\text{accretion velocity}, \text{external density})$$

- **Kinetic power**

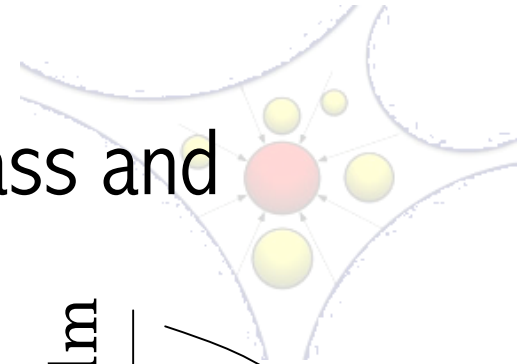
$$= f(\text{accretion velocity}, \text{external density})$$

**accretor mass**

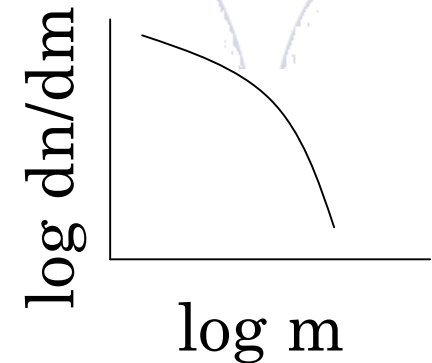
**accretor environment**



# Analytic Description of Accretor Mass and Environment

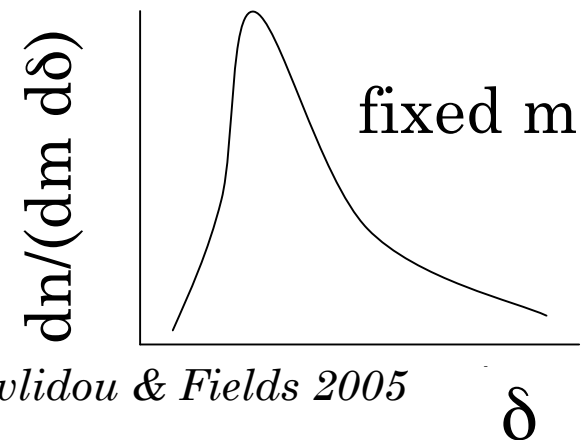


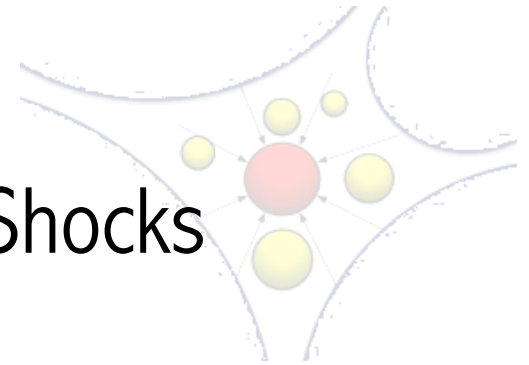
- ✓ Analytic description of **mass** distribution:  
**Press-Schechter** mass function



- ✓ Analytic Description of **mass+environment** distribution:  
**Double Distribution** of cosmic structures

- Parametrizes “environment” using local overdensity,  
 $\delta = \rho_{\text{local}} / \rho_{\text{cosmic}} - 1$
- Integrates back to Press-Schechter



