

RAMA: The Research Moored Array for African-Asian-Australian Monsoon Analysis  
and Prediction

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1 Capsule

2 A new moored buoy array in the historically data sparse Indian Ocean provides

3 measurements to advance monsoon research and forecasting.

4 Abstract

5

6         The Indian Ocean is unique among the three tropical ocean basins in that it is  
7 blocked at 25°N by the Asian land mass. Seasonal heating and cooling of the land sets the  
8 stage for dramatic monsoon wind reversals, strong ocean-atmosphere interactions, and  
9 intense seasonal rains over the Indian subcontinent, Southeast Asia, East Africa, and  
10 Australia. Recurrence of these monsoon rains is critical to agricultural production that  
11 supports a third of the world's population. The Indian Ocean also remotely influences the  
12 evolution of El Niño and the Southern Oscillation (ENSO), the North Atlantic Oscillation  
13 (NAO), North American weather, and Atlantic hurricane activity. Despite its importance  
14 in the regional and global climate system though, the Indian Ocean is the most poorly  
15 observed and least well understood of the three tropical oceans.

16         This article describes the **R**esearch **M**oored **A**rray for African-Asian-Australian  
17 **M**onsoon **A**nalysis and Prediction (RAMA), a new observational network designed to  
18 address outstanding scientific questions related to Indian Ocean variability and the  
19 monsoons. RAMA is a multi-nationally supported element of the Indian Ocean  
20 Observing System (IndOOS), a combination of complementary satellite and *in situ*  
21 measurement platforms for climate research and forecasting. The article describes the  
22 scientific rationale, design criteria, and progress towards implementing the array. Initial  
23 RAMA data are presented to illustrate how they contribute to improved documentation  
24 and understanding of phenomena in the region. Potential future applications of the data  
25 for societal benefit are also described.

26

26 1. Introduction

27 The Indian Ocean is unique among the three tropical ocean basins in that it is  
28 blocked at 25°N by the Asian land mass. Seasonal heating and cooling of the land sets the  
29 stage for dramatic monsoon wind reversals and intense seasonal rains over the Indian  
30 subcontinent, Southeast Asia, East Africa, and Australia. The Asian landmass also blocks  
31 the ocean to the north so that currents cannot carry heat from the tropics to higher  
32 northern latitudes as in the other oceans. Ocean-atmosphere interactions in the region are  
33 highly dynamic, involving seasonal current reversals associated with monsoon wind  
34 forcing and significant exchanges of heat across the air-sea interface. The Indian Ocean  
35 also receives heat from the Pacific via the Indonesian Throughflow (Gordon, 2001) while  
36 exporting heat to the Atlantic via the Agulhas Current system (de Ruijter et al, 1999).

37 Monsoon rains occur each year, supporting agricultural production that provides  
38 food for a third of the world's population. These rains are irregular, however, leading to  
39 years of drought or flood that have significant socio-economic consequences (Webster et  
40 al, 1998). Failure of Indian summer monsoon rains in 2002 (Waple and Larimore, 2003)  
41 and excessive rains in equatorial East Africa in late 2006 (Arguez, 2007) are recent  
42 examples of major disasters. Atmospheric teleconnections carry the effects of Indian  
43 Ocean climate anomalies to other regions of the globe, where they affect evolution of El  
44 Niño and the Southern Oscillation (ENSO) (McPhaden, 1999; Zhang, 2005), the North  
45 Atlantic Oscillation (Hoerling et al, 2001), Sahel rainfall (Giannini et al, 2003), Atlantic  
46 hurricane formation (Maloney and Hartmann, 2000), the atmospheric circulation of the  
47 North Pacific (Annamalai et al, 2007) and western U.S. weather (Higgins and Mo, 1997).  
48 Thus, the potential benefits of improved description, understanding and prediction of the

49 coupled ocean–atmosphere system in the Indian Ocean are enormous. However, the  
50 present lack of comprehensive data records for the Indian Ocean severely limits our  
51 knowledge of key physical processes and our ability to provide reliable monsoon  
52 forecasts even one season ahead.

53 Observational efforts in the Indian Ocean date back to the International Indian  
54 Ocean Expedition of the 1960s (Knauss, 1961; National Research Council, 2000) which  
55 was the first comprehensive study of Indian Ocean circulation and water mass properties.  
56 In the 1970s, the Indian Ocean Experiment (INDEX; Luyten and Roemmich, 1982) and  
57 the Summer Monsoon Experiment (MONEX; Krishnamurti, 1985) were designed to  
58 improve the description and understanding of seasonally varying oceanic and  
59 atmospheric circulation, respectively. The 10-year (1985-94) Tropical Ocean Global  
60 Atmosphere (TOGA) program (National Research Council, 1996) and the World Ocean  
61 Circulation Experiment (Siedler et al, 2001) with a field phase from 1990-98 were major  
62 international efforts to study ocean-atmosphere interactions and basin scale ocean  
63 circulation patterns related to climate. Short term, regional studies such as the Bay of  
64 Bengal Monsoon Experiment (Bhat et al, 2000), the Arabian Sea Monsoon Experiment  
65 (Sanjeeva Rao and Sikka, 2005) and the Joint Air-Sea Monsoon Experiment (Webster et  
66 al, 2002) complemented these broad scale programs by examining key physical  
67 processes controlling variability in specific areas of the Indian Ocean. In contrast to those  
68 studies which focussed on physical oceanography, ocean-atmosphere interactions, and  
69 climate variability, the Joint Global Ocean Flux Study Arabian Sea Expedition (Smith,  
70 2001) of the mid-1990s emphasized ocean biogeochemical cycles and biological  
71 productivity in relation monsoon forcing.

72           These programs laid a foundation for understanding oceanic and atmospheric  
73 variability related to the monsoons, but they did not leave a significant legacy of  
74 sustained ocean observations in the region (e.g. McPhaden et al, 1998). Development of a  
75 tropical Indian Ocean observing system progressed very slowly in part because Indian  
76 Ocean sea surface temperature (SST) anomalies on interannual time scales were most  
77 often within the range of observational errors (e.g. Annamalai and Murtugudde, 2004),  
78 because there was debate about whether ocean dynamics played a major role in their  
79 predictability, and because “the predictive relationship between Indian Ocean SST and  
80 monsoon rainfall have remained especially poorly characterized...” (Clark et al, 2000).  
81 By contrast, during TOGA and continuing through late 1990s as part of the follow-on  
82 Climate Variability and Predictability (CLIVAR) program, efforts in tropical ocean  
83 observing system development were focussed heavily on the Pacific and Atlantic where  
84 ocean-atmosphere interactions were better understood and seasonal prediction efforts  
85 were more advanced. The remoteness of the Indian Ocean from first world countries of  
86 North America, Europe, and Asia was also a factor in the relatively slow development of  
87 an Indian Ocean observing system compared to the Pacific and Atlantic.

88           The past 10 years have seen a rebirth of interest in the Indian Ocean stimulated by  
89 the prominence of the 1997 Indian Ocean Dipole (IOD<sup>1</sup>) event (Saji et al, 1999; Webster  
90 et al, 1999; Murtugudde et al, 2000). This event highlighted the dramatic nature and  
91 climatic consequences of ocean-atmosphere interactions in the Indian Ocean. The event  
92 also drew attention to the relative dearth of measurements in the Indian Ocean and the  
93 opportunities to develop a systematic and sustained ocean observing system there.

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<sup>1</sup> The Indian Ocean Dipole is also sometimes referred to as the Indian Ocean Dipole Zonal Mode, or IODZM.

94           This paper describes the rationale for and design of a sustained, basin-scale  
95 moored buoy array referred to as the *Research Moored Array* for African-Asian-  
96 Australian *Monsoon Analysis* and Prediction (RAMA<sup>2</sup>). The array complements other  
97 elements of the recently designed Indian Ocean Observing System (IndOOS) which  
98 collectively represents an Indian Ocean contribution to the Global Ocean Observing  
99 System (GOOS) (CLIVAR-GOOS Indian Ocean Panel, 2006; Meyers and Boscolo,  
100 2006). IndOOS in general, and RAMA in particular, address the need to establish a  
101 system for comprehensive, long term, high quality, real-time measurements in the Indian  
102 Ocean suitable for climate research and forecasting. The broad range of time scales and  
103 rapid changes that can occur in the Indian Ocean dictate the need for a moored buoy array  
104 providing time series data with high temporal resolution as an essential element of  
105 IndOOS. In this respect, RAMA is the Indian Ocean equivalent of the Tropical  
106 Atmosphere Ocean/Triangle Trans Ocean Buoy Network (TAO/TRITON; McPhaden et  
107 al, 1998; Kuroda and Amitani, 2000) and the Prediction and Research Moored Array in  
108 the Tropical Atlantic (PIRATA; Bourles et al, 2008), which respectively anchor basin  
109 scale observing systems in the tropical Pacific and Atlantic Oceans. RAMA is targeted at  
110 understanding and prediction of the east African, Asian, and Australian monsoons, but it  
111 will benefit nations outside the Indian Ocean region because of atmospheric  
112 teleconnections that influence the far field. In addition, real-time RAMA data will  
113 contribute to improved weather and marine forecasts, like those for tropical cyclones and  
114 storm surge. Even though the array is in the initial stages of implementation, the data are  
115 already finding practical applications as illustrated in the side bar on Australian farmers.

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<sup>2</sup> In Hindu mythology, Rama is an ancient king of India and the hero of the epic “Ramayana”.

116           The remainder of this paper is organized as follows. We first briefly review the  
117 range of phenomena that motivate the study of climate variability in the Indian Ocean  
118 region. Emphasis is on intraseasonal to interannual time scales and the upper ocean north  
119 of 30°S where RAMA will have its greatest impact. The reader is referred to Schott et al  
120 (2008) for a more complete account of Indian Ocean phenomenology including western  
121 boundary currents, the Indonesian Throughflow, major surface currents, and the cross-  
122 equatorial overturning circulation. We then describe IndoOS and the design criteria for  
123 RAMA, followed by a progress report on RAMA implementation. Examples of RAMA  
124 data are presented to illustrate their value for research applications, after which we  
125 conclude with a brief summary and discussion.

## 126       2. Indian Ocean Phenomenology

127           Hallmark attributes of the Indian Ocean climate system are the dramatic reversals  
128 of the surface winds propelled by seasonally varying land-ocean temperature contrasts.  
129 Equally dramatic are the seasonally varying wind-driven ocean currents that affect the  
130 evolution of SST (Fig. 1; see also Schott and McCreary, 2001 and Schott et al, 2008).  
131 Associated with monsoon wind variations are pronounced seasonal shifts in atmospheric  
132 convection and rainfall over East Africa, southern and eastern Asia, and northern  
133 Australia. Collectively, these variations are a regional manifestation of seasonal changes  
134 in the Hadley and Walker Circulations, which extend throughout the global tropics.

