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APPENDIX 4. PERTURBATIONS TO ELECTRON DENSITY MODELS WITH INPUT PARAMETER FORMS

The following perturbations to electron density models (irregularities) are available. The input parameter forms, which describe the perturbation, and the subroutine listings are given on the pages shown.

a. Do-nothing perturbation (ELECT1)	126
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To add other perturbations to electron density models the user must write a subroutine to modify the normalized electron density (X) and its gradient ($\partial X/\partial r$, $\partial X/\partial \theta$, $\partial X/\partial \varphi$) as a function of position in spherical polar coordinates (r , θ , φ).

The restrictions on electron density models also apply to perturbations. Again, the coordinates r , θ , φ refer to the computational coordinate system, which may not be the same as geographic coordinates. In particular, they are geomagnetic coordinates when the earth-centered dipole model of the earth's magnetic field is used.

The input to the subroutine is through blank common (see Table 3) for the position (r , θ , φ) and through common block /XX/ (see Table 8) for the unperturbed electron density and its gradient. The output is through common block /XX/. It is useful if the name of the subroutine suggests the perturbation model to which it corresponds. It should have an entry point ELECT1 so that it may be called by an electron density subroutine. Any parameters needed by the subroutine should be input into W151 through W199 of the W array. (See Table 2.)

If no perturbation is wanted, the following subroutine should be used.

```
C      SUBROUTINE ELECT1                                ELEC001
      USE WHEN AN ELECTRON DENSITY PERTURBATION IS NOT WANTED ELEC002
      COMMON /XX/ MODX(2),X(6)                          ELEC003
      COMMON /WW/ ID(10),W0,W(400)                      ELEC004
      EQUIVALENCE (PERT,W(150))                        ELEC005
      DATA (MODX(2)=6H NONE )                          ELEC006
      PERT=0.                                           ELEC007
      RETURN                                           ELEC008
      END                                               ELEC009-
```

INPUT PARAMETER FORM FOR SUBROUTINE TORUS

A perturbation to an ionospheric electron density model consisting of an East-West irregularity with an elliptical cross section above the equator

$$N = N_0 (1 + \Delta)$$

$$\Delta = C_0 \exp \left\{ - \left[\frac{(R_0 + H_0)(\theta - \pi/2) \cos \beta + (R - R_0 - H_0) \sin \beta}{A} \right]^2 - \left[\frac{(R - R_0 - H_0) \cos \beta - (R_0 + H_0)(\theta - \pi/2) \sin \beta}{B} \right]^2 \right\}$$

R_0 is the radius of the earth.

R, θ, φ give the position in spherical polar coordinates.

$N_0(R, \theta, \varphi)$ is any ionospheric electron density model.

Specify:

$C_0 =$ _____. (W151)

Semi-major axis of ellipse, $A =$ _____ km (W152)

Semi-minor axis of ellipse, $B =$ _____ km (W153)

Tilt of ellipse, $\beta =$ _____ degrees (W154)

Height of torus from ground, $H_0 =$ _____ km (W155)

(W150: = 1. to use perturbation, = 0. to ignore perturbation)

```

SUBROUTINE TORUS
COMMON /CONST/ PI,PIT2,PID2,DUM(5)
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX
COMMON R(6) /WW/ ID(10),W0,W(400)
EQUIVALENCE (EARTH,W(2)),(C0,W(151)),(A,W(152)),(B,W(153)),
1 (BETA,W(154)),(H0,W(155))
REAL LAMBDA
DATA (PDPP=0.), (MODX(2)=6H TORUS)
ENTRY ELECT1
IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0.) RETURN
IF (C0.EQ.0.) RETURN
R0=EARTH+H0
Z=R(1)-R0
LAMBDA=R0*(R(2)-PID2)
SINBET=SIN(BETA)
COSBET=COS(BETA)
P=LAMBDA*COSBET+Z*SINBET
Y=Z*COSBET-LAMBDA*SINBET
DELTA=C0*EXP(-(P/A)**2-(Y/B)**2)
DEL1=DELTA+1.
PDPR=-2.*DELTA*(P*SINBET/A**2+Y*COSBET/B**2)
PDPT=-2.*DELTA*(P*R0*COSBET/A**2-Y*R0*SINBET/B**2)
PXPR=PXPR*DEL1+X*PDPR
PXPTH=PXPTH*DEL1+X*PDPT
PXPPH=PXPPH*DEL1+X*PDPP
X=X*DEL1
RETURN
END
TOR 001
TOR 002
TOR 003
TOR 004
TOR 005
TOR 006
TOR 007
TOR 008
TOR 009
TOR 010
TOR 011
TOR 012
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TOR 021
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TOR 025
TOR 026
TOR 027
TOR 028-

