

APPENDIX 3. ELECTRON DENSITY SUBROUTINES WITH INPUT PARAMETER FORMS

The following electron density models are available. The input parameter forms, which describe the model, and the subroutine listings are given on the pages shown.

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A further source of versatility in this ray tracing program is the ease with which specific ionospheric models, suited to the users needs, may be introduced. To add electron density models not included in the program, the user must write a subroutine that calculates the normalized electron density (X) and its gradient ($\partial X/\partial r$, $\partial X/\partial \theta$, $\partial X/\partial \varphi$) as a function of position in spherical coordinates (r , θ , φ). ($X = 80.5 \times 10^{-6} N/f^2$, where N is the electron density in cm^{-3} and f is the wave frequency in MHz.)

Both X and its gradient must be continuous functions of position. The formulas for $\partial X/\partial r$, $\partial X/\partial \theta$, and $\partial X/\partial \varphi$ must be consistent with the variation of X with r , θ , and φ . Otherwise, the program will run slowly and give incorrect results.

The coordinates r , θ , φ refer to the computational coordinate system, which may not be the same as geographic coordinates. In particular, they are geomagnetic coordinates when the earth-centered dipole model of the earth's magnetic field is used.

The input to the subroutine (r , θ , φ) is through blank common. (See

Table 3.) The output is through common block /XX/. (See Table 8.) It is useful if the name of the subroutine suggests the model to which it corresponds. The subroutine should have an entry point ELECTX so that other subroutines in the program can call it. Any parameters needed by the subroutine should be input into W101 through W149 of the W array. (See Table 2.) If the model needs massive amounts of data, these should be read in by the subroutine following the example of TABLEX. As in the already existing electron density subroutines, provision should be made for perturbations to the electron density model (irregularities) by having the statement

```
IF(PERT.NE.0.) CALL ELECT1
```

before the RETURN statement at the end of the subroutine.

	SUBROUTINE TABLEX	TABX001
C	CALCULATES ELECTRON DENSITY AND GRADIENT FROM PROFILES HAVING	TABX002
C	THE SAME FORM AS THOSE USED BY CROFTS RAY TRACING PROGRAM	TABX003
C	MAKES AN EXPONENTIAL EXTRAPOLATION DOWN USING THE BOTTOM TWO POINTS	TABX004
C	NEEDS SUBROUTINE GAUSEL	TABX005
	DIMENSION HPC(250),FN2C(250),ALPHA(250),BETA(250),GAMMA(250),	TABX006
	1 DELTA(250),SLOPE(250),MAT(4,5)	TABX007
	COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,K,DUM(2)	TABX008
	COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	TABX009
	COMMON R(6) /WW/ ID(10),W0,W(400)	TABX010
	EQUIVALENCE (EARTH,W(2)),(F,W(6)),(READFN,W(100)),(PERT,W(150))	TABX011
	REAL MAT,K	TABX012
	DATA (MODX(1)=6HTABLEX)	TABX013
	ENTRY ELECTX	TABX014
	IF (READFN.EQ.0.) GO TO 10	TABX015
	READFN=0.	TABX016
	READ 1000, NOC,(HPC(I),FN2C(I),I=1,NOC)	TABX017
1000	FORMAT (I4/(F8.2,E12.4))	TABX018
	PRINT 1200, (HPC(I),FN2C(I),I=1,NOC)	TABX019
1200	FORMAT (1H1,14X,6HHEIGHT,4X,16HELECTRON DENSITY/(1X,F20.10,E20.10))	TABX020
	A=0.	TABX021
	IF (FN2C(1).NE.0.) A=ALOG (FN2C(2)/FN2C(1))/(HPC(2)-HPC(1))	TABX022
	FN2C(1)=K*FN2C(1)	TABX023
	FN2C(2)=K*FN2C(2)	TABX024
	SLOPE(1)=A*FN2C(1)	TABX025
	SLOPE(NOC)=0.	TABX026
	NMAX=1	TABX027
	DO 6 I=2,NOC	TABX028
	IF (FN2C(I).GT.FN2C(NMAX)) NMAX=I	TABX029
	IF (I.EQ.NOC) GO TO 4	TABX030
	FN2C(I+1)=K*FN2C(I+1)	TABX031
	DO 3 J=1,3	TABX032
	M=I+J-2	TABX033
	MAT(J,1)=1.	TABX034
	MAT(J,2)=HPC(M)	TABX035
	MAT(J,3)=HPC(M)**2	TABX036
3	MAT(J,4)=FN2C(M)	TABX037
	CALL GAUSEL (MAT,4,3,4,NRANK)	TABX038
	IF (NRANK.LT.3) GO TO 60	TABX039
	SLOPE(I)=MAT(2,4)+2.*MAT(3,4)*HPC(I)	TABX040
4	DO 5 J=1,2	TABX041
	M=I+J-2	TABX042
	MAT(J,1)=1.	TABX043
	MAT(J,2)=HPC(M)	TABX044
	MAT(J,3)=HPC(M)**2	TABX045
	MAT(J,4)=HPC(M)**3	TABX046
	MAT(J,5)=FN2C(M)	TABX047
	L=J+2	TABX048
	MAT(L,1)=0.	TABX049
	MAT(L,2)=1.	TABX050
	MAT(L,3)=2.*HPC(M)	TABX051
	MAT(L,4)=3.*HPC(M)**2	TABX052
5	MAT(L,5)=SLOPE(M)	TABX053
	CALL GAUSEL (MAT,4,4,5,NRANK)	TABX054
	IF (NRANK.LT.4) GO TO 60	TABX055
	ALPHA(I)=MAT(1,5)	TABX056
	BETA(I)=MAT(2,5)	TABX057
	GAMMA(I)=MAT(3,5)	TABX058
6	DELTA(I)=MAT(4,5)	TABX059
	HMAX=HPC(NMAX)	TABX060
	NH=2	TABX061
10	H=R(1)-EARTH	TABX062
	F2=F*F	TABX063
	PXPR=0.	TABX064
	IF (H.GE.HPC(1)) GO TO 12	TABX065
11	NH=2	TABX066
	X=0.	TABX067