135           In boreal winter, Northeast Monsoon winds converge with Southeast trade winds  
136 in the Intertropical Convergence Zone (ITCZ) located between 5°-12°S. These wind  
137 systems drive a surface ocean circulation that features broad, westward flowing currents  
138 on either side of the equator—the Northeast Monsoon Current and South Equatorial



139 Current—and a southward flowing Somali Current along the east coast of Africa.  
140 Sandwiched between the two westward currents is the eastward flowing Equatorial  
141 Countercurrent, in geostrophic balance with meridional density contrasts in the  
142 Seychelles-Chagos thermocline ridge region (Hermes and Reason, 2008; Yokoi et al,  
143 2008). This thermocline ridge, formed by wind stress curl driven upwelling in the ITCZ,  
144 rises to within 20 m of the surface and results in a zonal band of very thin mixed layers.  
145 SST variations here are sensitive to surface heat fluxes and vertical turbulent heat  
146 exchanges with the thermocline because of the low heat capacity of these thin mixed  
147 layers (Duvell and Vialard, 2007). Seasonal time scale SST anomalies in this region can  
148 moreover remotely influence subsequent development of summer monsoon rainfall along  
149 the western Ghats (Vecchi and Harrison, 2004; Izumo et al, 2008).

150 Wind and circulation patterns in boreal summer are radically different from those  
151 in boreal winter. Southwest Monsoon winds drive the currents north of the equator to the  
152 east and weaken both the countercurrent and the thermocline ridge. Intense wind-driven  
153 coastal upwelling during summer off the Horn of Africa leads to cooling and, through the  
154 supply of nutrient rich thermocline water, high biological productivity in the Arabian Sea.  
155 This upwelling is fed by a cross equatorial circulation cell that is unique to the Indian  
156 Ocean, with source waters formed in the southeastern Indian Ocean (~20°-30°S) flowing  
157 northward at thermocline depth to the Arabian Sea (Schott et al, 2004). SSTs in Bay of  
158 Bengal on the other hand remain relatively high during summer, stoking the growth of  
159 cyclones that often have devastating impacts on surrounding countries. These high SSTs  
160 result in part from a strongly salt stratified upper ocean capped by shallow fresh water  
161 mixed layers that trap surface heat fluxes and inhibit mixing with the thermocline (Shenoi

162 et al, 2002). River runoff and open ocean rainfall supply the fresh water to needed to  
163 maintain these thin mixed layers.

164         During the transition seasons between the Northeast and Southwest Monsoons,  
165 winds along the equator are westerly. These winds force the eastward flowing “Wyrтки  
166 Jets” (Wyrтки, 1973) in April-May and October-November of each year. The Wyrтки Jets  
167 are important in transporting mass from west to east along the equator and play an  
168 important role in the seasonal heat balance of the basin. They are also dynamically  
169 linked to variability off of Sumatra and Java through equatorial and coastal waveguide  
170 processes (Wijffels and Meyers, 2004).

171         Embedded within the seasonally varying monsoons are energetic intraseasonal  
172 oscillations (or ISOs) on weekly to monthly time scales. The best known form of ISO is  
173 the Madden-Julian Oscillation (MJO), an eastward propagating wave-like phenomenon in  
174 the atmosphere with periods of roughly 30-60 days (Madden and Julian, 1994; Zhang,  
175 2005). MJO convection is spawned over the Indian Ocean and subsequently propagates  
176 eastward as part of a planetary scale fluctuation in upper tropospheric winds. Variability  
177 in surface winds and deep atmospheric convection is most energetic in regions of warm  
178 SST ( $\geq 27^{\circ}\text{C}$ ), where interaction with the oceanic mixed layer plays an important role in  
179 organizing the MJO (Waliser et al, 1999). Along the equator, intraseasonal wind forcing  
180 generates energetic eastward propagating oceanic Kelvin waves (Han, 2005; Fu, 2007)  
181 which, upon encountering Sumatra, continue poleward as coastally trapped waves. The  
182 majority of coastally trapped intraseasonal wave energy propagating southeastward off  
183 Sumatra and Java eventually leaks into the Indonesian Seas through the Lombok Strait  
184 where it affects Throughflow transports (Syamsudin et al, 2004).

185           The MJO is strongest in boreal winter and spring in the southern tropics, with a  
186 secondary peak during boreal summer north of the equator (Zhang, 2005). During boreal  
187 winter, ocean-atmosphere interactions are particularly strong with large intraseasonal  
188 SST variations in the vicinity of the shallow Seychelles-Chagos thermocline ridge (Duvel  
189 and Vialard, 2007; Vialard et al, 2008a,b). In boreal summer, the MJO spins off cloud  
190 and rain bands near the equator that propagate poleward over the Bay of Bengal,  
191 contributing to alternating “active” periods of enhanced rainfall and “break” periods of  
192 reduced rainfall over the Indian Subcontinent (Madden and Julian, 1994). The number  
193 and duration of active/break periods in a particular year determines the net seasonal  
194 monsoon rainfall. For this reason intraseasonal oscillations are often referred to as the  
195 “building block of the monsoons.”

196           There are a variety of significant multi-time scale interactions involving the MJO.  
197 For instance, the MJO strongly modulates variability associated with the Wyrski Jets  
198 (Han et al, 2004; Masumoto et al, 2005; Sengupta et al, 2007). The diurnal cycle is  
199 important in modulating MJO SST variability and ocean feedbacks to the atmosphere  
200 (Woolnough et al, 2007). Synoptic time scale cyclones often develop in association with  
201 the MJO (Liebmann et al, 2004; Bessafi and Wheeler, 2006; Seiki and Takayabu, 2007).  
202 Positive IOD events tend to suppress the MJO and as well as higher frequency  
203 atmosphere fluctuations as a result of the reduced convection over the eastern Indian  
204 Ocean (Shinoda and Han, 2005). Farther afield, propagation of the MJO into the western  
205 Pacific affects the evolution of ENSO (McPhaden, 1999; McPhaden et al, 2006; Zhang,  
206 2005). The MJO likewise influences the development of Atlantic hurricanes (Maloney  
207 and Hartmann, 2000) and winter storms along the west coast of the U.S. (Higgins and

208 Mo, 1997). There are other notable intraseasonal variations in the Indian Ocean region  
209 besides the MJO, such as an oscillation in sea level with a spectral peak at periods near  
210 90 days that Han (2005) attributes to a wind-forced basin scale oceanic resonance.

211 The Indian Ocean is characterized by considerable interannual variability, the  
212 most prominent mode of which is the response to remote forcing from ENSO (Yamagata  
213 et al, 2004; Schott et al, 2008). An eastward shift in the ascending branch of Walker  
214 circulation into the central Pacific during El Niño leads to anomalous subsidence,  
215 suppressed convection, high atmospheric surface pressure, and anomalous easterlies over  
216 the Indian Ocean. El Niño's impact on precipitation includes reduced Indian summer  
217 monsoon rainfall, reduced rainfall in South Africa and Indonesia, and enhanced rainfall in  
218 equatorial east Africa. In boreal spring following the peak of El Niño, basin scale  
219 warming occurs due to a combination of increased surface heat fluxes and, south of the  
220 equator, downwelling Rossby waves forced by anomalous ENSO-induced surface wind  
221 stresses (Xie et al, 2002; Yu et al, 2005; Schott et al, 2008). The amplitude of this  
222 warming is relatively small compared to that in the tropical Pacific, but it leads to  
223 increased tropical cyclone activity in the southwest (Xie et al, 2002) and increased  
224 rainfall during the following summer over much of the basin (Yang et al, 2007). .  
225 Oceanic and atmospheric anomalies of opposite sign are evident in the Indian Ocean  
226 region associated with La Niña events.

227 Another prominent mode of interannual variability in the Indian Ocean is the  
228 IOD. Positive IOD events are characterized by anomalously cold SSTs and suppressed  
229 atmospheric convection off Java and Sumatra, warm SSTs and enhanced convection off  
230 east Africa, easterly wind anomalies along the equator, and a weak boreal fall season

231 Wyrтки Jet (Fig. 2). The IOD is a mode of coupled ocean-atmosphere variability that, like  
232 ENSO, develops via feedbacks between zonal wind stress, SST and thermocline depth  
233 anomalies. Unlike ENSO, it is shorter lived and confined mostly to the boreal fall  
234 season. Annual mean winds along the equator are westerly in Indian Ocean, tilting the  
235 thermocline down to the east. Thus, it is only during the normal September to November  
236 upwelling season off the coast of Java and Sumatra when the thermocline is brought close  
237 enough to the surface that ocean-atmosphere feedbacks can take hold. After that, the  
238 strong seasonality associated with the onset of the Northeast Monsoon overwhelms any  
239 SST anomalies in this upwelling region that may have developed during the previous  
240 summer and fall.

241         The most recent positive IOD events of significant amplitude occurred in 1994,  
242 1997, and 2006 (Fig. 2e). These events typically lead to above normal rainfall in East  
243 Africa, India, and Southeast Asia and dry conditions in Indonesia and Australia  
244 (Yamagata et al, 2004). Far field impacts of the IOD have also been reported on the  
245 seasonal climate of East Asia, Brazil, and Europe (*op. cit.*). Negative IOD events develop  
246 with anomalies and climatic impacts roughly opposite to those of positive events.

247         There is a tendency for positive IOD events to co-occur with El Niño and negative  
248 events with La Niña (Meyers et al, 2007). These co-occurrences have complicated the  
249 identification of IOD vis-à-vis ENSO climate impacts and have raised the question of  
250 whether the IOD is fundamentally tied to ENSO (e.g., Chang et al, 2006). However, co-  
251 occurrences with ENSO do not account for all IOD events and the dynamical ocean  
252 response to ENSO vs. purely IOD wind forcing is not identical (Yu et al, 2005). Thus,  
253 while ENSO may be an important triggering mechanism, the IOD appears to exist as an

254 independent mode of climate variability (Schott et al, 2008).

255           In addition to interannual variability, the Indian Ocean also experiences longer  
256 term decadal variations and trends. For example, there is a decadal modulation in the  
257 frequency of IOD events (Ashok et al, 2004), in the relationship between ENSO and the  
258 IOD (Schott et al, 2008), and in the relationship between ENSO and Indian summer  
259 monsoon rainfall (Krishna-Kumar et al, 1999). Decadal changes in Indian Ocean  
260 circulation have been documented, including a decrease in strength of the Indonesian  
261 Throughflow since 1976 (Wainwright et al., 2008) and an intensification around 2000 of  
262 the subtropical cell (Lee and McPhaden, 2008), which links upwelling in the Seychelles-  
263 Chagos thermocline ridge to source waters formed between 20°-30°S in the southeastern  
264 basin. Superimposed on these decadal variations are significant multi-decadal warming  
265 trends in SST (Cane et al, 1997) and increases in upper ocean heat content (Levitus et al,  
266 2005) that may plausibly be attributed to anthropogenic greenhouse gas forcing. The SST  
267 trend has been linked to drought in the African Sahel (Giannini et al., 2003) and in the  
268 northern hemisphere mid-latitudes (Hoerling and Kumar, 2003). Latent heat fluxes have  
269 increased steadily since the early 1980s over the Indian Ocean (Yu and Weller, 2007)  
270 suggesting that changes in surface fluxes are a response to, rather than the cause of, this  
271 SST trend. It is likely therefore that ocean dynamics play a role in producing the observed  
272 SST trend, although the precise mechanisms remain uncertain (Alory et al, 2007).

273           In summary, there is a broad spectrum of phenomena in the Indian Ocean, ranging  
274 from diurnal to decadal time scales, that contributes to the observed variability.  
275 Quantitative understanding of these phenomena, and how they interact with one another,  
276 is undermined though by a sparseness of data. Unlike in the Pacific Ocean where

277 systematic observations as part of the ENSO observing system were initiated in the early  
278 1980s, there are no high quality multi-decade in situ data records in the upper Indian  
279 Ocean except for those from a few ship-of-opportunity expendable bathythermograph  
280 (XBT) lines (Feng et al, 2001; Feng and Meyers, 2003). These limitations make it  
281 difficult to assess with confidence whether the Indian Ocean-atmosphere system is  
282 changing, or may change, as a result of greenhouse gas forcing (Harrison and Carson,  
283 2007).

