Space Weather Highlights 06 - 12 April 2009

SWO PRF 1754 14 April 2009

Solar activity was very low. No flares were observed. The visible disk was spotless.

No proton events were observed at geosynchronous orbit.

The greater than 2 MeV electron flux at geosynchronous orbit was at high levels 11 - 12 April.

The summary period began with solar wind speeds at about 400 km/s. Velocities slowly decreased to about 300 km/s by late on 08 April. At about 08/2015 UTC, solar wind speeds indicated a fairly sharp rise in velocity to about 475 km/s by 09/0300 UTC. From that time forward, wind speeds gradually increased to near 570 km/s by 11/1700 UTC and subsequently, speeds gradually decreased to about 500 km/s by the end of the summary period. From late on 08 April to late on 09 April, the Bz component of the IMF varied between +7 to -7 nT. Otherwise, Bz did not vary much beyond +/- 3 nT. This increase in wind speed and IMF variability was due to a recurrent coronal hole high speed stream. The geomagnetic field responded with quiet to unsettled periods at middle latitudes, while high latitudes observed quiet to active conditions with the one isolated minor storm period observed midday on 11 April. The period ended with quiet levels across all latitudes.

Space Weather Outlook 15 April - 11 May 2009

Solar activity is expected to be very low.

No proton events are expected at geosynchronous orbit.

The greater than 2 MeV electron flux at geosynchronous orbit is expected to increase to high levels during 08 - 12 May. Normal flux levels are expected during the rest of the period.

Geomagnetic field activity is expected to be at mostly quiet levels through 20 April. Activity is expected to increase to quiet to unsettled levels during 21 - 22 April due to a recurrent coronal hole high-speed stream. Activity is expected to decrease to quiet levels during 23 April to 05 May. Activity is expected to increase to quiet to unsettled levels during 06 - 09 May with active levels possible on the 6th, all due to another recurrent coronal hole high speed stream. Activity is expected to decrease to quiet levels during 10 - 11 May.



	Duly Solar Dala											
	Radio	Sun	Sunspot	Sunspot X-ray		Flares						
	Flux	spot	Area	Area Background		-ray F	lux		Optical			
Date	10.7 cm	No.	(10 ⁻⁶ hemi.)	С	М	Х	S	1	2	3	4
06 April	69	0	0	<a1.0< td=""><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td></a1.0<>	0	0	0	0	0	0	0	0
07 April	70	0	0	<a1.0< td=""><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td></a1.0<>	0	0	0	0	0	0	0	0
08 April	70	0	0	<a1.0< td=""><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td></a1.0<>	0	0	0	0	0	0	0	0
09 April	70	0	0	<a1.0< td=""><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td></a1.0<>	0	0	0	0	0	0	0	0
10 April	69	0	0	<a1.0< td=""><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td></a1.0<>	0	0	0	0	0	0	0	0
11 April	69	0	0	<a1.0< td=""><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td></a1.0<>	0	0	0	0	0	0	0	0
12 April	69	0	0	<a1.0< td=""><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td><td>0</td></a1.0<>	0	0	0	0	0	0	0	0

Daily Solar Data

Daily Particle Data

		oton Fluence	.)	Electron Fluence (electrons/cm ² -day-sr)				
Date	>1 MeV	>10 MeV	>100 MeV	$>.6 \mathrm{MeV}$ $>2 \mathrm{MeV}$ $>4 \mathrm{MeV}$				
06 April	5.4e+05	2.0e+04	4.6e+03	1.1e+06				
07 April	4.6e+05	1.9e+04	4.3e+03	8.8e+05				
08 April	5.5e+05	1.9e+04	4.3e+03	8.0e+05				
09 April	7.1e+05	1.9e+04	4.2e+03	9.5e+04				
10 April	1.2e+06	1.9e+04	4.1e+03	1.3e+07				
11 April	1.8e+06	1.9e+04	4.1e+03	6.4e+07				
12 April	1.1e+06	1.8e+04	4.4e+03	2.4e+08				