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IF(FN2C(1).EQ.0.) GO TO 50
X=FN2C(1)*EXP(A*(H-HPC(1)))/F2
PXPR=A*X
GO TO 50
12 IF (H.GE.HPC(NOC)) GO TO 18
NSTEP=1
IF (H.LT.HPC(NH-1)) NSTEP=-1
15 IF (HPC(NH-1).LE.H.AND.H.LT.HPC(NH)) GO TO 16
NH=NH+NSTEP
GO TO 15
16 X=(ALPHA(NH)+H*(BETA(NH)+H*(GAMMA(NH)+H*DELTA(NH)))/F2
PXPR=(BETA(NH)+H*(2.*GAMMA(NH)+H*3.*DELTA(NH)))/F2
GO TO 50
18 X=FN2C(NOC)/F2
50 IF (PERT.NE.0.) CALL ELECT1
RETURN
60 PRINT 6000, I,HPC(I)
6000 FORMAT(4H THE,I4,55H TH POINT IN THE ELECTRON DENSITY PROFILE HAS T
1HE HEIGHT,F8.2,40H KM, WHICH IS THE SAME AS ANOTHER POINT.)
CALL EXIT
END

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TABX068
TABX069
TABX070
TABX071
TABX072
TABX073
TABX074
TABX075
TABX076
TABX077
TABX078
TABX079
TABX080
TABX081
TABX082
TABX083
TABX084
TABX085
TABX086
TABX087
TABX088