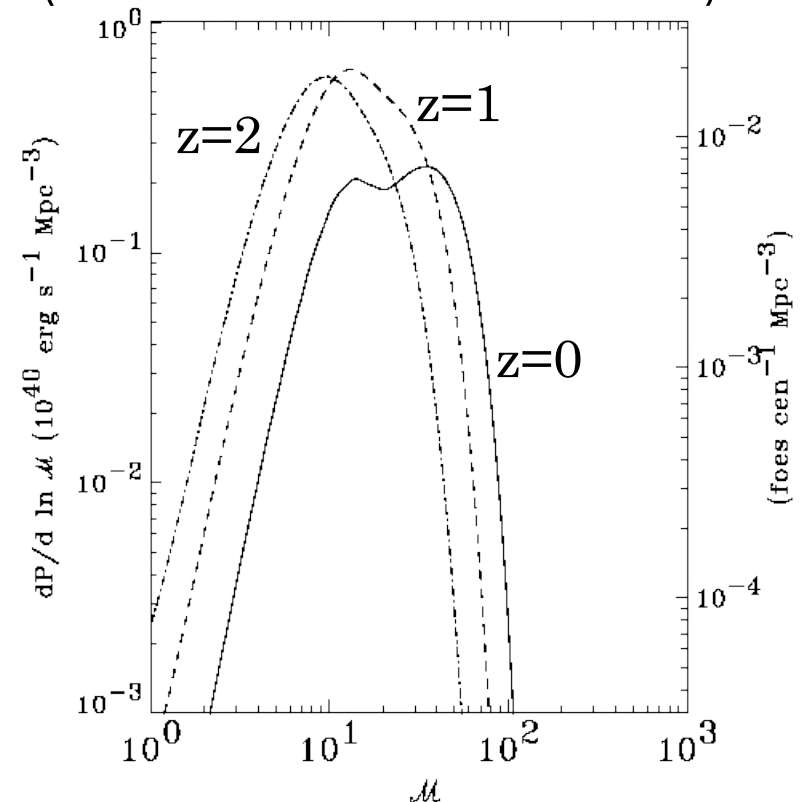
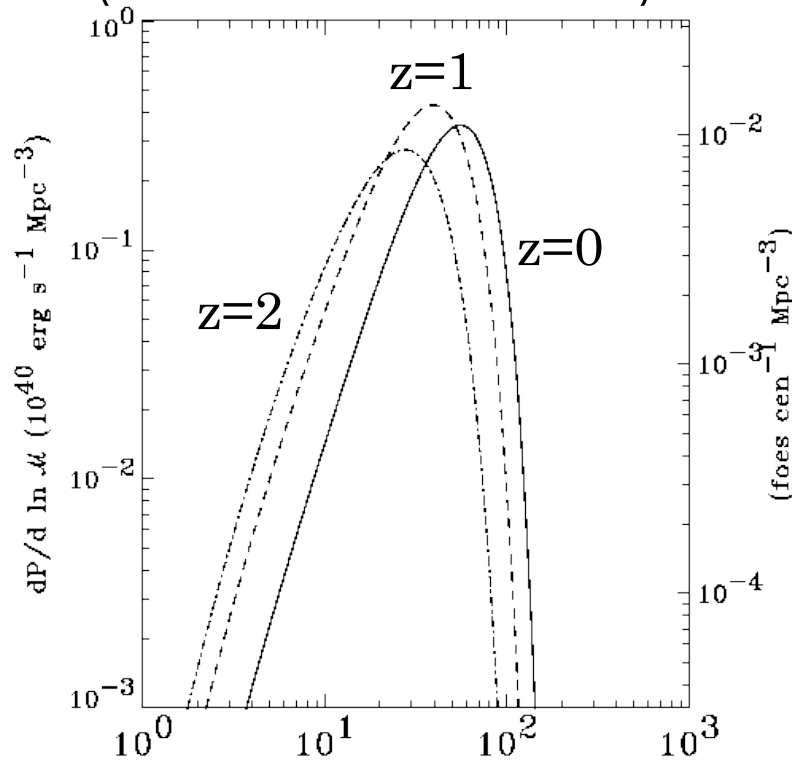


