

Caustics in galactic halos

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Outline :

1. Introduction.
2. What are caustics ? Caustics with dark matter.
3. Choosing initial conditions.
4. What do they look like ? The structure of caustics.
5. Conclusion.

- Caustics are regions of high density.
- The existence of caustics requires :
 - (a) *Collisionless matter.*
 - (b) *Low velocity dispersion.*
- Caustics have implications for :
 - (a) *Axion searches* : Peaks in the energy spectrum of microwave photons from $a \rightarrow \gamma$ conversion (*Duffy et al, 2005*).
 - (b) *WIMP direct detection* : Plateaux in the recoil energy spectrum (*Vergados, 2001; Green, 2001; Gelmini and Gondolo, 2001; Ling et al, 2004; Bernabei et al, 2005*)

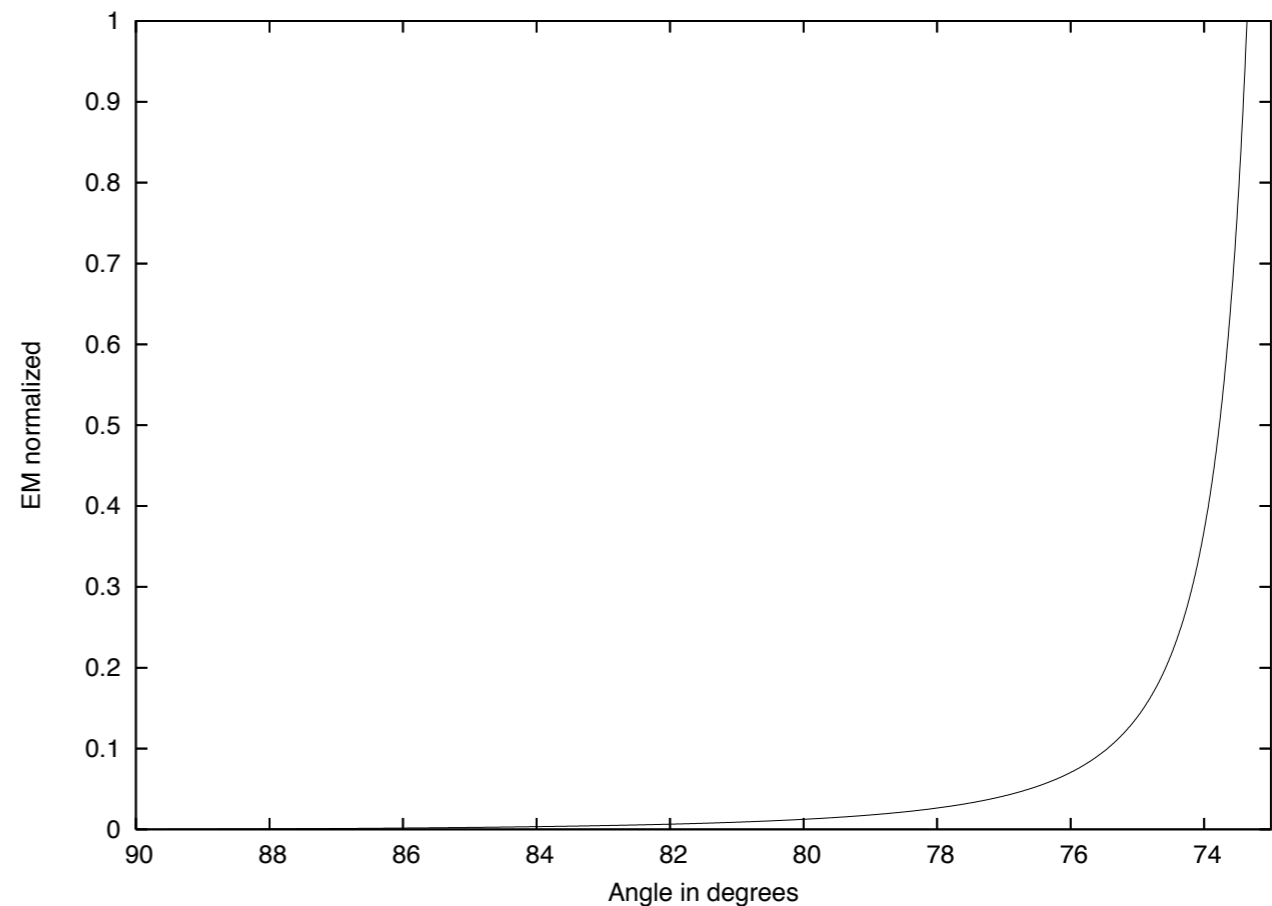
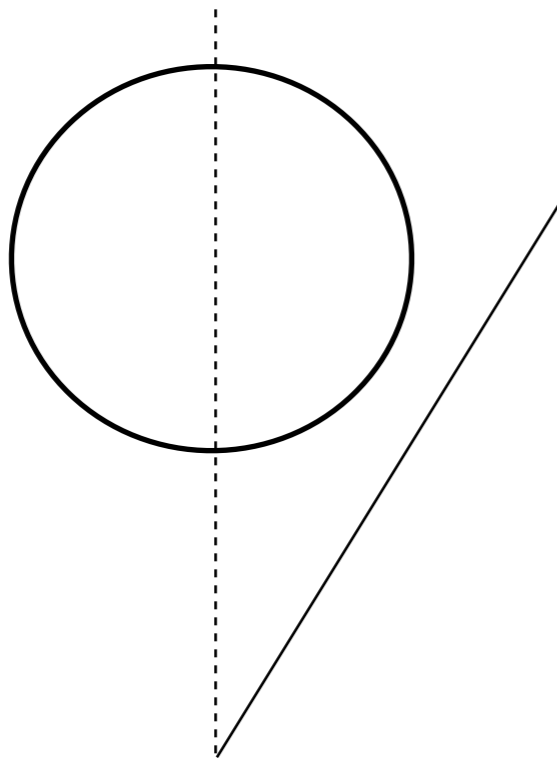
Some previous work on detection of caustics...

I. By particle annihilation (*Hogan, 2001*) :

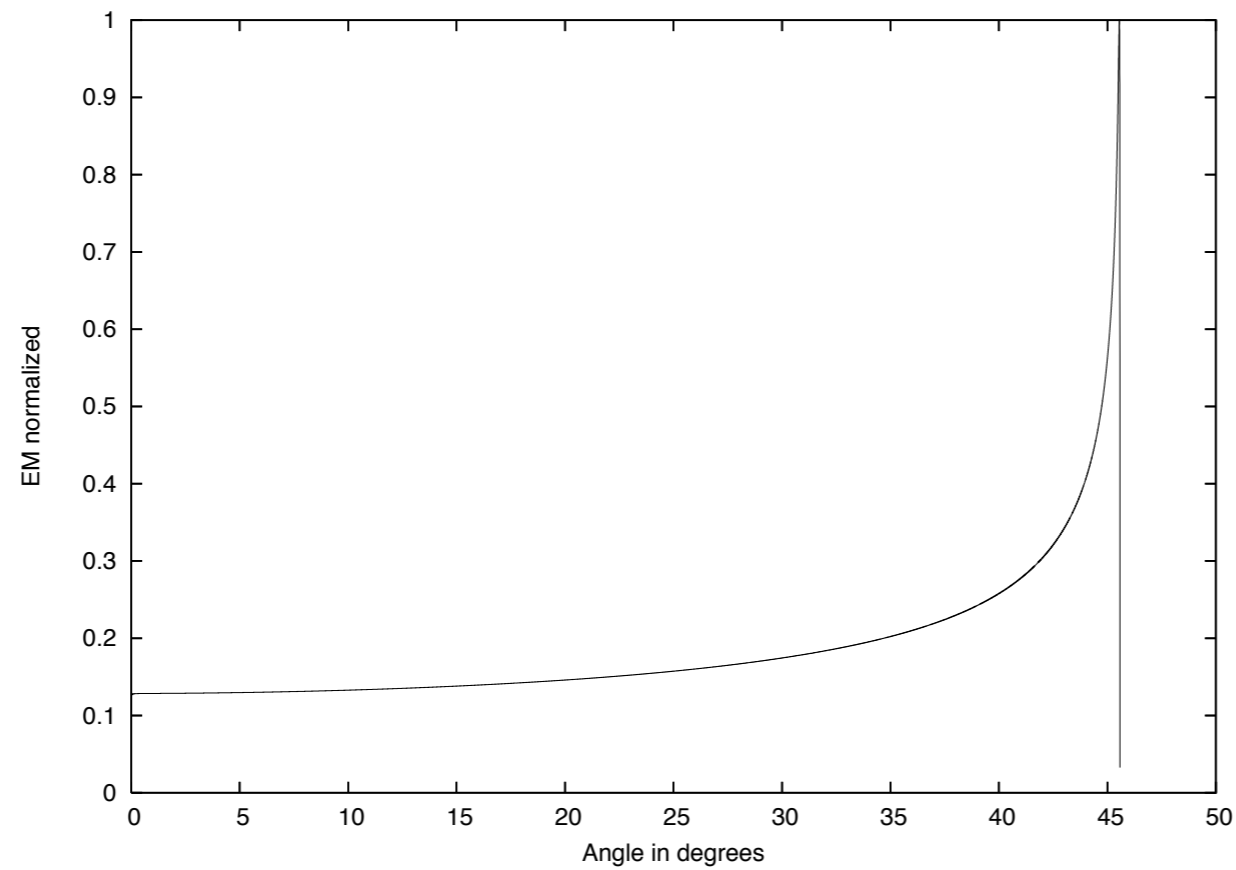
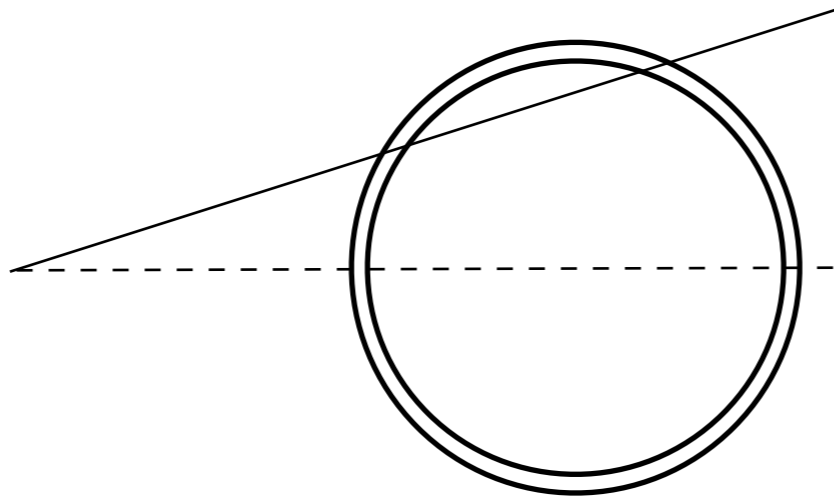
(a) In caustic rings.