284         Data limitations also constrain our ability to develop, initialize, and validate  
285 coupled ocean-atmosphere forecast models for monsoon prediction. Experimental  
286 forecasting with these models is in its infancy, and there are preliminary indications that  
287 skillful seasonal forecasts in the Indian Ocean region may be possible at 2-3 season lead  
288 times based on ENSO and IOD influences (Luo et al, 2007; Cherchi et al, 2007). Present  
289 levels of skill are limited by poor initialization of the subsurface ocean (Wajsowicz,  
290 2005), systematic errors in ocean and atmospheric models, and the general inability of  
291 either atmospheric general circulation models or coupled ocean-atmosphere models to  
292 accurately simulate intraseasonal variability like the MJO (Slingo et al, 1996; Zhang,  
293 2005). There is evidence to suggest that elements of this intraseasonal variability may be  
294 predictable up to 30 days in advance (Webster and Hoyos, 2004; Miura et al, 2007), so  
295 the fact that this variability is not well represented in dynamical prediction models  
296 represents a major challenge in both climate and weather forecasting.

### 297         3. RAMA as a contribution to IndoOS

#### 298             3.1 IndoOS

299           The International GOOS program and the Climate Variability and Predictability  
300 (CLIVAR) component of the World Climate Research Program (WRCP) established an  
301 Indian Ocean Panel (IOP) in 2004 to design, and guide the implementation of, a basin-  
302 scale, integrated Indian Ocean observing system for climate research and forecasting. The  
303 IOP focused on developing a strategy for *in situ* measurements to complement existing  
304 and planned satellite missions for surface winds, sea level, SST, rainfall, salinity, and  
305 ocean color. The resulting system, referred to as IndOOS (Fig. 3), is based on proven  
306 technologies, including moorings, Argo floats, ship-of-opportunity measurements,  
307 surface drifters, and tide gauge stations (CLIVAR-GOOS Indian Ocean Panel, 2006;  
308 Meyers and Boscolo, 2006). Transmission of data to shore in real-time via satellite relay,  
309 where feasible, was given high priority to promote use of the data in climate analysis and  
310 forecast products. Network design emphasized the measurement of physical climate  
311 variables but recognized that as the science matures and technology advances,  
312 widespread inclusion of biogeochemical measurements to support studies of the ocean  
313 carbon cycle and ecosystem dynamics would also be possible.

314           Embedded in this basin scale observing system are regionally focused national  
315 observing systems. These include Indian efforts in the Bay of Bengal and Arabian Sea,  
316 the Australian Integrated Marine Observing System (IMOS; Meyers and Proctor, 2008)  
317 and the Longterm Ocean Climate Observations (LOCO; Ridderinkhof and De Ruijter,  
318 2003) for western boundary currents in the Mozambique Channel. In addition, IndOOS  
319 provides a long-term, broad scale spatial and temporal context for short duration,  
320 geographically focused process studies, such as the Mirai Indian Ocean cruise for the  
321 Study of the MJO convection Onset (MISMO; Yoneyama et al, 2008) , the Validation of



322 the Aeroclipper System under Convective Occurrences (VASCO)-Cirene program  
323 (Duvel et al, 2008; Vialard et al, 2008a), and International Nusantara Stratification and  
324 Transport (INSTANT) program to study the Indonesian Throughflow (Gordon, 2005).

### 325 3.2 RAMA

326 A key element of IndOOS is the basin-scale moored buoy array, which we call  
327 RAMA (Fig. 4). Some of the Indian Ocean programs described in the Introduction  
328 collected 1-3 year long time series records either near the equator (Knox, 1976; Reppin et  
329 al, 1997) or in the Arabian Sea (Rudnick et al, 1997). Despite these noteworthy efforts  
330 though, there has been no plan until now for a coordinated, multi-national, basin-scale  
331 sustained mooring array like TAO/TRITON in the Pacific and PIRATA in the Atlantic.

332 RAMA addresses the need for such an array. It is designed specifically for  
333 studying large-scale ocean-atmosphere interactions, mixed-layer dynamics, and ocean  
334 circulation related to the monsoons on intraseasonal to decadal time scales. The planned  
335 array consists mainly of 38 surface moorings eight subsurface moorings (see box). Five  
336 of the eight subsurface moorings are Acoustic Doppler Current Profiler (ADCP)  
337 moorings to provide long time series measurements of currents in the upper 300-400 m.  
338 Four of these ADCP moorings are located along the equator where geostrophy breaks  
339 down and direct current measurements are necessary. A fifth ADCP mooring is also  
340 located in the upwelling zone off the coast of Java where the SST anomalies associated  
341 with the IOD first develop. This mooring is near the northern terminus of the frequently  
342 repeated Australia-to-Indonesia XBT line (Fig. 3) providing upper ocean temperature  
343 observations at weekly intervals. The primary focus of the array is the upper 500 m,  
344 where the ocean and atmosphere most immediately communicate with one another and

345 where intraseasonal-to-decadal time scale variability is most pronounced. However, in  
346 addition to the 38 surface and five ADCP moorings, three subsurface moorings along the  
347 equator at 77°E, 83°E, and 93°E are designed to monitor ocean currents down to 4000 m  
348 (Murty et al, 2006).

349 The array is intended to cover the major centers of ocean–atmosphere interaction  
350 in the open ocean away from western boundary current regions and the Indonesian  
351 marginal seas. These regions include the Arabian Sea and the Bay of Bengal; the  
352 equatorial waveguide where wind-forced intraseasonal and semi-annual current variations  
353 are prominent; the eastern and western poles of the IOD; the thermocline ridge between  
354 5-12°S, where wind-induced upwelling and Rossby waves affect SST; the southwestern  
355 tropical Indian Ocean, where ocean dynamics and air–sea interaction affect cyclone  
356 formation (Xie et al., 2002); and the southeastern basin where source waters of the cross  
357 equatorial and subtropical circulation cells are formed (Schott et al, 2004). Numerical  
358 model design studies have assessed the adequacy of the array to achieve its purposes in  
359 the context of other observing system components. Alternative sampling strategies have  
360 also been evaluated with the conclusion that the proposed array configuration is  
361 scientifically sound and cost-effective (Vecchi and Harrison, 2007; Oke and Schiller,  
362 2007).

363 Surface heat and moisture fluxes are important in determining mixed layer  
364 temperature and salinity variability. However, surface heat and moisture flux  
365 climatologies are poorly known in the Indian Ocean (Yu and McCreary, 2004;  
366 Yaremchuk, 2006; Yu et al, 2007) and in several regions annual means in net surface heat  
367 flux from presently available climatologies typically differ by 30-40 W m<sup>-2</sup> (Fig. 5).

368 Hence, in each of the key regions described above, plans call for at least one specially  
369 instrumented surface mooring as a surface flux reference site.

370 Surface mooring data are telemetered to shore in real-time from the via the  
371 Service Argos satellite relay system. Service Argos then places these data on the Global  
372 Telecommunications System (GTS) for transmission to operational weather, climate, and  
373 ocean forecasting centers. Internally recorded data from all moorings are post-processed  
374 and quality controlled on recovery, after which they are then posted on World Wide Web  
375 for public distribution (see box on RAMA Moorings). The RAMA data policy is based  
376 on the principle of free, open, and timely access to all data from the array.

#### 377 4. Progress Towards Implementation of RAMA

378 At the end of 2008, RAMA was 47% complete, with 22 of the 46 mooring sites  
379 occupied (Fig. 4). Nations that have provided mooring equipment, ship time, personnel,  
380 and/or logistic support so far include Japan, India, the United States, Indonesia, China,  
381 France and eight African countries that comprise the Agulhas and Somali Current Large  
382 Marine Ecosystems (ASCLME) Project. Contributing organizations and their year of  
383 initial involvement are described in the Implementation Time Line side box. Given  
384 current and projected national resource commitments, the array could be fully  
385 implemented by the end of 2012. The CLIVAR-GOOS Indian Ocean Panel and the  
386 CLIVAR Tropical Moored Buoy Implementation Panel provide scientific and technical  
387 guidance, respectively, for the development of RAMA.

388 The array will require a reliable, regular supply of ship time to fully implement  
389 since the surface moorings have a design lifetime of one year and must be replaced  
390 annually. Making reasonable assumptions about ship speeds, carrying capacity, and ports

391 of call around the Indian Ocean, We estimate that a minimum of approximately 150 days  
392 of dedicated ship time per year will be required to maintain the array once complete. For  
393 perspective, this requirement is about half that needed to maintain the 70 mooring  
394 TAO/TRITON array in the Pacific and about twice that needed to maintain the smaller 18  
395 mooring PIRATA array in the Atlantic. The actual number of sea days required per year  
396 to maintain the array may be higher than the absolute minimum though since research  
397 cruises sponsored by partner countries often address national scientific priorities in  
398 addition to those of the RAMA mission.

399 The greatest impediment to implementation, assuming adequate financial  
400 resources and ship time can be found, is vandalism by fishing vessels. Surface buoys are  
401 effectively Fish Aggregation Devices (FADs) that attract fish and, consequently,  
402 fishermen. Vandalism occurs primarily in pursuit of tuna and affects TAO/TRITON and  
403 PIRATA as well as RAMA. Strategies to mitigate vandalism include engineering design  
404 improvements to the moorings and outreach to the fishing community. Data losses can  
405 also be minimized by scheduling cruises to repair or replace moorings at yearly or more  
406 frequent intervals.

#### 407 5. Data and Research Applications

408 RAMA, even in the initial stages of development, is providing valuable data for  
409 describing and understanding variability in the Indian Ocean. For example, a pronounced  
410 semi-annual cycle in upper ocean temperature, salinity and zonal velocity is evident in  
411 the first 3-years of data from near equatorial moorings at 90°E (Fig. 6). Hase et al (2008)  
412 relate this variability to remote zonal wind forcing in the central Indian Ocean associated  
413 with the monsoon transitions. Upward phase propagation at semi-annual periods in both

414 temperature and zonal velocity is presumably the signature of wind forced equatorial  
415 Kelvin waves. This vertical propagation is consistent with, but more sharply defined than,  
416 that evident in equatorial time series collected during INDEX in the 1970s (McPhaden,  
417 1982; Luyten and Roemmich, 1982). The subsurface salinity maximum centered near  
418 100-150 m at 90°E (Fig. 6b) is due to eastward transport along the equator of high  
419 salinity water originating in the southern hemisphere subtropics and the Arabian Sea  
420 (Taft and Knauss, 1967; Schott et al, 2004). Semi-annual increases in maximum salinity  
421 values in this subsurface layer are most likely the response to increases in eastward  
422 velocity associated with the semi-annual Wyrтки Jets.

423         Superimposed on these semi-annual variations are energetic 30-50 day period  
424 oscillations, which presumably reflect the effects of wind (e.g., Fig. 7a), heat flux, and  
425 fresh water forcing on intraseasonal scales. Intraseasonal oscillations are largest for  
426 temperature in the thermocline (Fig. 6a) and largest for salinity and zonal velocity in the  
427 surface mixed layer (Figs. 6b and 7b). Meridional velocity is most strongly influenced on  
428 intraseasonal time scales by 10-20 day period oscillations (Fig. 7c) which are evident not  
429 only in the upper 400 m but also at depths greater than 2000 m from the deep ocean  
430 moorings along the equator. Sengupta et al. (2004) identified these oscillations as wind  
431 forced mixed Rossby-gravity waves.