# Two Models for Cosmic Accretion Shocks

**Press-Schechter–based**  
(no environmental effects)

**Double-Distribution–based**  
(includes environmental effects)

Kinetic Power

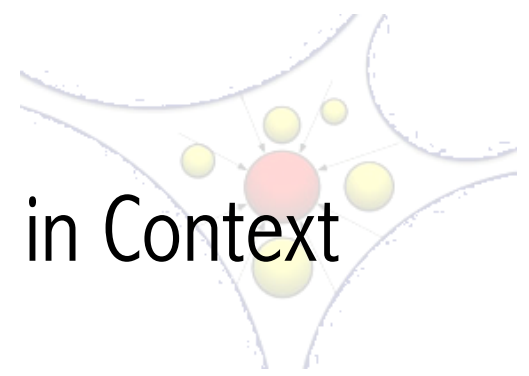


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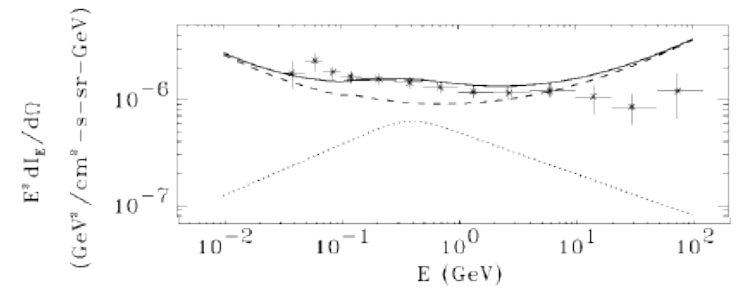
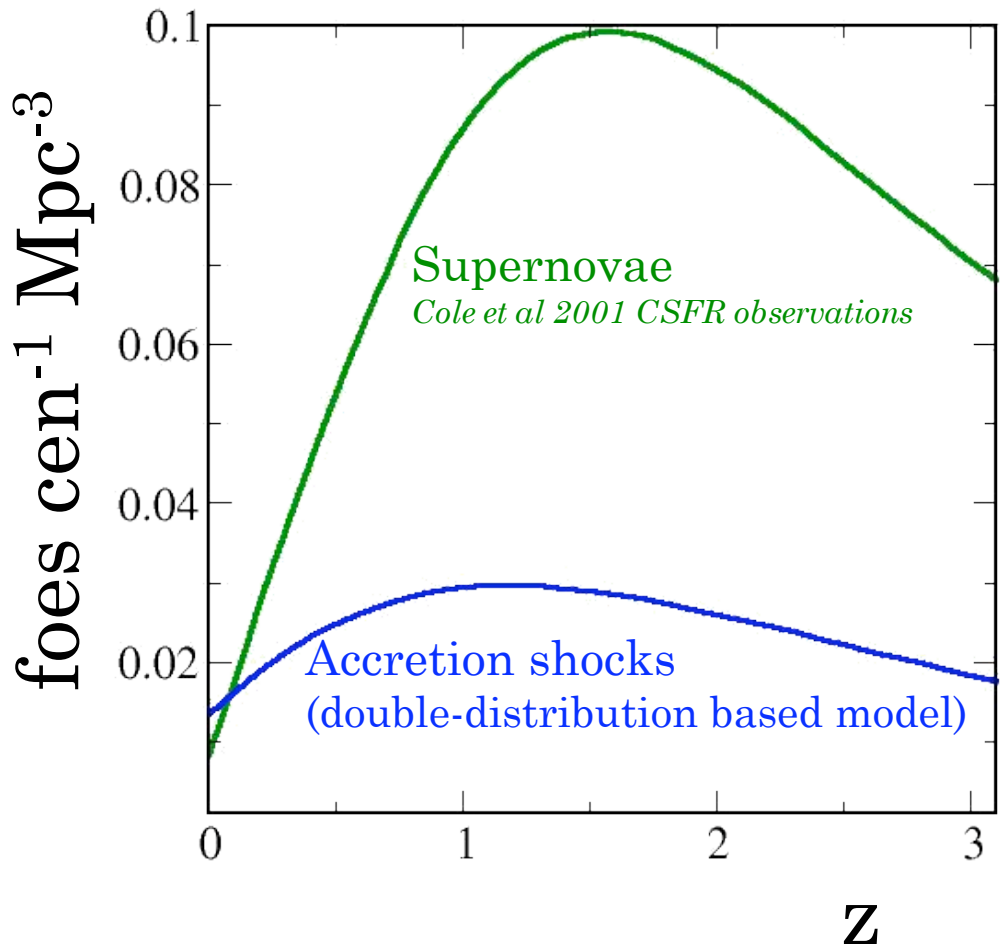
**Mach**