(*Bergstrom et al, 2001; Pieri and Branchini, 2005*)

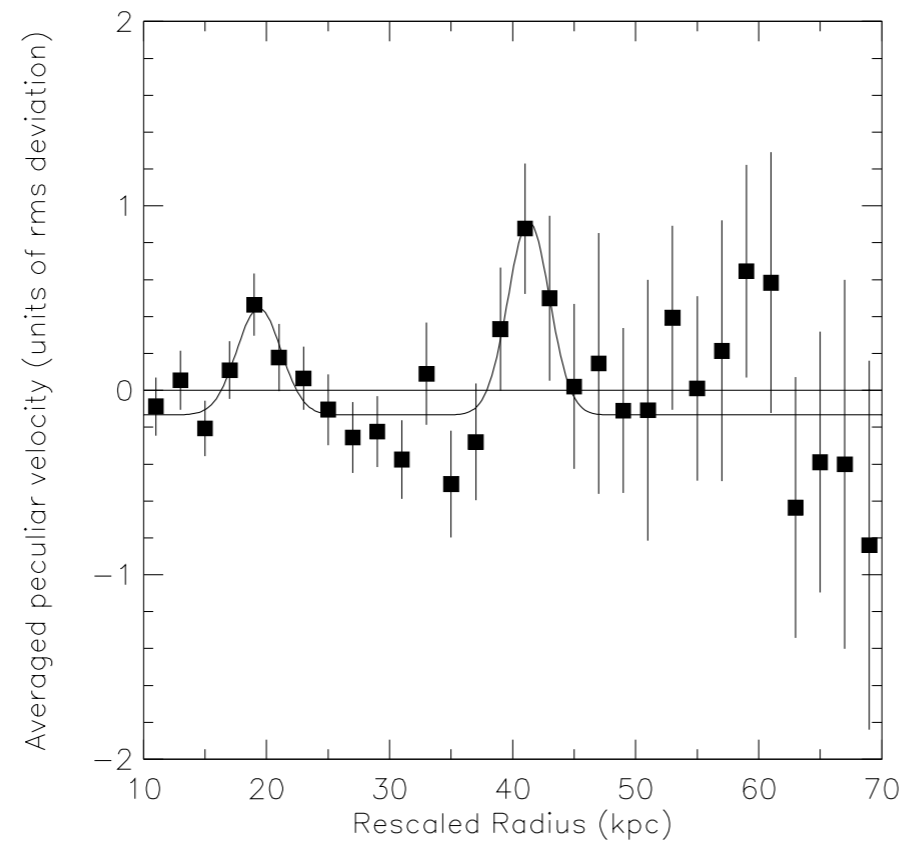
$$EM = \int_{l_{os}} \rho^2(\lambda) d\lambda$$



(b) In caustic spheres (*Mohayaei and Shandarin, 2005*)



3. Using rotation curve data (*Kinney and Sikivie, 2003*).



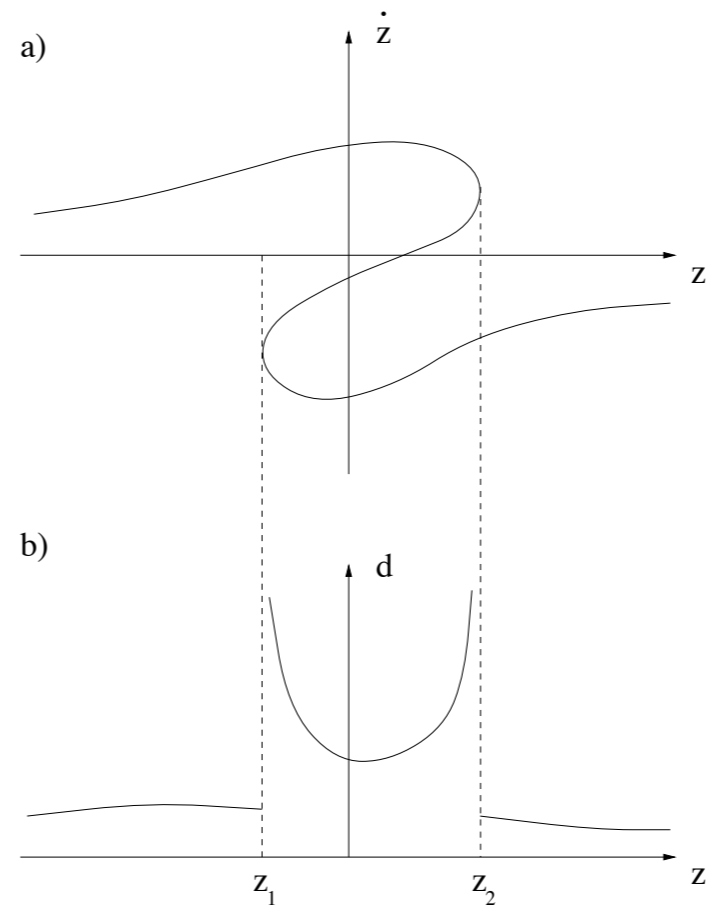
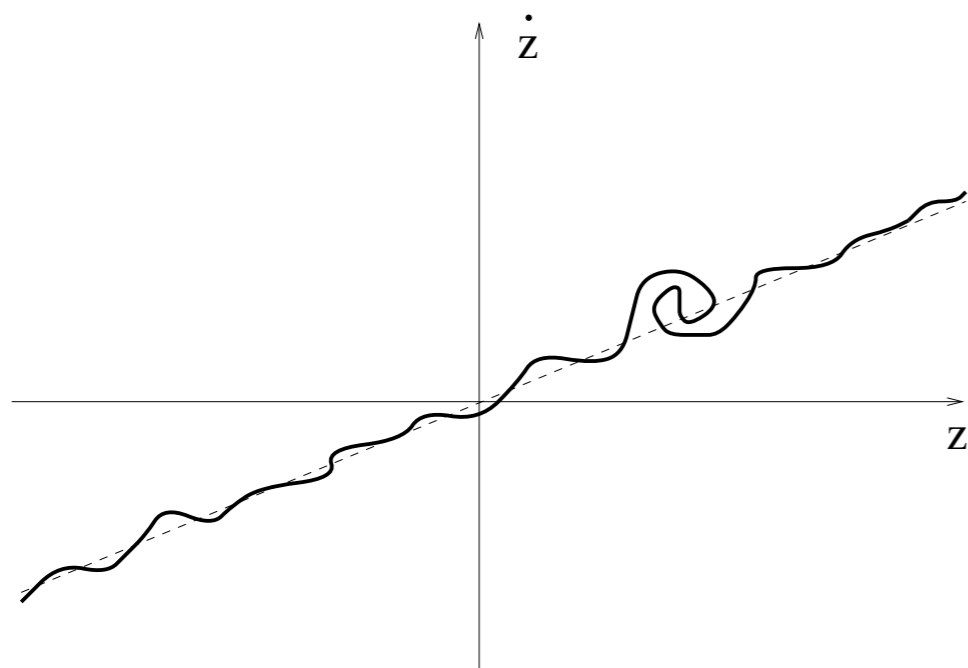
4. By gravitational lensing

(*Hogan, '99; Charmousis et al, 2003; Gavazzi et al, 2005*)

Dark matter caustics are of two kinds :

I. Outer Caustics :

- (a) First few typically occur at scales of 100's of kpc.
- (b) They are topological spheres surrounding galaxies.