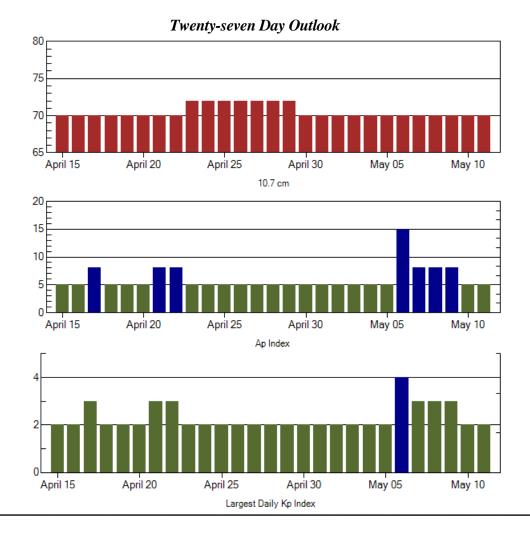
Daily Geomagnetic Data

	Middle Latitude			High Latitude		Estimated
	F	Fredericksburg		College]	Planetary
Date	А	K-indices	Α	K-indices	Α	K-indices
06 April	2	0-0-0-2-1-1-1-1	1	0-0-0-0-1-0-1	3	1-0-0-1-1-1-0-1
07 April	2	0-0-1-1-1-0-1-1	1	0-0-1-1-0-0-0-0	2	0-0-1-1-1-0-0-1
08 April	4	0-1-1-1-2-1-2-1	3	1-2-0-1-1-0-1-1	5	0-0-1-1-2-2-2-2
09 April	9	2-3-3-1-2-2-2-2	12	1-3-3-4-2-3-2-1	12	3-4-3-2-1-3-2-2
10 April	6	2-1-2-2-2-1-2-1	8	1-1-3-3-3-2-2-0	8	2-2-2-2-1-1-3-2
11 April	7	3-1-2-1-2-2-1-2	16	2-2-2-5-4-4-1-1	9	3-2-2-2-2-1-3
12 April	4	2-2-2-0-0-1-1-2	3	1-2-2-1-1-0-0-1	8	2-2-2-1-1-2-2-3

Alerts and Warnings Issued

Date & Time of I	ssue Type of Alert or Warning D	Date & Time of Event UTC
11 Apr 0146	WARNING: Geomagnetic $K = 4$	11 Apr 0147 - 1600
11 Apr 1556	ALERT: Electron 2MeV Integral Flux >= 1000pf	u 11 Apr 1540
11 Apr 0135	ALERT: Geomagnetic $K = 4$	11 Apr 0128
12 Apr 1215	ALERT: Electron 2MeV Integral Flux >= 1000pf	u 12 Apr 0900





	Radio Flux	Planetary	Largest		Radio Flux	Planetary	Largest
Date	10.7 cm	A Index	Kp Index	Date	10.7 cm	A Index	Kp Index
15 Apr	70	5	2	29 Apr	72	5	2
16	70	5	2	30	70	5	2
17	70	8	3	01 May	70	5	2
18	70	5	2	02	70	5	2
19	70	5	2	03	70	5	2
20	70	5	2	04	70	5	2
21	70	8	3	05	70	5	2
22	70	8	3	06	70	15	4
23	72	5	2	07	70	8	3
24	72	5	2	08	70	8	3
25	72	5	2	09	70	8	3
26	72	5	2	10	70	5	2
27	72	5	2	11	70	5	2
28	72	5	2				



	Energetic Events													
	Time		X-ray Optical Information			ı	Peak	Sweep Freq						
Date		1/2	Integ	Imp/	Location	Rgn	Radio Flux	Intensity						
	Begin Max	Max	Class Flux	Brtns	Lat CMD	#	245 2695	II IV						

No Events Observed

		Optical						
	Time	X-ray	Imp /	Location	Rgn			
Date	Begin Max End	Class.	Brtns	Lat CMD				
06 April	No Flares Observed							
07 April	No Flares Observed							
08 April	No Flares Observed							
09 April	No Flares Observed							
10 April	No Flares Observed							
11 April	No Flares Observed							
12 April	No Flares Observed							

Region Summary										
Location		Sunspot C	Character	ristics			Flaı	es		
Н	elio Area	Extent	Spot	Spot	Mag	X-ray		Optic	al	_
Date (° Lat ° CMD) Lo	n (10 ⁻⁶ he	mi) (helio)	Class	Count	Class	C M X	S 1	2	3	4

No active regions.



