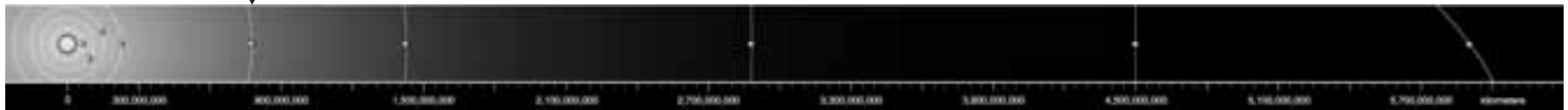
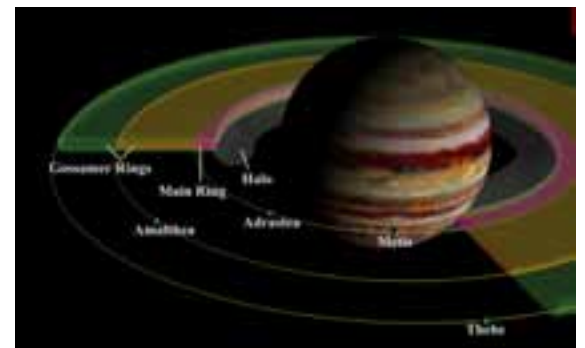
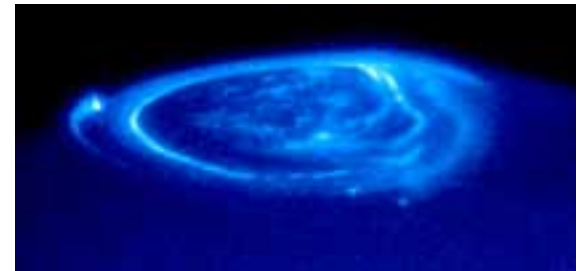
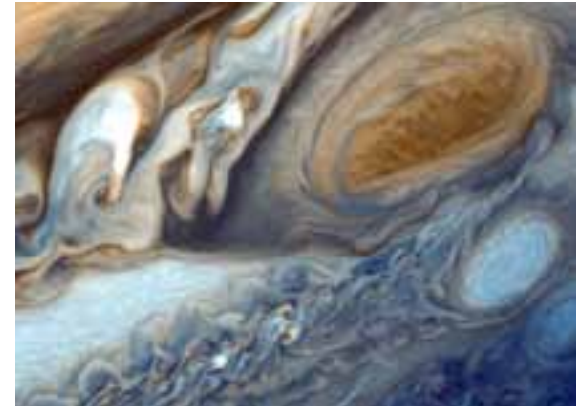
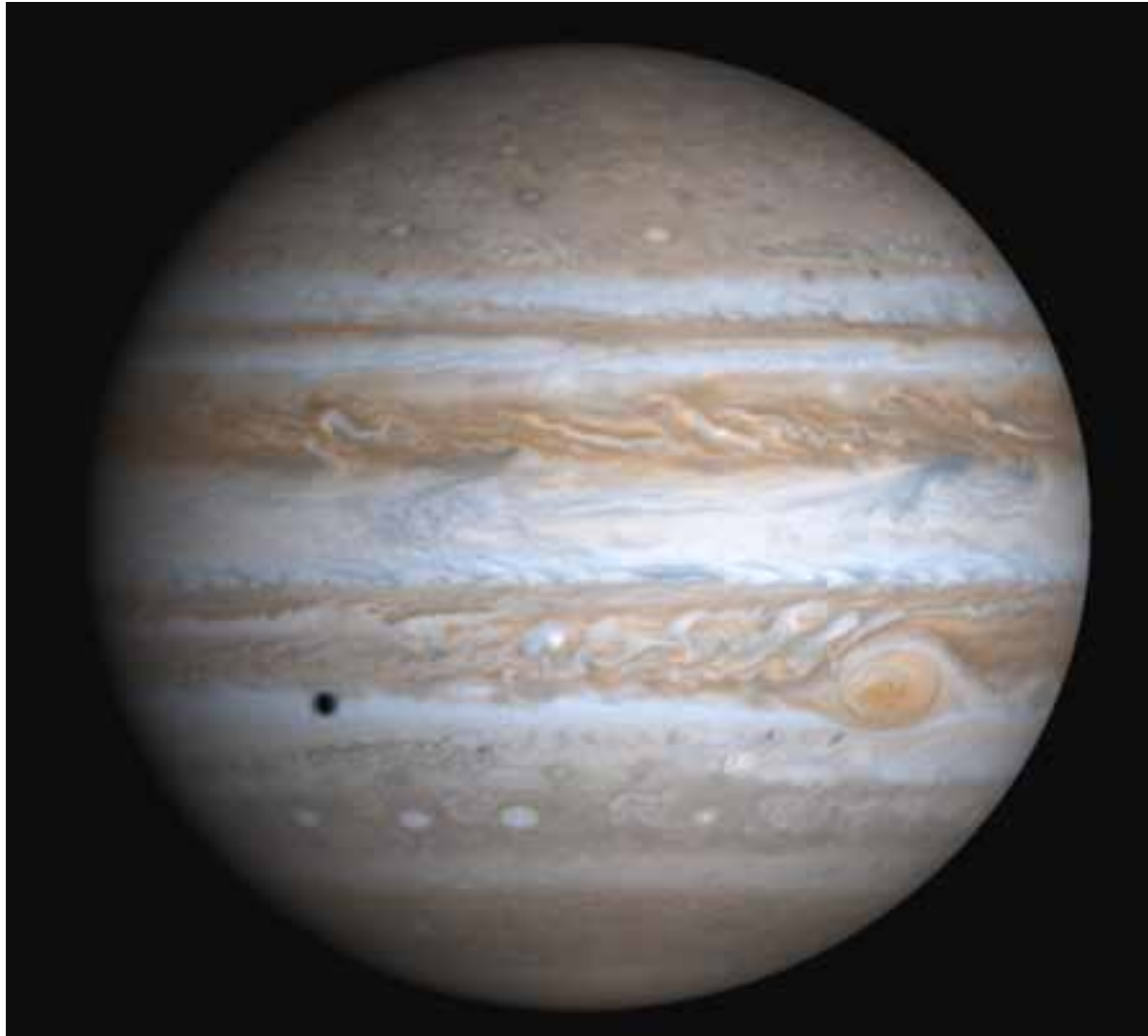


