

## UARS measurements of ClO and NO<sub>2</sub> at 40 and 46 km and implications for the model "ozone deficit"

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**Abstract.** We present the first diurnally-resolved measurements of ClO and NO<sub>2</sub> at 3.1 and 1.4 hPa, approximately 40 and 46 km. Comparisons between the measurements of ClO and a zero-dimensional box model utilizing standard photochemistry and constrained by Upper Atmosphere Research Satellite (UARS) tracer measurements show that photochemical models overpredict ClO by a factor of two or more at 3.1 hPa and by a factor of 1.5-2.0 at 1.4 hPa. NO<sub>2</sub> is well simulated by the model at 3.1 hPa, but systematically overestimated at 1.4 hPa. The model overestimate of ClO results in an overestimate of ozone loss in the upper stratosphere of 10-40%.

### Introduction

Because of the short lifetime of ozone (O<sub>3</sub>) in the upper stratosphere, transport can be neglected and the assumption can be made that production and loss balance when averaged over 24-hours [e.g., Eluszkiewicz and Allen, 1993, Fig. 1]. Despite the relative simplicity afforded by the neglect of transport, early attempts to model O<sub>3</sub> in this region predicted 40% to 60% lower ozone than was measured [e.g., Froidevaux et al., 1985, Jackman et al., 1986], a fact which became known as the "ozone deficit" or the "40-km problem".

Most analyses of the discrepancy between modeled and measured upper stratospheric O<sub>3</sub> used atmospheric observations to constrain the radicals and/or radical precursors in a model and either compare the resulting O<sub>3</sub> prediction with observations or compare production and loss of O<sub>3</sub>. Analyses constrained by LIMS measurements generally underpredict O<sub>3</sub>. The most recent studies find that model O<sub>3</sub> is ~10 to 25% less than measurements [Eluszkiewicz and Allen, 1993, Natarajan and Callis, 1989, Siskind et al., 1995] in the 40 and 45 km region. Analyses constrained by ATMOS measurements [Minschwaner et al., 1993, Natarajan and Callis, 1989] predict model O<sub>3</sub> in agreement with observations at 40 km, with a 10-20% deficit at 46 km. Analyses incorporating UARS HALOE measurements [Cruzen et al., 1995] predict model O<sub>3</sub> within a few percent of observations at 40 and 46 km.

In this paper, we investigate issues related to the ozone deficit at ~40 and 46 km using measurements of ClO and

NO<sub>2</sub> made throughout the day. This represents an advance because ClO and NO<sub>2</sub> are two of the three radicals that *directly* determine the rates of the ozone-destroying catalytic cycles.

### UARS NO<sub>2</sub> and ClO Measurements

Measurements of ClO were made by the Microwave Limb Sounder (MLS), and minor adjustments have been applied to the version 3 data according to Waters et al. [1996]. Over the altitude range considered here, the accuracy of the ClO data is estimated to be the root-sum-of-squares of an 8% scaling uncertainty and a 75-pptv bias uncertainty (all uncertainties discussed in this paper are ~68% confidence limits, i.e., "1σ"). This translates into accuracy estimates of ±20-25% for these data.

Measurements of the NO<sub>2</sub> were made by the Cryogenic Limb Array Etalon Spectrometer (CLAES). Estimation of the accuracy of the CLAES NO<sub>2</sub> measurement is more problematic because the version 7 CLAES NO<sub>2</sub> data are not validated. However, the HALOE instrument, also aboard UARS, measures sunrise and sunset NO and NO<sub>2</sub>, the sum of which is NO<sub>x</sub>. These measurements can be used to calibrate the CLAES NO<sub>2</sub> measurements throughout the day [see the detailed discussion of this procedure in Dessler, A. E. et al., "A test of the partitioning between ClO and ClONO<sub>2</sub> using simultaneous UARS measurements of ClO, NO<sub>2</sub>, and ClONO<sub>2</sub>", *J. Geophys. Res.*, submitted, 1996]. Briefly, we increase CLAES NO<sub>2</sub> measurements throughout the day by a factor (1.16 to 1.67, depending on latitude and altitude) such that CLAES NO<sub>2</sub> measurements made just after sunset agree with HALOE