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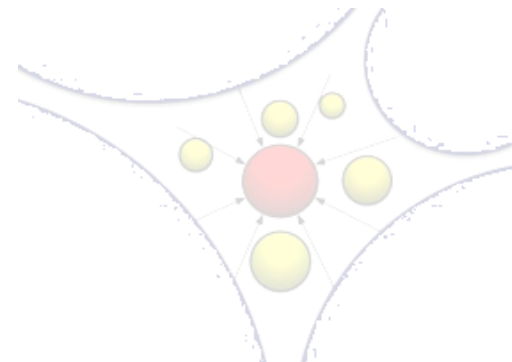


# Energetics of Cosmic Accretion Shocks in Context



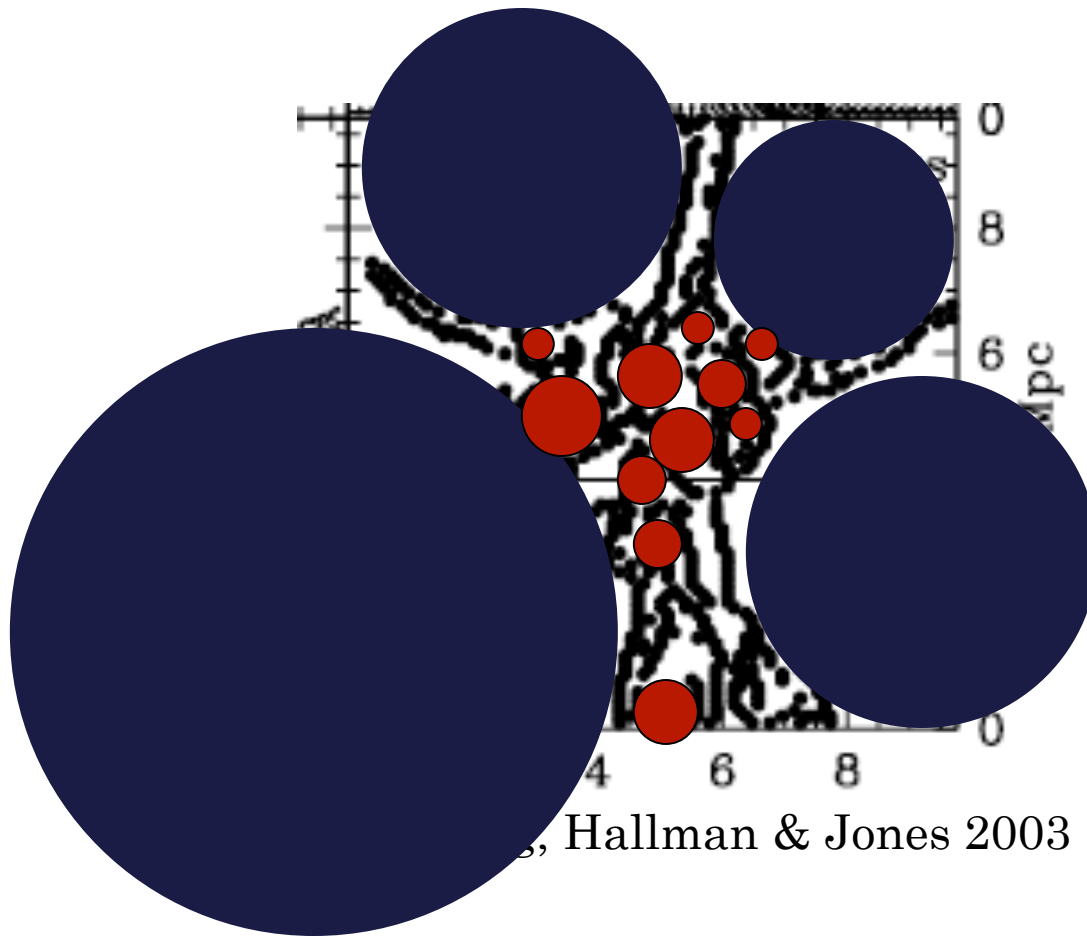
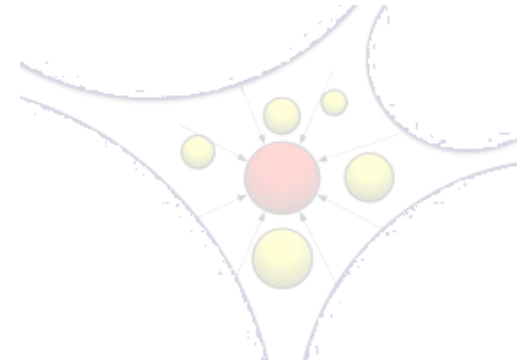
**The Extragalactic Gamma-Ray Background**  
Data points: Sreekumar et al 1998,  
Models: Stecker & Salamon 1996 (blazars)  
Pavlidou & Fields 2002 (galaxies)