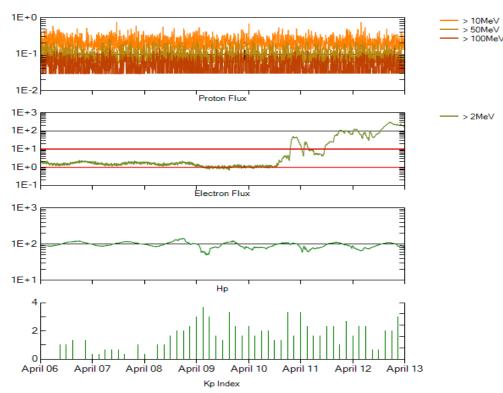
Of the observed monthly mean values											
Sunspot Numbers Radio Flux Geomagne								gnetic			
	Observed	values	<u>Ratio</u>	Smooth	values	*Penticton	Smooth	Planetary	Smooth		
Month	SEC	RI	RI/SEC	SEC	RI	10.7 cm	Value	Ap	Value		
				2	007						
April	6.9	3.7	0.54	16.0	9.9	72.4	75.2	9	8.5		
May	19.4	11.7	0.60	14.2	8.7	74.5	74.2	9	8.4		
June	20.0	12.0	0.60	12.8	7.7	73.7	73.2	7	7.8		
July	15.6	10.0	0.64	11.6	7.0	71.6	72.5	8	7.4		
August	9.9	6.2	0.63	10.2	6.1	69.2	71.8	7	7.6		
Septembe	r 4.8	2.4	0.50	9.9	5.9	67.1	71.5	9	7.8		
				10.0				0			
October	1.3	0.9	0.70	10.0	6.1	65.5	71.5	9	7.9		
November		1.7	0.68	9.4	5.7	69.7	71.1	5	7.8		
December	: 16.2	10.1	0.62	8.1	5.0	78.6	70.5	4	7.8		
				•	000						
т	F 1	2.4	0.67		008	70.1	70.0	6			
January	5.1	3.4	0.67	6.9	4.2	72.1	70.0	6	7.7		
February	3.8	2.1	0.55	5.9	3.6	71.2	69.6	9	7.6		
March	15.9	9.3	0.58	5.3	3.3	72.9	69.5	10	7.4		
April	4.9	2.9	0.59	5.3	3.3	70.3	69.6	9	7.1		
May	5.7	2.9	0.51	5.7	3.5	68.4	69.7	6	6.9		
June	4.2	3.1	0.74	5.2	3.2	65.9	69.2	7	6.8		
<i>v</i> and		0.11	0.71	0.2	0.2	0017	07.2	,	0.0		
July	1.0	0.5	0.50	4.5	2.7	65.8	68.8	6	6.6		
August	0.0	0.5	**	4.4	2.6	66.4	68.6	5	6.2		
Septembe		1.1	0.73	3.7	2.2	67.1	68.4	5	5.8		
October	5.2	2.9	0.56			68.3		6			
November	r 6.8	4.1	0.60			68.6		3			
December	: 1.3	0.8	0.62			69.2		2			
				•	000						
Ionnom	20	15	0.54	2	009	60.8		2			
January Fahruary	2.8	1.5	0.54			69.8 70.0		3			
February Marah	2.5	1.4	0.56			70.0		4 4			
March	0.7	0.7	1.00			69.2		4			

Recent Solar Indices (preliminary) Of the observed monthly mean values

<u>NOTE</u>: All smoothed values after September 2002 and monthly values after March 2003 are preliminary estimates. The lowest smoothed sunspot index number for Cycle 22, RI = 8.0, occurred in May 1996. The highest smoothed sunspot number for Cycle 23, RI= 120.8, occurred April 2000. *After June 1991, the 10.7 cm radio flux data source is Penticton, B.C. Canada. Prior to that, it was Ottawa.