Table 1. Inputs to Zero-Dimensional Model

	Source*	0°N		35°N		55°N	
		3.1 hPa	1.4 hPa	3.1 hPa	1.4 hPa	3.1 hPa	1.4 hPa
O <sub>3</sub> (ppmv)	MLS	7.6	3.9	7.1	4.4	6.5	4.7
	HALOE	6.5	3.1	6.5	3.9	6.4	4.2
NO <sub>y</sub> (ppbv)	HALOE NO <sub>x</sub> (see text)	15.4	12.9	12.1	10.9	12.0	10.7
Cl <sub>y</sub> (ppbv)	HALOE HCl and MLS ClO (see text)	3.1	3.4	3.2	3.2	3.1	3.0
H <sub>2</sub> O (ppmv)	Avg. of MLS and HALOE	5.0	5.8	5.8	5.9	5.8	6.1
CH <sub>4</sub> (ppmv)	HALOE	0.96	0.47	0.64	0.50	0.55	0.40
Temp (K)	Avg. of MLS and CLAES	248	266	248	260	241	252

\*These data were obtained between Feb. 12 and March 16, 1993 and averaged over the latitude bands 0°±5°, 35°N±5°, and 55°N±5°, except for 0° HALOE data which are averaged over 0°±10°. The data have been interpolated onto the 3.1 and 1.4 hPa pressure surfaces.

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NO<sub>x</sub> measurements at sunset. In this way, the calibration from HALOE is combined with the diurnal cycle information of CLAES to produce an adjusted NO<sub>2</sub> data set. Combining the uncertainty in the HALOE NO and NO<sub>2</sub> data ( $\pm 15\%$  each [Gordley *et al.*, 1996]) with the uncertainty in the magnitude of the CLAES NO<sub>2</sub> diurnal cycle ( $\pm 25\%$  [Dessler *et al.*, *submitted*]) as a root-sum-of-squares yields a total accuracy estimate for the adjusted NO<sub>2</sub> data of  $\pm 34\%$ .

Model diurnal profiles of ClO and NO<sub>2</sub> are obtained from a zero-dimensional model incorporating the chemical integrator and photolysis code from the Goddard three-dimensional Chemical Transport Model (3D CTM) [Kawa *et al.*, 1995]. This model includes the photochemistry listed in the JPL-94 report [DeMore *et al.*, 1994]. The model is initialized with the UARS measurements listed in Table 1. The initial estimates of total odd nitrogen ( $\text{NO}_y = \text{NO}_x + \text{HNO}_3 + 2\text{xN}_2\text{O}_5 + \text{minor species}$ ) and total inorganic chlorine ( $\text{Cl}_y = \text{HCl} + \text{ClO} + \text{HOCl} + \text{minor species}$ ) are determined from a combination of model calculations and UARS data. Model simulations and previous measurements [Russell *et al.*, 1988] of NO<sub>y</sub> constituent species at these altitudes indicate that at sunset, NO<sub>x</sub> makes up greater than 94% of NO<sub>y</sub>. Based on this, we assume that NO<sub>y</sub> is 1.00 to 1.06 times sunset NO<sub>x</sub> measured by HALOE, depending on latitude and altitude. The model Cl<sub>y</sub> is determined by summing sunset HCl (measured by HALOE) and sunset ClO and making a small adjustment (<10%) to account for HOCl. Cl<sub>y</sub> estimated in this way is in good agreement with the Cl<sub>y</sub> estimate from the Goddard 2D model and with 1994 ATMOS measurements [Michelsen *et al.*, 1996]. Initial values for constituents not listed in Table 1 are taken from the 3D CTM run discussed in Kawa *et al.* [1995] at noon local time on March 1, 1992, at 180° longitude and the appropriate latitude and pressure. Model-derived diurnal cycles are obtained after

running the model for four days, which at these altitudes closely approximates running the model to steady state.

Figures 1 and 2 compare the diurnal cycles of UARS ClO and NO<sub>2</sub> with model-derived ClO and NO<sub>2</sub>. The UARS measurements were obtained between Feb. 12 and March 16, 1993, and at pressures of 3.1 and 1.4 hPa, corresponding approximately to altitudes of 40 and 46 km, respectively. Data from the latitude bands 0° $\pm$ 5°, 35°N $\pm$ 5°, and 55°N $\pm$ 5° are sorted by local solar time into one-hour and three-hour bins for the 3.1- and 1.4-hPa data, respectively. These bins are then averaged to obtain the diurnal cycle. Note that while the CLAES and HALOE NO<sub>2</sub> data are retrieved on the 3.1 and 1.4 hPa pressure surfaces, the MLS ClO data are actually retrieved on the 4.6, 2.2, and 1.0 hPa pressure surfaces and interpolated onto the 3.1 and 1.4 hPa surfaces. Because of the shape of the ClO profile, this could potentially introduce a low bias in the interpolated MLS ClO data not likely to exceed  $\sim 20\%$ .