# Jupiter



# Jupiter



The most massive planet in our solar system, with four planet-sized moons and many smaller moons, Jupiter forms a kind of miniature solar system. Jupiter resembles a star in composition. In fact, if it had been about eighty times more massive, it would have become a star rather than a planet.

On January 7, 1610, using his primitive telescope, astronomer Galileo Galilei saw four small “stars” near Jupiter. He had discovered Jupiter’s four largest moons, now called Io, Europa, Ganymede, and Callisto. Collectively, these four moons are known today as the Galilean satellites.

Galileo would be astonished at what we have learned about Jupiter and its moons in the past 30 years. Io is the most volcanically active body in our solar system. Ganymede is the largest planetary moon and is the only moon in the solar system known to have its own magnetic field. A liquid ocean may lie beneath the frozen crust of Europa. Icy oceans may also lie deep beneath the crusts of Callisto and Ganymede. In 2003 alone, astronomers discovered 23 new moons orbiting the giant planet, giving Jupiter a total moon count of 62 — the most in the solar system. The numerous small outer moons may be asteroids captured by the giant planet’s gravity.

Jupiter’s appearance is a tapestry of beautiful colors and atmospheric features. Most visible clouds are composed of ammonia. Water exists deep below and can sometimes be seen through clear spots in the clouds. The planet’s “stripes” are dark belts and light zones created by strong east–west winds in Jupiter’s upper atmosphere. Within these belts and zones are storm systems that have raged for years. The Great Red Spot, a giant spinning storm, has been observed for more than 300 years.