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SUBROUTINE GAUSEL (C,NRD,NRR,NCC,NSF)
C***** SAME AS SUBROUTINE GAUSSEL WRITTEN BY L. DAVID LEWIS *****
DIMENSION C(NRD,NCC),L(128,2)
C      BITS = 2.**-18
DATA (BITS=3.8146972656E-6)
NR=NRR
NC=NCC
IF(NC.LT.NR.OR.NR.GT.128.OR.NR.LE.0) CALL EXIT
C
C      INITIALIZE.
NSF=0
NRM=NR-1
NRP=NR+1
D=1.
LSD=1
DO 1 KR=1,NR
L(KR,1)=KR
1 L(KR,2)=0
IF(NR.EQ.1) GO TO 42
C
C      ELIMINATION PHASE.
DO 41 KP=1,NRM
KPP=KP+1
PM=0.
MPN=0
C
C      SEARCH COLUMN KP FROM DIAGONAL DOWN FOR MAX PIVOT.
DO 2 KR=KP,NR
LKR=L(KR,1)
PT=ABS(C(LKR,KP))
IF(PT.LE.PM) GO TO 2
PM=PT
MPN=KR
LMP=LKR
2 CONTINUE
C
C      IF MAX PIVOT IS ZERO, MATRIX IS SINGULAR.
IF(MPN.EQ.0) GO TO 9
NSF=NSF+1
IF(MPN.EQ.KP) GO TO 3

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GAUS001
GAUS002
GAUS003
GAUS004
GAUS005
GAUS006
GAUS007
GAUS008
GAUS009
GAUS010
GAUS011
GAUS012
GAUS013
GAUS014
GAUS015
GAUS016
GAUS017
GAUS018
GAUS019
GAUS020
GAUS021
GAUS022
GAUS023
GAUS024
GAUS025
GAUS026
GAUS027
GAUS028
GAUS029
GAUS030
GAUS031
GAUS032
GAUS033
GAUS034
GAUS035
GAUS036
GAUS037
GAUS038
GAUS039
GAUS040

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C		GAUS041
C	NEW ROW NUMBER KP HAS MAX PIVOT.	GAUS042
	LSD=-LSD	GAUS043
	L(KP,2)=L(MPN,1)=L(KP,1)	GAUS044
	L(KP,1)=LMP	GAUS045
C		GAUS046
C	ROW OPERATIONS TO ZERO COLUMN KP BELOW DIAGONAL.	GAUS047
3	MKP=L(KP,1)	GAUS048
	P=C(MKP,KP)	GAUS049
	D=D*P	GAUS050
	DO 41 KR=KPP, NR	GAUS051
	MKR=L(KR,1)	GAUS052
	Q=C(MKR,KP)/P	GAUS053
	IF(Q.EQ.0.) GO TO 41	GAUS054
C		GAUS055
C	SUBTRACT Q * PIVOT ROW FROM ROW KR.	GAUS056
	DO 4 LC=KPP, NC	GAUS057
	R=Q*C(MKP,LC)	GAUS058
	C(MKR,LC)=C(MKR,LC)-R	GAUS059
4	IF(ABS(C(MKR,LC)).LT.ABS(R)*BITS) C(MKR,LC)=0.	GAUS060
41	CONTINUE	GAUS061
C		GAUS062
C	LOWER RIGHT HAND CORNER.	GAUS063
42	LNR=L(NR,1)	GAUS064
	P=C(LNR, NR)	GAUS065
	IF(P.EQ.0.) GO TO 9	GAUS066
	NSF=NSF+1	GAUS067
	D=D*P*LSD	GAUS068
	IF(NR.EQ.NC) GO TO 8	GAUS069
C		GAUS070
C	BACK SOLUTION PHASE.	GAUS071
	DO 61 MC=NRP, NC	GAUS072
	C(LNR, MC)=C(LNR, MC)/P	GAUS073
	IF(NR.EQ.1) GO TO 61	GAUS074
	DO 6 LL=1, NRM	GAUS075
	KR=NR-LL	GAUS076
	MR=L(KR,1)	GAUS077
	KRP=KR+1	GAUS078
	DO 5 MS=KRP, NR	GAUS079
	LMS=L(MS,1)	GAUS080
	R=C(MR, MS)*C(LMS, MC)	GAUS081
	C(MR, MC)=C(MR, MC)-R	GAUS082
5	IF(ABS(C(MR, MC)).LT.ABS(R)*BITS) C(MR, MC)=0.	GAUS083
6	C(MR, MC)=C(MR, MC)/C(MR, KR)	GAUS084
61	CONTINUE	GAUS085
C		GAUS086
C	SHUFFLE SOLUTION ROWS BACK TO NATURAL ORDER.	GAUS087
	DO 71 LL=1, NRM	GAUS088
	KR=NR-LL	GAUS089
	MKR=L(KR,2)	GAUS090
	IF(MKR.EQ.0) GO TO 71	GAUS091
	MKP=L(KR,1)	GAUS092
	DO 7 LC=NRP, NC	GAUS093
	Q=C(MKR, LC)	GAUS094
	C(MKR, LC)=C(MKP, LC)	GAUS095
7	C(MKP, LC)=Q	GAUS096
71	CONTINUE	GAUS097
C		GAUS098
C	NORMAL AND SINGULAR RETURNS. GOOD SOLUTION COULD HAVE D=0.	GAUS099
8	C(1,1)=D	GAUS100
	GO TO 91	GAUS101
9	C(1,1)=0.	GAUS102
91	RETURN	GAUS103
	END	GAUS104-