The disagreement shown in Fig. 1 between model-predicted and UARS-measured ClO abundance is significant, with modeled ClO a factor of two higher than measured ClO at 3.1 hPa and a factor of 1.5-2.0 higher at 1.4 hPa. These differences greatly exceed the uncertainty of the measurements. Because Cl<sub>y</sub> in this model is constrained by UARS measurements, we conclude that differences between modeled and measured ClO reflect problems with our understanding of how Cl<sub>y</sub> is partitioned at these altitudes between its major reservoirs, ClO and HCl. Such problems with model predictions of ClO in the mid and upper stratosphere have been seen elsewhere [e.g., Michelsen *et al.*, 1996, Stachnik *et al.*, 1992].

Figure 2 shows general agreement between measured NO<sub>2</sub> and the model. At 3.1 hPa, our UARS adjusted daytime NO<sub>2</sub> data agree within  $\sim 20\%$  with our model estimate of NO<sub>2</sub>, within uncertainty of the data. There is a slight tendency for model NO<sub>2</sub> to be higher than the UARS ad-

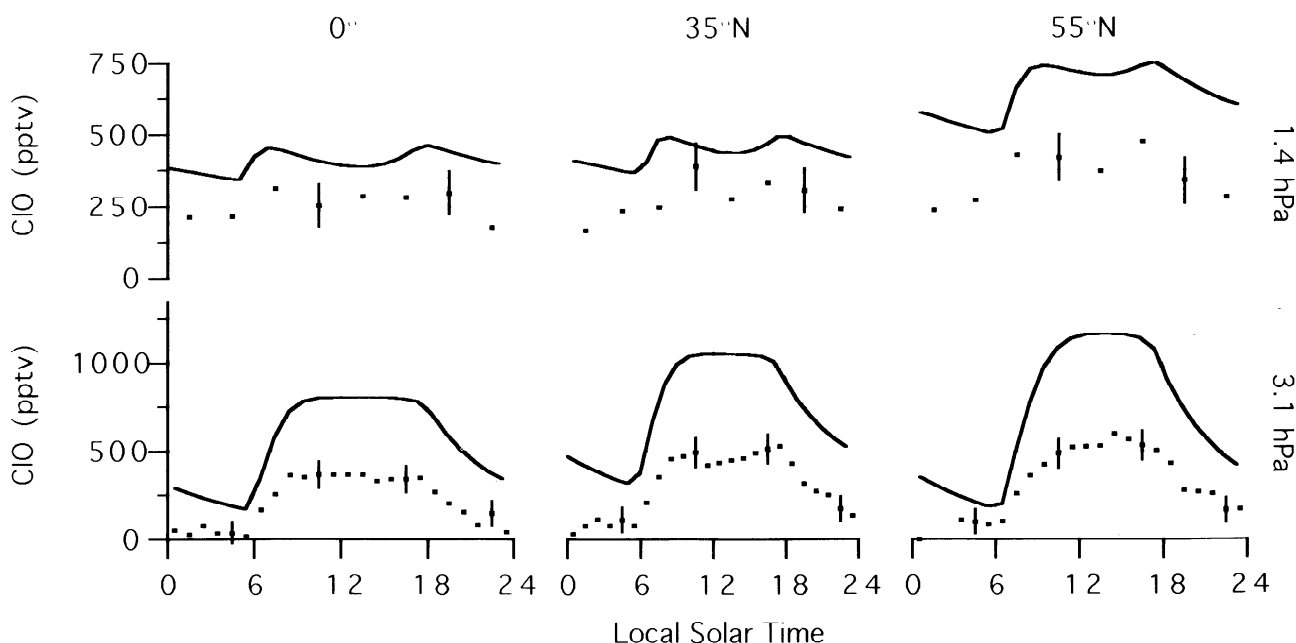


Figure 1. ClO (mixing ratio) vs. local solar time (hr) at 3.1 and 1.4 hPa, and 0°, 35°N, and 55°N latitude. The solid line is ClO from the profile model, the dots are ClO measurements from MLS. Error bars are the root-sum-of-squares of a 8% scaling uncertainty and a 75-pptv bias uncertainty ( $\sim \pm 20\text{-}25\%$ ).