432         Interannual variability associated with the 2006 IOD event was captured by the  
433 moored array as illustrated in the time series from 0°, 80.5°E for two contrasting periods:  
434 October-November 2004 and October-November 2006 (Fig. 8). Compared to late 2004,  
435 which was near normal in terms of IOD activity (Fig. 2e), zonal surface winds and the  
436 zonal mixed layer currents they forced along the equator largely flowed in the opposite

437 direction, i.e. to the west, in late 2006 (Fig. 8a, b). The westward currents drained the  
438 eastern basin of upper ocean mass, which may have contributed to periods of thermocline  
439 shoaling at 0°, 80.5°E (Fig. 8c). Also, though SST was only slightly warmer than usual at  
440 0°, 80.5°E ( $\leq 1^\circ\text{C}$ ) as would be expected for a mooring located outside the IOD SST  
441 index regions, these elevated temperatures favored enhanced convection and rainfall in  
442 the central basin (Fig. 2). The dramatic drop in mixed layer salinity of over 1 psu in late  
443 2006 relative to late 2004 at 0°, 80.5°E (Fig. 8d) is consistent with this enhanced rainfall,  
444 as well as with equatorward flow of low salinity water from the Andaman Sea (Murty et  
445 al, 2008). It is also interesting to note that the meridional component of mixed layer  
446 velocity in Fig. 8b exhibits very regular biweekly oscillations as observed in the 0°, 90°E  
447 ADCP data (Fig. 7c) and in the deep ocean moorings along the equator (Sengupta et al,  
448 2004).

449 Data from the mooring at 8°S, 67°E in the region of the Seychelles-Chagos  
450 thermocline ridge illustrate ocean-atmosphere interactions associated with the passage of  
451 tropical cyclone Dora (Duvel et al, 2008; Vialard et al, 2008a). Dora began as a tropical  
452 disturbance northeast of the buoy in the vicinity of Diego Garcia (7°S, 72°E) in late  
453 January 2007. As this disturbance migrated in south-southwesterly direction, it intensified  
454 to named tropical storm strength on January 30 and to tropical cyclone strength on  
455 February 1. Dora reached peak intensity of over 100 kts ( $51 \text{ m s}^{-1}$ ) on February 3 when it  
456 was centered near 18°S, 67°E. Afterwards, it continued to track to the southwest,  
457 eventually dissipating by February 12.

458 Dora passed near the buoy in the early stages of its development (January 25-29)  
459 before it became a named tropical storm. Its effects in the real-time daily averaged data

460 can be seen in the sudden change in wind speed and direction, increase in precipitation,  
461 decrease in atmospheric pressure, and decrease in incoming shortwave radiation at the  
462 buoy site in late January 2007 (Fig. 9, left panel). Also evident in late January are abrupt  
463 cooling and freshening of the surface mixed layer associated with decreased surface heat  
464 fluxes and increased fresh water fluxes. Internally recorded 1- to 10-minute data  
465 recovered from the buoy show even greater detail during the last week of January 2007  
466 (Fig. 9, right panel). One sees for example a very sharp drop in surface salinity and  
467 temperature associated with a nighttime rain event on January 24. Also, SST dropped in a  
468 very stepwise fashion by 2°C over the week. Periods of very low insolation on rainy (and  
469 presumably cloudy) days are evident. There is also a pronounced solar semi-diurnal tide  
470 in atmospheric surface pressure. These data are being analyzed to quantitatively address  
471 the relative roles of surface fluxes, vertical turbulent mixing, and horizontal advection in  
472 the heat and salt balances for this particular period.

473 RAMA mooring data can also be used to identify deficiencies in currently  
474 available surface flux products as a stimulus to their possible improvement. For example,  
475 surface heat fluxes were calculated from data at three of RAMA mooring sites for  
476 comparison with numerical weather prediction (NWP) model flux products, the  
477 International Satellite Cloud Climatology Project (ISCCP) radiation product (Zhang et al,  
478 2004) and an objectively analyzed (OA) turbulent heat flux product (OA-Flux; Yu and  
479 Weller, 2007). Results (Fig. 10) indicate that that ISCCP overestimates solar radiation  
480 and underestimates longwave radiation at these sites. The NWP products significantly  
481 overestimate latent heat fluxes such that the net heat flux into the ocean is underestimated  
482 by 40-60 W m<sup>-2</sup> at 0°, 80.5°E. This heat flux deficit, if accumulated in a 50 m thick mixed

483 layer over three months, would translate into a temperature error of  $\sim 2^{\circ}\text{C}$ , which is  
484 equivalent to the seasonal range of SST at this location. The OA-Flux product slightly  
485 underestimates latent heat fluxes at all three locations, but is a significant improvement  
486 on the NWP turbulent fluxes.

## 487 6. Summary and Discussion

488 RAMA addresses a long-standing need for a sustained moored buoy array in the  
489 Indian Ocean for climate studies. It will take several more years to complete and will  
490 require coordinated resource contributions (financial, human, and ship time) from several  
491 countries. Implementation of RAMA will result in a globe-girdling network of tropical  
492 moored buoy arrays that includes TAO/TRITON in the Pacific and PIRATA in the  
493 Atlantic. Like these other arrays, we can expect that RAMA will fundamentally advance  
494 our understanding of large-scale ocean dynamics, ocean-atmosphere interactions, and  
495 climate variability in the Indian Ocean region.

496 We can also expect that RAMA will contribute to operational activities as a  
497 component of GOOS, the Global Climate Observing System (GCOS), and the Global  
498 Earth Observing System of Systems (GEOSS). Real-time RAMA data available on the  
499 GTS are, for example, already being incorporated into weather forecasts and seasonal  
500 climate forecasts produced by operational centers. RAMA can likewise provide valuable  
501 data for operational ocean state estimation and for oceanic and atmospheric reanalysis  
502 products. The array will provide data for validating satellite retrievals and products based  
503 on them such as for surface winds, SST, rainfall, and mixed layer currents. Finally,  
504 RAMA data will be valuable for quantitatively assessing the performance of oceanic and  
505 atmospheric dynamical models.



506 RAMA moorings, and in particular those specially instrumented reference sites  
507 for air-sea heat, moisture and momentum fluxes, are a contribution to the Ocean  
508 Sustained Interdisciplinary Timeseries Environment observation System (OceanSITES),  
509 a worldwide network of deep water stations providing high temporal resolution data for  
510 ocean research and environmental forecasting (<http://www.oceansites.org/>). In addition,  
511 RAMA is capable of accommodating biogeochemical sensors to support programs such  
512 as the International Ocean Carbon Coordination Project (IOCCP; <http://www.ioccp.org/>)  
513 and the Sustained Indian Ocean Biogeochemical and Ecological Research (SIBER)  
514 program (Hood et al, 2008). RAMA will provide information on oceanic variability for  
515 the Joint Aerosol-Monsoon Experiment (JAMEX; Lau et al, 2008), which is a multi-  
516 national study scheduled for 2007-11 to investigate the effects of aerosols on ocean-  
517 atmosphere-land interactions that govern the Asian monsoon water cycle. In the wake of  
518 the December 2004 Asian tsunami, discussions are also underway with organizations  
519 involved in developing the Indian Ocean tsunami warning system on how best to  
520 coordinate implementation efforts with IndOOS. A joint RAMA and tsunami mooring  
521 cruise has already been conducted aboard the Indonesian Research Vessel Baruna Jaya III  
522 in September 2007 (<http://www.noaanews.noaa.gov/stories2007/s2919.htm>). In the  
523 longer term, it may be beneficial to consider development of a multi-hazard moored buoy  
524 platform for both tsunami warnings and climate studies.

525 RAMA mooring stations provide convenient focal points around which to conduct  
526 short term process studies such as MISMO, which took place in October-December 2006  
527 centered at 0°, 80.5°E, and VASCO-Cirene which took place in January-February 2007  
528 near the mooring at 8°S, 67°E in the Seychelles-Chagos thermocline ridge region. These

529 process studies benefited from RAMA because they examined ocean-atmosphere  
530 interactions associated with the MJO for which high resolution moored time series data  
531 are especially valuable (e.g. Vialard et al, 2008b). RAMA in turn can benefit from  
532 process studies, since new knowledge gained from programs like MISMO and VASCO-  
533 Cirene can feedback into design specifications for the moored array. Moreover, the  
534 VASCO-Cirene field campaign provided an opportunity to deploy the flux reference site  
535 mooring at 8°S, 67°E.

536         Neither RAMA nor IndOOS will address all the observational needs for  
537 understanding and predicting climate variability in the region since we still lack a  
538 sustained, in situ observing system for the atmosphere over the Indian Ocean. This lack  
539 of atmospheric data limits our understanding of convective processes, atmospheric  
540 circulation, and interactions of the atmosphere with the ocean. Regular cruises to service  
541 RAMA moorings can help mitigate this problem though by providing platforms of  
542 opportunity for routine meteorological observations.

543         In summary, full implementation of RAMA promises significant scientific and  
544 societal benefits. Challenges that we must overcome include securing the necessary  
545 resources to complete the array and forging long-lasting multi-national partnerships to  
546 sustain it. We know that success is possible though because it has been achieved in the  
547 past for similar mooring programs of ambitious scope in the tropical Pacific and Atlantic.  
548

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964 SIDE BOX 1. RAMA Helps Australian Farmers

965 “No summer rain, then enough to sow in April/May on the arable country, but the  
966 plains still missed out. This meant feeding sheep all through autumn and a very low  
967 lambing percentage. Then...” Susan Carn is recounting the interplay between weather  
968 patterns and management of her farm the past season. She and her husband Ben raise  
969 sheep in the low rainfall area of the Flinders Ranges in South Australia. “So looking back  
970 ... I'm really glad I stuck to my guns and told my husband to stop sowing as things were  
971 not adding up for a good season.”

972 Susan is also the science liaison for the BestPrac farming group in the region. She  
973 knows a lot about climate and is always on the look-out for new streams of information to  
974 support the decisions they must make throughout the year. The Indian Ocean Dipole  
975 affects rainfall in their region and Susan was using RAMA data provided by the Indian  
976 Ocean Panel as an indicator of evolving Dipole conditions. The Panel was happy to hear  
977 this report: “Thank you for sending me the TRITON info. Back in July I showed it to my  
978 BestPrac group who all thought it scary but very useful!”

979 Modern farmers have many sources of information available to them for  
980 improving their bottom lines. Climate information is just one of the sources, but farmers  
981 and climate researchers are forging new links to exploit that information. The BestPrac  
982 Group meets regularly to share experiences that can enhance their farm management  
983 skills. An important part of each meeting is to review the latest seasonal forecasts from  
984 the Australian Bureau of Meteorology along with climate indices such as the Southern  
985 Oscillation Index and the Indian Ocean Dipole Index. Data from sources like RAMA  
986 moorings are also factored into the deliberations. This information helps the farmers to

987 make informed decisions on a range of practices such as cropping programs, fertilizer and  
988 spray applications, and stocking rates (i.e., the number of animals to keep on the  
989 property). The group can then better cope with climate related risks and capitalize on  
990 opportunities. Their motto is "Hope is not a plan!"