```

INPUT PARAMETER FORM FOR SUBROUTINE DTORUS

A perturbation to an ionospheric electron density model consisting of two east-west irregularities with elliptical cross sections above the equator. Since the model is expressed in spherical coordinates and does not depend on longitude, the perturbation is actually a torus circling the earth above the equator.

$$\begin{aligned}
 N &= N_0 (1 + \Delta) \\
 \Delta &= C_1 \exp \left\{ - \left[\frac{(r_0 + H_1)(\theta - \pi/2) \cos \beta + (r - r_0 - H_1) \sin \beta}{A_1} \right]^2 \right. \\
 &\quad \left. - \left[\frac{(r - r_0 - H_1) \cos \beta - (r_0 + H_1)(\theta - \pi/2) \sin \beta}{B_1} \right]^2 \right\} \\
 &+ C_2 \exp \left\{ - \left[\frac{(r_0 + H_2)(\theta - \pi/2 + \delta\theta) \cos \beta + (r - r_0 - H_2) \sin \beta}{A_2} \right] \right. \\
 &\quad \left. - \left[\frac{(r - r_0 - H_2) \cos \beta - (r_0 + H_2)(\theta - \pi/2 + \delta\theta) \sin \beta}{B_2} \right]^2 \right\}
 \end{aligned}$$

$\delta\theta$ = Northward angular displacement of the lower blob from the upper one

$$= \frac{H_1 - H_2}{\tan \beta (r_0 + H_2)}$$

r_0 is the radius of the earth.

r, θ, φ are spherical (earth-centered) polar coordinates.

$N_0(r, \theta, \varphi)$ is any electron density model.

Specify:

use perturbation _____ (W150 = 1.)

ignore perturbation _____ (W150 = 0.)

Fractional perturbation electron density at the center of the upper blob, $C_1 =$ _____ (W151).

That of the lower blob, $C_2 =$ _____ (W156).

Height (above ground) of the center of the upper blob,

$H_1 =$ _____ km (W155).

That of the lower blob, $H_2 =$ _____ km (W159).

Angle (with a horizontal southward vector) of the line joining the
blob centers, $\theta =$ _____ $\frac{\text{rad}}{\text{deg}}$ (W154).

Semi-axis of the upper blob, to the $1/e$ perturbation contour, in the
direction of the line joining the blobs, $A_1 =$ _____ km (W152).

That of the lower blob, $A_2 =$ _____ km (W157).

Semi-axis of the upper blob in the direction normal to the line joining
the blobs, $B_1 =$ _____ km (W153).

That of the lower blob, $B_2 =$ _____ km (W158).

```
SUBROUTINE DTORUS
COMMON /CONST/ PI,PIT2,PID2,DUM(5)
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT
COMMON R(6) /WW/ ID(10),W0,W(400)
EQUIVALENCE (EARTH,W(2)),(C1,W(151)),(A1,W(152)),(B1,W(153)),
1 (BETA,W(154)),(H1,W(155)),(C2,W(156)),(A2,W(157)),(B2,W(158)),
2 (H2,W(159))
REAL LAMBDA1,LAMBDA2
DATA (MODX(2)=6HDTORUS),(PDPP=0.)
ENTRY ELECT1
IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0.) RETURN
IF (C1.EQ.0.) RETURN
R1=EARTH+H1
R2=EARTH+H2
Z1=R(1)-R1
Z2=R(1)-R2
LAMBDA1=R1*(R(2)-PID2)
LAMBDA2=R2*(R(2)-PID2+(H1-H2)/R2/TANF(BETA))
SINBET=SIN(BETA)
COSBET=COS(BETA)
P1=LAMBDA1*COSBET+Z1*SINBET
P2=LAMBDA2*COSBET+Z2*SINBET
Y1=Z1*COSBET-LAMBDA1*SINBET
Y2=Z2*COSBET-LAMBDA2*SINBET
DELTA1=C1*EXP(-(P1/A1)**2-(Y1/B1)**2)
DELTA2=C2*EXP(-(P2/A2)**2-(Y2/B2)**2)
DEL1=1.+DELTA1+DELTA2
PDPR1=-2.*DELTA1*(P1*SINBET/A1**2+Y1*COSBET/B1**2)
PDPR2=-2.*DELTA2*(P2*SINBET/A2**2+Y2*COSBET/B2**2)
PDPT1=-2.*DELTA1*(P1*R1*COSBET/A1**2-Y1*R1*SINBET/B1**2)
PDPT2=-2.*DELTA2*(P2*R2*COSBET/A2**2-Y2*R2*SINBET/B2**2)
PXPR=PXPR*DEL1+X*(PDPR1+PDPR2)
PXPTH=PXPTH*DEL1+X*(PDPT1+PDPT2)
PXPPH=PXPPH*DEL1*PDPP
X=X*DEL1
RETURN
END
```