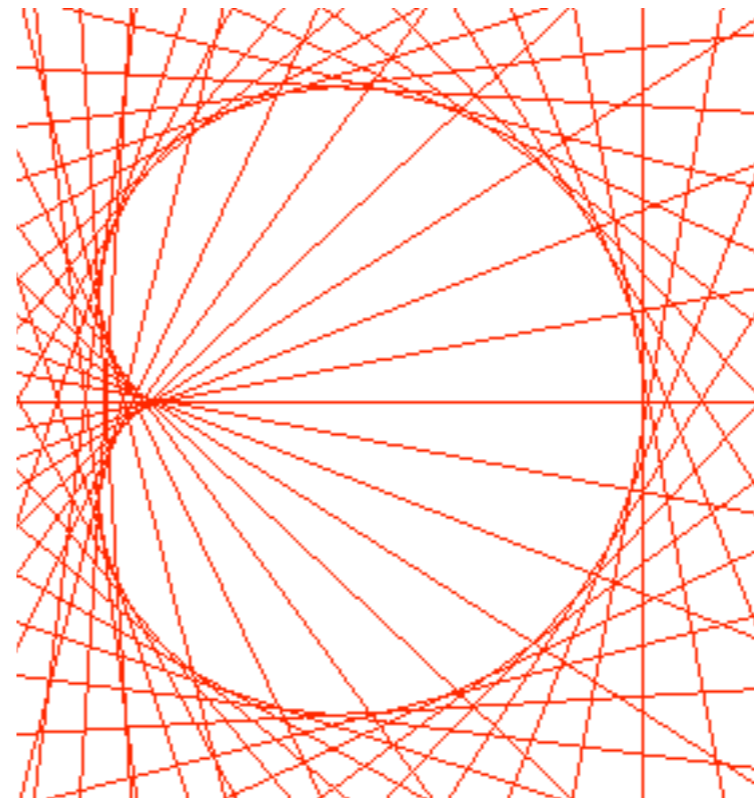
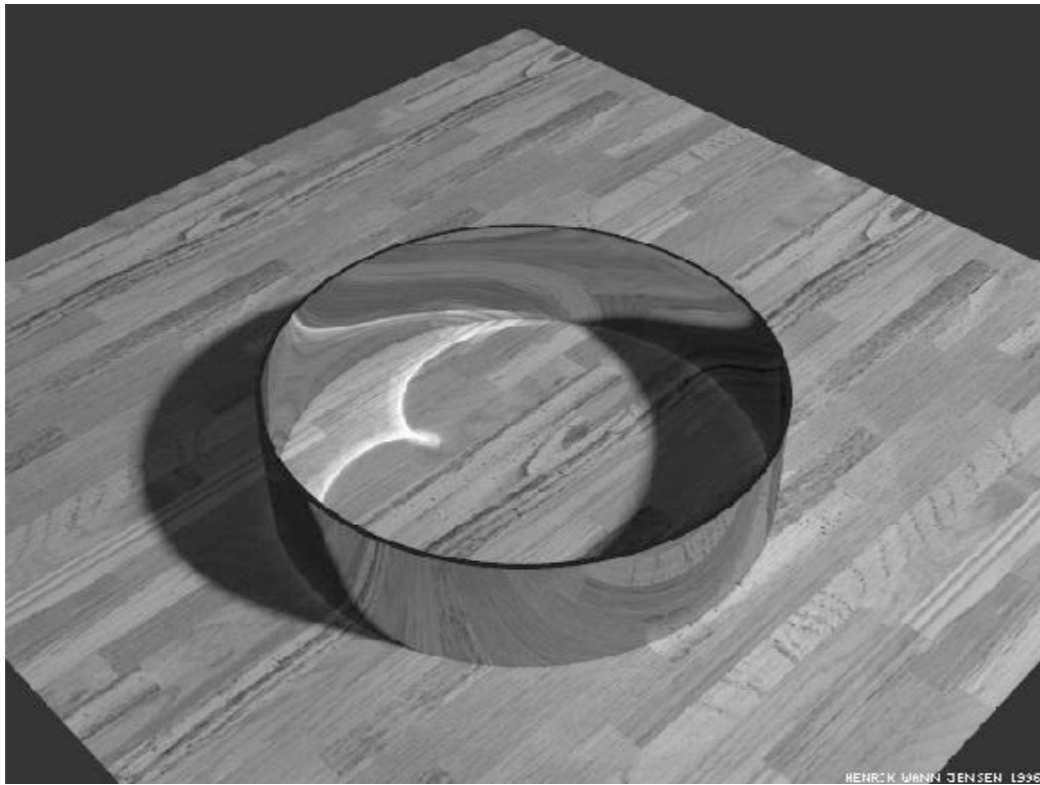


(Sikivie '99)

2. Inner Caustics :

- (a) First few typically occur at 10's of kpc.
- (b) The structure of inner caustics depends on the angular momentum distribution of the dark matter.

Let's first look at a simple example with light ...



(courtesy Henrik Jensen)

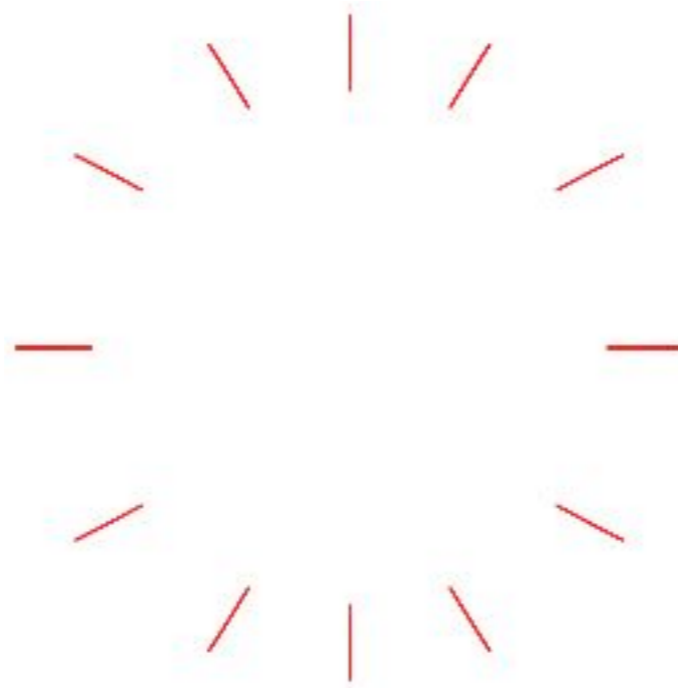
$$S = \tan 2\theta - \frac{\sin \theta - y}{\cos \theta - x} = 0$$
$$\frac{\partial S}{\partial \theta} = 0$$

Catastrophes with light -

For a poster showing the geometry of the catastrophes, see

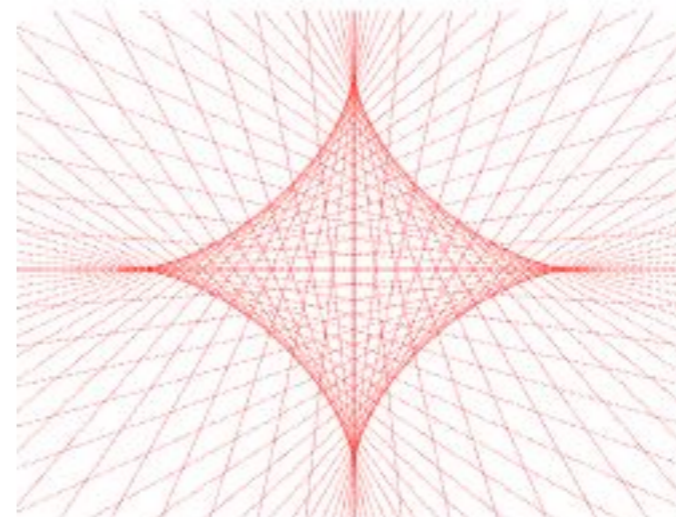
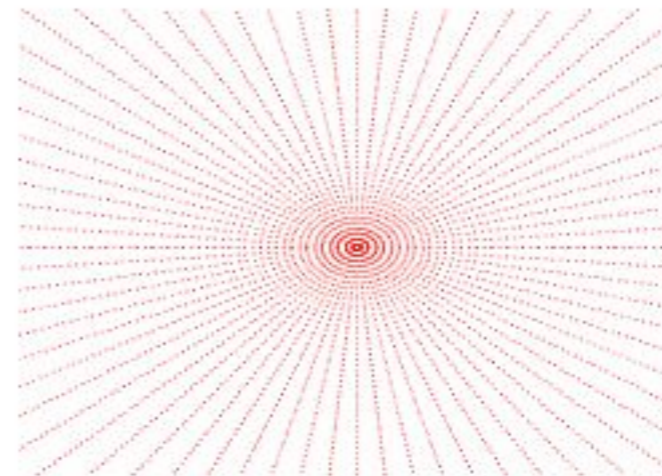
www.phy.bris.ac.uk/research/theory/Berry/pictures/poster1.pdf

Now let's do the same thing with dark matter -



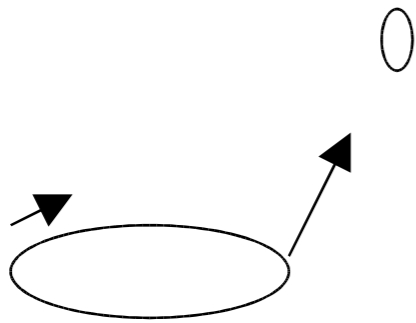
$$\begin{aligned}x &= x(t, \theta, \phi) \\y &= y(t, \theta, \phi) \\z &= z(t, \theta, \phi)\end{aligned}$$

$$\mathbf{J} = \begin{vmatrix} \frac{\partial x}{\partial \theta} & \frac{\partial x}{\partial \phi} & \frac{\partial x}{\partial t} \\ \frac{\partial y}{\partial \theta} & \frac{\partial y}{\partial \phi} & \frac{\partial y}{\partial t} \\ \frac{\partial z}{\partial \theta} & \frac{\partial z}{\partial \phi} & \frac{\partial z}{\partial t} \end{vmatrix} = 0$$



- The caustic is the set of points (x,y,z) where $J=0$.
- The structure of the caustic depends on the initial conditions, i.e. on the spatial distribution of angular momentum.

The initial conditions : Tidal Torque theory -



$$\mathbf{r}(t, \mathbf{q}) = a(t)\mathbf{q} - D(t)\nabla\varphi(\mathbf{q}) \quad (\text{Zeldovich})$$

$$\varphi(\mathbf{q}) = \varphi(\bar{\mathbf{q}}) + (q_i - \bar{q}_i) \frac{\partial\varphi(\bar{\mathbf{q}})}{\partial q_i} + \frac{1}{2}(q_i - \bar{q}_i) \frac{\partial^2\varphi(\bar{\mathbf{q}})}{\partial q_i \partial q_j} (q_j - \bar{q}_j) + \dots$$

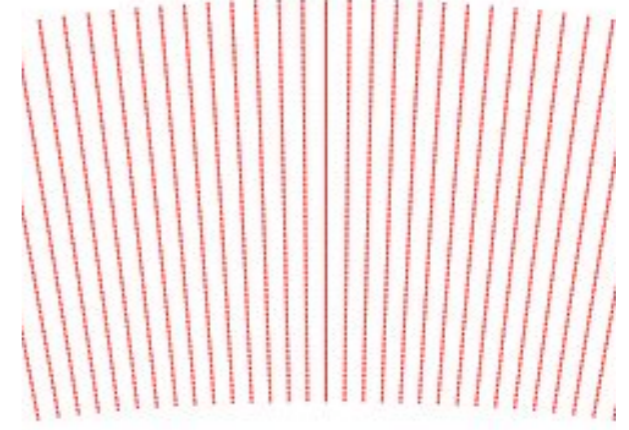
(White '84)

$$\vec{v}(r) = M \vec{r}$$

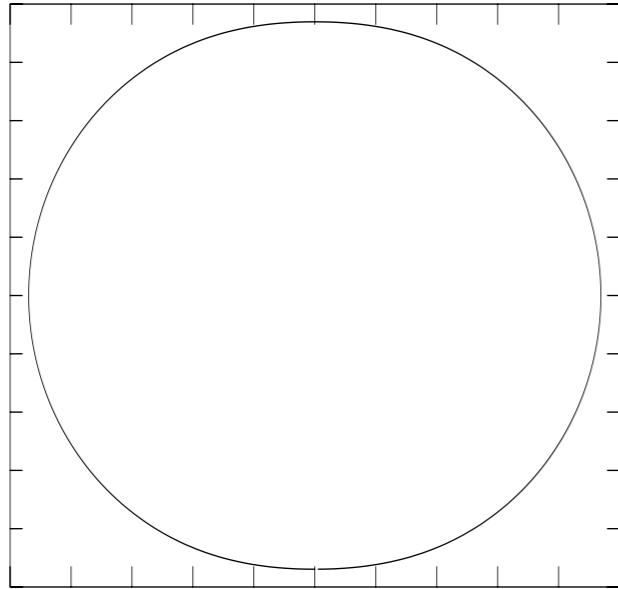
In general, $M = S + A + T$

For TTT, $M = M^T$

Infall of a shell ($M = M^T$)