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A BestPrac meeting with members immersed in study. Susan and Ben are on the left.  
Photo credit: John Squires.



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BestPrac members fat-scoring sheep to determine if they need supplementary feeding.  
Photo credit: John Squires.



1001 SIDE BOX 2: RAMA moorings

1002 Four types of moorings are presently used in RAMA: surface moorings, surface  
1003 moorings with enhanced measurement capabilities for comprehensive air-sea fluxes ( i.e.,  
1004 “flux reference site” moorings), ADCP moorings, and deep-ocean moorings. Each  
1005 mooring type is described briefly below.

1006 The surface moorings consist of both Autonomous Temperature Line Acquisition  
1007 System (ATLAS) moorings and two different types of Triangle Trans Ocean Buoy  
1008 Network (TRITON) moorings. These moorings have a design lifetime of one year and so  
1009 must be serviced annually. Most of the surface moorings are ATLAS moorings supplied  
1010 by NOAA’s Pacific Marine Environmental Laboratory (PMEL). These taut line surface  
1011 moorings are anchored to the ocean floor in depths of typically 2500-5000 m.

1012 Measurements on the surface float include air temperature, relative humidity, wind  
1013 velocity, downwelling shortwave radiation, and rain rate at heights of 3-4 m above mean  
1014 sea level. Sea surface temperature and conductivity (which together yield salinity) are  
1015 measured from the buoy at a nominal depth of 1 m. Sensors on the mooring line measure  
1016 ocean temperature (12 depths between 10 m and 500 m), conductivity (5 depths between  
1017 10 m and 100 m), mixed layer velocity (at 10 m depth) and pressure (at two depths). A  
1018 detailed description of the individual sensors, including their heights and depths relative  
1019 to mean sea level, is contained in the electronic supplement. Daily averages of all data  
1020 and several hourly samples per day of most meteorological variables are transmitted to  
1021 shore in real-time via Service Argos. These data are placed on the Global  
1022 Telecommunications System (GTS) for use in operational weather, climate, and ocean  
1023 forecasting. Higher temporal resolution data (at 1- to 10-minute intervals in most cases)

1024 are internally recorded and available after mooring recovery.

1025           Some ATLAS moorings are enhanced with additional instrumentation to more  
1026 precisely define surface heat, moisture and momentum fluxes. These flux reference site  
1027 moorings include sensors for downwelling longwave radiation and barometric pressure.  
1028 Additional sensors may also be deployed in the upper 140 m of the ocean for finer  
1029 vertical resolution measurements of temperature, salinity and velocity. Additional  
1030 information on ATLAS moorings can be found at  
1031 [http://www.pmel.noaa.gov/tao/proj\\_over/mooring.shtml](http://www.pmel.noaa.gov/tao/proj_over/mooring.shtml) .

1032           The Japan Marine-Earth Science and Technology Agency (JAMSTEC) has  
1033 deployed TRITON moorings and, since February 2008, mini-TRITON (m-TRITON)  
1034 moorings at 1.5°S, 90°E and 5°S, 95°E. TRITON moorings are taut-line moorings  
1035 designed to be functionally equivalent to ATLAS moorings in terms of sensor payloads,  
1036 temporal resolution, and data accuracy. There are some differences (e.g. TRITON  
1037 moorings are deployed with more salinity sensors and measurements are made to 750 m  
1038 rather than 500 m), but these differences do not affect comparability of the basic data  
1039 sets.

1040           After one-year inter-comparison with the TRITON mooring at 1.5°S, 90°E,  
1041 JAMSTEC began to deploy m-TRITON moorings at 1.5°S, 90°E and 5°S, 95°E in  
1042 February 2008. The m-TRITON sensor and signal processing system are based on those  
1043 for TRITON, but the mooring system itself is slack line rather than taut-line. Slack line  
1044 moorings allow for substantial vertical excursions of subsurface sensors due to variable  
1045 currents and winds. Thus, all m-TRITON subsurface sensors are equipped to measure  
1046 pressure so that the shape of the mooring line can be determined, after which temperature

1047 and salinity data can be interpolated to standard depths.

1048           Hourly averages of all TRITON and m-TRITON data are transmitted to shore in  
1049 real-time via Service Argos and placed on the GTS. Higher temporal resolution data (at  
1050 1- to 10-minute intervals in most cases) are internally recorded and available after  
1051 mooring recovery. TRITON and m-TRITON sensor specifications, including heights and  
1052 depths relative to mean sea level, are described in the electronic supplement. More  
1053 information on TRITON and m-TRITON moorings can be found at  
1054 [http://www.jamstec.go.jp/jamstec/TRITON/real\\_time/overview.php/po.php](http://www.jamstec.go.jp/jamstec/TRITON/real_time/overview.php/po.php) and  
1055 <http://www.jamstec.go.jp/iorgc/iomics/index.html>, respectively.

1056           Acoustic Doppler Current Profiler (ADCP) moorings are deployed at several  
1057 locations in the array. The ADCP is positioned with its acoustic beams pointing upward  
1058 in a subsurface float typically located at depths 300-400 m below the surface. Velocity  
1059 profiles are measured at hourly intervals with 8 m vertical resolution. Backscatter from  
1060 the ocean surface interferes with velocity retrievals in the upper 30-40 m of the water  
1061 column, so subsurface ADCPs do not measure effectively in this depth range. Hence the  
1062 need for current meters in the mixed layer on surface moorings. ADCP data are available  
1063 upon mooring recovery. More information on ADCP specifications can be found in the  
1064 electronic supplement.

1065           Deep ocean moorings maintained by India's National Institute of Oceanography  
1066 (NIO) have a subsurface float nominally at 100 m depth, below which 6 mechanical  
1067 current meters are attached to the mooring line for velocity measurements down to a  
1068 depth of approximately 4000 m. Beginning in 2003, upward-pointing ADCPs similar to  
1069 those used by PMEL and JAMSTEC were added to the near surface float for additional

1070 velocity measurements. Data at 1-2 hourly intervals are available from these moorings on  
1071 recovery. More technical information on these moorings can be found in the electronic  
1072 supplement and in Murty et al (2006).

1073 A web portal with pointers to data from all moorings is found at  
1074 [http://www.incois.gov.in/Incois/iogoos/home\\_indoos.jsp](http://www.incois.gov.in/Incois/iogoos/home_indoos.jsp). Subsets of the data are  
1075 available from <http://www.pmel.noaa.gov/tao/>,  
1076 [http://www.jamstec.go.jp/jamstec/TRITON/real\\_time/php/top.php](http://www.jamstec.go.jp/jamstec/TRITON/real_time/php/top.php), and  
1077 [http://www.nio.org/data\\_info/deep-sea\\_mooring/oos-deep-sea-currentmeter-](http://www.nio.org/data_info/deep-sea_mooring/oos-deep-sea-currentmeter-moorings.htm)  
1078 [moorings.htm](http://www.nio.org/data_info/deep-sea_mooring/oos-deep-sea-currentmeter-moorings.htm).  
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1079 SIDE BOX 3: RAMA Implementation Time Line

1080           RAMA traces its roots to Indian and Japanese national efforts initiated in 2000.  
1081 JAMSTEC deployed an ADCP mooring at 0°, 90°E in 2000 (Masumoto et al, 2005) and  
1082 two TRITON moorings at 1.5°S, 90°E and 5°S, 95°E in 2001 (Hase et al, 2008). NIO  
1083 also began subsurface mooring deployments to sample the deep ocean along the equator  
1084 in 2000 (Sengupta et al, 2004; Murty et al, 2006). These efforts, which have continued  
1085 without interruption, were precursors to RAMA and incorporated into the array design.

1086           Then, in October-November 2004, NOAA/PMEL in collaboration with NIO and  
1087 the Indian Ministry of Earth Sciences (MoES) deployed four ATLAS moorings and one  
1088 ADCP mooring near the equator between 80°-90°E. PMEL and the Indonesian Agency  
1089 for the Assessment and Application of Technology (BPPT) and the Ministry for Marine  
1090 Affairs and Fisheries (DKP) occupied sites at 4°N, 90°E and 8°N, 89°E in November  
1091 2006. The French-lead VASCO-Cirene experiment (Duvel et al, 2008; Vialard et al,  
1092 2008a) provided an opportunity in January 2007 to establish the flux reference site  
1093 ATLAS mooring at 8°S, 67°E. In November 2007, a Chinese ADCP mooring was  
1094 deployed off the coast of Java as part of a collaboration between the Chinese First  
1095 Institute of Oceanography (FIO), BPPT, and DKP. Also in November 2007, PMEL and  
1096 Indian technicians deployed two ATLAS moorings in the Bay of Bengal (12°N and  
1097 15°N, 90°E) on a cruise lead by the Indian National Center for Ocean Information  
1098 Services (INCOIS). In November 2008, the Agulhas and Somali Current Large Marine  
1099 Ecosystems (ASCLME) Project, a consortium of eight African countries, joined the  
1100 RAMA fold by deploying two ATLAS moorings along 55°E (8°S and 12°S) from the

1101 Norwegian RV Dr Fridtjof Nansen. The ASCLME includes Kenya, Tanzania,  
1102 Mozambique, South Africa, Madagascar, Mauritius, Seychelles, and Comoros.  
1103           Efforts to build and sustain the array will continue within a framework of formal  
1104 bilateral agreements that are either approved or under development between agencies in  
1105 the various partner countries. Our expectation based on current and projected resource  
1106 commitments is that RAMA will be completed by the end of 2012. The precise  
1107 implementation time line cannot be determined in detail at this time though since not all  
1108 formal agreements are finalized yet.

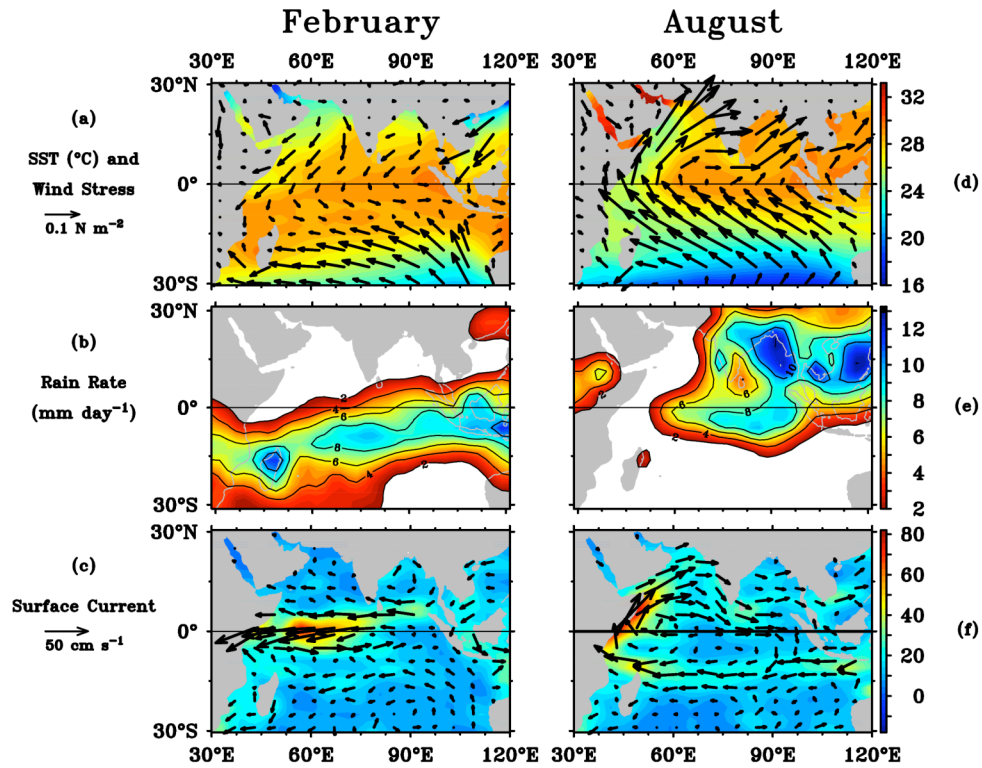


Figure 1. Monthly means for February of (a) surface wind stress and SST; (b) rain rate; and (c) surface current velocity (vectors) and scalar speed (color shading). Corresponding monthly means for August are in (d)-(f). Rain rate is contoured only for values  $\geq 2$  mm day<sup>-1</sup>. SSTs are based on Reynolds et al. (2002), winds are from the ECMWF-ERA40 reanalysis dataset ([http://data.ecmwf.int/data/d/era40\\_daily/](http://data.ecmwf.int/data/d/era40_daily/)), currents are from Cutler and Swallow (1984), and rain rate is from an analysis of satellite and in situ raingauge data (Janowiak and Xie, 1999).