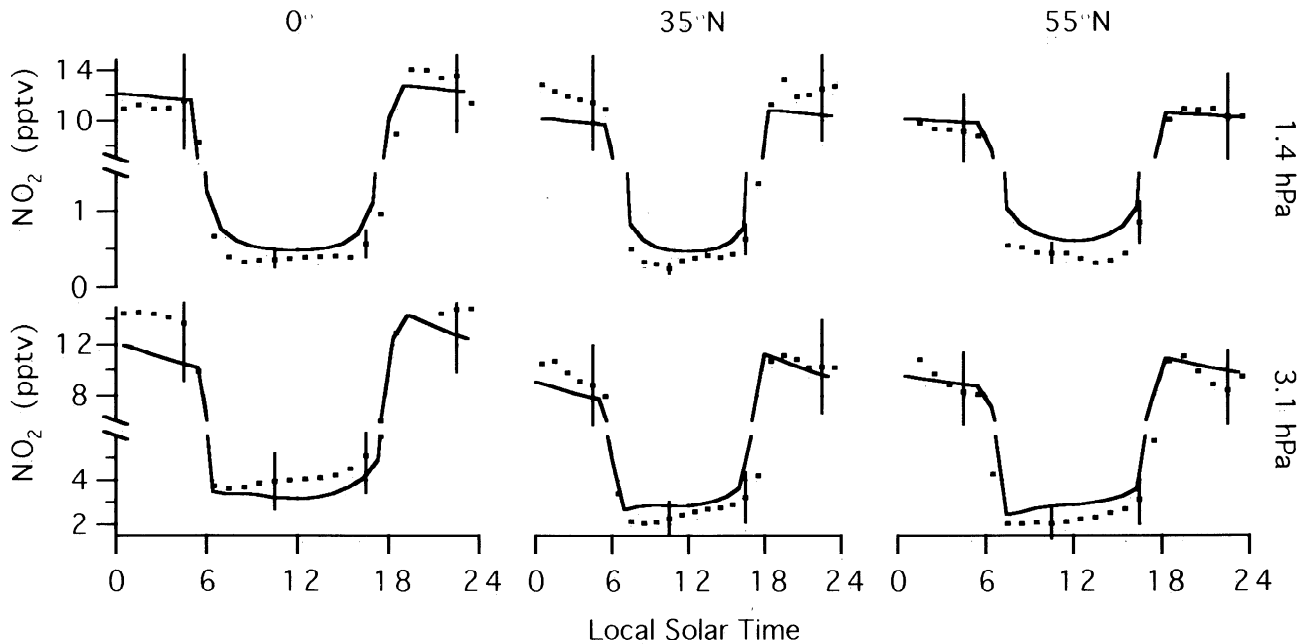


Figure 2. NO<sub>2</sub> (mixing ratio) vs. local solar time (hr) at 3.1 and 1.4 hPa, and 0°, 35°N, and 55°N latitude. The solid line is NO<sub>2</sub> from the profile model, the dots are the NO<sub>2</sub> measurements (see text). Error bars are  $\pm 34\%$ .

justed NO<sub>2</sub> data. This is consistent with the previously discussed model overestimate of ClO, which causes NO<sub>2</sub> abundance to be increased through the enhancement of the reaction between ClO and NO. Taking this into account, the agreement at 3.1 hPa is excellent. At 1.4 hPa, the nighttime UARS adjusted NO<sub>2</sub> data agree within  $\sim 20\%$  with our model's nighttime estimate. During the day, however, the UARS adjusted NO<sub>2</sub> data is on average  $\sim 50\%$  lower than the model NO<sub>2</sub>, a difference in excess of our estimate of the uncertainty of the UARS adjusted NO<sub>2</sub> data. This disagreement is not likely due to an overestimate in the model's NO<sub>y</sub> because NO<sub>y</sub> in the model is constrained by UARS measurements. About one fifth of the daytime discrepancy can be attributed to errors in the model's ClO abundance. The remainder of the difference is just beyond the limit of our uncertainty in the NO<sub>2</sub> measurement, suggesting an error in the model partitioning of NO<sub>y</sub>. However, because the uncertainty in the measurement ( $\pm 34\%$ ) is only slightly smaller than the unattributed disagreement between model and measurement ( $\sim \pm 40\%$ ), there is no evidence that a model partitioning error for NO<sub>2</sub> exceeds a few percent.