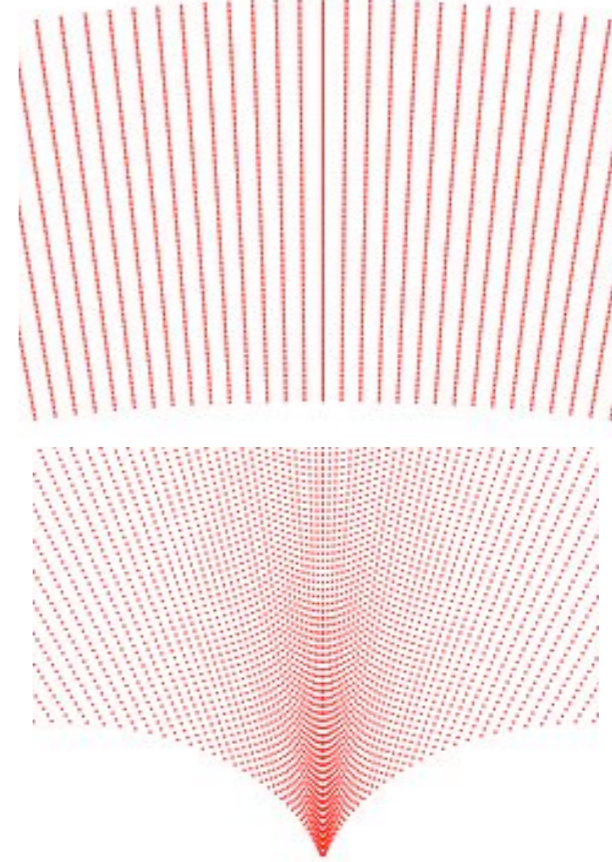
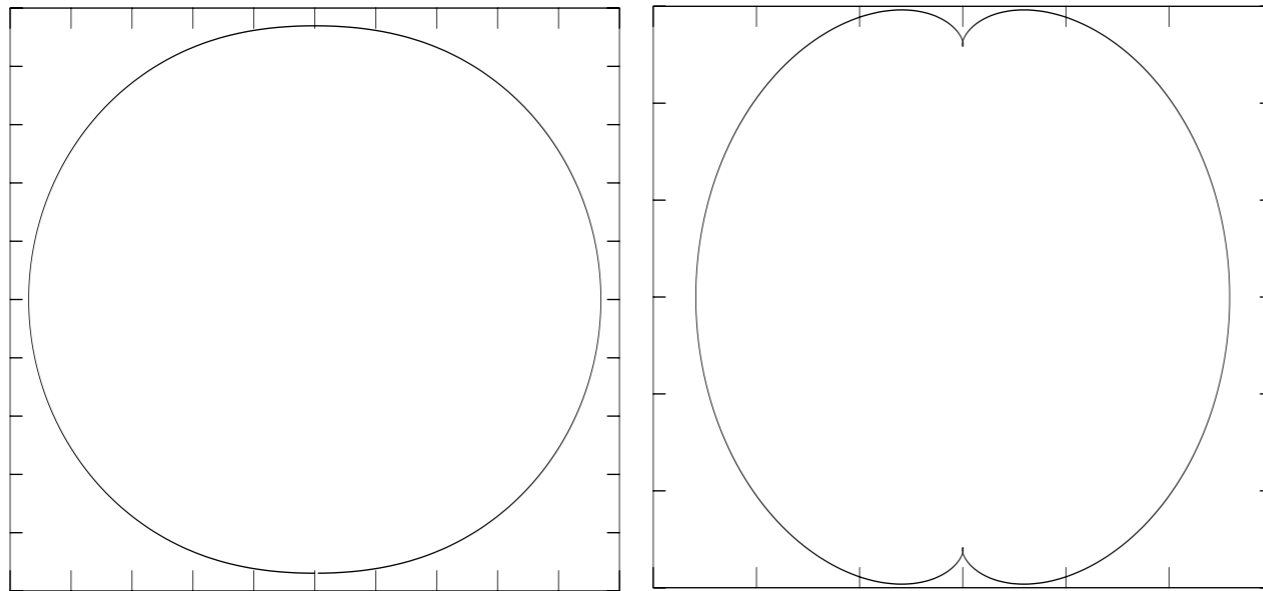
↑
z



→
x

Infall of a shell ($M = M^T$)

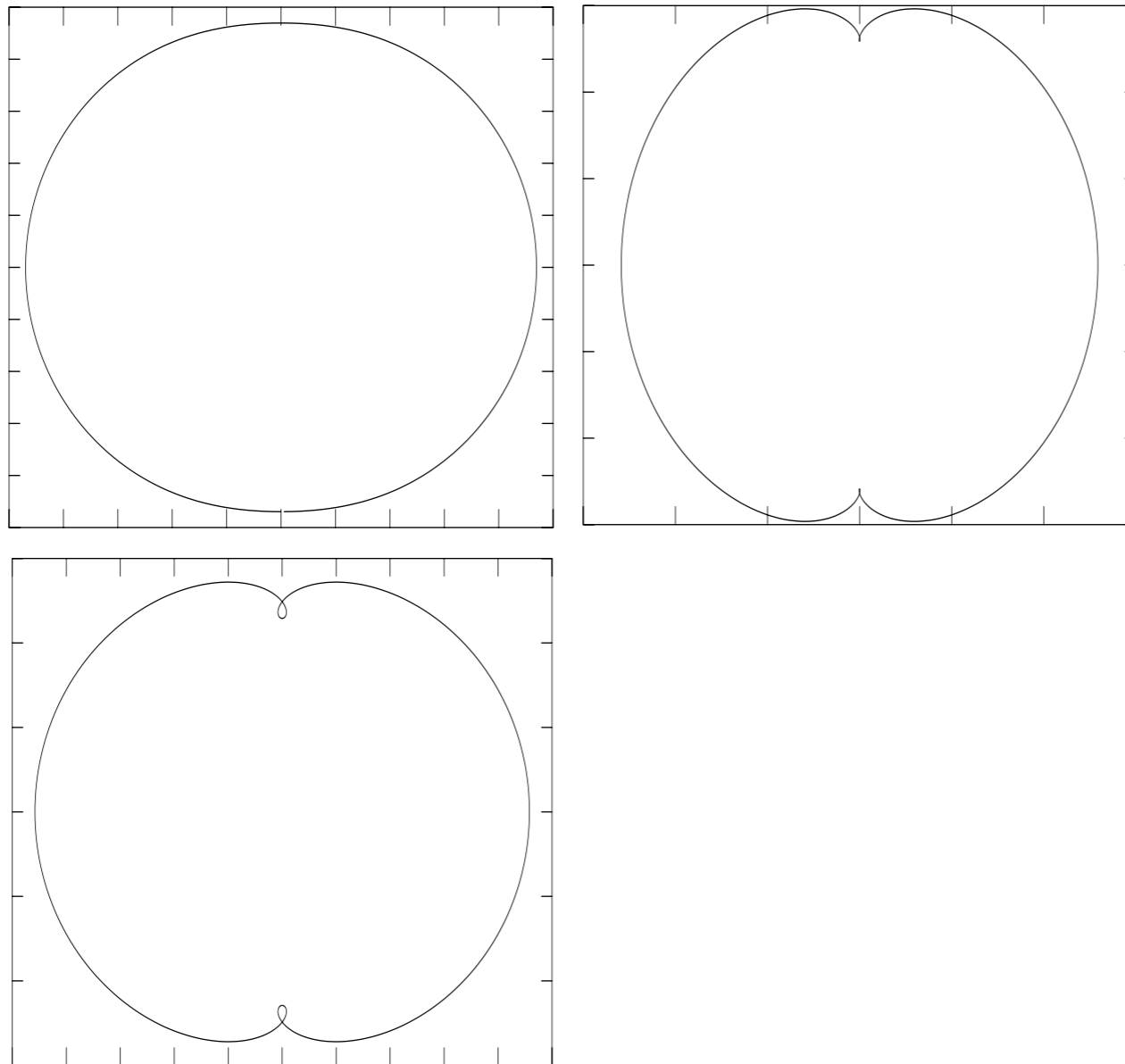
↑
z



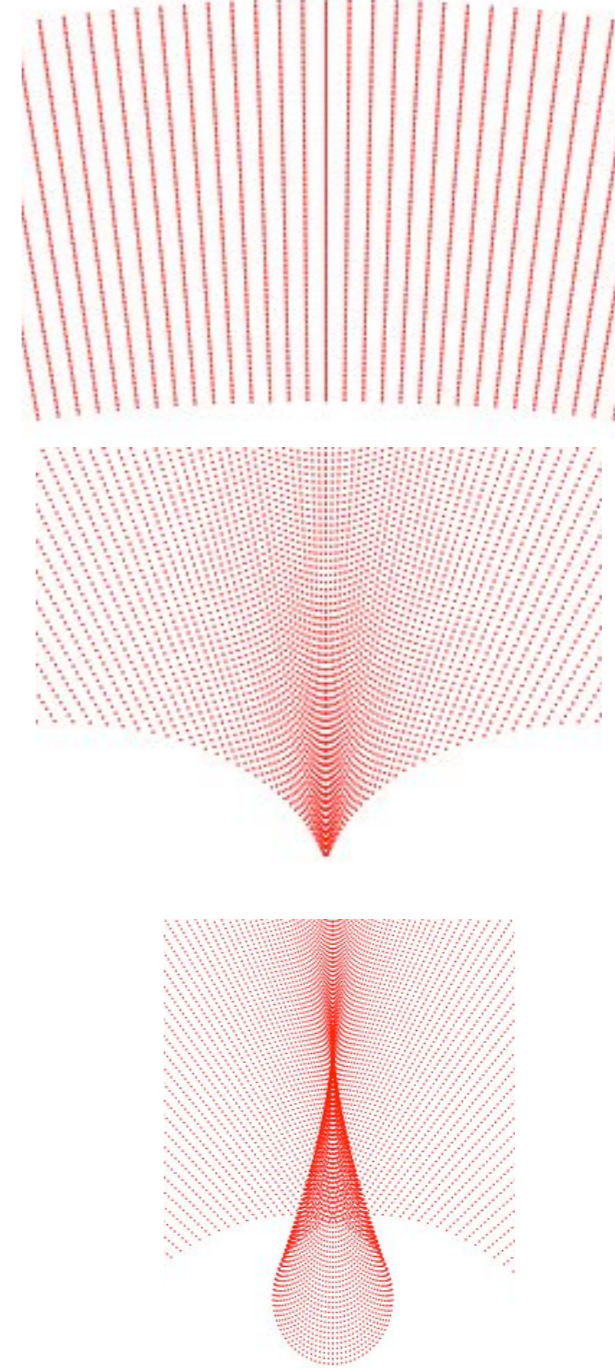
→
x

Infall of a shell ($M = M^T$)