INPUT PARAMETER FORM FOR SUBROUTINE CHAPX

An ionospheric electron density model consisting of a Chapman layer with tilts, ripples, and gradients

$$f_N^2 = f_c^2 \exp\left(\alpha(1-z-e^{-z})\right)$$

$$z = \frac{h - h_{\max}}{H}$$

$$f_c^2 = f_{c0}^2 \left(1 + A \sin\left(2\pi\left(\theta - \frac{\pi}{2}\right)/B\right) + C\left(\theta - \frac{\pi}{2}\right)\right)$$

$$h_{\max} = h_{\max 0} + E\left(\theta - \frac{\pi}{2}\right) R_0$$

f_N is the plasma frequency

h is the height above the ground

R_0 is the radius of the earth in km

and θ is the colatitude in radians.

Specify:

Critical frequency at the equator, $f_{c0} =$ _____ MHz (W101)

Height of the maximum electron density at the equator, $h_{\max 0} =$ _____ km (W102)

Scale height, $H =$ _____ km (W103)

$\alpha =$ _____ (W104, 0.5 for an α Chapman layer, 1.0 for a β Chapman layer)

Amplitude of periodic variation of f_c^2 with latitude, $A =$ _____ (W105)

Period of variation of f_c^2 with latitude, $B =$ _____ $\frac{\text{rad}}{\text{deg}}$ (W106)
km

Coefficient of linear variation of f_c^2 with latitude, $C =$ _____ rad^{-1} (W107)

Tilt of the layer, $E =$ _____ $\frac{\text{rad}}{\text{deg}}$ (W108)

	SUBROUTINE CHAPX	CHAP001
	CHAPMAN LAYER WITH TILTS, RIPPLES, AND GRADIENTS	CHAP002
C	W(104) = 0.5 FOR AN ALPHA-CHAPMAN LAYER	CHAP003
C	= 1.0 FOR A BETA-CHAPMAN LAYER	CHAP004
	COMMON /CONST/ PI,PIT2,PID2,DUM(5)	CHAP005
	COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	CHAP006
	COMMON R(6) /WW/ ID(10),W0,W(400)	CHAP007
	EQUIVALENCE (THETA,R(2))	CHAP008
	EQUIVALENCE (EARTHTR,W(2)),(F,W(6)),(FC,W(101)),(HM,W(102)),	CHAP009
	1 (SH,W(103)),(ALPHA,W(104)),(A,W(105)),(B,W(106)),(C,W(107)),	CHAP010
	2 (E,W(108)),(PERT,W(150))	CHAP011
	DATA (MODX(1)=6H CHAPX)	CHAP012
	ENTRY ELECTX	CHAP013
	THETA2=THETA-PID2	CHAP014
	HMAX=HM+EARTHTR*E*THETA2	CHAP015
	H=R(1)-EARTHTR	CHAP016
	Z=(H-HMAX)/SH	CHAP017
	D=0.	CHAP018
	IF (B.NE.0.) D=PIT2/B	CHAP019
	TEMP=1.+A*SIN(D*THETA2)+C*THETA2	CHAP020
	EXZ=1.-EXP(-Z)	CHAP021
	X=(FC/F)**2*TEMP*EXP(ALPHA*(EXZ-Z))	CHAP022
	PXPR=-ALPHA*X*EXZ/SH	CHAP023
	PXPTH=X*(D*A*SIN(PID2-D*THETA2)+C)/TEMP-PXPR*EARTHTR*E	CHAP024
	IF (PERT.NE.0.) CALL ELECT1	CHAP025
	RETURN	CHAP026
	END	CHAP 27-

