## A Roadmap for Nuclear Physics: Objectives and Performance Targets

	2006	200	8	2010	2012	2014	2016		
Medium Energy Nuclear Physics	Obtain first polarized high energy proton-proton data studying the proton spin (2006) NSF  Determine the strange of content of the proton (2)			proton, ne using high		NSF	<ul> <li>○ Demonstrate progress in realizing a quantitative understanding of the quark substructure of the proton, neutron, and simple nuclei</li> <li>by comparison of precision measurements of their</li> </ul>		
		Begin search for an moment of Radium-2	electric dipole	Determine gluon contribution to proton spin (2010) NSF		n measurements to exotic mesons to understanding of NSF c confinement (2013)	fundamental properties with theoretical		
Heavy Ion Nuclear	Begin studies of rare p the formation of hot, do matter <b>(2004)</b>		of the Experimen	high transverse-r through hot, dens dominated by glu its at the Relativistic Heavy	y Ion Collider (RHIC) at Bro	ookhaven			
Physics		in the laboratory using colliding beams of a nuclei (2007)	ng fundament	nal Laboratory will have established the basic characteristics of a mental new form of nuclear matter that exists at extremely high tratures and densities (2008)			samples of hot, dense nuclear matter (2015)		
Low Energy Nuclear Physics	Establish reaction rates for understanding how light elements are created in supernovae (2006)  Develop three-dimensional computer simulations for the behavior of supernovae, including		Complete measurements in new regions of nuclear structure and develop — the nuclear many-body theory to predict nuclear properties (2008) NSF	Begin studies of nuclei at the new GRETINA gam revolutionizing detector  Launch next-generation studying decay of the next-generation studying studying decay of the next-generation stu	technology (2010) NSF neutron experiments	0 <u>9</u> ii r : ii - li r	O Demonstrate progress in investigating new regions of nuclear structure, study interactions in nuclear matter like those occurring in neutron stars, and determining the reactions that created the nuclei of		
	incorporate reaction dyr  Quantify ne	e and explosion, which the relevant nuclear amics (2006)			Begin a high-precision electric dipole moment which will test new theo fundamental particle in	search for the of the neutron, ories of NSF teractions (2013)	atomic elements inside stars and supernovae (2015)  Demonstrate progress in determining the fundamental properties of neutrinos and		
	neutrinos from the sun, cosmic-ray interactions, and nuclear reactors (2006) HEP			Establish an electron neutrino mass (2011) HEP	Begin experiments to lo neutrinoless double be essential information a scale of neutrino mass	ta decay to provide bout the absolute	fundamental symmetries by using neutrinos from the sun and nuclear reactors and by using radioactive decay measurements (2015)		
Future Fac (Cross cut and multiple object targets):	cilities Upg d support to u cos tives and he	to upgrade of the CEBAF - a sotopes that ar cost-effective way to double the energy of the existing    powerful resear isotopes that ar becay Under the powerful resear isotopes that ar becay Under the energy of the existing    powerful resear isotopes that ar becay Under the energy of the existing    powerful resear isotopes that are the energy of the existing    powerful resear isotopes that are the energy of the existing    powerful resear isotopes that are the energy of the existing    powerful resear isotopes that are the energy of the existing    powerful resear isotopes that are the energy of the existing    powerful resear isotopes that are the energy of the existing    powerful resear isotopes that are the energy of the existing    powerful resear isotopes that are the energy of the existing    powerful resear isotopes that are the energy of the existing    powerful resear isotopes that are the energy of the existing    powerful resear isotopes that are the energy of the existing    powerful resear isotopes that are the energy of the existing    powerful resear isotopes that are the energy of the existing    powerful resear isotopes that are the energy of the existing    powerful resear isotopes that are the energy of the existing    powerful research is the existing    powerful research is the existing    powerful research is the existing    powerful r		Upgraded CEBAF operations begin (2012)  ccelerator (RIA): Begin construction of the world's most ch facility dedicated to producing and exploring new rare enot found naturally on Earth. (2009)  rground Detector: Begin construction to enable measurements of neutrino of whether the neutrino and its anti-particle are identical. (2007) NSF HEP			(2013)		

Interdependencies: (Descriptions)

Broadly with **ASCR** on computational developments, both hardware and software, affecting all facets of basic research and advanced instrumentation.

NSF = with NSF HEP = with HEP =Key Intermediate Objective from DOE Strategic Plan
 =Long Term Success Measure from PART

This timeline is for planning purposes only and does not constitute financial or contractual commitments by the Federal Government.