### The Upper Stratospheric Model Ozone Deficit

In this section, we perform a simple calculation of O<sub>3</sub> production and loss at 40 and 46 km using the UARS measurements of ClO and NO<sub>2</sub>. The instantaneous production rate of O<sub>3</sub> can be calculated using the expression,

$$P = 2J_{O_2}[O_2] \quad (1)$$

where  $J_{O_2}$  is the photolysis rate constant for O<sub>2</sub> ( $s^{-1}$ ) and  $[O_2]$  is the number density of O<sub>2</sub> (molecules  $cm^{-3}$ ). The instantaneous loss rate of O<sub>3</sub> is the sum of the rates of the limiting steps of the catalytic cycles and the rate of O<sub>3</sub>+O,

$$L = 2k_{ClO+O}[ClO][O] + 2k_{NO_2+O}[NO_2][O] + 2k_{HO_2+O}[HO_2][O] + 2k_{O_3+O}[O_3][O] \quad (2)$$

where  $k_{x+y}$  is the rate constant for the reaction between species  $x$  and  $y$  ( $cm^3/molecule/sec$ ). The concentration of O atoms is determined from a steady-state relation,

$$[O] = \frac{J_{O_3}[O_3]}{k_{O_2+O}[O_2][M]} \quad (3)$$

Photolysis rates for O<sub>2</sub> and O<sub>3</sub> ( $J_{O_2}$  and  $J_{O_3}$ ) are calculated based on pressure, temperature, overhead O<sub>3</sub>, and solar zenith angle. The photolysis rate routine is taken from the Goddard 3D CTM [Kawa *et al.*, 1995]. Overhead O<sub>3</sub> is determined from zonally averaged MLS O<sub>3</sub> measurements (205 GHz, version 3) obtained between Feb. 26 and March 5, 1993. Note that the results are essentially unchanged if HALOE O<sub>3</sub> measurements are used for the overhead column. The radiation calculations are for clear-sky conditions, and a surface albedo of 0.3. All photolysis cross sections and rate constants used in this analysis are taken from JPL-94 [DeMore *et al.*, 1994].

To determine the balance between O<sub>3</sub> production and loss, we first calculate instantaneous production and loss rates every 15 minutes for an entire day using March 1 solar conditions. At each time step, ClO and NO<sub>2</sub> either from UARS measurements or the zero-dimensional model, HO<sub>2</sub> from the zero-dimensional model (constrained by UARS measurements of H<sub>2</sub>O, O<sub>3</sub>, and NO<sub>y</sub>), O atom concentration determined from Eq. (3), average O<sub>3</sub> (from either MLS or HALOE) and temperature from UARS (see Table 1) are combined in Eqs. (1) and (2) to determine the instantaneous production and loss rates. 24-hour averages are determined by averaging the instantaneous rates.

Table 2 lists the ratios of 24-hour average production to 24-hour average loss of O<sub>3</sub> (hereafter referred to as P/L). Because of the short lifetime of O<sub>3</sub> at these altitudes, average production should equal average loss; in other words, P/L should equal one. P/L values less than (greater than) one imply that the model can balance production and loss given less (more) O<sub>3</sub> than specified, leading to a model

**Table 2.** Ratio of 24-hour Average Production to 24-hour Average Loss

	0°N	35°N	55°N	0°N	35°N	55°N
	UARS ClO and NO <sub>2</sub> , MLS O <sub>3</sub>			Model ClO and NO <sub>2</sub> , MLS O <sub>3</sub>		
3.1 hPa	0.88	1.10	1.19	0.78	0.78	0.80
1.4 hPa	0.85	0.92	1.01	0.74	0.75	0.77
	UARS ClO and NO <sub>2</sub> , HALOE O <sub>3</sub>			Model ClO and NO <sub>2</sub> , HALOE O <sub>3</sub>		
3.1 hPa	1.06	1.22	1.22	0.95	0.86	0.81
1.4 hPa	1.13	1.08	1.16	0.98	0.87	0.87

