Lecture 2 Conservation laws and planetary waves

- Potential vorticity
- Orographic forcing of Rossby waves
- Barotropic Instability
- •Baroclinic Instability
- Fronts and wave breaking
- Midlatitude storm tracks

Planetary Balance and conservation

If the time scale $\boldsymbol{\tau}$ of the motion is larger than :

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\tau > f^{-1} > N^{-1}
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 τ > 20hs > 10min And for scales large such U < L/f, L<~ a (earth radius).

The flow is in *hydrostatic balance* Buoyancy ~vertical component of pressure force. And in *geostrophic balance*: coriolis force ~ horizontal component of pressure force.

Hydrostatic Balance



Geostrophic Balance

If the Rossby Number is much smaller than 1 ($R_0 = U/(fL) << 1$)

$$-u2\Omega\sin\phi_0 = \frac{C_p\theta_0}{a\cos\phi_0}\frac{\partial\pi}{\partial\lambda}$$
$$v2\Omega\sin\phi_0 = \frac{C_p\theta_0}{a}\frac{\partial\pi}{\partial\phi}$$

Where *u* and *v* are zonal and meridional velocity components and ϕ_0 a particular latitude

Conservation of Potential Vorticity

Full systemQuasi-geostrophicShallow water
$$\frac{\partial Q}{\partial t} + u \frac{\partial Q}{\partial x} + v \frac{\partial Q}{\partial y} + w \frac{\partial Q}{\partial z} = 0$$
 $\frac{\partial q}{\partial t} + u_s \frac{\partial q}{\partial x} + v_s \frac{\partial q}{\partial y} = 0$ $\frac{\partial q_s}{\partial t} + u_s \frac{\partial q_s}{\partial x} + v_s \frac{\partial q_s}{\partial y} = 0$ $Q = (f + \zeta) \theta_z + \eta \theta_y + \mu \theta_x$ $q = (\zeta + f_0 + \beta y + \frac{f_0}{\rho_r} \frac{\partial}{\partial z} \left\{ \frac{\rho_r}{N^2} \phi_z \right\}$ $q_s = (\frac{\zeta + f}{H + h})$ $\zeta = \frac{\partial v}{\partial x} - \frac{\partial u}{\partial y}$ $\zeta = \frac{\partial v}{\partial x} - \frac{\partial u}{\partial y} = \nabla^2 \phi$ $\chi^2 = \frac{\partial v}{\partial x} - \frac{\partial u}{\partial z}$ $\mu = \frac{\partial w}{\partial x} - \frac{\partial v}{\partial z}$ $N^2 = \frac{g}{\theta_r} \frac{\partial \theta}{\partial z}$ H

Rossby Waves



Possibility of stationary because:





Fig. 4.7 A cylindrical column of air moving adiabatically, conserving potential vorticity.

³ Named for the German meteorologist Hans Ertel. A more general form of Ertel's potential vorticity is discussed, for example, in Gill (1982). Potential vorticity is often expressed in the potential vorticity unit (PVU), where 1 PVU = 10^{-6} K kg⁻¹ m² s⁻¹.





Stationary response due to orography (shallow water global model)











Figure 6. The quasigeostrophic solution for steady flow over a semi-infinite ridge in a semi-infinite atmosphere. The plets show (top) total buoyancy (solid) and perturbation (dashed)buoyancy, (middle) meridional velocity perturbation, and (bottom) relative vorticity perturbation. The heavy contour indicates the mountain profile.





Barotropic and Baroclinic Instabilities

The major source of midlatitudes weather











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Zonal Mean Flow and Potential Temperature for the Baroclinic Simulations

Elements of Baroclinic Instability





Fig. 8.1 A schematic picture of cyclogenesis associated with the arrival of an upper-level positive vorticity perturbation over a lower-level baroclinic region. (a) Lower-level cyclonic vorticity induced by the upper-level vorticity anomaly. The circulation induced by the vorticity anomaly is shown by the solid arrows, and potential temperature contours are shown at the lower boundary. The advection of potential temperature by the induced lower-level circulation leads to a warm anomaly slightly east of the upper-level vorticity anomaly. This in turn will induce a cyclonic circulation as shown by the open arrows in (b). The induced upper-level circulation will reinforce the original upper-level anomaly and can lead to amplification of the disturbance. (After Hoskins et al., 1985.)



Cut-off cyclone

PV anomaly PT anomaly

Cut-off anticyclone

b











Fig. 10.13 The observed mean energy cycle for the Northern Hemisphere. Numbers in squares are energy amounts in units of 10^5 Jm^{-2} . Numbers next to arrows are energy transformation rates in units of W m⁻². $B(\overline{p})$ represents a net energy flux into the Southern Hemisphere. Other symbols are defined in the text. (Adapted from Oort and Peixoto, 1974.)

Baroclinic effects:

Atmospheric Fronts

Transfer of momentum and heat poleward

Breaking upper and lower waves







Potential Vorticity of the Shallow Water Global Model for three different amplitude of the forcing







Collective effects of baroclinic waves (daily weather)

Produce

Storm tracks and quasi stationary patterns affecting the basic flow (westerlies)

Observed Precipitation \underline{V}_{925} and Z_{500}



Precipitation: Xie-Arkin 1979/80 – 1998/99, V₉₂₅: ERA40 1962-01, Z₅₀₀: ERA40 1962-01, CI=10 dam

Mark Rodwell

Time-space scale of atmospheric systems







Fig. 10.16 Meridional cross sections showing the relationship between the time mean secondary meridional circulation (continuous thin lines with arrows) and the jet stream (denoted by J) at locations (a) upstream and (b) downstream from the jet stream cores. (After Blackmon et al., 1977. Reproduced with permission of the American Meteorological Society.)



Surfaces of relative vorticity as inferred from regression analysis (-24h, 0h, +24h)

I. Orlanski and B. Gross: Baroclinic lifecycles in a storm track environment. JAS 2000

Mean intensity of the cyclones for DJF: 1979-95



Patterns quite well simulated, although peak intensity generally underestimated in HadAM3. Resolution?

Mean track density for DJF: 1979-95



Pacific storm track is too strong. Atlantic storm track too weak in HadAM3, possibly related to lack of systems coming off the eastern side of the Rockies.

Number of Cyclones DJFM 1982-2001



Experiments on Storm Track variability

A high resolution (9km and 18km) cloud resolving non-hydrostatic model (ZETANC*) was used to perform several idealized storm track simulations. The solution run for 220days and sensitivity to imposed SST were performed. The animation shows the column liquid water content.

The frontal rain band ahead of cold fronts displays a variety of cloud systems, from deep convective clouds in sub tropical latitudes to more stable stratiform clouds in the middle latitudes.



Things to remember from Lecture 2.

 Planetary waves are an essential part of the daily weather and climate.

 Land-sea contrast, topographic features etc can force planetary waves, Rossby waves.

 Its characteristics are; they propagate to the west and radiate energy to the east.

• Also these waves could be unstable, they can grow to large amplitudes from very small perturbations.

• The major planetary wave generation are due to barotropic and baroclnic instabilities. They are responsible for producing the daily weather. Most the time they transport heat and momentum to higher latitudes. Warming the polar regions and removing heat from the subtropics.

• The assemble of all these waves tend to produce wave active regions which we called the storm tracks.