INPUT PARAMETER FORM FOR SUBROUTINE VCHAPX

An ionospheric electron density model consisting of a Chapman layer with variable scale height

$$f_N^2 = f_c^2 \tau^{\frac{1}{2}} e^{\frac{1}{2}} (1 - \tau)$$

$$\tau = \left(\frac{h_{\max}}{h} \right)^\chi$$

h is the height above the ground.

Specify:

critical frequency, $f_c =$ _____ MHz (W101)

height of maximum electron density, $h_{\max} =$ _____ km (W102)

$\chi =$ _____ (W103)

<pre> SUBROUTINE VCHAPX CHAPMAN LAYER WITH VARIABLE SCALE HEIGHT COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX COMMON R(6) /WW/ ID(10),W0,W(400) EQUIVALENCE (EARTH,R(2)),(F,W(6)),(FC,W(101)),(HM,W(102)), 1 (CHI,W(103)),(PERT,W(150)) DATA (MODX(1)=6HVCHAPX) ENTRY ELECTX HMAX=HM X=PXPR*0. H=R(1)-EARTH IF (H.LE.0.) GO TO 50 TAU=(HM/H)**CHI X=(FC/F)**2*SQRT(TAU)*EXP(0.5*(1.-TAU)) PXPR=.5*X*(TAU-1.)*CHI/H 50 IF (PERT.NE.0.) CALL ELECT1 RETURN END </pre>	<pre> VCHA001 VCHA002 VCHA003 VCHA004 VCHA005 VCHA006 VCHA007 VCHA008 VCHA009 VCHA010 VCHA011 VCHA012 VCHA013 VCHA014 VCHA015 VCHA016 VCHA017 VCHA018- </pre>
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INPUT PARAMETER FORM FOR SUBROUTINE DCHAPT

An ionospheric electron density model consisting of a double, tilted α -Chapman layer

$$f_N^2 = f_{c1}^2 \exp \frac{1}{2} (1 - z_1 - e^{-z_1}) + f_{c2}^2 \exp \frac{1}{2} (1 - z_2 - e^{-z_2})$$

$$z_1 = \frac{h - h_{m1}}{H_1} \quad ; \quad z_2 = \frac{h - h_{m2}}{H_2}$$

$$f_{c1}^2 = f_{c10}^2 C(\theta - \pi/2)$$

$$f_{c2}^2 = f_{c20}^2 C(\theta - \pi/2)$$

$$h_{m1} = h_{m10} + R_o E \left(\frac{\pi}{180} \right) \left(\theta - \frac{\pi}{2} \right)$$

$$h_{m2} = h_{m20} + R_o E \left(\frac{\pi}{180} \right) \left(\theta - \frac{\pi}{2} \right)$$

Specify:

$$f_{c10} = \text{_____} \text{MHz (} f_{c1} \text{ at equator)} \quad \text{(W101)}$$

$$h_{m10} = \text{_____} \text{Km (} h_{m1} \text{ at equator)} \quad \text{(W102)}$$

$$H_1 = \text{_____} \text{Km} \quad \text{(W103)}$$

$$f_{c20} = \text{_____} \text{MHz (} f_{c2} \text{ at equator)} \quad \text{(W104)}$$

$$h_{m20} = \text{_____} \text{Km (} h_{m2} \text{ at equator)} \quad \text{(W105)}$$

$$H_2 = \text{_____} \text{Km} \quad \text{(W106)}$$

$$C = \text{_____} \text{rad}^{-1} \text{ (fractional change in } f_{c1}, f_{c2}, \text{ position for increases southward)} \quad \text{(W107)}$$

$$E = \text{_____} \text{deg (positive for upward tilt to the south)} \quad \text{(W108)}$$