↑
Z

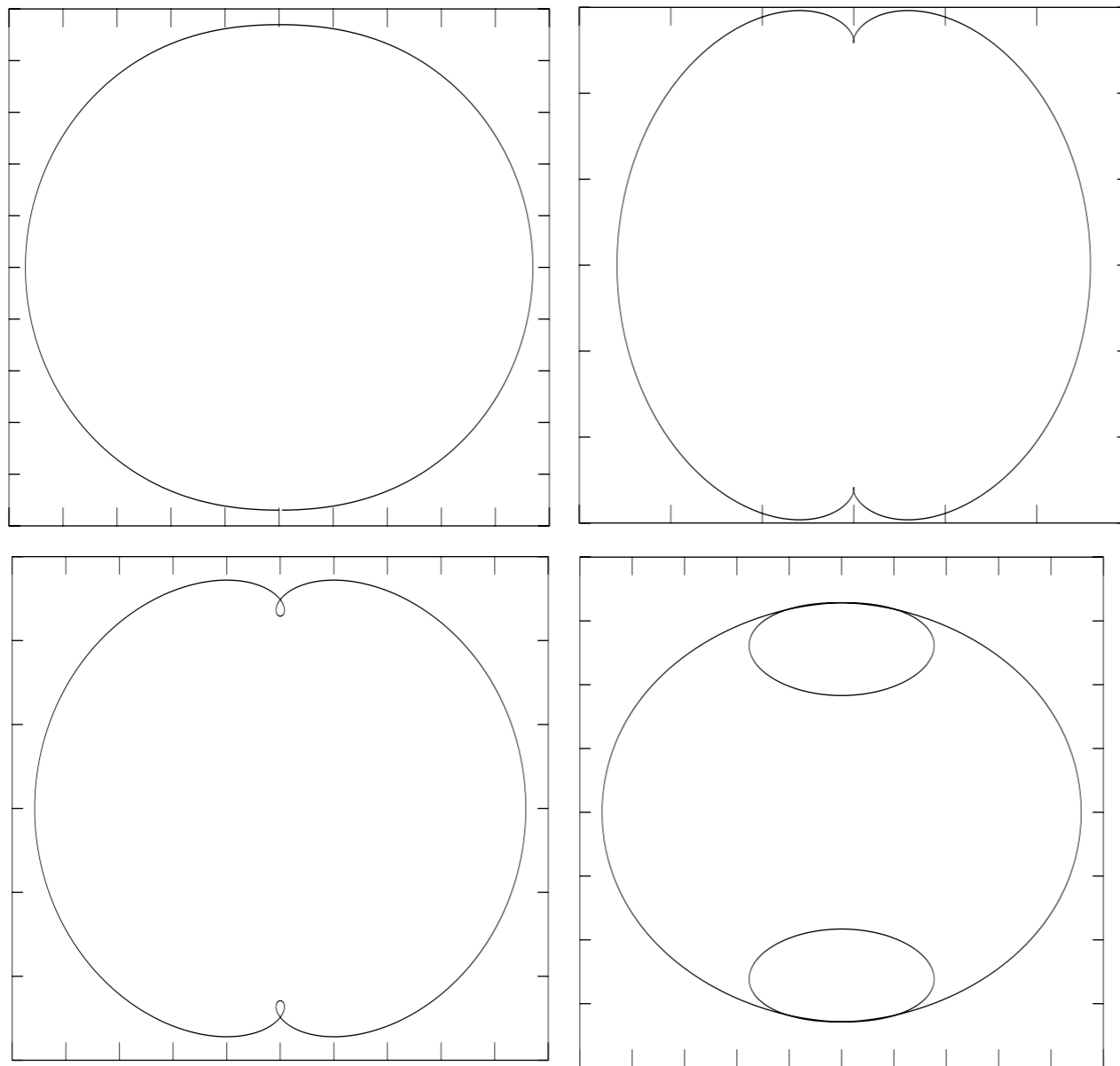


→
X

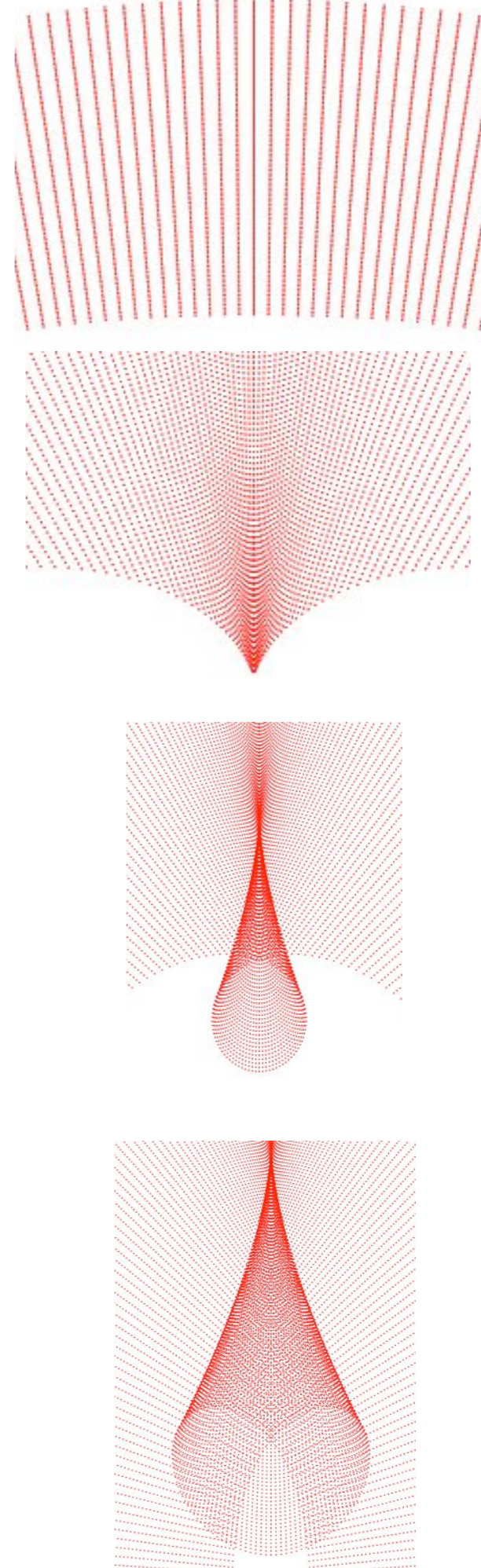


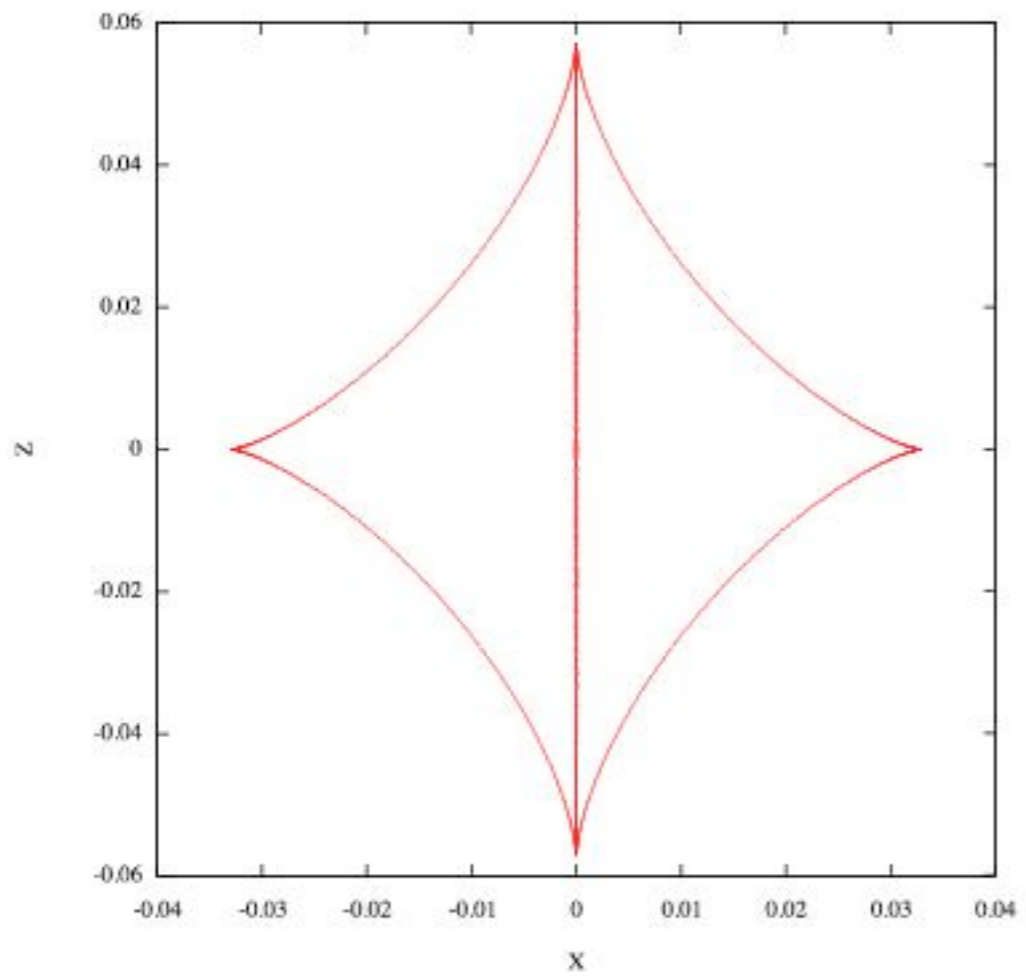
Infall of a shell ($M = M^T$)