DTOR001
DTOR002
DTOR003
DTOR004
DTOR005
DTOR006
DTOR007
DTOR008
DTOR009
DTOR010
DTOR011
DTOR012
DTOR013
DTOR014
DTOR015
DTOR016
DTOR017
DTOR018
DTOR019
DTOR020
DTOR021
DTOR022
DTOR023
DTOR024
DTOR025
DTOR026
DTOR027
DTOR028
DTOR029
DTOR030
DTOR031
DTOR032
DTOR033
DTOR034
DTOR035
DTOR036
DTOR037-

INPUT PARAMETER FORM FOR SUBROUTINE TROUGH

A perturbation to an ionospheric electron density model consisting of an increase in electron density near any latitude

$$N = (1 + \Delta) N_0 (R, \theta, \varphi) \qquad W = B \text{ for } \frac{\pi}{2} - \theta - \lambda \geq 0$$

$$\Delta = A \exp\left(-\left(\frac{\pi/2 - \theta - \lambda}{W}\right)^2\right) \qquad W = B \times C \text{ for } \frac{\pi}{2} - \theta - \lambda < 0$$

$N_0 (R, \theta, \varphi)$ is any ionospheric electron density model.

R, θ, φ give the position in spherical polar coordinates.

Specify:

Amplitude of the perturbation, $A =$ _____ (W151)

half width of the perturbation, $B =$ _____ degrees (W152)

latitude of the perturbation, $\lambda =$ _____ degrees (W153)

width factor for South of trough, $C =$ _____ (W154)

(W150: = 1. to use perturbation, = 0. to ignore perturbation)

C	<pre> SUBROUTINE TROUGH A PERTURBATION TO AN ELECTRON DENSITY MODEL COMMON /CONST/ PI,PIT2,PID2,DUM(5) COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX COMMON R(6) /WW/ ID(10),W0,W(400) EQUIVALENCE (A,W(151)),(E,W(152)),(ALAT,W(153)),(FACTOR,W(154)) DATA (MODX(2)=6HTROUGH) ENTRY ELECT1 IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0.) RETURN IF (A.EQ.0.) RETURN ANGLE=R(2)+ALAT-PID2 WIDTH=B IF (ANGLE.GT.0.) WIDTH=FACTOR*B ANGLE=ANGLE/WIDTH DELTA=A*EXP(-ANGLE**2) DEL1=DELTA+1. PXPR=PXPR*DEL1 PXPTH=PXPTH*DEL1-2.*X*ANGLE*DELTA/WIDTH PXPPH=PXPPH*DEL1 X=X*DEL1 RETURN END </pre>	<pre> TROU001 TROU002 TROU003 TROU004 TROU005 TROU006 TROU007 TROU008 TROU009 TROU010 TROU011 TROU012 TROU013 TROU014 TROU015 TROU016 TROU017 TROU018 TROU019 TROU020 TROU021 TROU022- </pre>
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INPUT PARAMETER FORM FOR SUBROUTINE SHOCK

A perturbation to an ionospheric electron density model consisting of an increase in electron density produced by a shock wave

$$N(R, \theta, \varphi) = N_0(R, \theta, \varphi) \left[1 + P \exp\left(-9 \left(\frac{\rho_c - \rho}{w}\right)^2\right) \right]$$

$$\rho_c = s(h - h_0) - w$$

$$\rho = R \left| \cos^{-1} \left[\cos(\varphi - \varphi_0) \cos(\lambda - \lambda_0) \right] \right|$$

$N_0(R, \theta, \varphi)$ is the ambient electron density specified by any electron density model.