**SWPC sunspot number was less than RI value, so a ratio could not be done.





Weekly Geosynchronous Satellite Environment Summary Week Beginning 06 April 2009

GOES-11 designated Primary Electron Satellite and GOES-10 Secondary: December 1, 2008 the GOES-12 Electron sensor began experiencing periods of noise and sensor is unreliable.

Protons plot contains the five-minute averaged integral proton flux (protons/cm²-sec -sr) as measured by GOES-11 (W135) for each of three energy thresholds: greater than 10, 50, and 100 MeV.

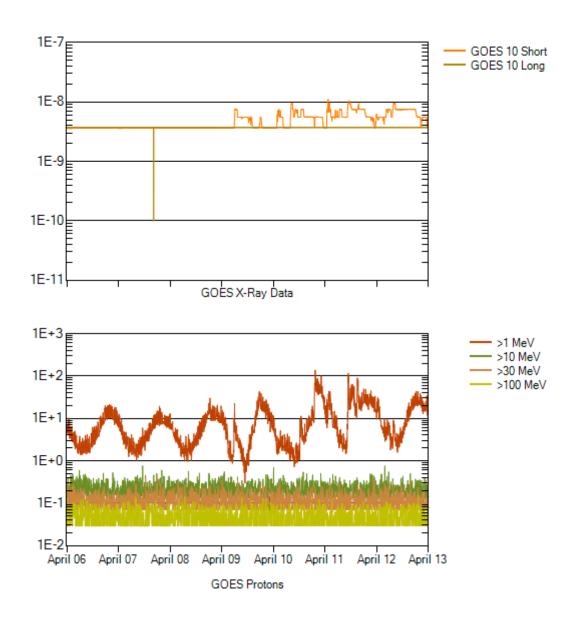
Electrons plot contains the five-minute averaged integral electron flux (electrons/cm²-sec -sr) with energies greater than 2 MeV at GOES-11 (W135).

Hp plot contains the five minute averaged magnetic field H - component in nanoteslas (nT) as measured by GOES-11. The H component is parallel to the spin axis of the satellite, which is nearly parallel to the Earth's rotation axis.

Kp plot contains the estimated planetary 3-hour K-index (derived by the Air Force Weather Agency) in real time from magnetometers at Meanook, Canada; Sitka, AK; Glenlea, Canada; St. Johns, Canada; Ottawa, Canada; Newport, WA; Fredericksburg, VA; Boulder, CO; Fresno, CA and Hartland, UK. These data are made available through cooperation from the Geological Survey of Canada (GSC), British Geological Survey (BGS) and the US Geological Survey. These may differ from the final Kp values derived from a more extensive network of magnetometers.

The data included here are those now available in real time at the SWPC and are incomplete in that they do not include the full set of parameters and energy ranges known to cause satellite operating anomalies. The proton and electron fluxes and Kp are "global" parameters that are applicable to a first order approximation over large areas. H parallel is subject to more localized phenomena and the measurements generally are applicable to within a few degrees of longitude of the measuring satellite.





Weekly GOES Satellite X-ray and Proton Plots

X-ray plot contains five-minute averaged x-ray flux (watts/ m^{2}) as measured by GOES 10 (W060) and GOES 11 (W135) in two wavelength bands, .05 -. 4 and .1 - .8 nm. The letters A, B, C, M and X refer to x-ray event levels for the .1 - .8 nm band.

Proton plot contains the five-minute averaged integral proton flux (protons/cm² –sec-sr) as measured by GOES-11 (W135) for each of the energy thresholds: >1, >10, >30 and >100 MeV. P10 event threshold is 10 pfu (protons/cm²-sec-sr) at greater than 10 MeV.