↑
Z

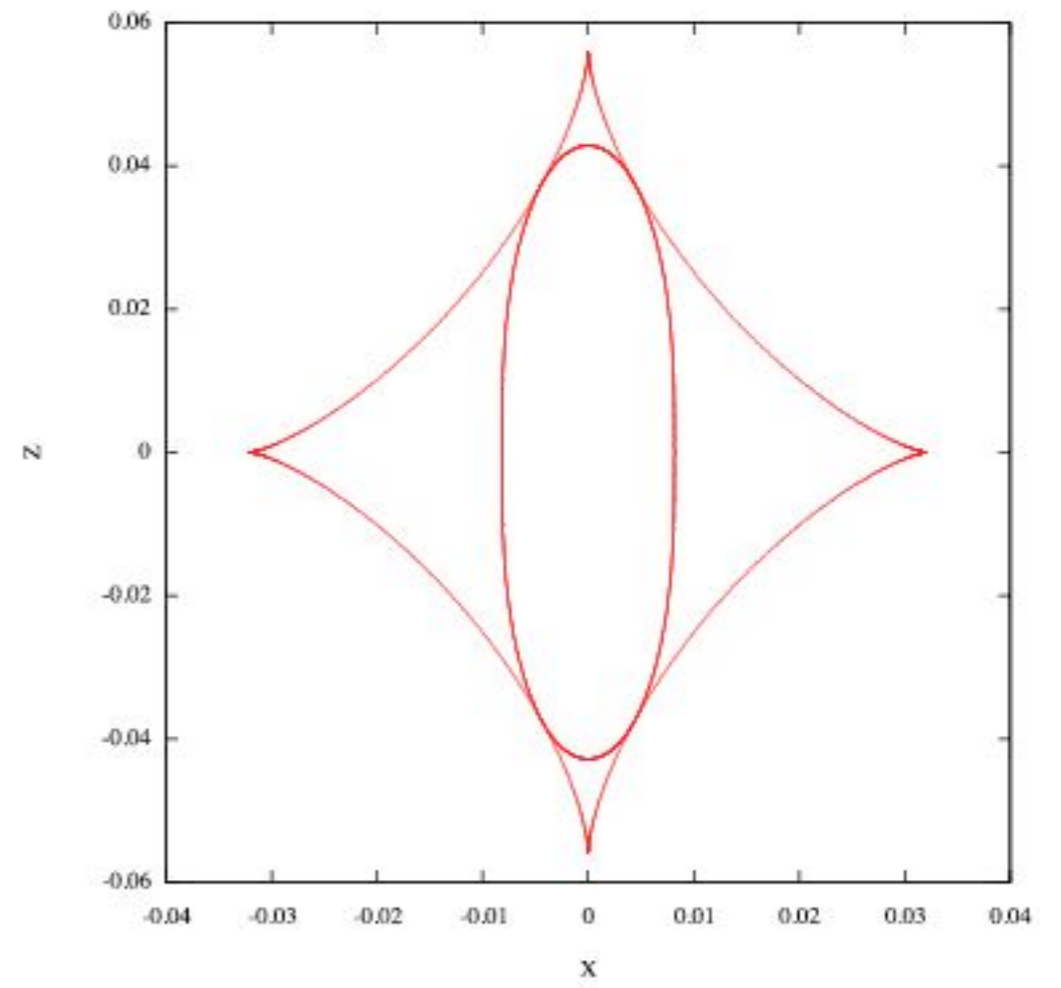


→
X

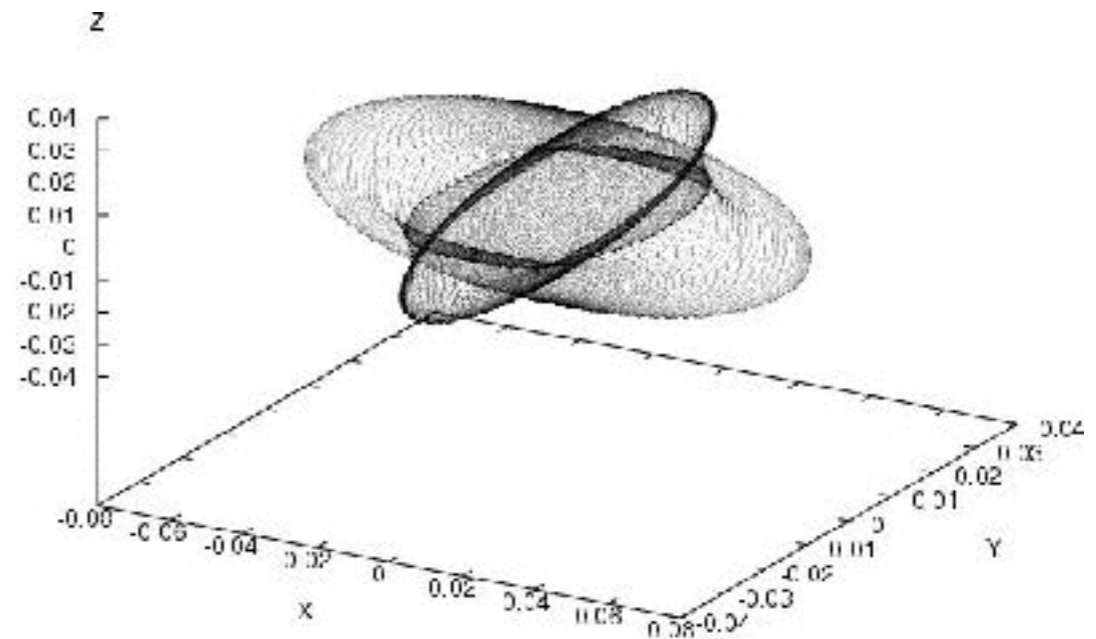
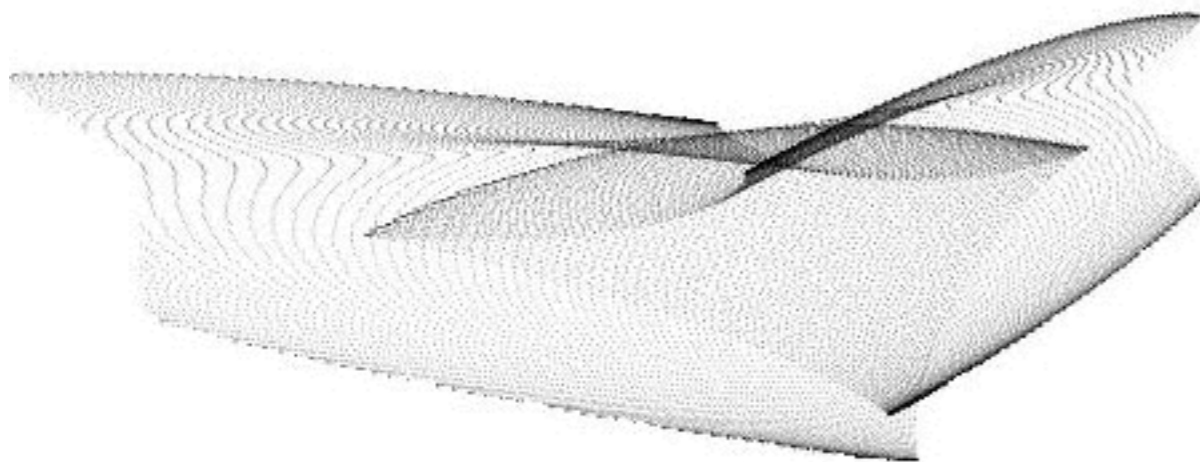
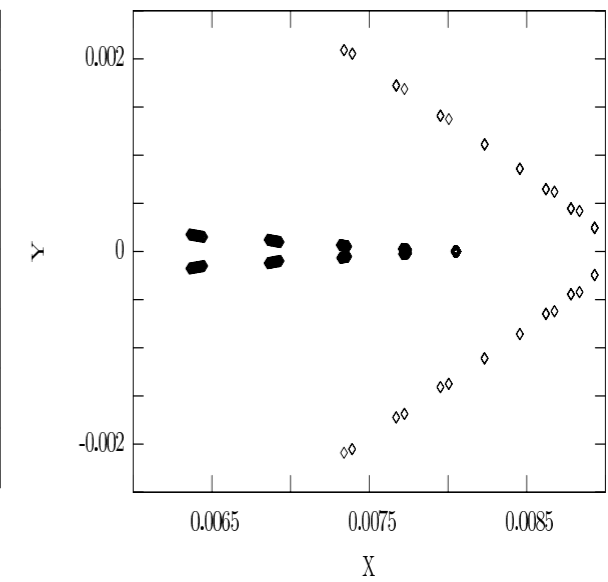
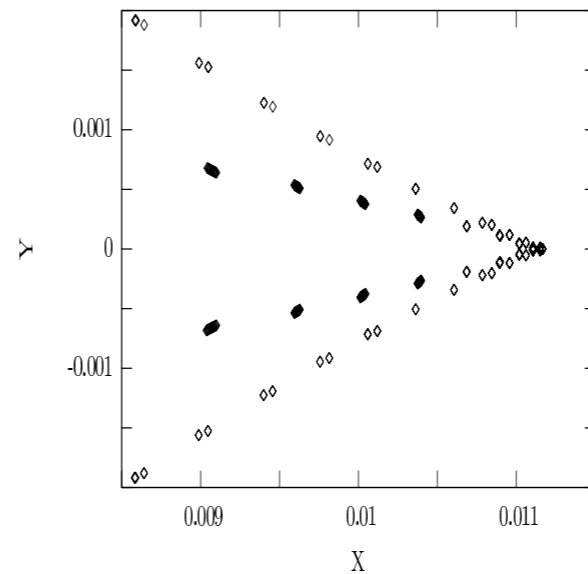
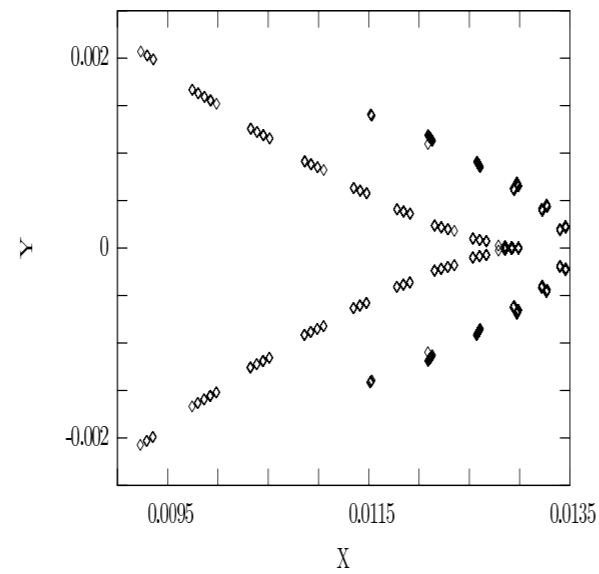
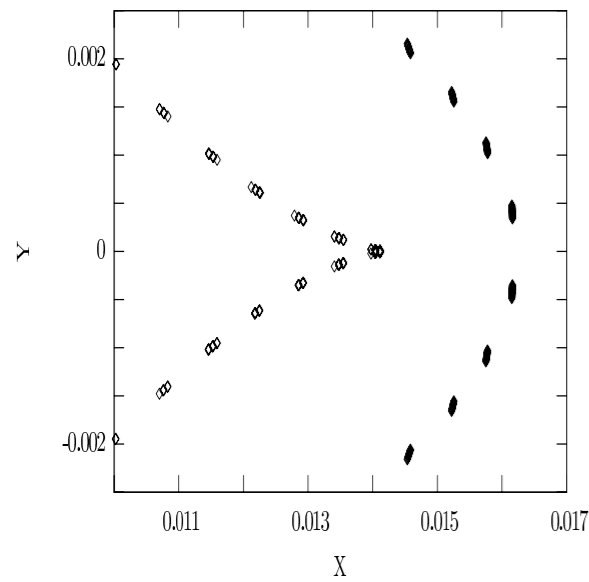