SUBROUTINE DCHAPT	DCHA001
TWO CHAPMAN LAYERS WITH TILTS	DCHA002
COMMON /CONST/ PI,PIT2,PID2,DUM(5)	DCHA003
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	DCHA004
COMMON R(6) /WW/ ID(10),WO,W(400)	DCHA005
EQUIVALENCE (EARTH,W(2)),(F,W(6)),(FC1,W(101)),(HM1,W(102)),	DCHA006
1 (SH1,W(103)),(FC2,W(104)),(HM2,W(105)),(SH2,W(106)),(C,W(107)),	DCHA007
2 (E,W(108)),(PERT,W(150))	DCHA008
DATA (MODX(1)=6HDCHAPT)	DCHA009
ENTRY ELECTX	DCHA010
EARTHE=EARTH*E	DCHA011
THETA2=R(2)-PID2	DCHA012
HMAX=HM1+EARTHE*THETA2	DCHA013
X=PXPR=PXPTH=0.	DCHA014
H=R(1)-EARTH	DCHA015
IF (H.LT.0.) GO TO 50	DCHA016
Z1=(H-HMAX)/SH1	DCHA017
EXPZ1=1.-FXP(-Z1)	DCHA018
TEMP=1.+C*THETA2	DCHA019
X=(FC1/F)**2*TEMP*EXP(.5*(EXPZ1-Z1))	DCHA020
PXPR=-0.5*X*EXPZ1/SH1	DCHA021
PXPTH=X*C/TEMP-PXPR*EARTHE	DCHA022
IF (FC2.EQ.0.) GO TO 50	DCHA023
Z2=(H-HM2-EARTHE*THETA2)/SH2	DCHA024
EXPZ2=1.-EXP(-Z2)	DCHA025
X2=(FC2/F)**2*TEMP*EXP(.5*(EXPZ2-Z2))	DCHA026
X=X+X2	DCHA027
PXPR2=-0.5*X2*EXPZ2/SH2	DCHA028
PXPR=PXPR+PXPR2	DCHA029
PXPTH=PXPTH+X2*C/TEMP-PXPR2*EARTHE	DCHA030
50 RETURN	DCHA031
END	DCHA032
	DCHA033-

INPUT PARAMETER FORM FOR SUBROUTINE LINEAR

An ionospheric electron density model consisting of a linear layer

$$N = 0 \quad \text{for } h \leq h_{\min}$$

$$N = A(h - h_{\min}) \quad \text{for } h > h_{\min}$$

The ray will penetrate if $h > h_{\max}$.

Specify:

A = _____ electrons/cm³/ km (W101)

h_{\max} = _____ km (W102)

h_{\min} = _____ km (W103)

	<pre> SUBROUTINE LINEAR LINEAR ELECTRON DENSITY MODEL COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,K,DUM(2) COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX COMMON R(6) /WW/ ID(10),W0,W(400) EQUIVALENCE (EARTH,W(2)),(F,W(6)),(FACT,W(101)),(HM,W(102)), 1 (HMIN,W(103)),(PERT,W(150)) REAL K DATA (MODX(1)=6HLINEAR) ENTRY ELECTX H=R(1)-EARTH HMAX=HM X=PXPR=0. IF (H.LE.HMIN) GO TO 50 PXPR=K*FACT/F**2 X=PXPR*(H-HMIN) 50 IF (PERT.NE.0.) CALL ELECT1 RETURN END </pre>	<pre> LINE001 LINE002 LINE003 LINE004 LINE005 LINE006 LINE007 LINE008 LINE009 LINE010 LINE011 LINE012 LINE013 LINE014 LINE015 LINE016 LINE017 LINE018 LINE019- </pre>
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INPUT PARAMETER FORM FOR SUBROUTINE QPARAB

An ionospheric electron density model consisting of a parabolic or a quasi-parabolic layer (concentric)

$$f_N^2 = f_c^2 \left[1 - \frac{h-h_{\max}}{Y_m} \cdot C^2 \right] \quad \text{if } f_N^2 > 0.$$

$$f_N^2 = 0. \quad \text{otherwise.}$$

$C = 1.$ for a parabolic layer

$$C = \frac{R_o + h_{\max} - Y_m}{R_o + h} \quad \text{for a quasi-parabolic layer}$$

where R_o is the radius of the earth.