ozone deficit (surplus). Table 2 shows an ozone deficit for all cases using model-derived ClO and NO<sub>2</sub> radical abundances, consistent with results from our 2D model and many previous model analyses. When using UARS-measured ClO and NO<sub>2</sub> radical abundances, P/Ls are considerably higher and are generally greater than one. P/L calculated using HALOE O<sub>3</sub> data is 0.01-0.28 higher than P/L calculated using MLS O<sub>3</sub>. Table 2 also shows that virtually any answer between a deficit of 25% and a surplus of 20% can be obtained depending on the choice of latitude, O<sub>3</sub> data set and various other parameters.

Deviations of P/L from one represent either errors in the parameters that are used to calculate P/L (e.g., errors in O<sub>3</sub> or k<sub>O<sub>3</sub>+O</sub> measurements, J<sub>O<sub>2</sub></sub> calculations, etc.) or problems in our understanding of O<sub>3</sub> photochemistry in this region of the atmosphere. Standard error propagation of the uncertainties of the individual parameters suggests that the uncertainty in our calculation of P/L is  $\sim\pm 0.3$  (1 $\sigma$ ). Deviations of P/L from one that are less than 0.3 can therefore be wholly explained by the stated uncertainties in the parameters used in the calculation. Deviations of P/L from one that are greater than 0.3 suggest problems in our understanding of the photochemistry of the upper stratosphere.

All of the cases shown in Table 2 and virtually all of the analyses in the current literature concern deviations of P/L from one that are less than 0.3 (deficits or surpluses less than 30%). Differences between the various literature estimates of P/L can often be explained by the choice of the parameters going into the calculations. For example, the calculation of *Crutzen et al.* [1995] showing production equaling loss (P/L  $\sim$  one) is likely attributable to their use of the lower HALOE O<sub>3</sub> values as well as their choice of a relatively low latitude (23°). Had they used the higher MLS O<sub>3</sub> values, our analysis suggests that their slight surplus would have likely become a 15 to 20% deficit (P/L  $\sim$  0.80-0.85), consistent with other recent analyses [*Eluszkiewicz and Allen*, 1993, *Siskind et al.*, 1995]. This underscores the large uncertainties associated with calculations of P/L.

## Conclusions

In this paper, we have investigated the photochemistry of O<sub>3</sub> at 3.1 and 1.4 hPa, approximately 40 and 46 km. For the first time, we have incorporated diurnally-resolved daytime measurements of both ClO and NO<sub>2</sub>. From our analysis, we draw the following conclusions:

(1) Comparisons between measurements of ClO and a zero-dimensional model using standard JPL-94 photochemistry show that models overpredict ClO by about a factor of two at 3.1 and a factor of 1.5-2.0 at 1.4 hPa. This discrepancy exceeds the uncertainty in the measurements and is likely due to problems with our understanding of Cl<sub>y</sub> partitioning. Comparisons between measurements of NO<sub>2</sub>

and the model show that NO<sub>2</sub> is well simulated at 3.1 hPa, but is systematically overestimated by the model at 1.4 hPa.

(2) Because of model overestimates of ClO, models incorporating standard JPL-94 photochemistry will produce a loss of O<sub>3</sub> whose magnitude is 10 to 40% larger than calculations utilizing measured ClO. Note that the systematic disagreement between measured and modeled NO<sub>2</sub> at 1.4 hPa does not have a significant effect on O<sub>3</sub> loss because NO<sub>2</sub>+O is not an important loss pathway for O<sub>3</sub> at 1.4 hPa.

(3) Because of the large uncertainty in the P/L calculation ( $\sim\pm 0.3$ , 1 $\sigma$ ), only extremely large deviations of P/L from one can be used as evidence of problems in our understanding of upper stratospheric photochemistry. Virtually all of the recent analyses of upper stratospheric O<sub>3</sub> produce surpluses or deficits that are within this uncertainty. This suggests that O<sub>3</sub> P/L calculations are no longer sensitive tests of our understanding of O<sub>3</sub> photochemistry. Instead, we believe that a better approach to testing the photochemistry of this region is to compare measurements and model calculations of the radicals involved in the rate limiting steps of the catalytic cycles that control ozone abundance.

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