R, θ, φ give the position in spherical polar coordinates.

$h = R - a$ is the height above the surface of the earth.

a is the radius of the earth.

$\lambda = \frac{\pi}{2} - \theta$ is the latitude.

Specify:

Relative increase in electron density, $P =$ _____ (W151).

Width of the disturbance, $w =$ _____ km (W152).

Latitude of the center of the disturbance, $\lambda_0 =$ _____ radians or degrees (W153).

Longitude of the center of the disturbance, $\varphi_0 =$ _____ radians or degrees (W154).

Slope measured from vertical - rate of increase of ρ_c with height, $s =$ _____ (W155).

Height to the bottom of the disturbance, $h_0 =$ _____ km (W156).

(W150: = 1. to use perturbation, = 0. to ignore perturbation)


```

SUBROUTINE SHOCK SHOC001
C A PERTURBATION TO AN ELECTRON DENSITY MODEL SIMULATING A SHOCK WAVE SHOC002
COMMON /CONST/ PI,PID2,PID2,DUM(5) SHOC003
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX SHOC004
COMMON R(6) /WW/ ID(10),W0,W(400) SHOC005
EQUIVALENCE (EARTH,W(2)),(P,W(151)),(WW,W(152)),(ALAT,W(153)), SHOC006
1 (ALON,W(154)),(S,W(155)),(H0,W(156)) SHOC007
REAL LAT,LON SHOC008
DATA (MODX(2)=6H SHOCK) SHOC009
ENTRY ELECT1 SHOC010
IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0.) RETURN SHOC011
IF (P.EQ.0..OR.WW.EQ.0.) RETURN SHOC012
H=R(1)-EARTH SHOC013
RHOC=S*(H-H0)-WW SHOC014
LON=R(3)-ALON SHOC015
LAT=PID2-R(2)-ALAT SHOC016
COSLON=COS(LON) SHOC017
COSLAT=COS(LAT) SHOC018
U=COSLON*COSLAT SHOC019
RHO=R(1)*ACOS(U) SHOC020
DIF=RHOC-RHO SHOC021
CON=-9./WW**2 SHOC022
CONS=P*EXP(CON*DIF**2) SHOC023
CONST=1.+CONS SHOC024
CON=2.*CON*CONS*DIF SHOC025
PXPR=PXPR*CONST+X*CON*(S-RHO/R(1)) SHOC026
CONS=R(1)*(1./SQRT(1.-U**2)) SHOC027
PXPTH=PXPTH*CONST+X*CON*CONS*COSLON*SIN(LAT) SHOC028
PXPPH=PXPPH*CONST-X*CON*CONS*COSLAT*SIN(LON) SHOC029
X=X*CONST SHOC030
RETURN SHOC031
END SHOC032-

```

INPUT PARAMETER FORM FOR SUBROUTINE WAVE

A perturbation to an ionospheric electron density model consisting of a "gravity-wave" irregularity traveling from north pole to south pole

$$N = N_0(1 + \Delta)$$

$$\Delta = \delta \exp\left\{-\left[(R - R_0 - z_0)/H\right]^2\right\} \cdot \cos\left\{2\pi\left[t' + (\pi/2 - \theta)\frac{R_0}{\lambda_x} + (R - R_0)/\lambda_z\right]\right\}$$

$$\frac{\partial N}{\partial t} = \frac{-2\pi}{\lambda_x} V_x N_0 \delta \exp\left\{-\left[(R - R_0 - z_0)/H\right]^2\right\} \cdot \sin 2\pi\left[t' + (\pi/2 - \theta)\frac{R_0}{\lambda_x} + (R - R_0)/\lambda_z\right]$$

R_0 is the radius of the earth.

R , θ , φ are the spherical (earth-centered) polar coordinates
(Δ is independent of φ).

$N_0(R, \theta, \varphi)$ is any electron density model.