→
Perturbation

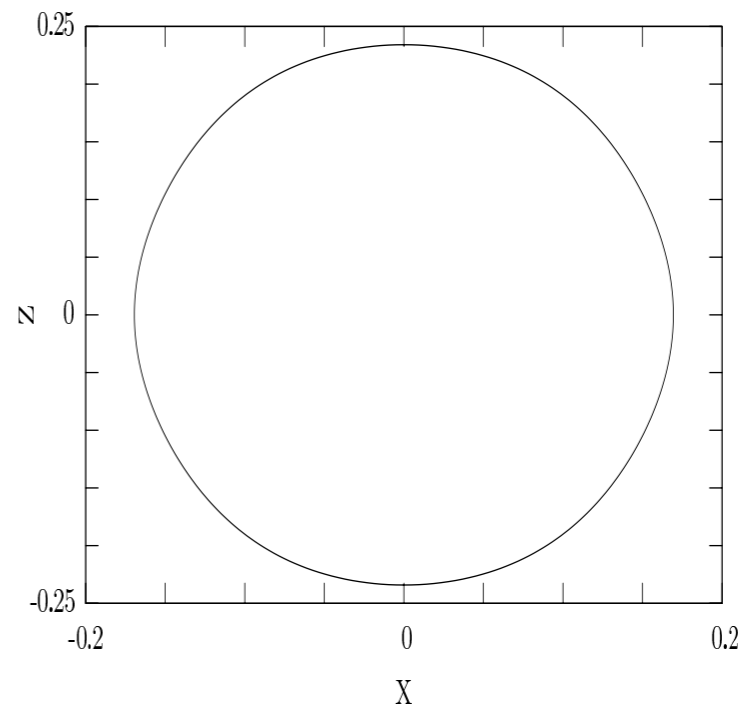
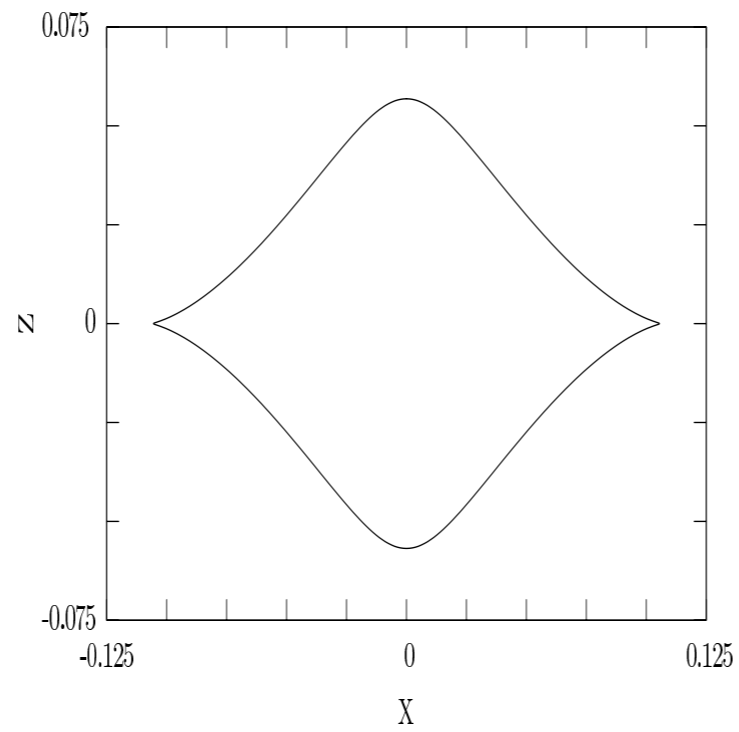
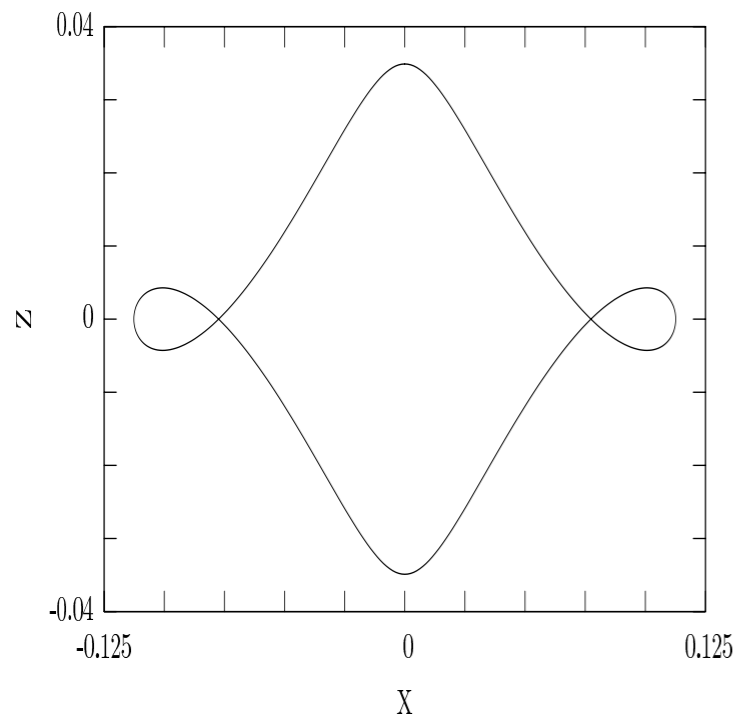
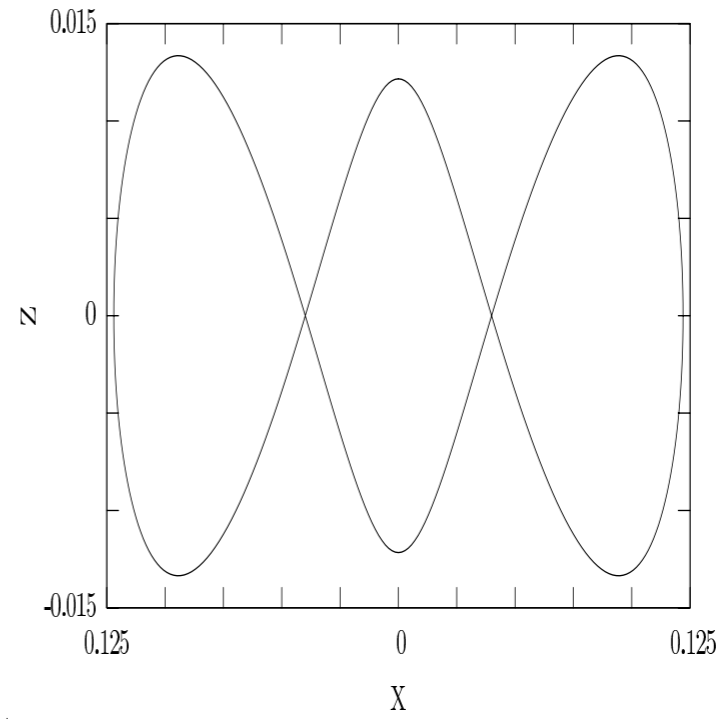
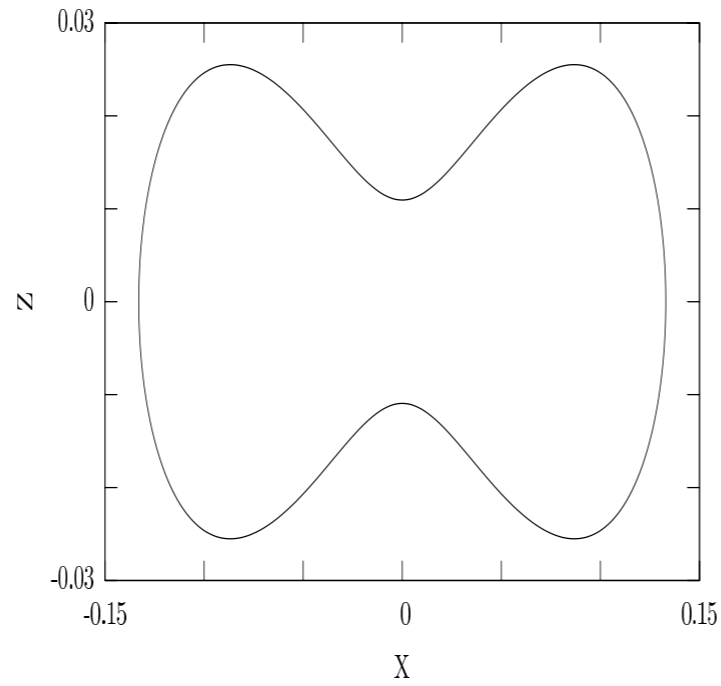
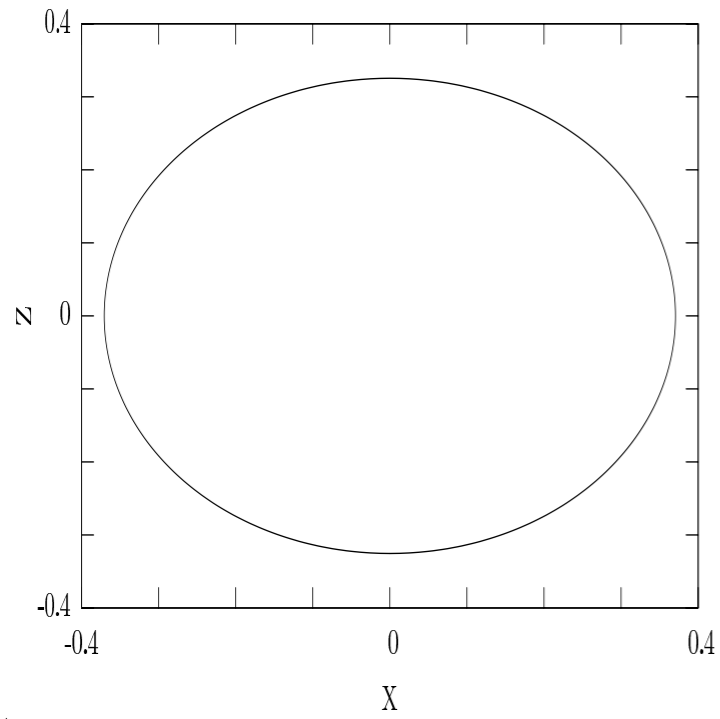