# Conclusions



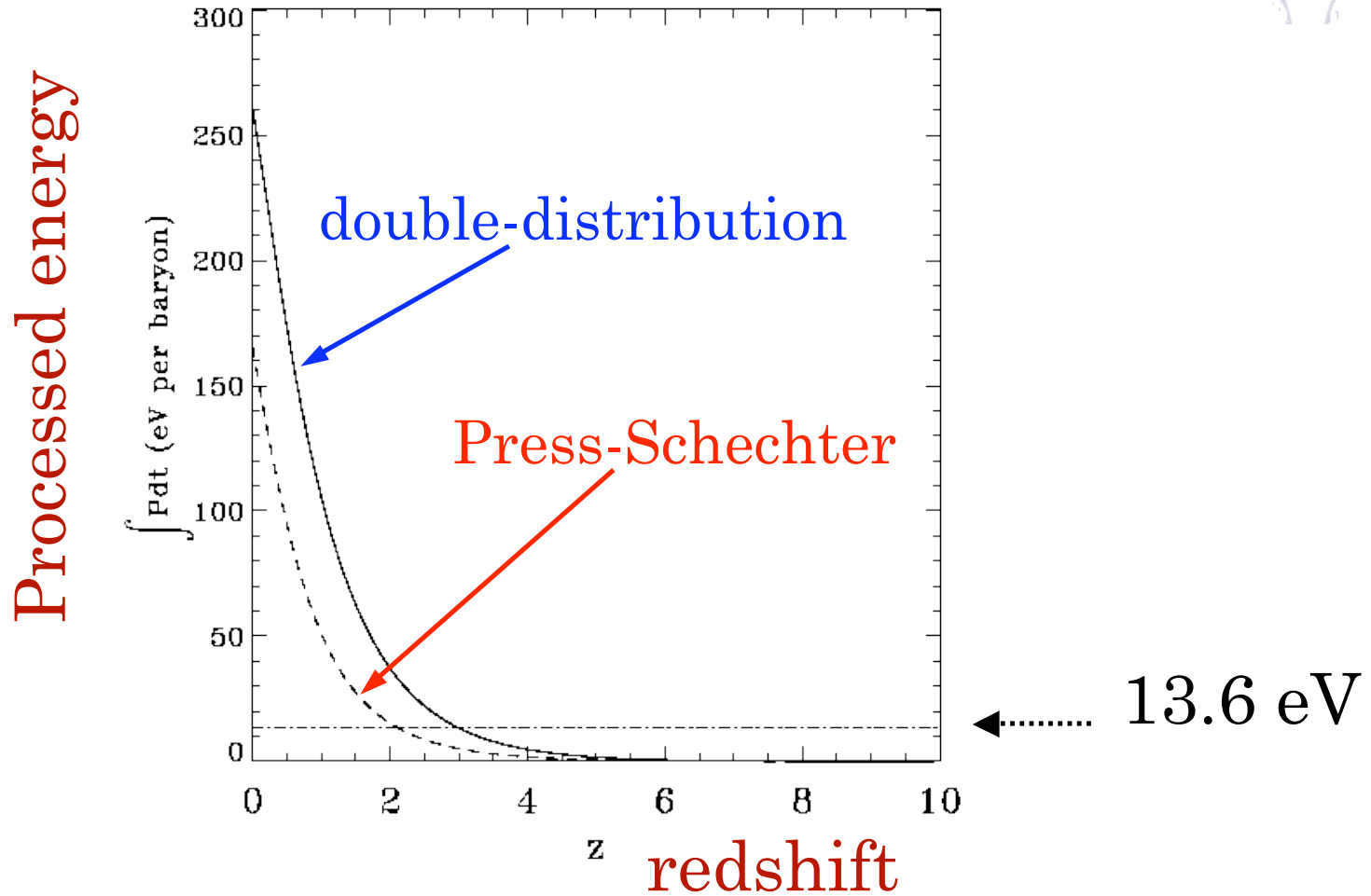
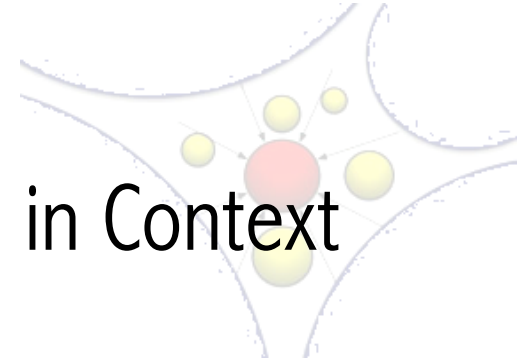
- ✓ Cosmic shocks: accretion, merger, filament (depending on driving mechanism)
- ✓ Double distribution of cosmic structures can be used to investigate effect of environment on cosmic accretion shocks
- ✓ Inclusion of environmental effects distributes energy processed by shocks among broader range of Mach numbers
- ✓ Energy processed by accretion shocks  $\sim 1/3$  of supernovae at high  $z$ , overtakes supernovae output in local universe