## November 2006 Anomalies

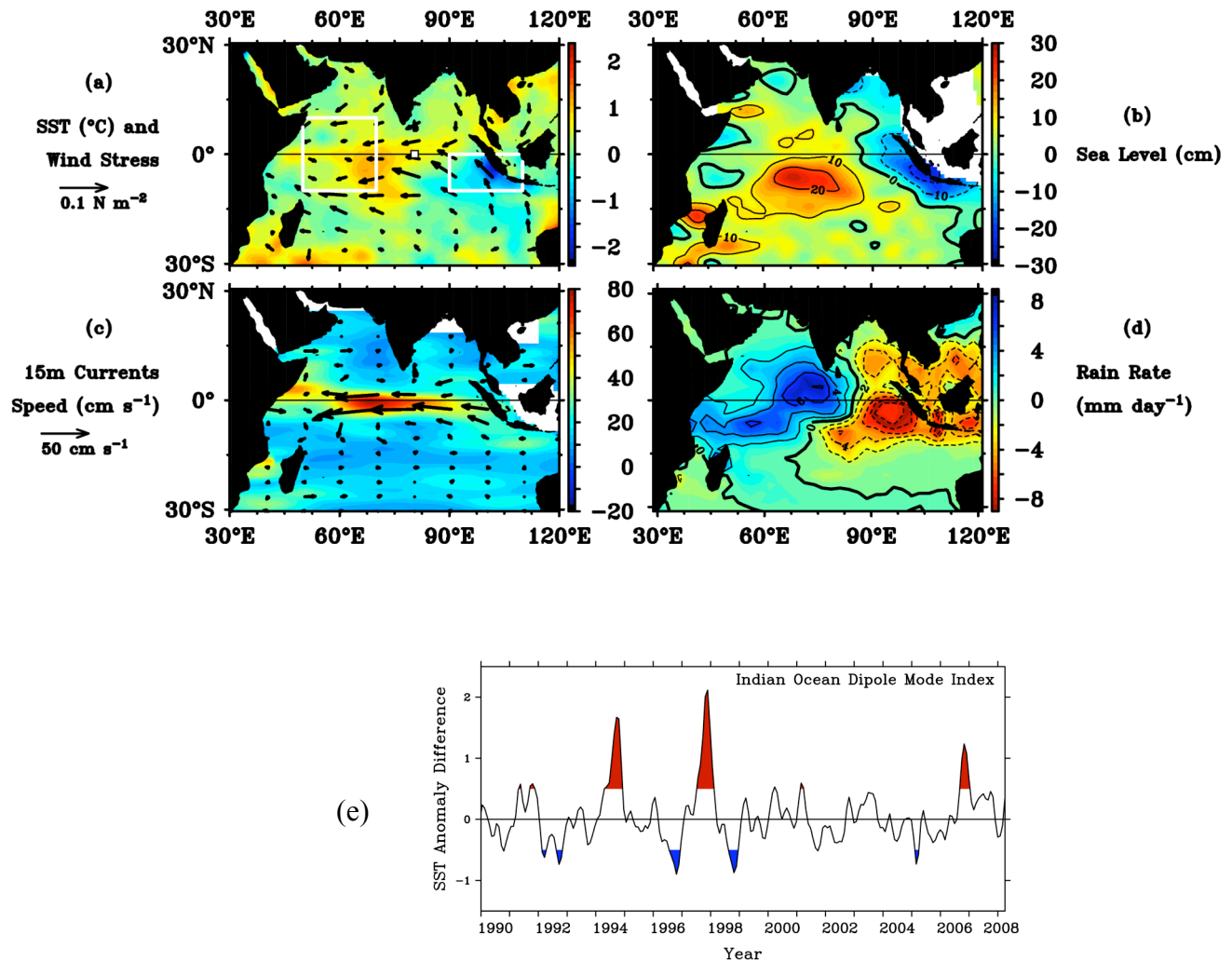


Figure 2. Monthly anomalies from the mean seasonal cycle in November 2006 for (a) SST (based on Reynolds et al., 2002) and QuikSCAT wind stress (<http://www.ifremer.fr/cersat/>); (b) Jason satellite altimeter sea level anomalies; (c) current velocity (vectors) and speeds (color shading) representative of flow at 15 m depth in the surface mixed layer (Bonjean and Lagerloef, 2002); (d) rain rate (Janowiak and Xie, 1999). Shown in (e) is the IOD index for 1990-2008, with the strongest positive events (index > 0.5°C) and negative events (index is < -0.5°C) highlighted in red and blue respectively. The index represents the difference between SST anomalies in the western minus the eastern basin regions outlined in (a). Also shown in (a) is the location of an ATLAS mooring at 0°, 80.5°E, time series of which are shown in Figure 8.



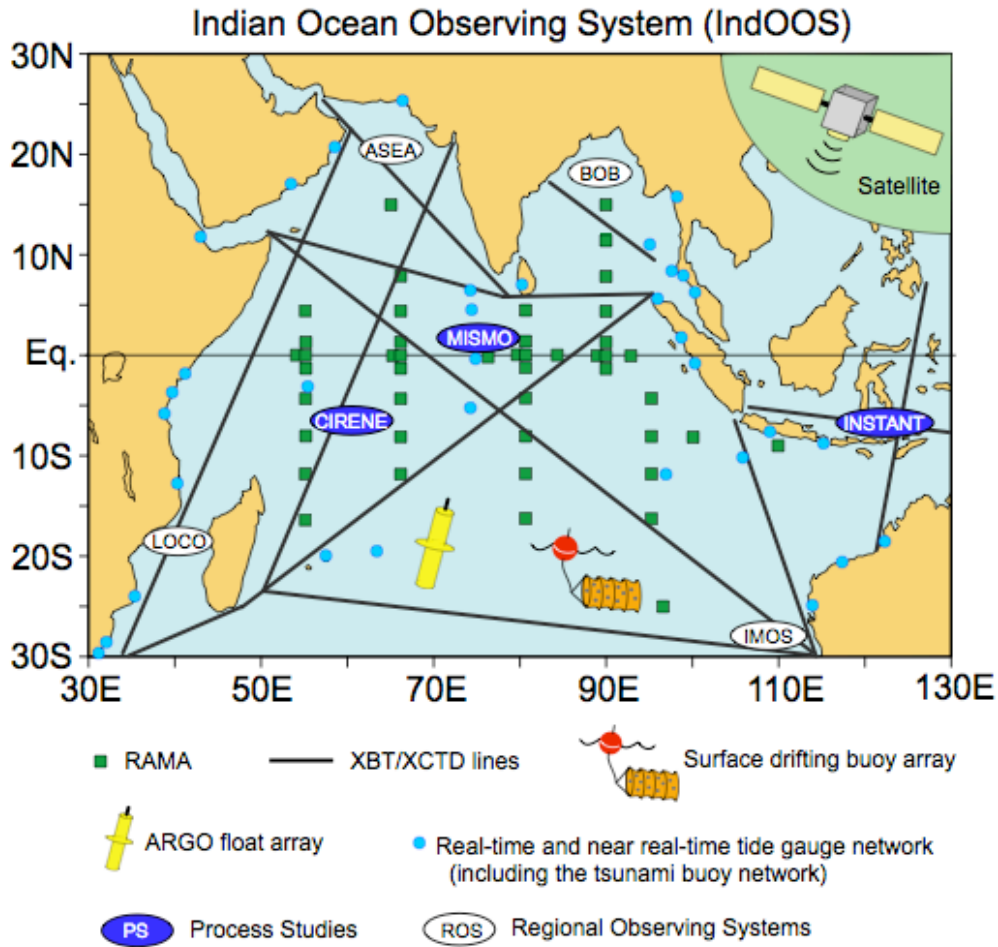


Figure 3. Schematic of the Indian Ocean Observing System (IndOOS). Green squares indicate the locations of RAMA moorings. Tide gauges are indicated by blue dots. Argo floats and surface drifters are indicated by a single symbol although many of each are spread throughout the basin (141 drifters and 296 Argo floats as of 31 August 2008). Expendable bathythermograph (XBT) and expendable conductivity/temperature/depth (XCTD) sections sampled by ships-of-opportunity are shown as black lines. Most of these lines are sampled 12-18 times per year at along-track intervals of approximately  $1^\circ$  (though the Australia-Sumatra line and the Australia-Mauritius-South Africa are sampled more frequently to measure details of ocean circulation). Nationally sponsored regional observing systems (ROS) are shown in ovals: IMOS, LOCO, Arabian Sea (ASEA), and Bay of Bengal (BOB). Process studies (PS) are shown in blue ovals: MISMO, VASCO-Cirene and INSTANT. The satellite in the upper right symbolizes the constellation of Earth-observing satellites for SST, surface winds, sea level, and other important oceanic and atmospheric parameters.

**Research Moored Array for African–Asian–Australian  
Monsoon Analysis and Prediction (*RAMA*)**

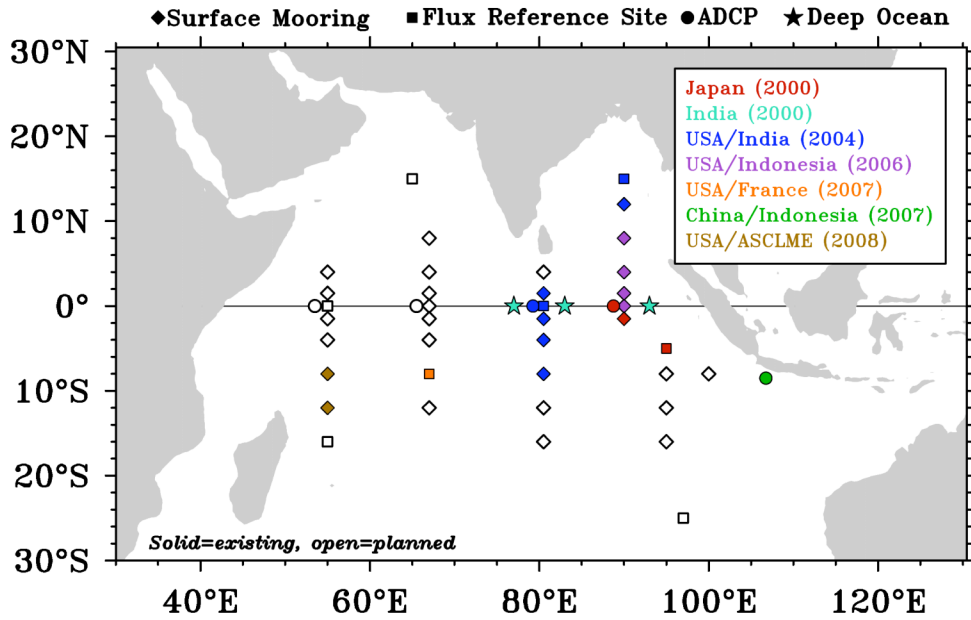


Figure 4. Schematic of RAMA as of December 2008. Solid symbols indicate those sites occupied so far. Color coding indicates national support, with year of first involvement shown in the upper right box. Open symbols indicate sites that are not yet instrumented. ASCLME is a consortium of eight African nations including Kenya, Tanzania, Mozambique, South Africa, Madagascar, Mauritius, Seychelles, and Comoros.