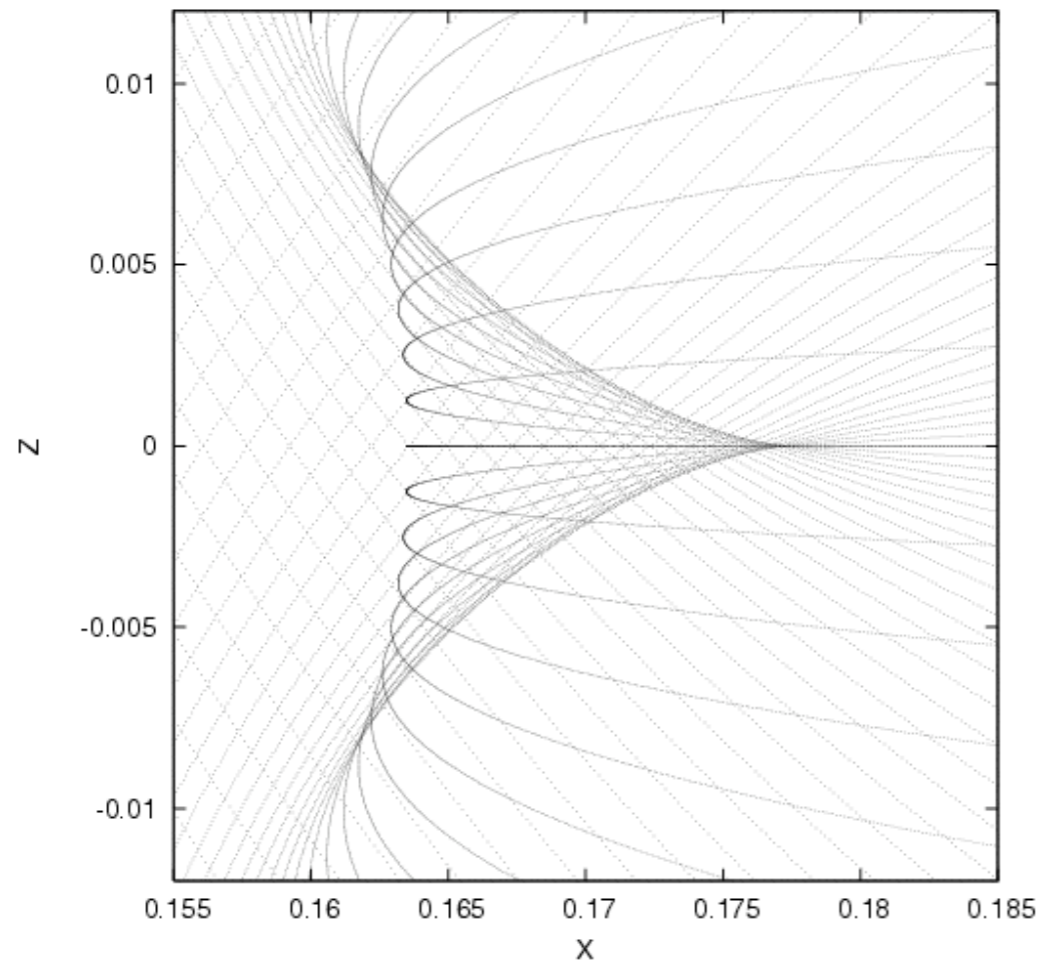
Different x-y cross sections of the caustic -



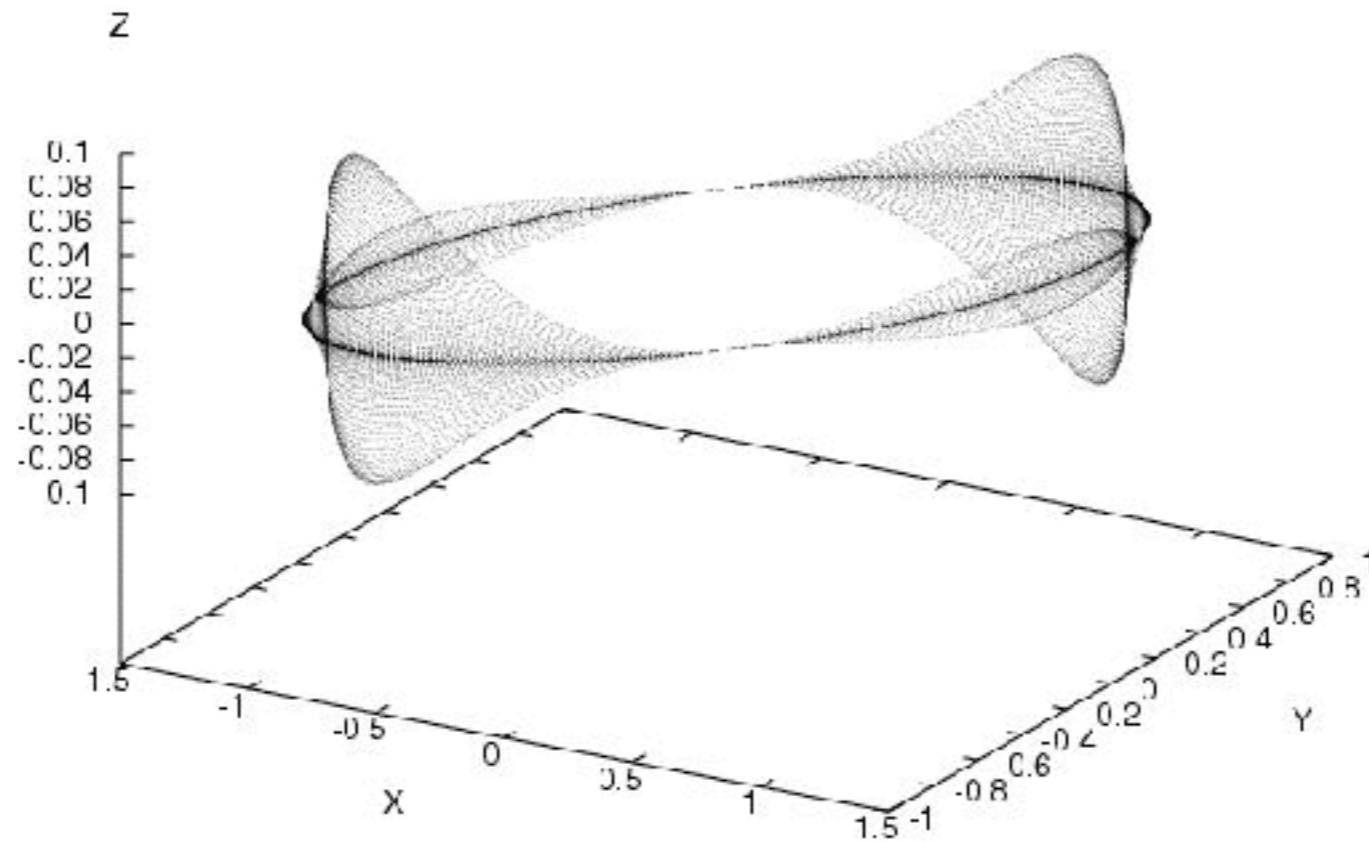
(hyperbolic umbilic catastrophe)

Infall of a shell : $\vec{v}(r) = M \vec{r}$ with $M = -M^T$





(elliptic umbilic catastrophe)



Conclusions:

- The continuous infall of dark matter forms caustics under two assumptions :
 1. The dark matter is collisionless.
 2. It has low velocity dispersion.
- Outer caustics are spherical shells surrounding galaxies. Inner caustics are made up of sections of the higher order catastrophes.
- The structure of inner caustics depends on the dark matter angular momentum.