The composition of Jupiter’s atmosphere is similar to that of the Sun — mostly hydrogen and helium. Deep in the atmosphere, the pressure and temperature increase, compressing the hydrogen gas into a liquid. At depths about a third of the way down, the hydrogen becomes metallic and electrically conducting. In this metallic layer, Jupiter’s powerful magnetic field is generated by electrical currents driven by Jupiter’s fast rotation. At the center, the immense pressure may support a solid core of ice–rock about the size of Earth.

Jupiter’s enormous magnetic field is nearly 20,000 times as powerful as Earth’s. Trapped within Jupiter’s magnetosphere (the area in which magnetic field lines encircle the planet from pole to pole) are swarms of charged particles. Jupiter’s rings and moons are embedded in an intense radiation belt of electrons and ions trapped in the magnetic field. The jovian magnetosphere, comprising these particles and fields, balloons 1 to 3 million kilometers (600,000 to 2 million miles) toward the Sun and tapers into a windsock-shaped tail extending more than 1 billion kilometers (600 million miles) behind Jupiter as far as Saturn’s orbit.

Discovered in 1979 by NASA’s Voyager 1 spacecraft, Jupiter’s rings were a surprise: a flattened main ring and an inner cloud-like ring, called the halo, are both composed of small, dark particles. A third ring, known as the gossamer ring because of its transparency, is actually three rings of microscopic debris from three small moons: Amalthea, Thebe, and Adrastea. Jupiter’s ring system may be formed by dust kicked up as interplanetary meteoroids smash into the giant planet’s four small inner moons. The main ring probably comes from the moon Metis. Jupiter’s rings are only visible when backlit by the Sun.

In December 1995, NASA’s Galileo spacecraft dropped a probe into Jupiter’s atmosphere, which collected the first direct measurements of Jupiter’s atmosphere. Following the release of the probe, the Galileo spacecraft began a multiyear study of Jupiter and the largest moons. As Galileo began its 29th orbit, the Cassini–Huygens spacecraft was nearing Jupiter for a gravity-assist maneuver on the way to Saturn. The two spacecraft made simultaneous observations of the magnetosphere, solar wind, rings, and Jupiter’s auroras.

## FAST FACTS

Namesake	King of the Roman gods
Mean Distance from the Sun	778.41 million km (483.68 million mi)
Orbit Period	11.8565 Earth years (4,330.6 Earth days)
Orbit Eccentricity (Circular Orbit = 0)	0.04839
Orbit Inclination to Ecliptic	1.305 deg
Inclination of Equator to Orbit	3.12 deg

Rotation Period	9.92 hr
Equatorial Radius	71,492 km (44,423 mi)
Mass	317.82 of Earth’s
Density	1.33 g/cm <sup>3</sup>
Gravity	20.87 m/sec <sup>2</sup> (68.48 ft/sec <sup>2</sup> )
Atmosphere Primary Components	hydrogen, helium
Effective Temperature	–148 deg C (–234 deg F)
Known Moons*	62
Rings	1 (three major components)

\*As of November 2005.

## SIGNIFICANT DATES

- 1610 — Galileo Galilei makes the first detailed observations of Jupiter using a telescope.
- 1973 — Pioneer 10 becomes the first spacecraft to cross the asteroid belt and fly past Jupiter.
- 1979 — Voyager 1 and 2 discover Jupiter’s faint rings, several new moons, and volcanic activity on Io’s surface.
- 1994 — Astronomers watch pieces of comet Shoemaker–Levy 9 collide with Jupiter.
- 1995–2003 — The Galileo spacecraft conducts extended observations of Jupiter and its moons and rings.

## ABOUT THE IMAGES



**1** A detailed, true-color image of Jupiter taken by the Cassini spacecraft. The Galilean moon Europa casts a shadow on the planet.

- 2** A Voyager 1 image of Jupiter’s Great Red Spot.
- 3** An image of Jupiter’s aurora, a sign of the interaction between Jupiter’s magnetic field and energy from the Sun.
- 4** A schematic of the components of Jupiter’s ring system.

## FOR MORE INFORMATION

[solarsystem.nasa.gov/planets/profile.cfm?Object=Jupiter](http://solarsystem.nasa.gov/planets/profile.cfm?Object=Jupiter)