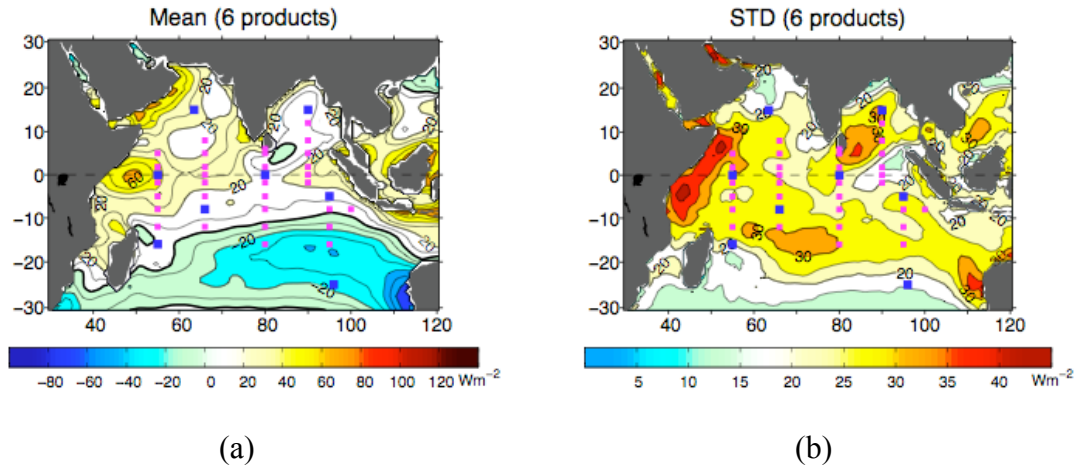


Figure 5. The (a) average and (b) standard deviation of long term mean net heat fluxes from six different products. The RAMA array is superimposed on these fields with flux reference sites highlighted as blue squares. The six products are: the National Centers for Environmental Prediction NCEP1 (Kalnay et al, 1996) and NCEP2 (Kanamitsu et al, 2002) fluxes; European Center for Medium Range Forecasting (ECMWF) operational fluxes (ECMWF, 1994); ECMWF reanalysis fluxes (ERA40; Uppala et al, 2005); the U.K. National Oceanography Center (NOC) fluxes (Josey et al, 1999); and an objectively analyzed turbulent surface heat flux product produced by Woods Hole Oceanographic Institution (OA-Flux; Yu and Weller, 2007) combined with radiation data from the International Satellite Cloud Climatology Project (ISSCP; Zhang et al, 2004). Long term means for NCEP1, NCEP2, OAFlux+ISSCP, and ECMWF are based on the period 1983-2004; ERA40 means were based on the period 1983-2001 and the NOC climatology was based the period on 1980-2005.

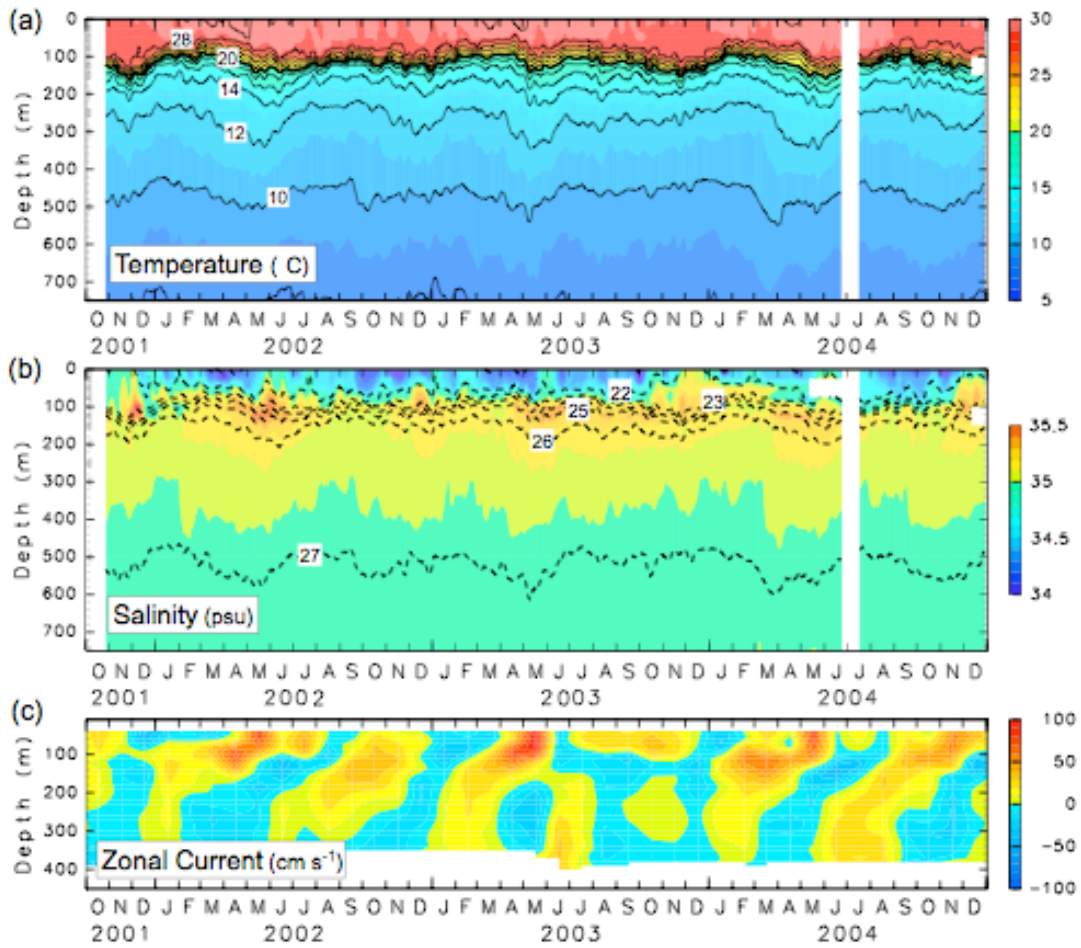


Figure 6. Time/depth sections of (a) temperature and (b) salinity from 1.5°S, 90°E and (c) zonal velocity from 0°, 90°E. Daily data in (a) and (b) have been smoothed with a 7-day running mean filter and in (c) with a monthly filter. Color intervals for the temperature, salinity and velocity are 1°C, 0.1, and 10 cm s<sup>-1</sup>, respectively. Thin black lines are at intervals of 2°C for temperature with 20°C isotherm shown as a thick line. Contours of the potential density at intervals of 1.0 kg m<sup>-3</sup> are superimposed on salinity sections (after Hase et al, 2008).

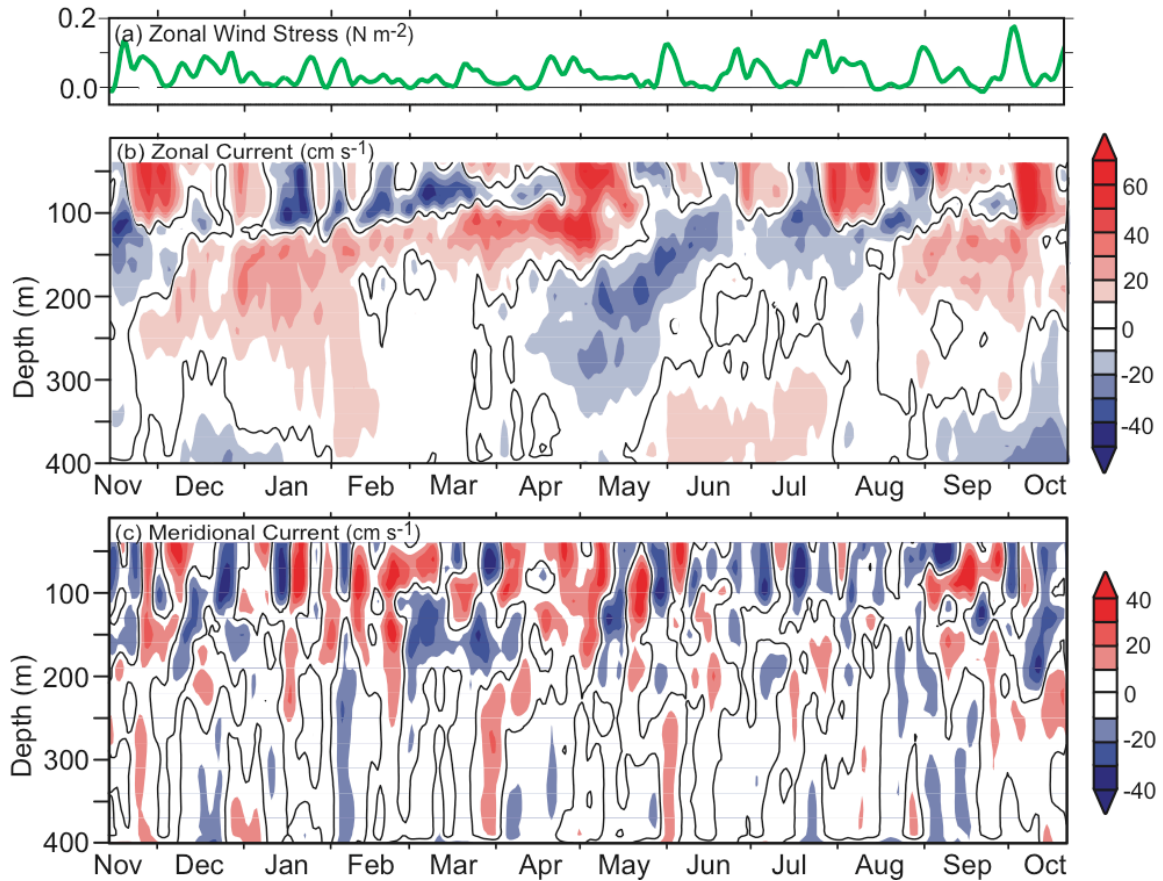


Figure 7. Time series of (a) zonal wind stress (in  $\text{N m}^{-2}$ ) at the sea surface averaged between  $80^{\circ}\text{E}$  and  $90^{\circ}\text{E}$  observed by the QuickSCAT satellite. Time-depth sections of (b) zonal current and (c) the meridional current observed at  $0^{\circ}$ ,  $90^{\circ}\text{E}$ . The eastward (westward) and northward (southward) currents are shaded in reds (blues), with black contours for zero velocity. Daily data have been low pass filtered with a 5-day period cut off (after Masumoto et al, 2005).