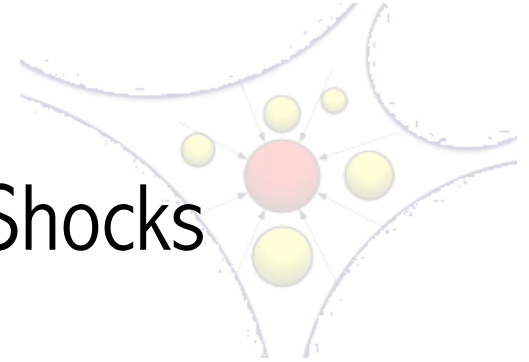


..., Hallman & Jones 2003

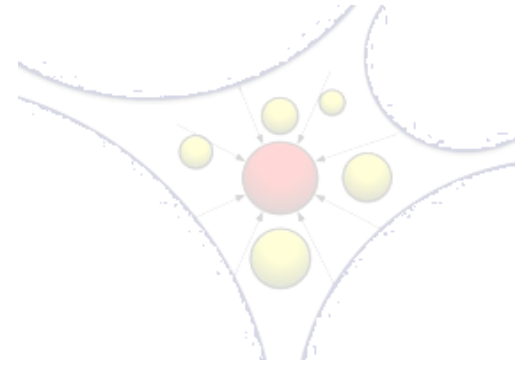
# Energetics of Cosmic Accretion Shocks in Context



# Two Models for Cosmic Accretion Shocks



1. **Press-Schechter based (no environmental effects)**
  - identical environment for all collapsed structures at given  $z$ :
  - All shock properties =  $f(\text{accretor mass})$
  - Distribution of accretor masses  $\equiv$  Press-Schechter mass function
2. **Double-Distribution based (includes environmental effects)**
  - distribution of possible environments for each accretor mass
  - All shock properties =  $f(\text{accretor mass, local overdensity})$
  - Distribution of accretor masses + local overdensities  $\equiv$  Double Distribution of Cosmic Structures



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