Specify:

the height of maximum wave amplitude, $z_0 = \underline{\hspace{2cm}}$ km (W151)

wave-amplitude "scale height," $H = \underline{\hspace{2cm}}$ km (W152)

wave perturbation amplitude, $\delta = \underline{\hspace{2cm}}$ [0. to 1.] (W153)

horizontal trace velocity, $V_x = \underline{\hspace{2cm}}$ km/sec (W154)
(needed only if Doppler shift is calculated)

horizontal wavelength, $\lambda_x = \underline{\hspace{2cm}}$ km (W155)

vertical wavelength, $\lambda_z = \underline{\hspace{2cm}}$ km (W156)

time in wave periods, $t' = \underline{\hspace{2cm}}$ [0. to 1.] (W157)

(W150: = 1. to use perturbation, = 0. to ignore perturbation)

	SUBROUTINE WAVE	WAVE001
C	PERTURBATION TO AN ALPHA-CHAPMAN ELECTRON DENSITY MODEL	WAVE002
	COMMON /CONST/ PI,PIT2,PID2,DUM(5)	WAVE003
	COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	WAVE004
	COMMON R(6) /WW/ ID(10),W0,W(400)	WAVE005
	EQUIVALENCE (EARTH,W(2)),(Z0,W(151)),(SH,W(152)),(DELTA,W(153)),	WAVE006
	1 (VSUBX,W(154)),(LAMBDAZ,W(155)),(LAMBDAZ,W(156)),(TP,W(157))	WAVE007
	REAL LAMBDAZ,LAMBDAZ	WAVE008
	DATA (MODX(2)=6H WAVE)	WAVE009
	ENTRY ELECT1	WAVE010
	IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0..) RETURN	WAVE011
	IF (DELTA.EQ.0..OR.SH.EQ.0..) RETURN	WAVE012
	H=R(1)-EARTH	WAVE013
	EXPQ=EXP(-((H-Z0)/SH)**2)	WAVE014
	TMP=PIT2*(TP+(PID2-R(2))*EARTH/LAMBDAZ+H/LAMBDAZ)	WAVE015
	SINW=SIN(TMP)	WAVE016
	COSW=SIN(PID2-TMP)	WAVE017
	CONS=1.0+DELTA*EXPQ*COSW	WAVE018
	IF (H.NE.0.) PXPR=PXPR*CONS-X*DELTA*EXPQ*(2.0/SH**2*(H-Z0)*COSW	WAVE019
	1 +PIT2/LAMBDAZ*SINW)	WAVE020
	PXPTH=PXPTH*CONS+X*DELTA*PIT2*EARTH/LAMBDAZ*SINW*EXPQ	WAVE021
	PXPPH=PXPPH*CONS	WAVE022
	PXPT=0.	WAVE023
	IF (VSUBX.NE.0.) PXPT=-PIT2*VSUBX/LAMBDAZ*X*DELTA*EXPQ*SINW	WAVE024
	X=X*CONS	WAVE025
	RETURN	WAVE026
	END	WAVE027-

INPUT PARAMETER FORM FOR SUBROUTINE WAVE 2
 PERTURBATION TO AN IONOSPHERIC ELECTRON DENSITY MODEL

A "gravity-wave" irregularity traveling from north pole to south pole - same as WAVE 1, but with Gaussian amplitude variations in latitude and longitude, and provision for a horizontal "group velocity "

$$N = N_0 (1 + AC)$$

$$A = \delta \exp - \left(\frac{r - r_0 - z_0}{H} \right)^2 \cdot \exp - \left(\frac{\theta - \theta_0(t)}{\Theta} \right)^2 \cdot \exp - \left(\frac{\varphi}{\Phi} \right)^2$$

$$C = \cos 2\pi \left[t' + (\pi/2 - \theta) \frac{r_0}{\lambda_x} + (r - r_0)/\lambda_z \right]$$

$$\theta_0 = \theta_\infty + V_g t / r_0$$

r_0 is the radius of the earth.

r, θ, φ are spherical (earth-centered) polar coordinates.

$N_0(r, \theta, \varphi)$ is any electron density model.

Specify:

use perturbation _____ (W150 = 1.)

ignore perturbation _____ (W150 = 0.)

the height of maximum wave amplitude, $z_0 =$ _____ km (W151)

wave-amplitude "scale height," $H =$ _____ km (W152)

wave perturbation amplitude, $\delta =$ _____ (0 to 1) (W153)

horizontal trace velocity, $V_x =$ _____ km/sec. (W154)
 (needed only if Doppler shift is calculated)

horizontal wavelength, $\lambda_x =$ _____ km (W155)

vertical wavelength, $\lambda_z =$ _____ km (W156)

time in wave periods, $t' =$ _____ (W157)

amplitude "scale distance" in latitude, $\Theta =$ _____ degrees (W159)

amplitude "scale distance" in longitude, $\Phi =$ _____ degrees (W160)

latitude of maximum amplitude at $t = 0$, $\theta_\infty =$ _____ degrees (W158)

southward group velocity, $V_g =$ _____ km/sec (W161)
 (needed even if Doppler shift is not calculated)