Specify:

Critical frequency, $f_c =$ _____ Mc/s (W101)

Height of maximum electron density, $h_{\max} =$ _____ km. (W102)

Semi-thickness, $Y_m =$ _____ km. (W103)

Type of profile:

Plain parabolic _____ (W104 = 0.)

Quasi-parabolic _____ (W104 = 1.)

	SUBROUTINE QPARAB	PARA001
C	PLAIN PARABOLIC OR QUASI-PARABOLIC PROFILE	PARA002
C	W(104) = 0. FOR A PLAIN PARABOLIC PROFILE	PARA003
C	= 1. FOR A QUASI-PARABOLIC PROFILE	PARA004
	COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	PARA005
	COMMON R(6) /WW/ ID(10),W0,W(400)	PARA006
	EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(FC,W(101)),(HM,W(102)),	PARA007
	1 (YM,W(103)),(QUASI,W(104)),(PERT,W(150))	PARA008
	DATA (MODX(1)=6HQPAB)	PARA009
	ENTRY ELECTX	PARA010
	HMAX=HM	PARA011
	PXPR=0.	PARA012
	H=R(1)-EARTH	PARA013
	FCF2=(FC/F)**2	PARA014
	CONST=1.	PARA015
	IF (QUASI.EQ.1.) CONST=(EARTH+HM-YM)/R(1)	PARA016
	Z=(H-HM)/YM*CONST	PARA017
	X=MAX1F(0.,FCF2*(1.-Z*Z))	PARA018
	IF (X.EQ.0.) GO TO 50	PARA019
	IF (QUASI.EQ.1.) CONST=(EARTH+HM)*(EARTH+HM-YM)/R(1)**2	PARA020
	PXPR=-2.*Z*FCF2/YM*CONST	PARA021
50	IF (PERT.NE.0.) CALL ELECT1	PARA022
	RETURN	PARA023
	END	PARA024-

INPUT PARAMETER FORM FOR SUBROUTINE BULGE

An analytic ionospheric electron density model which represents the general latitude variation of the equatorial ionosphere (afternoon, equinox, sunspot maximum) - see the center panel of figure 3.18b, page 133 of Davies (1965).

The model is an alpha Chapman layer with parameters which vary with geomagnetic latitude.

$$f_N^2 = f_c^2 e^{\frac{1}{2}(1-z-e^{-z})}$$

$$\text{where } z = \frac{h - h_{\max}}{H}$$

f_N is the plasma frequency

f_c is the critical frequency

h_{\max} is the height of the maximum electron density

H is the scale height

h is height

f_c , h_{\max} , H vary with geomagnetic latitude in the following way:

if $h < 100$ km, $h_{\max} = 350$ km, $f_c = 15$ Mc/s

For $h \geq 100$ km,

$h_{\max} = 350$ if $\lambda \geq 24^\circ$

$h_{\max} = 430 + 80 \cos\left(\frac{180}{24} \lambda\right)$ if $\lambda < 24^\circ$

λ is the geomagnetic latitude in degrees

$$f_c = \sqrt{50 \left(\frac{\lambda}{8}\right)^2 \exp\left(2 - \left|\frac{\lambda}{8}\right|\right) + 40}$$