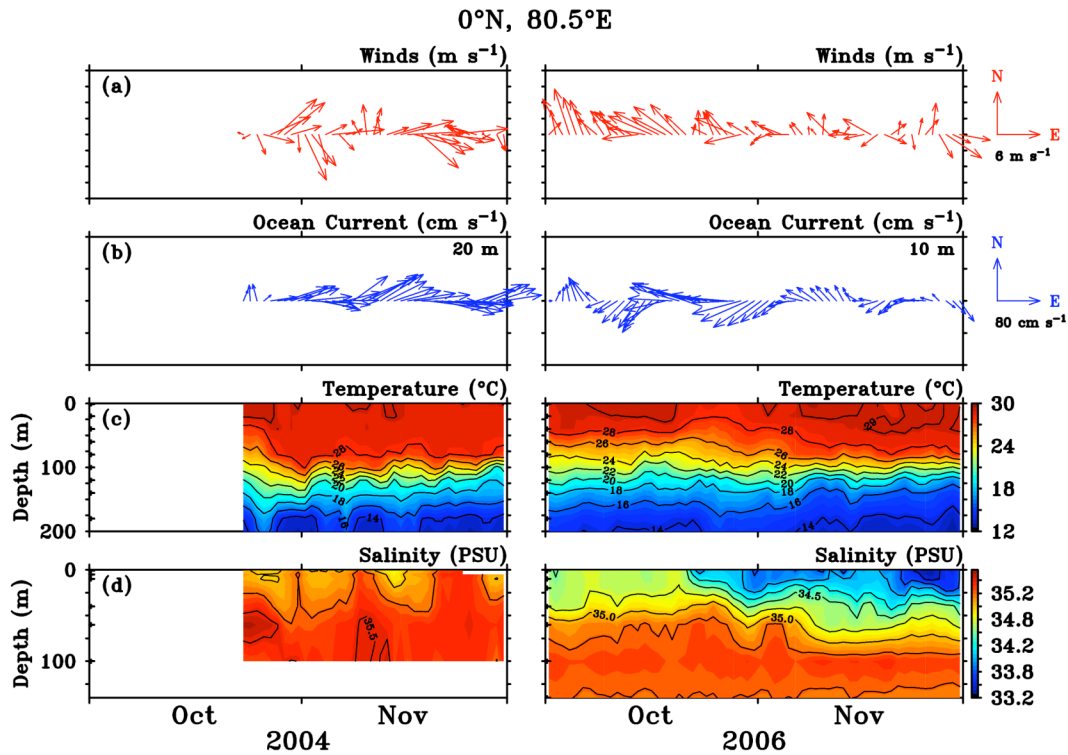


Figure 8. Daily averaged data at 0°, 80.5°E for October-November 2004 (a neutral IOD year) and October-November 2006 (a positive IOD year, as illustrated in Fig. 2). Shown are (a) wind vectors, (b) velocity vectors in the mixed layer (20 m depth in 2004 and 10 m depth in 2006), (c) temperature and (d) salinity. Start of the time series in mid-October 2004 is coincident with the first deployment at this site.

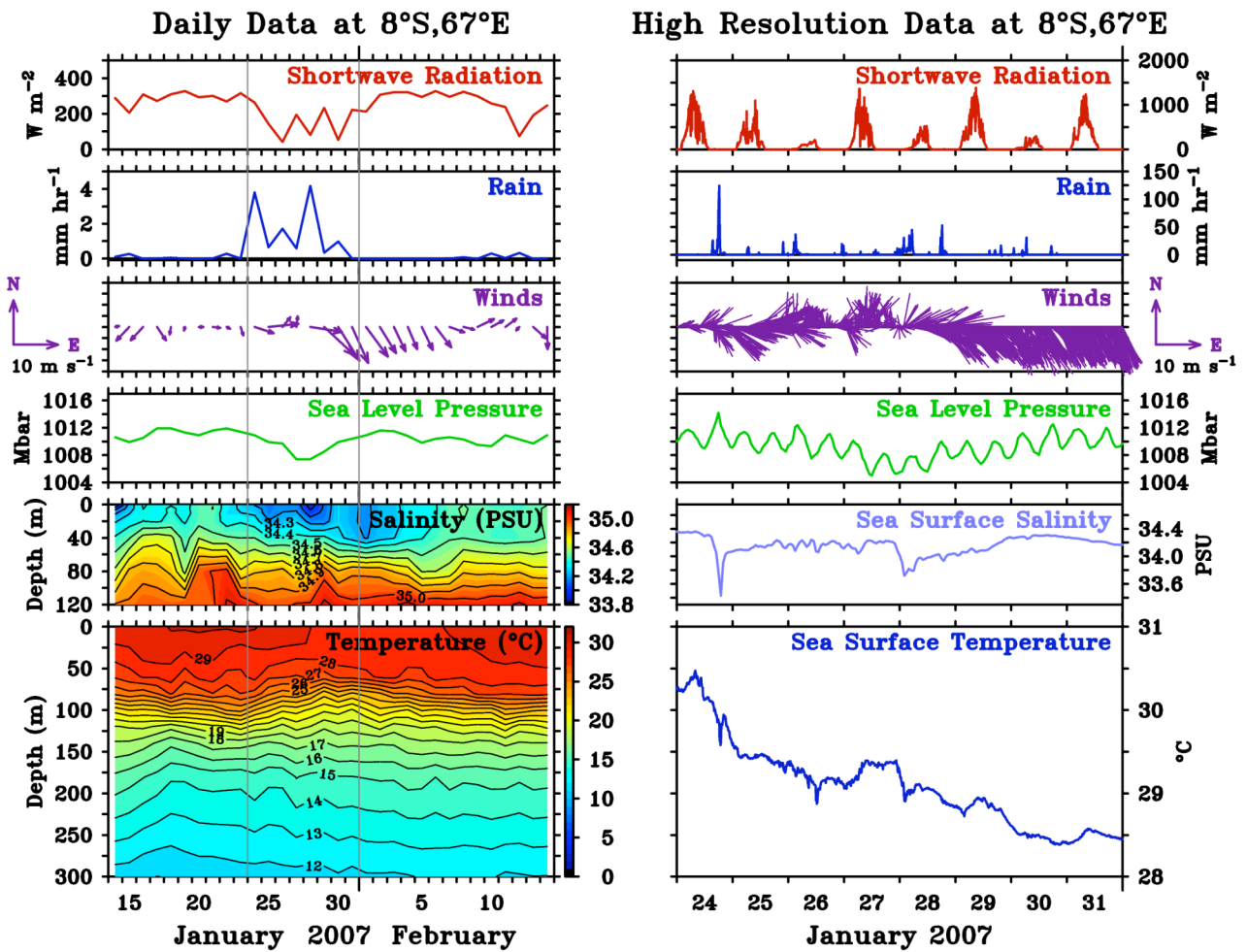


Figure 9. Time series data from 8°S, 67°E in the Seychelles-Chagos thermocline ridge region during January-February 2007. Left panel shows daily averaged real-time data whereas the right panel shows hourly averages based internally recorded data for selected variables during the week of January 24-31. The period highlighted on the right (delineated in the left panel by light grey lines) coincides with the passage nearby of tropical cyclone Dora in the early stages of its development. Arrow heads have been left off the vector winds in left panel for clarity.

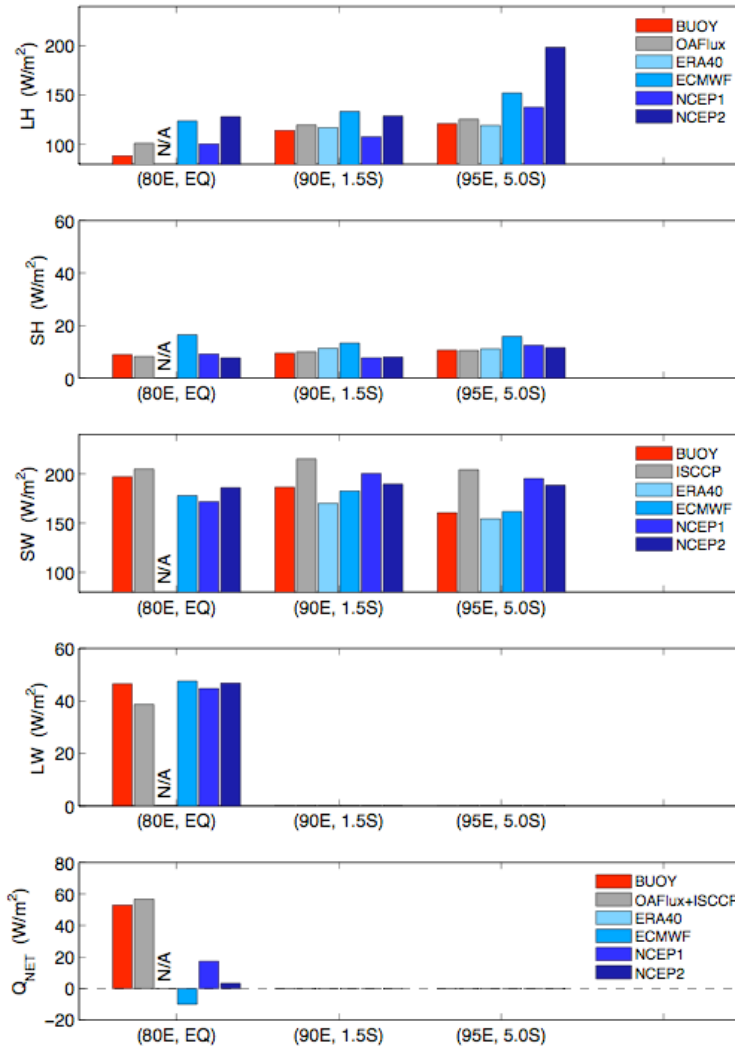


Figure 10. Comparison of surface heat flux components at three RAMA sites with those computed from six different research and operational surface flux products: NCEP1, NCEP2, ECMWF, ERA40, OA-Flux, and ISCCP (see Figure 5 caption for product definitions and references). Fluxes are, from top to bottom, latent heat flux (LH), sensible heat flux (SH), net shortwave radiation (SW), net longwave radiation (LW), and net surface heat flux ( $Q_{net}$ ). Direct measurements of downwelling longwave radiation were made only at  $0^\circ$ ,  $80.5^\circ\text{E}$ . Hence, LW and  $Q_{net}$  (which represents the sum of the four individual components) are shown only for this site. Turbulent fluxes were computed with the COARE version 3.0 flux algorithm (Fairall et al, 2003) using daily averaged data. No warm layer cool skin corrections were applied. Shortwave radiation was adjusted for 6% albedo. Upwelling longwave radiation at  $0^\circ$ ,  $80.5^\circ\text{E}$  was computed assuming blackbody radiation from the sea surface with an emissivity of 0.97. Comparisons were based on 70 days of overlapping data at  $0^\circ$ ,  $80.5^\circ\text{E}$  between 23 Oct 2004 to 31 Dec 2004; 310 days of overlapping data at  $1.5^\circ\text{S}$ ,  $90^\circ\text{E}$  between 23 Oct 2001 and 31 Aug 2002; and 287 days of overlapping data at  $5^\circ\text{S}$ ,  $95^\circ\text{E}$  between 26 Oct 2001 and 31 Aug 2002.