	SUBROUTINE WAVE2	WAV2001
C	PERTURBATION TO AN ANY ELECTRON DENSITY MODEL	WAV2002
	COMMON /CONST/ PI,PIT2,PID2,DUM(5)	WAV2003
	COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT	WAV2004
	COMMON R(6) /WW/ ID(10),W0,W(400)	WAV2005
	EQUIVALENCE (EARTH,R,W(2)),(Z0,W(151)),(SH,W(152)),(DELTA,W(153)),	WAV2006
	1 (VSUBX,W(154)),(LAMBDAZ,W(155)),(LAMBDAZ,W(156)),(TP,W(157)),	WAV2007
	2 (TH00,W(158)),(THC,W(159)),(PHIC,W(160)),(VGX,W(161))	WAV2008
	REAL LAMBDAZ,LAMBDAZ	WAV2009
	DATA (MODX(2)=6H WAVE2)	WAV2010
	ENTRY ELECT1	WAV2011
	IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0..) RETURN	WAV2012
	IF (DELTA.EQ.0..OR.SH.EQ.0..) RETURN	WAV2013
	H=R(1)-EARTH	WAV2014
	TH0=TH00+LAMBDAZ*TP*VGX/VSUBX/EARTH	WAV2015
	EXPR=EXP(-((H-Z0)/SH)**2)	WAV2016
	EXPTH=EXP(-((R(2)-TH0)/THC)**2)	WAV2017
	EXPPHI=EXP(-((R(3)/PHIC)**2)	WAV2018
	WW=PIT2*(TP+(PID2-R(2))*EARTH/LAMBDAZ+H/LAMBDAZ)	WAV2019
	SINW=SIN(WW)	WAV2020
	COSW=COS(WW)	WAV2021
	E=DELTA*EXPR*EXPTH*EXPPHI	WAV2022
	CONS=1.0+E*COSW	WAV2023
	PXPR=PXPR*CONS-X*E*2.*(COSW*(H-Z0)/SH**2+PI/LAMBDAZ*SINW)	WAV2024
	PXPTH=PXPTH*CONS+2.*E*(X*PI*EARTH*SINW/LAMBDAZ-(R(2)-TH0)/	WAV2025
	1 THC**2*COSW)	WAV2026
	PXPPH=PXPPH*CONS-X*2.*E*R(3)/PHIC**2*COSW	WAV2027
	PXPT=-PIT2*VSUBX*E/LAMBDAZ*SINW+2.0*E*VGX/EARTH*COSW*(R(2)-TH0	WAV2028
	1 -LAMBDAZ*TP/EARTH)/THC	WAV2029
	X=X*CONS	WAV2030
	RETURN	WAV2031
	END	WAV2032-