In all cases H is determined by the constraint that

$$f_N = 2 \text{ Mc/s at } 100 \text{ km.}$$

	SUBROUTINE BULGE	BULG001
C	ANALYTICAL MODEL OF THE VARIATION OF THE EQUATORIAL F2 LAYER	BULG002
C	IN GEOMAGNETIC LATITUDE (EQUATORIAL BULGE AND ANOMALY)	BULG003
C	SEE FIGURE 3.18B, PAGE 133 IN DAVIES (1965).	BULG004
C	THIS MODEL HAS NO VARIATION IN GEOMAGNETIC LONGITUDE.	BULG005
	COMMON /CONST/ PI,PID2,PID2,DUM(5)	BULG006
	COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	BULG007
	COMMON R(6) /WW/ ID(10),W0,W(400)	BULG008
	EQUIVALENCE (EARTH,W(2)),(F,W(6)),(PERT,W(150))	BULG009
	DATA (MODX(1)=6H BULGE)	BULG010
	ENTRY ELECTX	BULG011
	H=R(1)-EARTH	BULG012
	PHMPH=PFC2PTH=0.	BULG013
	HMAX=350.	BULG014
	FC2=225.	BULG015
	IF(H.LT.100.) GO TO 2	BULG016
C	EQUATORIAL BULGE	BULG017
	BULLAT=7.5*(PID2-R(2))	BULG018
	IF(ABS(BUL LAT).GE.PI) GO TO 1	BULG019
	HMAX=430.+80.*COS(BULLAT)	BULG020
	PHMPH=600.*SIN(BUL LAT)	BULG021
C	EQUATORIAL ANOMALY	BULG022
1	ANMLAT=22.5*(PID2-R(2))/PI	BULG023
	POW=2.-ABS(ANM LAT)	BULG024
	FC2=50.*ANM LAT**2*EXP(POW) + 40.	BULG025
	PFC2PTH=-1125./PI*POW*ANMLAT*EXP(POW)	BULG026
C	FORCING PLASMA FREQ AT 100 KM TO BE 2 MHZ IN ORDER TO CALCULATE SH	BULG027
2	ALPHA=2.*ALOG(FC2/4.)+1.	BULG028
	Z100=-ALOG(ALPHA)	BULG029
	DO 3 I=1,5	BULG030
3	Z100=-ALOG(ALPHA-Z100)	BULG031
	SH=(100.-HMAX)/Z100	BULG032
	Z=(H-HMAX)/SH	BULG033
	EXZ=1.-EXP(-Z)	BULG034
	X=FC2*EXP(.5*(EXZ-Z))/F**2	BULG035
	PXPR=-0.5*X*EXZ/SH	BULG036
	PXPTH=-PXPR*(1.-Z/Z100)*PHMPH+(1.-Z*EXZ/(Z100*(1.-EXP(-Z100))))	BULG037
1	*X/FC2*PFC2PTH	BULG038
	IF (PERT.NE.0.) CALL ELECT1	BULG039
	RETURN	BULG040
	END	BULG041-

INPUT PARAMETER FORM FOR SUBROUTINE EXPX

An exponential electron density profile

$$N = N_0 e^{a(h-h_0)}$$

h is the height above the ground.

Specify:

the electron density at the height h_0 , $N_0 =$ _____ cm^{-3} (W101)

the reference height, $h_0 =$ _____ km (W102)

the exponential increase of N with height, $a =$ _____ km^{-1} (W103)

	SUBROUTINE EXPX	EXPX001
C	EXPONENTIAL ELECTRON DENSITY MODEL	EXPX002
	COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,K,DUM(2)	EXPX003
	COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	EXPX004
	COMMON R(6) /WW/ ID(10),W0,W(400)	EXPX005
	EQUIVALENCE (EARTH,W(2)),(F,W(6)),	EXPX006
	1 (N0,W(101)),(H0,W(102)),(A,W(103)),(PERT,W(150))	EXPX007
	REAL N, NO *K	EXPX008
	DATA (MODX(1)=4HEXPX),(HMAX=350.)	EXPX009
	ENTRY ELECTX	EXPX010
	H=R(1)-EARTH	EXPX011
	N=NO * EXP(A*(H-H0))	EXPX012
	X=K*N/F**2	EXPX013
	PXPR=A*X	EXPX014
	IF (PERT.NE.0.) CALL ELECT1	EXPX015
	RETURN	EXPX016
	END	EXPX017-