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SUBROUTINE DOPPLER                                DOPP001
C   COMPUTES DN/DT                                FROM PROFILES HAVING THE SAME FORM DOPP002
C   AS THOSE USED BY SUBROUTINE TABLE X          DOPP003
C   MAKES AN EXPONENTIAL EXTRAPOLATION DOWN USING THE BOTTOM TWO POINTS DOPP004
C   NEEDS SUBROUTINE GAUSEL                       DOPP005
DIMENSION HPC(250),FN2C(250),ALPHA(250),BETA(250),GAMMA(250), DOPP006
1  DELTA(250),SLOPE(250),MAT(4,5)                DOPP007
COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,K,DUM(2)    DOPP008
COMMON /XX/ MODX(2),XDUM,PXPR,PXPTH,PXPPH,X,HMAX DOPP009
COMMON R(6) /WW/ ID(10),W0,W(400)               DOPP010
EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(READFN,W(151)) DOPP011
REAL MAT,K                                       DOPP012
DATA (MODX(2)=7HDOPPLER)                         DOPP013
ENTRY ELECT1                                     DOPP014
IF (READFN.EQ.0.) GO TO 10                       DOPP015
READFN=0.                                        DOPP016
READ 1000, NOC,(HPC(I),FN2C(I),I=1,NOC)          DOPP017
1000 FORMAT (I4/(F8.2,E12.4))                    DOPP018
PRINT 1200, (HPC(I),FN2C(I),I=1,NOC)            DOPP019
1200 FORMAT(1H1,14X,6HHEIGHT,4X,16H DN/DT      / (1X,F20.10,E20.1)) DOPP020
A=0.                                             DOPP021
IF (FN2C(1).NE.0.) A=ALOG (FN2C(2)/FN2C(1))/ (HPC(2)-HPC(1)) DOPP022
FN2C(1)=K*FN2C(1)                               DOPP023
FN2C(2)=K*FN2C(2)                               DOPP024
SLOPE(1)=A*FN2C(1)                              DOPP025
SLOPE(NOC)=0.                                    DOPP026
NMAX=1                                           DOPP027
DO 6 I=2,NOC                                     DOPP028
IF (FN2C(I).GT.FN2C(NMAX)) NMAX=I               DOPP029
IF (I.EQ.NOC) GO TO 4                           DOPP030
FN2C(I+1)=K*FN2C(I+1)                           DOPP031
DO 3 J=1,3                                       DOPP032
M=I+J-2                                          DOPP033
MAT(J,1)=1.                                      DOPP034
MAT(J,2)=HPC(M)                                  DOPP035
MAT(J,3)=HPC(M)**2                               DOPP036
MAT(J,4)=FN2C(M)                                 DOPP037
3 CALL GAUSEL (MAT,4,3,4,NRANK)                  DOPP038
IF (NRANK.LT.3) GO TO 60                         DOPP039
SLOPE(I)=MAT(2,4)+2.*MAT(3,4)*HPC(I)            DOPP040
4 DO 5 J=1,2                                     DOPP041
M=I+J-2                                          DOPP042
MAT(J,1)=1.                                      DOPP043
MAT(J,2)=HPC(M)                                  DOPP044
MAT(J,3)=HPC(M)**2                               DOPP045
MAT(J,4)=HPC(M)**3                               DOPP046
MAT(J,5)=FN2C(M)                                 DOPP047
L=J+2                                             DOPP048
MAT(L,1)=0.                                      DOPP049
MAT(L,2)=1.                                      DOPP050
MAT(L,3)=2.*HPC(M)                               DOPP051
MAT(L,4)=3.*HPC(M)**2                           DOPP052
5 MAT(L,5)=SLOPE(M)                              DOPP053
CALL GAUSEL (MAT,4,4,5,NRANK)                    DOPP054
IF (NRANK.LT.4) GO TO 60                         DOPP055
ALPHA(I)=MAT(1,5)                                DOPP056
BETA(I)=MAT(2,5)                                  DOPP057
GAMMA(I)=MAT(3,5)                                DOPP058
6 DELTA(I)=MAT(4,5)                              DOPP059
HMAX=AMAX1(HMAX,HPC(NMAX))                       DOPP060
NH=2                                              DOPP061
10 H=R(1)-EARTH R                                DOPP062
F2=F*F                                           DOPP063
IF (H.GE.HPC(1)) GO TO 12                       DOPP064
11 NH=2                                           DOPP065

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IF(FN2C(1).EQ.0.) GO TO 50	DOPP066
X=FN2C(1)*EXP(A*(H-HPC(1)))/F2	DOPP067
GO TO 50	DOPP068
12 IF (H.GE.HPC(NOC)) GO TO 18	DOPP069
NSTEP=1	DOPP070
IF (H.LT.HPC(NH-1)) NSTEP=-1	DOPP071
15 IF (HPC(NH-1).LE.H.AND.H.LT.HPC(NH)) GO TO 16	DOPP072
NH=NH+NSTEP	DOPP073
GO TO 15	DOPP074
16 X=(ALPHA(NH)+H*(BETA(NH)+H*(GAMMA(NH)+H*DELTA(NH)))/F2	DOPP075
GO TO 50	DOPP076
18 X=FN2C(NOC)/F2	DOPP077
50 CONTINUE	DOPP078
RETURN	DOPP079
60 PRINT 6000, I,HPC(I)	DOPP080
6000 FORMAT(4H THE,I4,55HTH POINT IN THE DN/DT PROFILE HAS	DOPP081
THE HEIGHT,F8.2,40H KM, WHICH IS THE SAME AS ANOTHER POINT.)	DOPP082
CALL EXIT	DOPP083
END	DOPP084
	DOPP085